

# ESP Status. Calibration, Degradation, Validation

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# ESP Status (1)

- **Fully operational, correctable, and predictable. No degradation, except for the Al filters is detected.** We do:
- Regular tests of ESP variables (changes of dark count-rates, gain, visible light, filter degradation) based on daily and weekly calibrations.
- Prompt data reduction and submitting EVE internal Reports, e.g. related to the Filter Scans, FOV maps, Cruciform maps, etc. These reports allow us updating ESP Level1 program for calculation of absolute irradiances with better accuracy, to calculate locations of the flares, etc.
- Regular workshops to discuss current and future works (the last one was on Aug 23-24) @ LASP, including technical issues and science with ESP.

## ESP Status (2)

- ESPR calibration @SURF BL-2 in Sep was successful. The results will be ready upon receiving SURF fluxes. A fast look shows very small changes compared to the previous ESPR calibration.
- ESP QD data were used as near-real time proxies for GOES XRS-A,B with small latency in Level-0CS.
- Level-0CS updated to get ESP irradiance-like data (all channels) with small latency.
- ESP Ch9 (30.4 nm) channel is validated as a proxy for SEM 30.4 nm channel.
- We improve ESP Level1 Program. Current work is to implement filter degradation 'live' updates from daily/weekly calibrations.

# How Filter Degradation is Determined

- We use ESP daily and weekly calibrations to calculate ratios between effective counts measured with different filters.
- Time difference between two filter measurements is as small as one minute to provide smallest signal fluctuations related to possible changes of the solar EUV flux.
- We work on a new algorithm, which can provide accurate measurements of the ratios for the calibration performed during rapid changes of the solar irradiance, e.g. after a solar flare\*.

\* Time for each calibration is fixed, thus, we may catch some flare(s).

# What ESP Tests During Daily/Weekly Calibrations

- ESP provides **absolute** solar irradiance in different (5) spectral bands centered at 3.6, 19.0, 25.4, 29.9, and 36.4 nm.
- As an advanced version of SOHO/SEM, ESP electronics has more test options than SEM electronics, such as
- Test / correction for the electronics gain, including changes of the electronics related to the TID;
- Test / correction for the thermal / shunt\_resistance changes of the detectors;
- Test / correction for the visible light / filter\_pinholes;
- A filter wheel with three science aluminum filters, one prime (F3) and two spare (F2 and F4). This option allows us to measure a filter degradation as a ratio of effective counts between the tested and reference filters

# ESP Calibrations

ESP has a Filter Wheel with 5 filters: Dark, three Al filters (F2, F3, F4) and one Fused Silica Filter (F5).

- Before DOY = 249 (Sep 5):

