

Future Observations of Earth's Radiation budget and the science they enable

Maria Hakuba 2022 Sun Climate Symposium, May 19, 2022



Outline

- Libera science overview
- Libera's sub-band shortwave measurement
- Space Ball concept overview
- Space Ball feasibility simulations

Like mother (CERES), like daughter (Libera)



Libera, NASA's first Earth Venture Continuity Mission

Overarching Science Goals



OG1: Provide seamless continuity of the Clouds and the Earth's Radiant Energy System (CERES) ERB Climate data record (CDR).

 Measurement of TOT, SW and LW with same characteristics as CERES to prevent gap in ERB Climate data record.

OG2: Advance the development of a self-contained, innovative & affordable observing system.

• Wide field-of-view camera for Scene ID and split-SW ADM development.

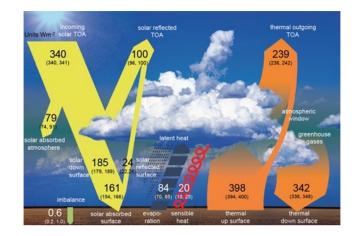
OG3: Provide new and enhanced capabilities that support extending ERB science goals.

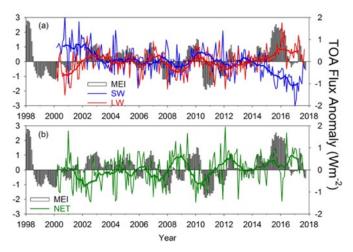
• Additional split-SW channel to derive shortwave near-IR and visible irradiance.

Benefits of continuity

Science objective 1: Use extended CDR to identify and quantify processes responsible for ERB variability on various times scales.

- What will the unprecedented >1.5°C future bring?
- Trend vs. internal variability
- Time-evolving temperature patterns impact global mean ERB
- Constraining climate models, forcing and feedback





ADMs & camera application

OG2: Development of a self-contained, innovative & affordable observing system

Demonstrate feasibility of separating Libera from complex imagers so to fly on SmallSats

Science objective 2:

- Explore utility of scene identification from a small and cost-effective camera. (Sebastian Schmidt)
- Develop angular distribution models (ADM) to facilitate shortwave near-IR and visible radianceto-flux conversion. (Jake Gristey)

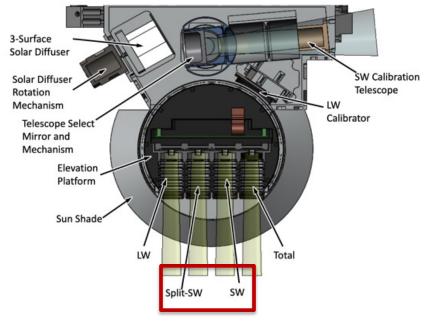


Monochromatic (555 nm) wide field of view (WFOV) camera provides images at ~1 km pixel resolution.

Libera's shortwave sub-band measurement

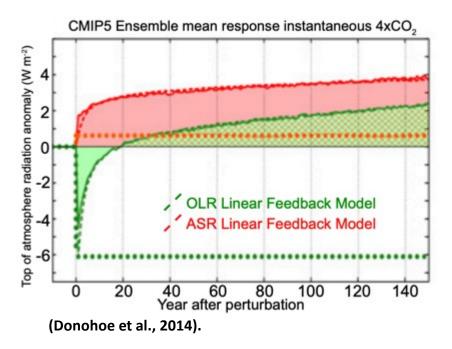
OG3: Provide new and enhanced capabilities that support extending ERB science goals.

