Sources for Alfvén Waves Above the Auroral Oval

C. C. Chaston¹, T. D. Phan¹, C. W. Carlson¹, J. P. McFadden¹, J. W. Bonnell¹, L.M. Peticolas¹, F. S. Mozer¹ M. Wilber¹, G. Parks¹, R. E. Ergun², R. J. Strangeway³, E. Lund⁴, M. Andre⁵, Y. Khoyaintsev⁵, M. Acuna⁶, M. Goldstein⁶, V. Genot⁷, H. Reme⁷, I. Dandouras⁷, A. Fazakerley⁸, A. Balogh⁹

¹Space Science Laboratory, University of California, Berkeley, California, USA
²Laboratory for Atmospheric and Space Physics, University of Colorado, Boulder, Colorado, USA
³Institute for Geophysical and Planetary Physics, University of California, Los Angeles, California, USA
⁴University of New Hampshire, NH, USA
⁵Swedish Institute of Space Physics, Uppsala, Sweden
⁶NASA Goddard Space flight Center, Greenbelt, Maryland, USA
⁷CESR, Toulouse, France
⁸Mullard Space Science Laboratory, Dorking, United Kingdom
⁹Imperial College, London, United Kingdom

In this presentation we examine FAST and Cluster spacecraft observations to identify sources for the multi-scale Alfvén waves commonly observed above the auroral oval. Several case studies are explored in detail which show that surface Alfvén waves on the magnetopause and reconnection may provide large Poytning fluxes of Earthward directed Alfvén waves on multiple scales extending from ion gyro-radii up to those appropriate for MHD. From conjunctive observations from FAST and Cluster it is demonstrated that these waves populate the flux tubes on which they are observed with ionospheric plasmas through field-aligned electron and transverse ion acceleration.