Filter positions for **daily**: 3, 4, 5, 1, 3: **Ratio 4/3**

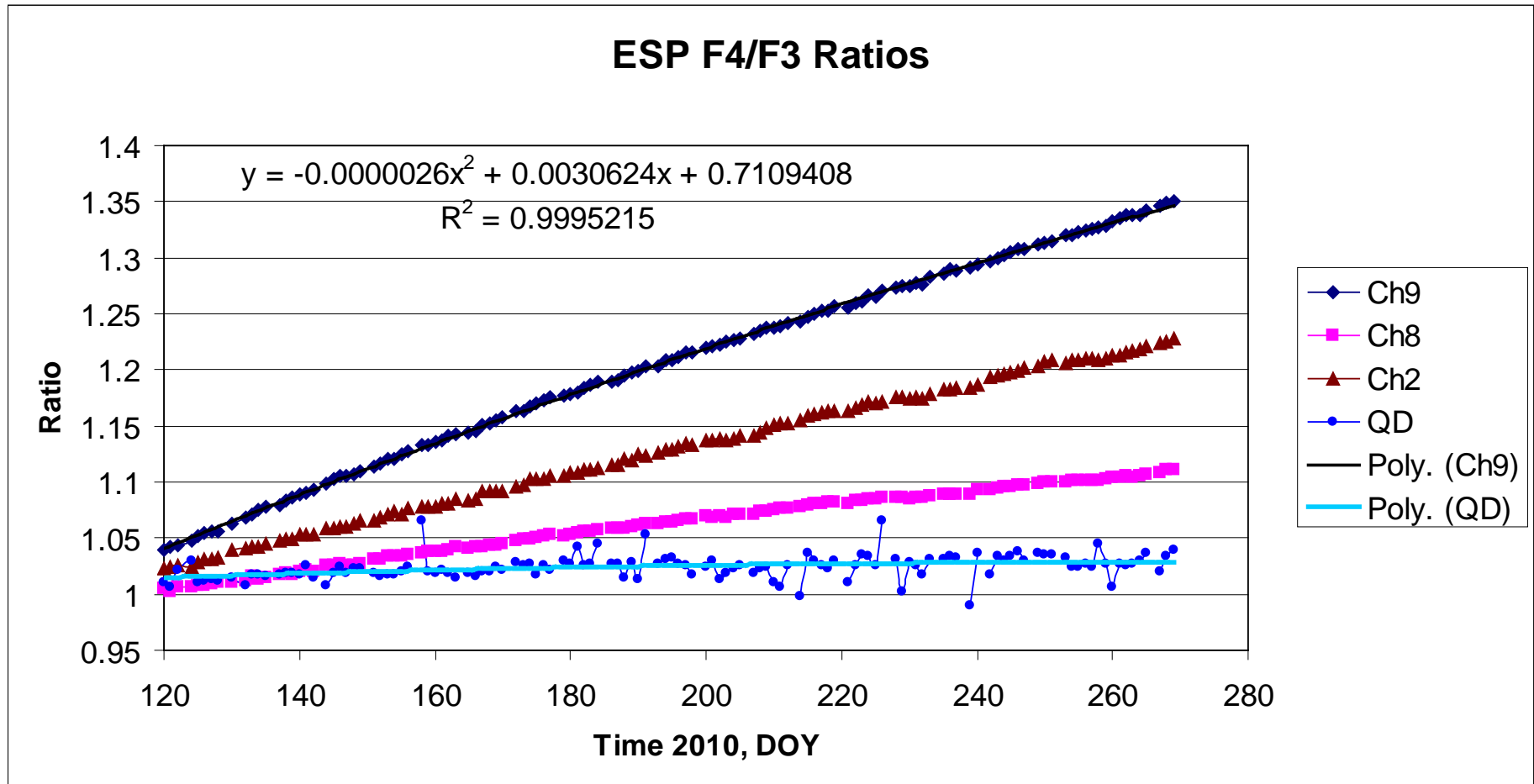
Filter positions for **weekly**: 3, 1, 2, 3: **Ratio 2/3**

- After/on DOY = 249:

Filter positions for **daily**: 3, 4, 1, 3: **Ratio 4/3**

Filter positions for **weekly**: 3, 4, 5, 1, 2, 3: **Ratios 4/3 and 2/3.**

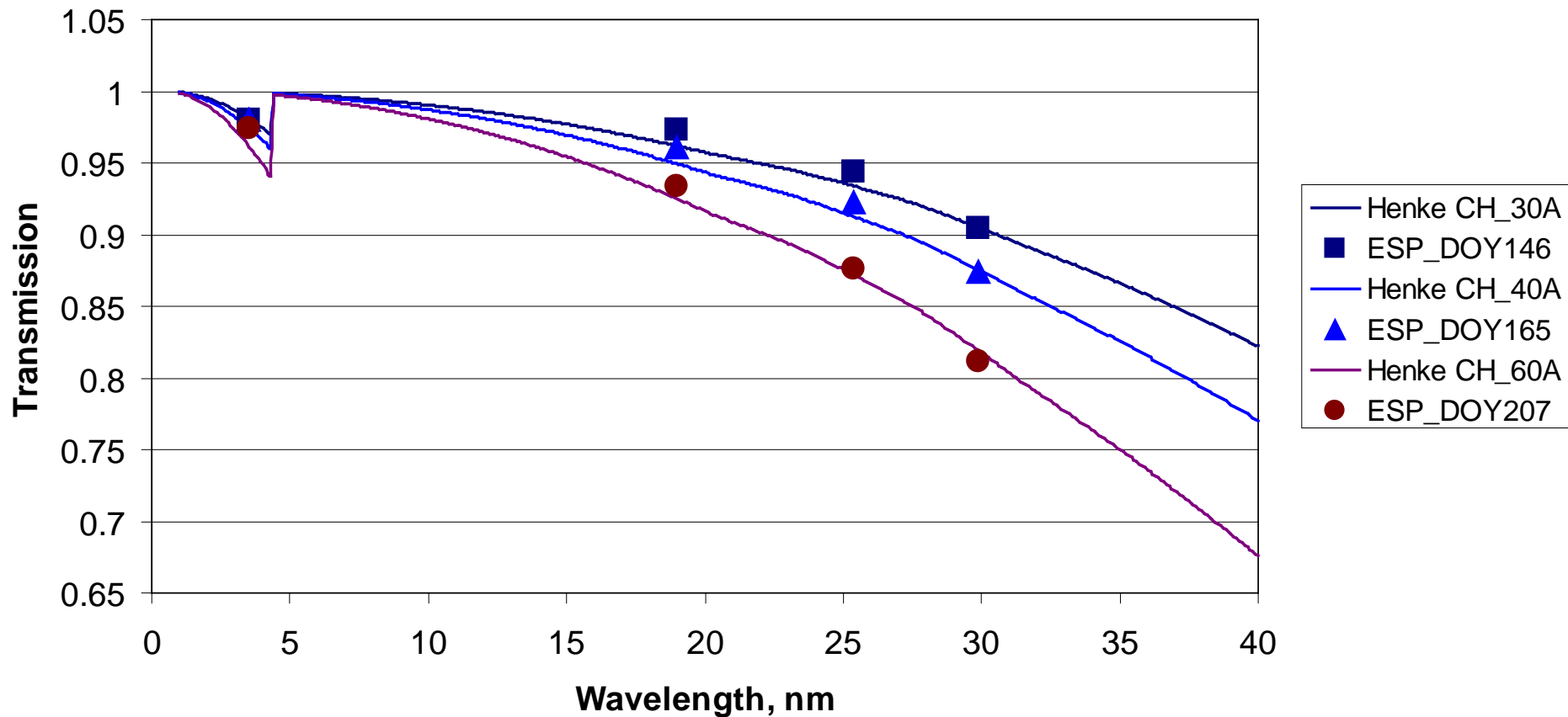
# ESP Filter 3 (prim sci) Degradation



Degradation of ESP AI filter is due to deposition of hydrocarbon, which reduces the transmission of the filter. The change of transmission is wavelength dependent.  
Predicted Ch9 ratio for DOY=365 is 1.45 0.03

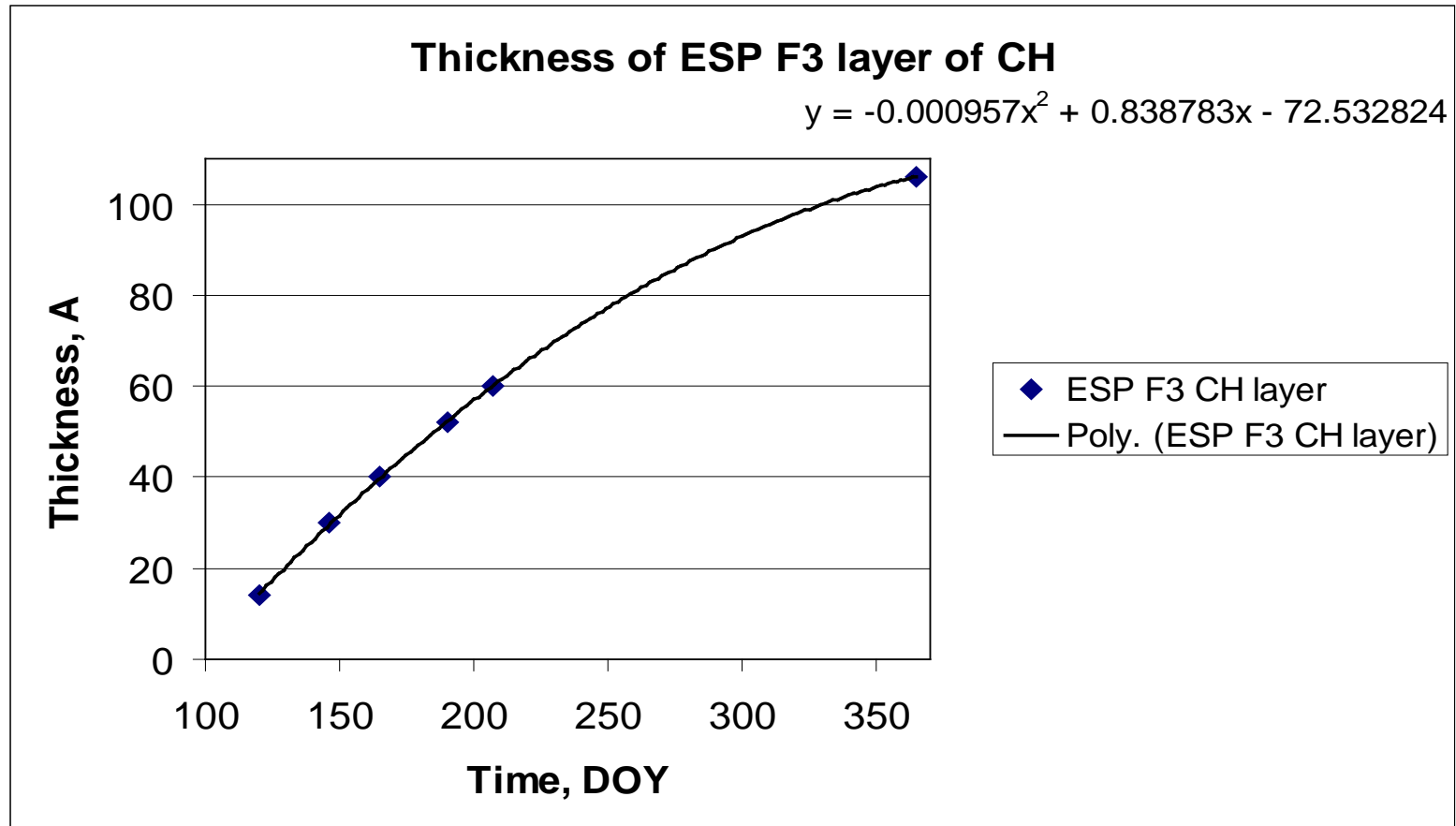
# A Comparison of a Henke-Model and Data

