

# MEASURED DEGRADATION IN SOLAR EUV SPECTROMETERS SOHO-CELIAS-SEM AND SDO-EVE

**Andrew Jones**

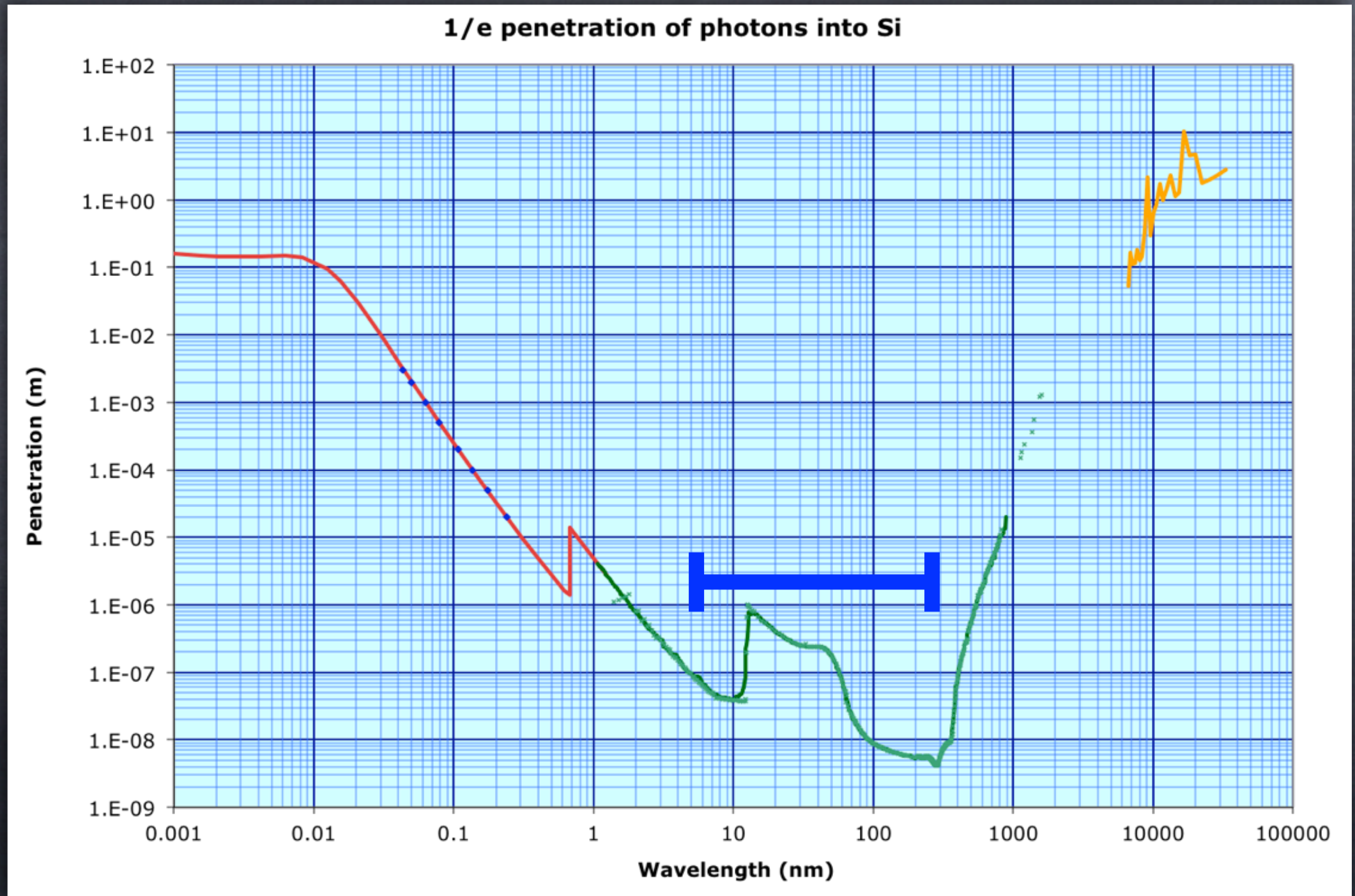
Laboratory for Atmospheric and Space Physics  
University of Colorado, Boulder  
[andrew.jones@lasp.colorado.edu](mailto:andrew.jones@lasp.colorado.edu)





# THE PROBLEM IN EUV

## ☀ Photon penetration depth !





# EUV DEGRADATION

## ☀ Contamination

- ☀ Hydrocarbons

- ☀ Siloxanes (silicones, RTV)