- Libera's fourth channel measures VIS at the same accuracy as the total SW radiance (0.17%)
- Retrieval of VIS, NIR, SW irradiance at TOA and surface:
 - VIS at 0.3-0.7 μm
 - SW at 0.3-5 μm
 - Conversion to VIS irradiance using WFOV camera radiances at 555nm (J. Gristey, S. Schmidt)
 - NIR = SW VIS
- Goals:
 - NIR & VIS signatures of processes that control the absorption of solar radiation & SW climate feedbacks.
 - Better understand the hemispheric symmetry of planetary albedo.
 - Quasi-spectral model evaluation to reveal processrelated biases

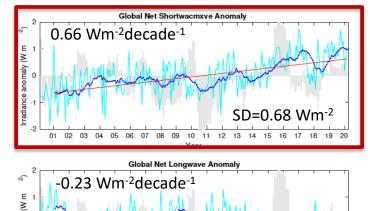


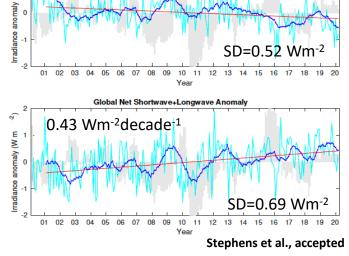
Libera Instrument Details

Motivation: Absorption of solar radiation warms our planet

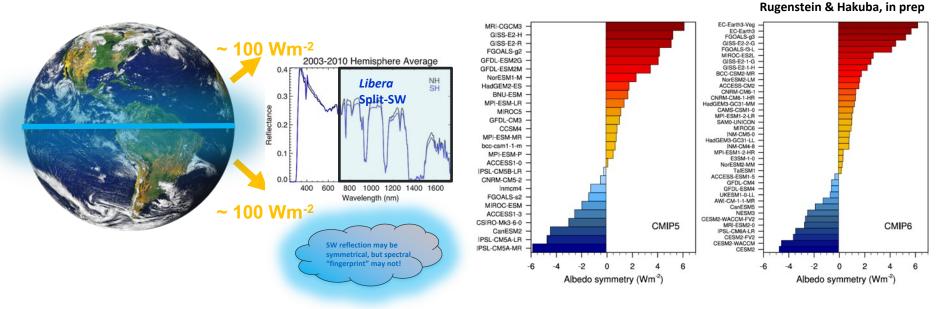


Climate model simulations under different future scenarios suggest global warming on decadal to centennial time scales is largely sustained by shortwave absorption (positive climate feedbacks).





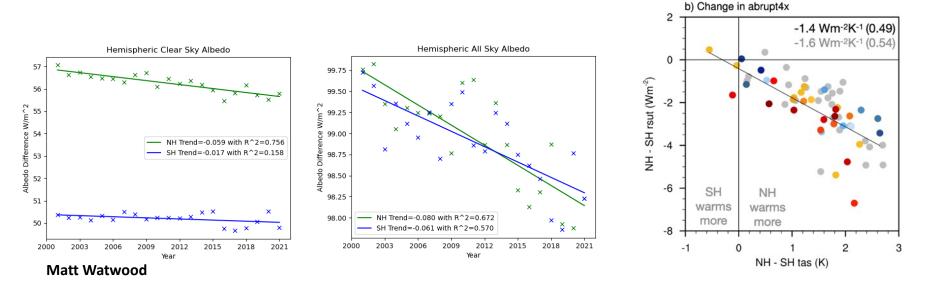
Motivation: Hemispheric albedo symmetry



- Average and decreasing trends in albedo are symmetric across hemispheres
- Surface albedo asymmetry compensated by cloud asymmetry over extra-tropics, SH storm-tracks are 10% cloudier
- 30% of models are symmetric within $\pm\,1$ Wm^-2
- 70 (80)% of models are symmetric within ± 4 Wm⁻². (An asymmetry of 4 Wm⁻² in the hemispheric energy budget would imply 3 times more heat transport across equator!)