ESP F3 Degradation With a CH Deposition





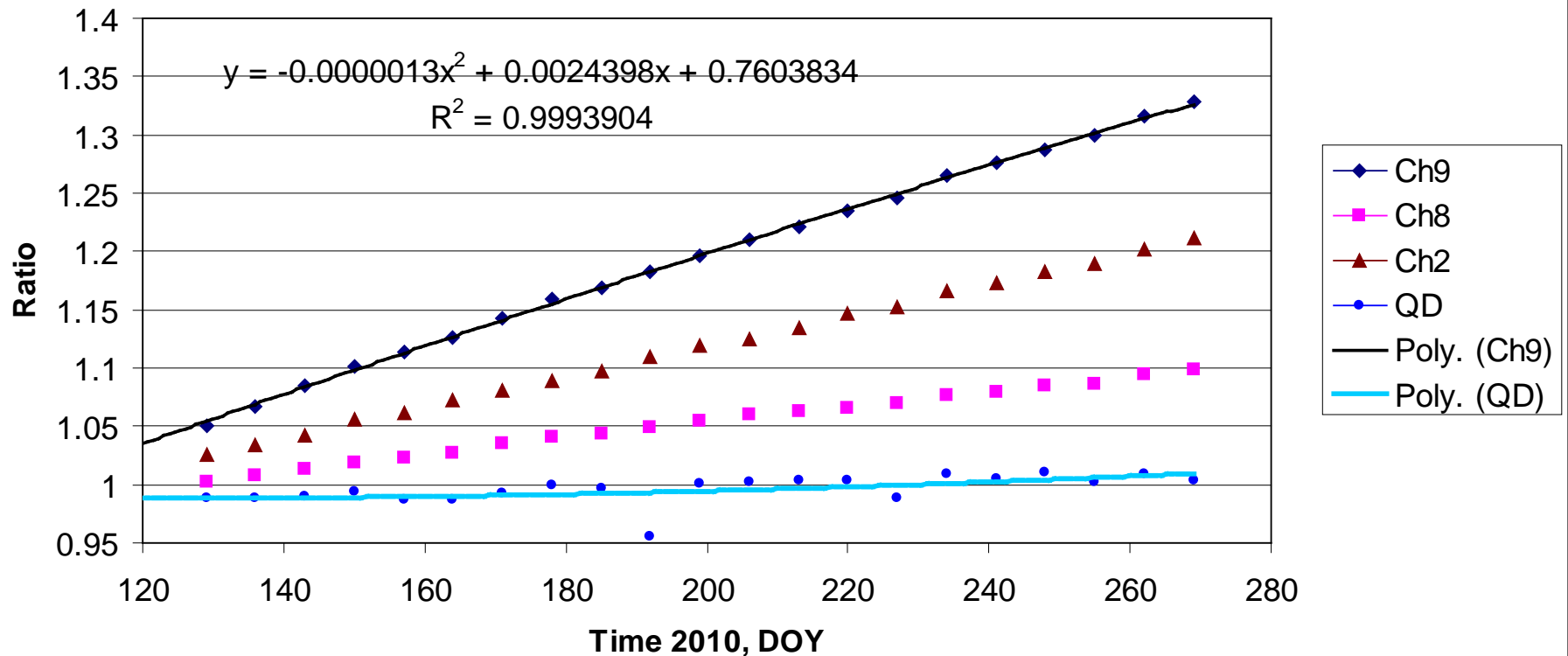
# Thickness of F3 CH Layer: a Model Based on Measurements