- ☀ Fuel

## ☀ Detector Degradation

## ☀ How do we track the Degradation?

- ☀ Rocket Underflights

- ☀ Multiple optical filters with different exposures

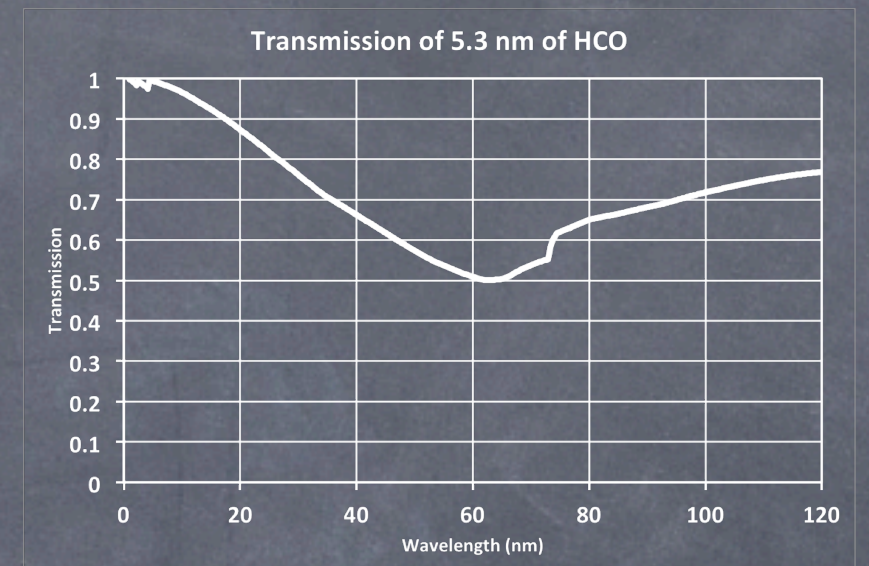
- ☀ Pre-flight Calibration



# COMMON CONTAMINANTS

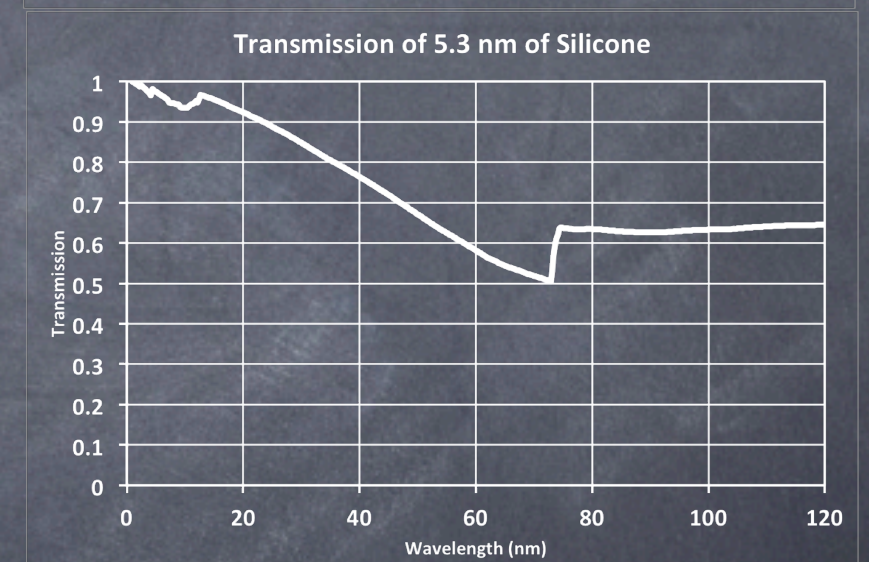
## ☀ Hydrocarbons (C H O)

☀  $\sim 1.2 \mu\text{g}/\text{cm}^2$  gives 50% attenuation  $\sim 62 \text{ nm}$



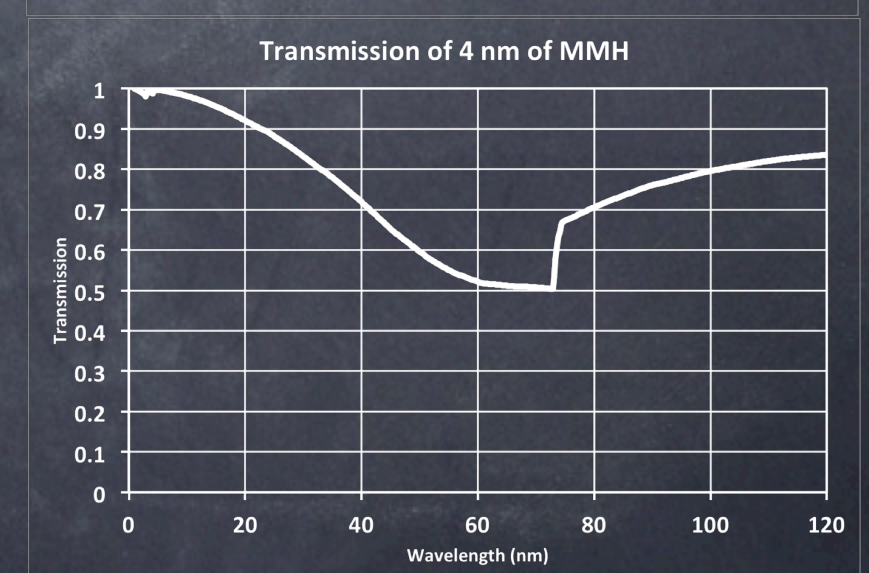
## ☀ Siloxanes (Silicones) (Si O C<sub>3</sub> H<sub>6</sub>)

☀  $\sim 1.2 \mu\text{g}/\text{cm}^2$  gives 50% attenuation  $\sim 72 \text{ nm}$



## ☀ Propellant (C H<sub>6</sub> N<sub>2</sub>)

☀  $\sim 0.9 \mu\text{g}/\text{cm}^2$  gives 50% attenuation  $\sim 72 \text{ nm}$