Predicted thickness of CH layer for DOY=365 is about 107 A

# Another Evidence of F3 Degradation

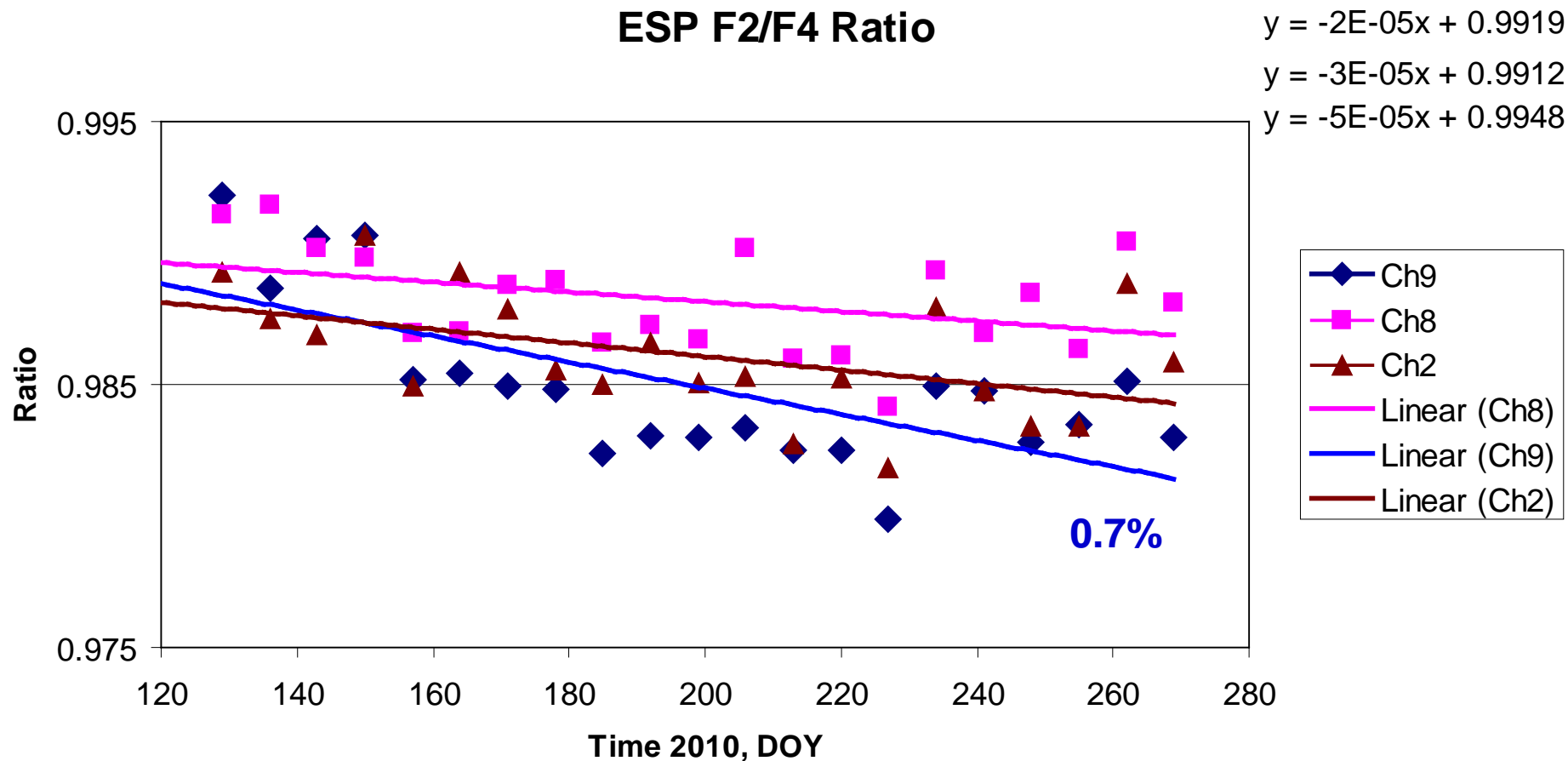
ESP F2/F3 Ratio



Data from weekly calibrations (using F2) show F2/F3 ratios similar to F4/F3 ratios. However, they are smaller than F4/F3 ratios.

# An Evidence of F2 Small Degradation

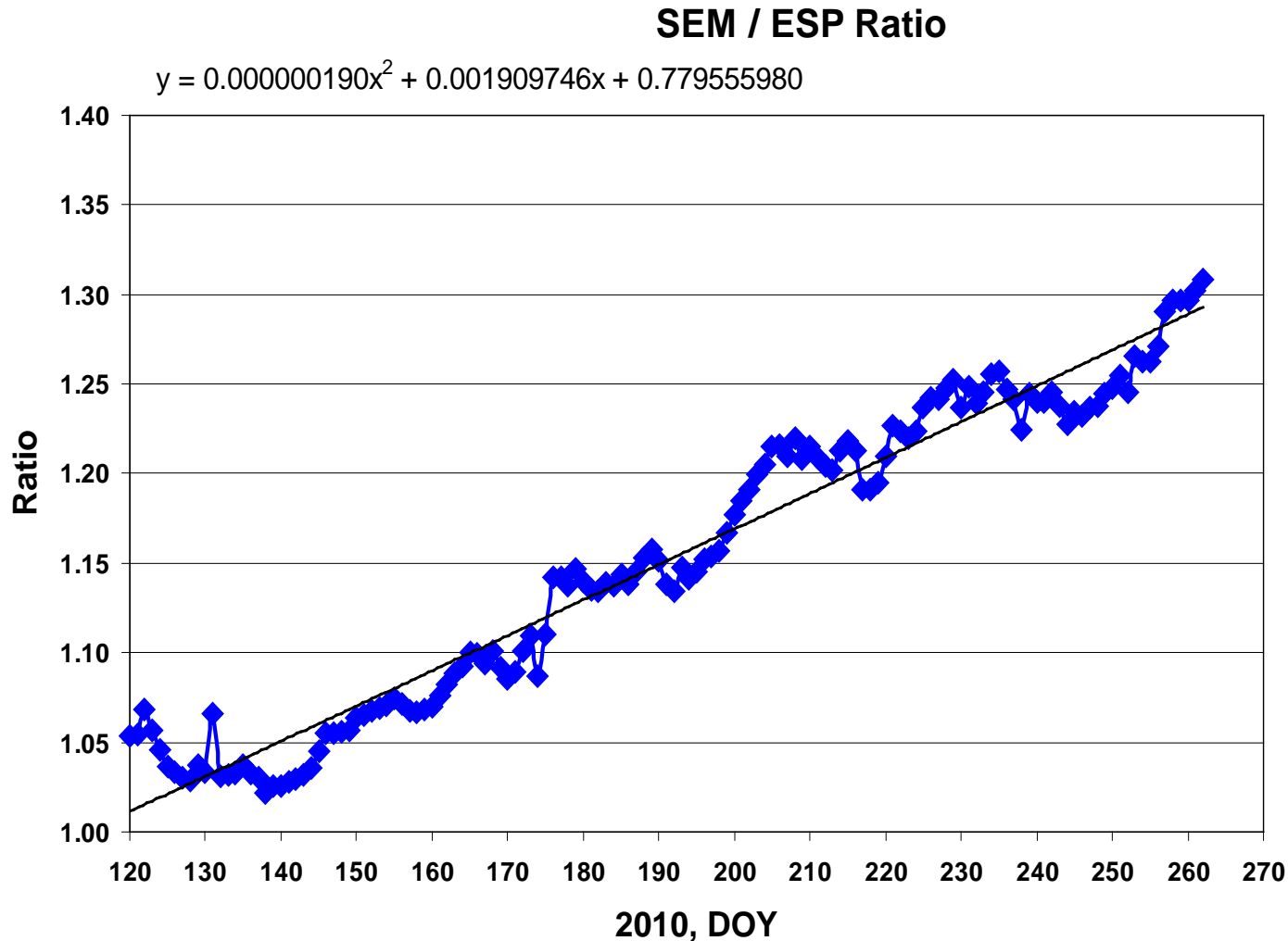
ESP F2/F4 Ratio



These slopes for the first-order ratios show some degradation of 'reference' F2, which was exposed to the Sun for 1 min/week (compared to 6 min/week for F4)