# INSTRUMENTS

## ☀ SOHO-SEM:

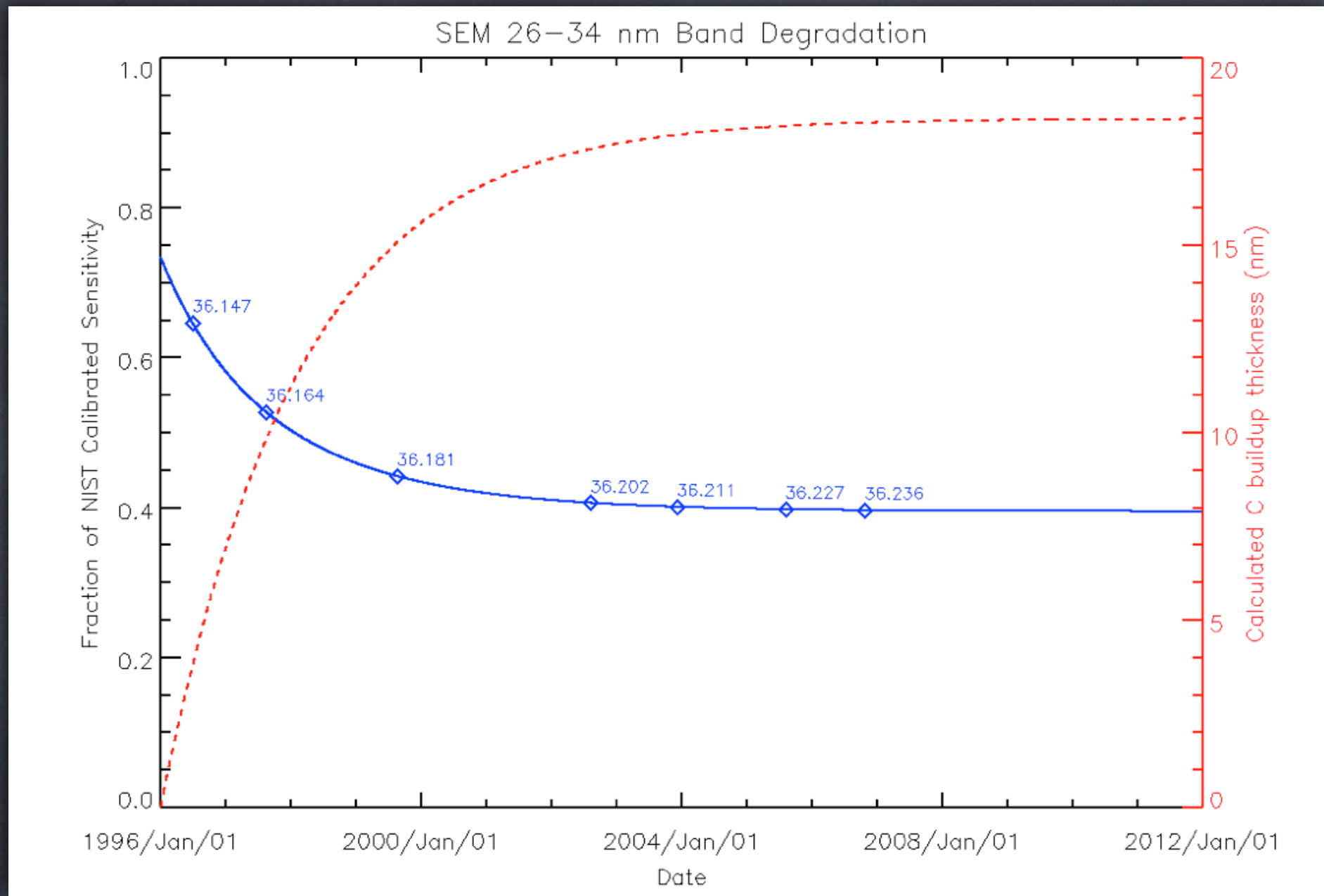
- ☀ Transmission grating spectrophotometer
- ☀ Photodiode detectors
- ☀ Single Al entrance filter
- ☀ Sounding rocket calibration under flight program

## ☀ SDO-EVE:

- ☀ Reflection (MEGS) and transmission (ESP) gratings
- ☀ CCD and photodiode detectors
- ☀ Filter wheels for all channels
- ☀ Sounding rocket calibration under flight program