# ESP Validation: SEM / ESP (30.4 nm) Comparison

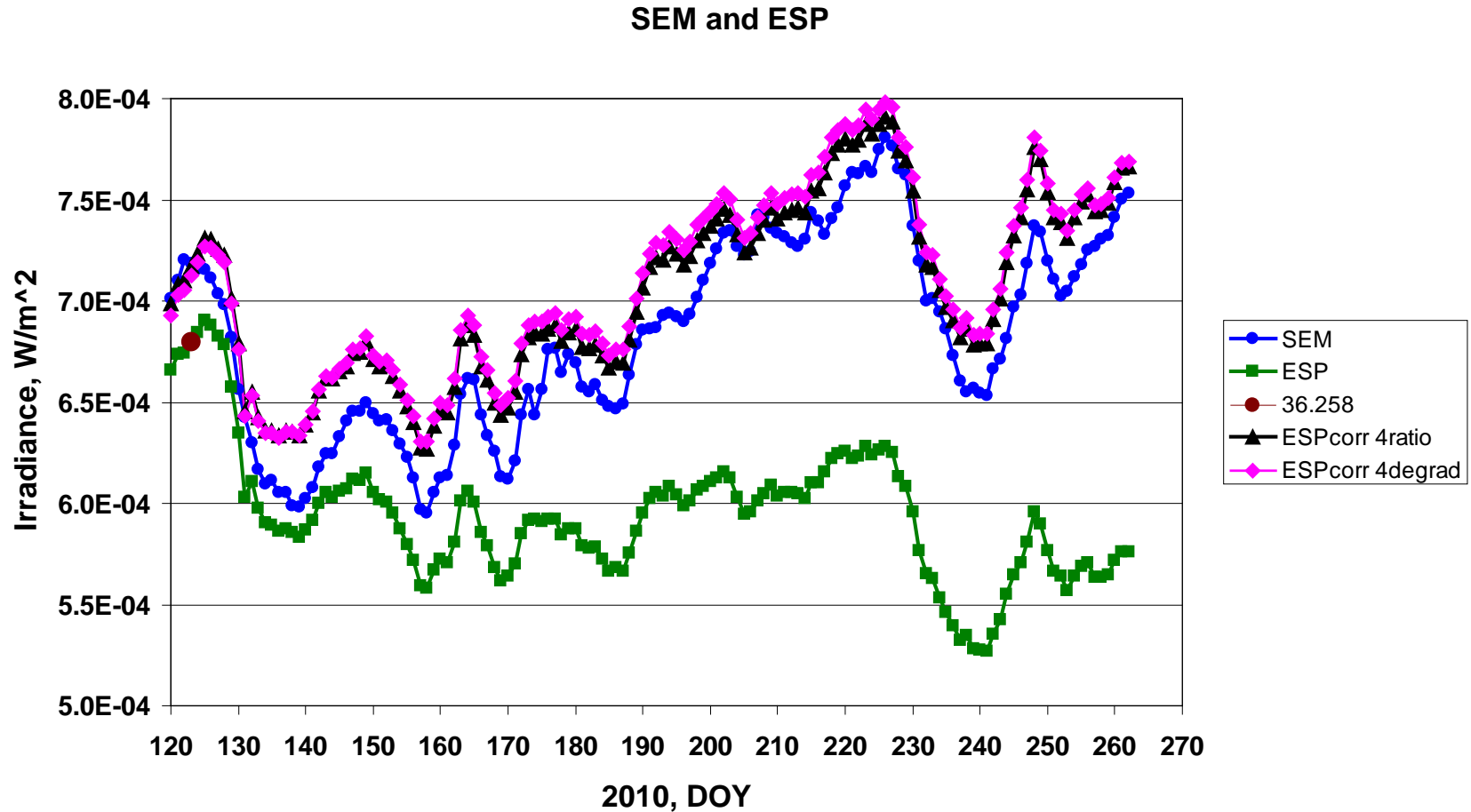
## That Revealed a Filter Degradation Issue



**The change of  
SEM/ESP ratio**

**was discussed  
at the EVE/AIA  
Workshop at  
Stanford and at  
COSPAR (2010).  
We pointed out  
that one of  
explanations of  
this change is  
ESP degradation**