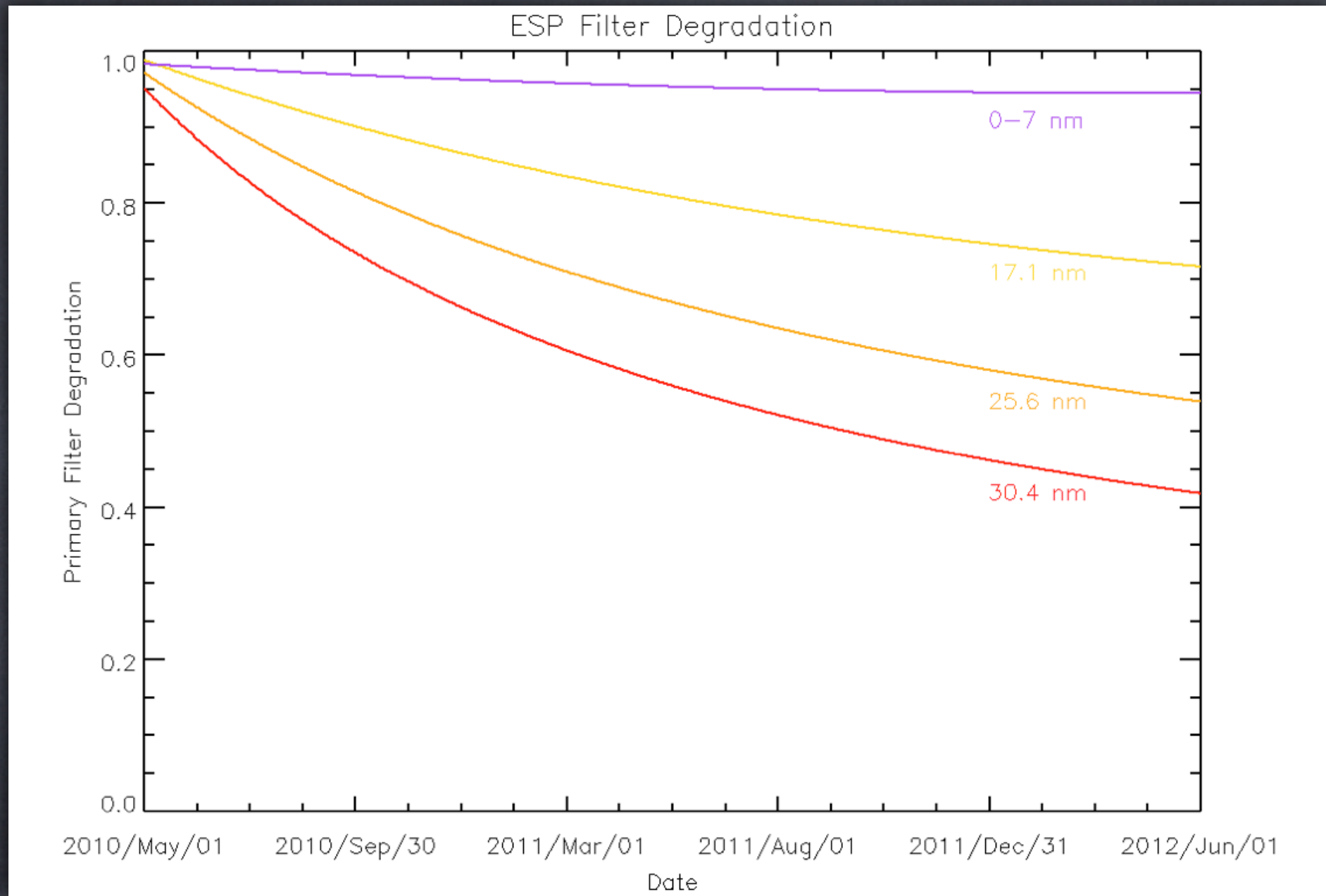
# SEM FILTER DEGRADATION



- ☀ Degradation tracked with sounding rockets
  - ☀ Modeled as C buildup
  - ☀ First-light degradation ~ 9 nm C



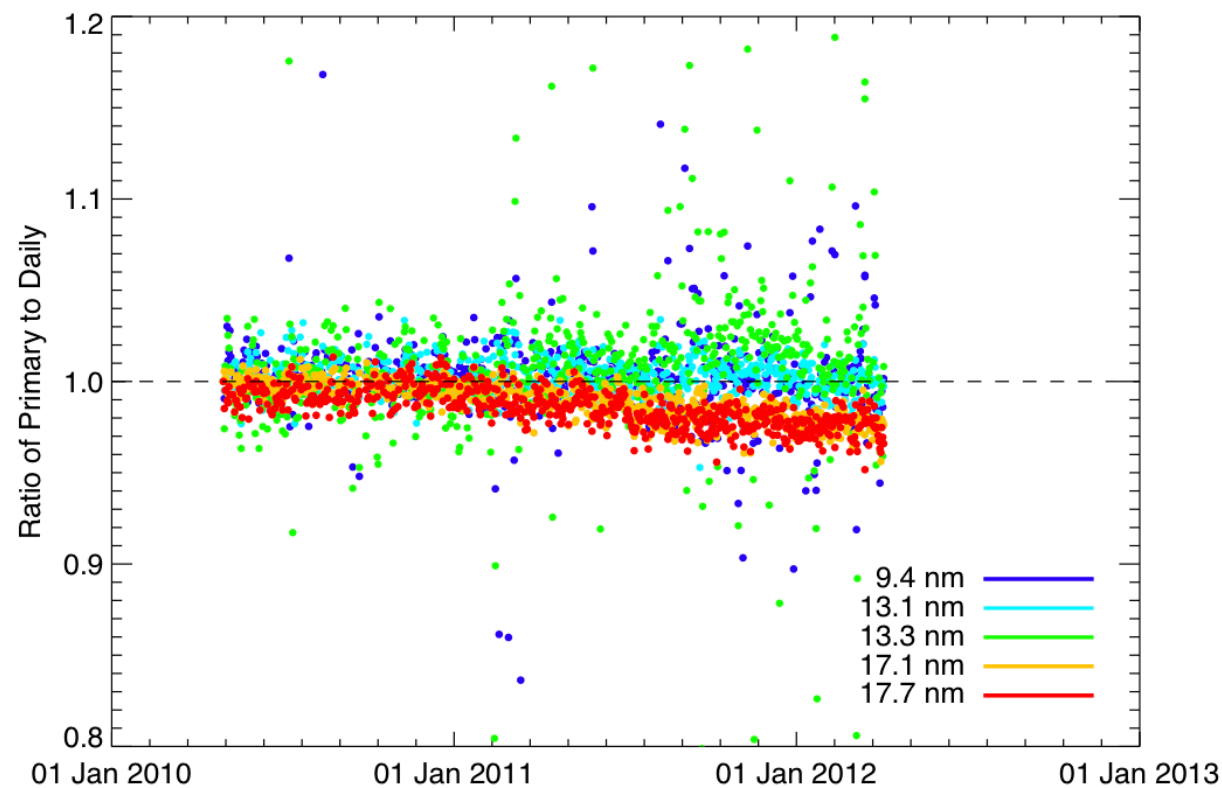
# ESP FILTER DEGRADATION



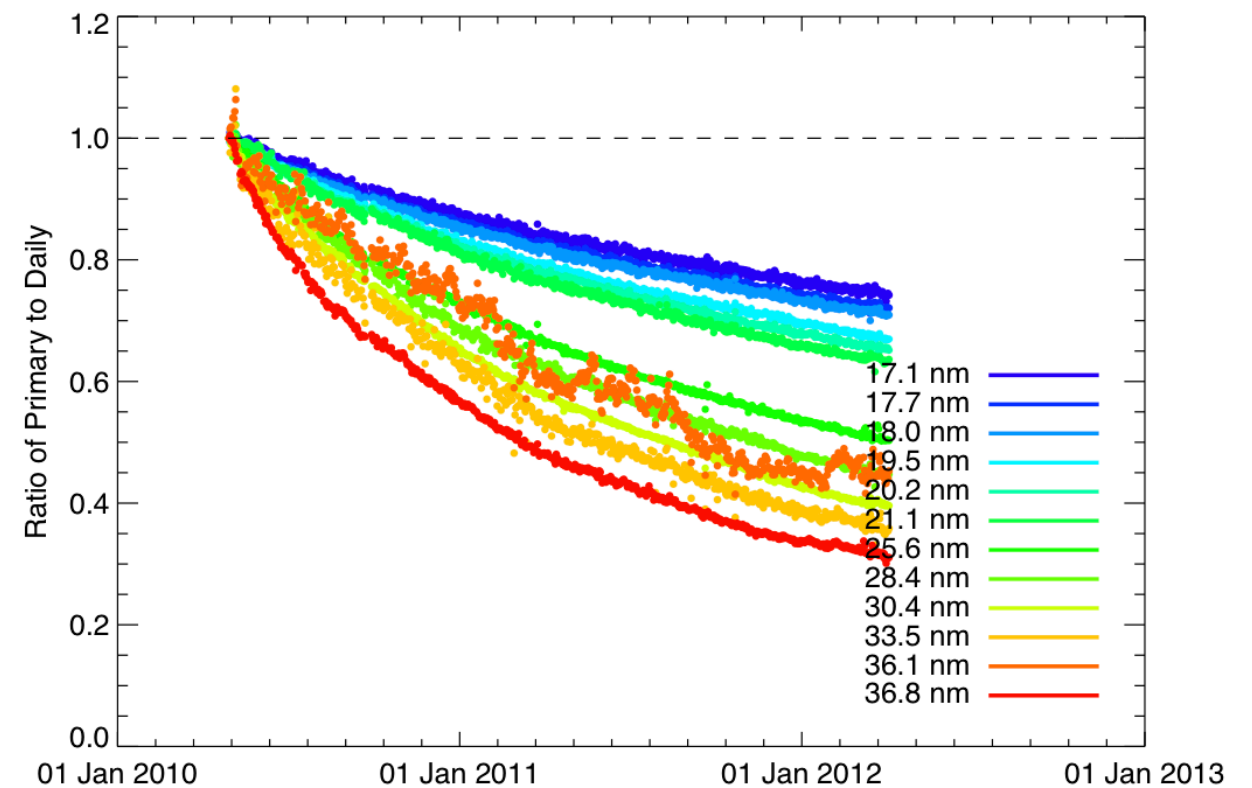
- ☀ Degradation tracked with sounding rockets and redundant filters
  - ☀ Consistent with absorption of  $\sim 25$  nm C