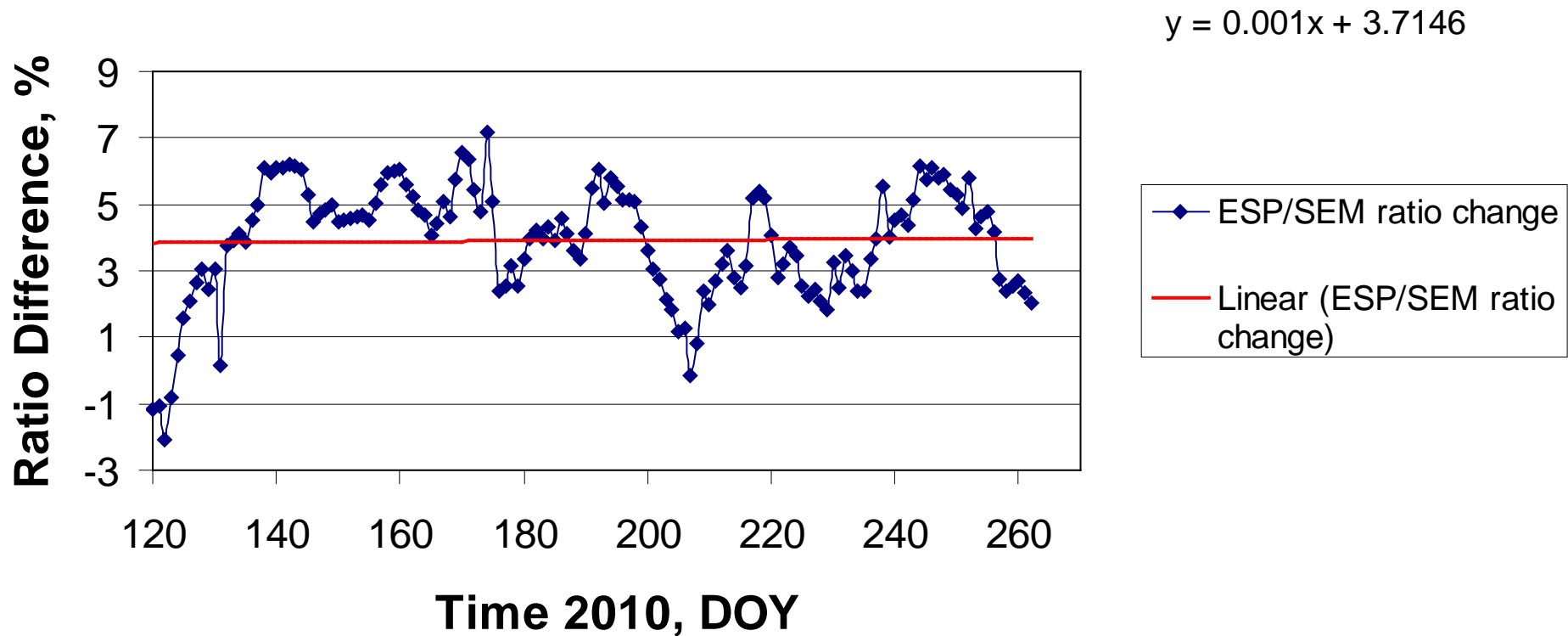
# ESP Validation: Ch9 as SEM Proxy



Black: ESP corrected using a fit to the SEM/ESP ratio; Red: corrected using a fit to the filter degradation curve.

# ESP Validation: Ch9 Provides Accurate ( 2.5%) SEM Proxy!

## ESP/SEM Irradiance Ratio Change



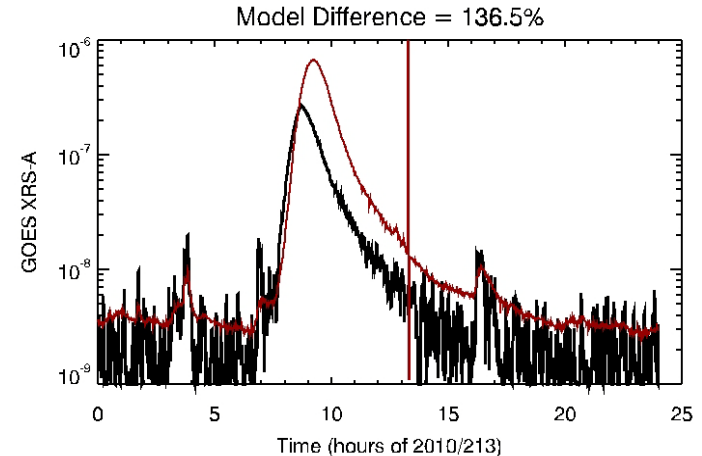
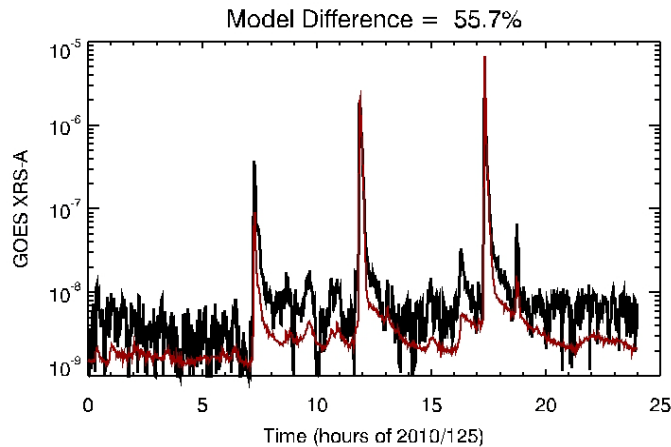
Practically horizontal linear fit (no trend) shows that both SEM and ESP filter degradations were calculated correctly and accurately. STD deviations around the trend line are about 2.5%.

# Acknowledgments

- This work was supported by the University of Colorado **award 153-5979**

# Backup Slide: GOES Proxies

$$XRS - A(0.05 - 0.4nm) = 10^{[-18.94 + 3.606 \cdot \log_{10}(ESP_{0-7nm})]}$$



$$XRS - B(0.1 - 0.8nm) = 10^{[-11.86 + 2.259 \cdot \log_{10}(ESP_{0-7nm})]}$$