# MEGS FILTER DEGRADATION



A1: C/Zr/C filter



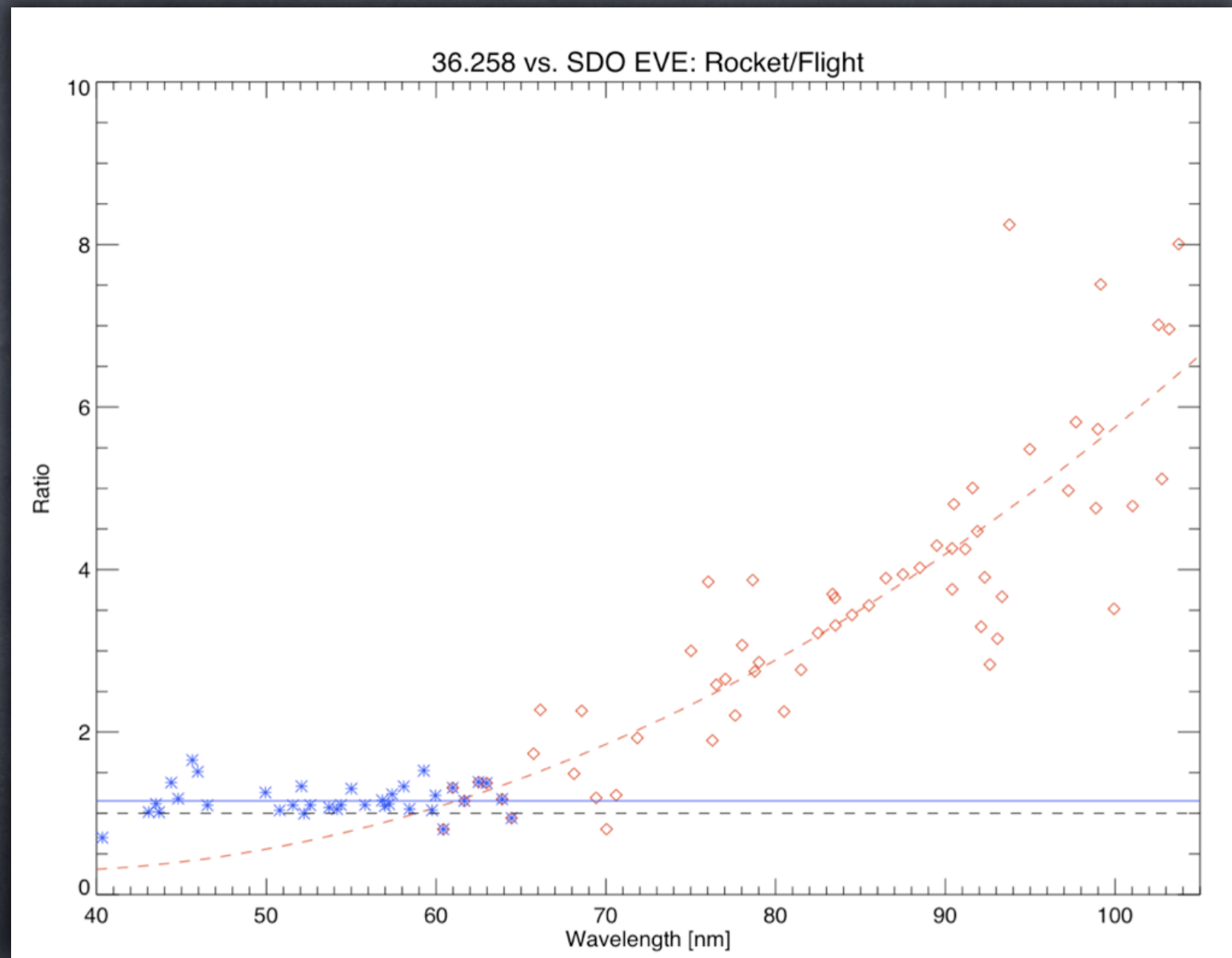
A2: Al/Ge/C filter

A2 consistent with hydrocarbon absorption and  
very similar to ESP

A1 <1% degradation even at 17 nm !



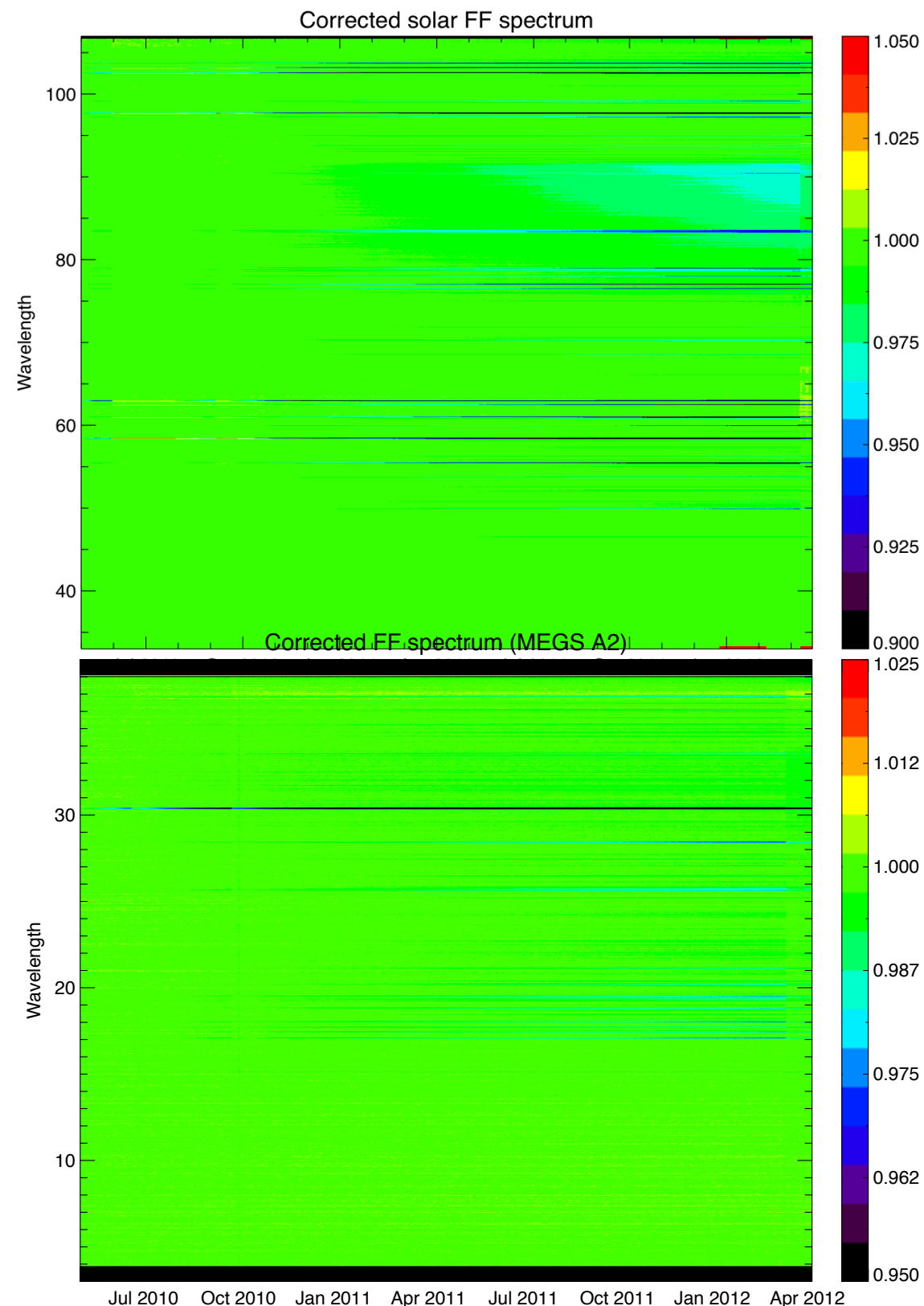
# SDO-EVE CCD FIRST-LIGHT DEGRADATION



Consistent with Si absorption  
due to dead-layer at the top of the CCD



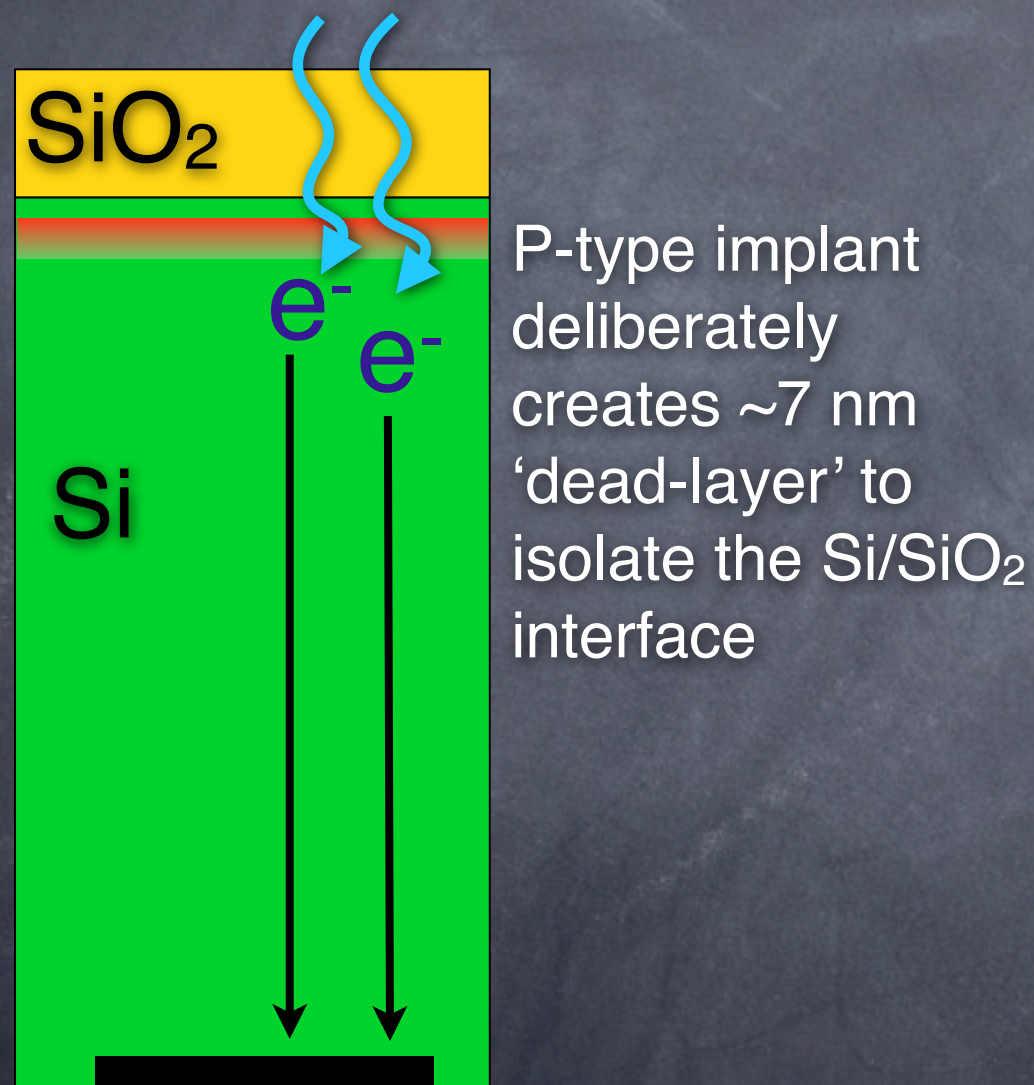
# SDO-EVE CCD FLAT FIELDS



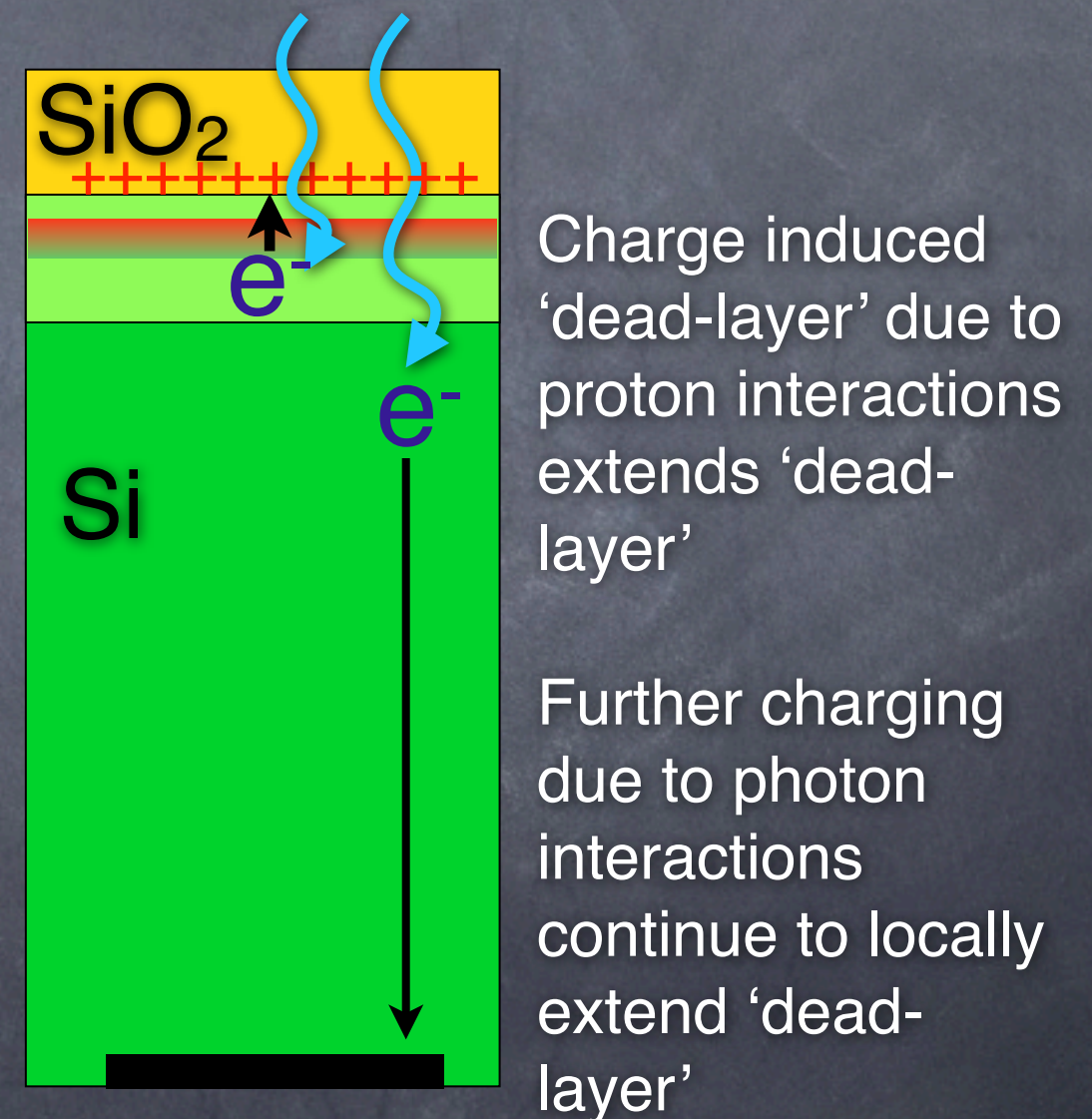
- ☀ Burn-in due to local surface charging
- ☀ Some recovery from bakeouts



# CCD BACK-SIDE CHARGE QE REDUCTION MECHANISM



Charge Collection



Charge Collection



# LESSONS LEARNED

- ☀ No epoxies / silicones in the optical cavity or with path to optical cavity
  - ☀ Well... only clean ones (Stycast 2850, Scotchweld 2216)
- ☀ No oil pumps—anywhere anytime
- ☀ Meticulous cleanliness program: at instrument and spacecraft level
- ☀ Treat like a UHV system, bake all items and test for contaminants
  - ☀ Flex cables are easier to clean than harnesses
  - ☀ Careful with coatings (no organics, temperature limits etc.)



# LESSONS LEARNED -2

- ☀ Constant N2 purge (we use filtered N2 dewar boil-off)
  - ☀ Instrument
  - ☀ Optics / subassembly storage
  - ☀ Make sure the tubing is clean too
- ☀ Constant monitoring (witness samples, RGA, TQCM...)
- ☀ Consider radiation damage in risk assessments
- ☀ Fly calibration rocket as soon as possible after start of mission



# NASA SUGGESTIONS FROM SSI VALIDATION WORKSHOP

- ☀ Have vents towards anti-Sun side and out past any radiator plate
- ☀ Provide means to warm up optics and detectors during flight
- ☀ Fly a cold trap / cold plate near sensitive optics
- ☀ Fly a TQCM (Thermoelectric Quartz Crystal Microbalance) to monitor contamination deposition rate real-time during flight
  - ☀ Important for early orbit commissioning and any propulsion use



# FUTURE WORK

- ☀ Understand difference in Zr and Al filter behavior
- ☀ Continue to refine EVE and SEM degradation models
- ☀ Collaboration with other instruments / missions



# RESOURCES

- ☀ CXRO-Henke
- ☀ HOCS-Palik
- ☀ NIST: X-Ray Form Factor, Attenuation and Scattering Tables (FFAST)
- ☀ NIST: Photon Cross Section Database (XCOM)