Evolution of hemispheric albedo symmetry

- CMIP models suggest hemispheric albedo is evolving to a-symmetry in response to asymmetric surface warming
- Observations indicate a positive trend in NH-SH clear-sky albedo of -0.4 Wm⁻²/decade

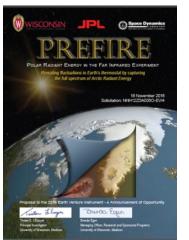


Rugenstein & Hakuba, in prep

Motivation: The future is spectral

- Spectral signatures of the TOA fluxes reveal insights into processes that shape the ERB and changes to it
- Measurements are on the way:

Spectral OLR – 2022 (L'Ecuyer)



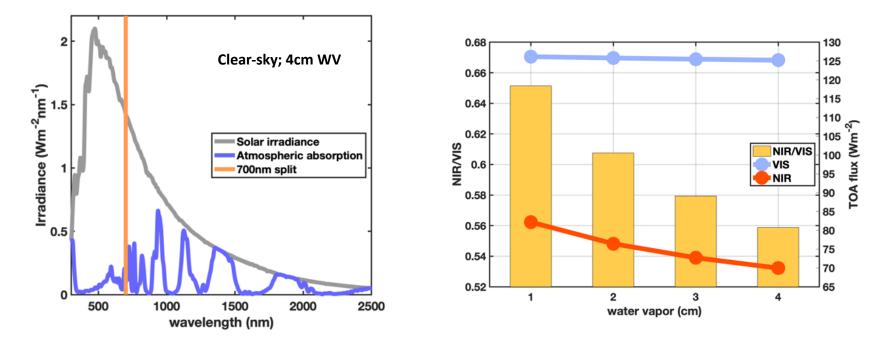
CLARREO-PF - 2022 spectral SW (Shea)



X-element linear array of absolute radiometers to enhance spatial & spectral resolution (Coddington)

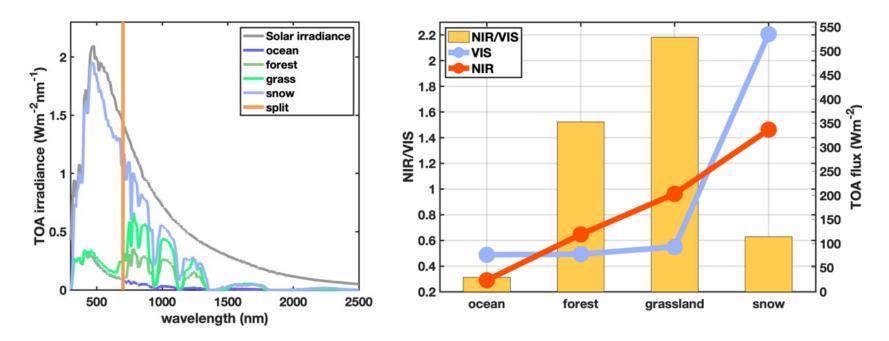


Spectral nature of shortwave radiation (MODTRAN calc. by Bruce Kindel)



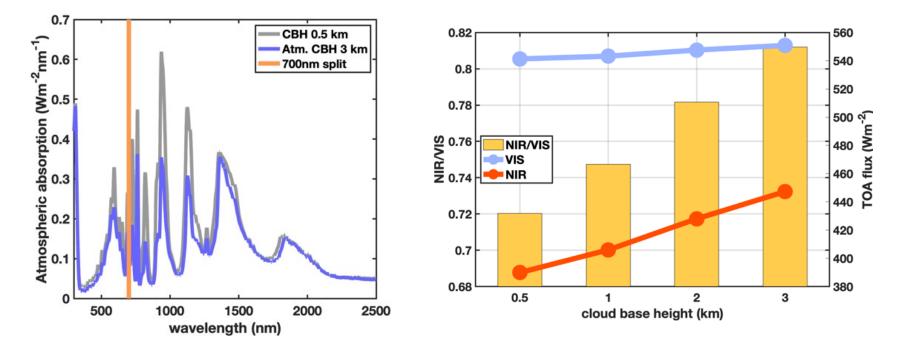
Water vapor absorption acts mostly in the NIR; ~6 Wm⁻²/cm NIR absorption C-C 7%/K ~ 0.2cm/K \longrightarrow ~ 1 Wm⁻² NIR absorption

Spectral nature of shortwave radiation



Different surface types associated with very different NIR/VIS ratios (~NDVI); Libera's NIR/VIS may help track land cover change and impact on ERB

Spectral nature of shortwave radiation



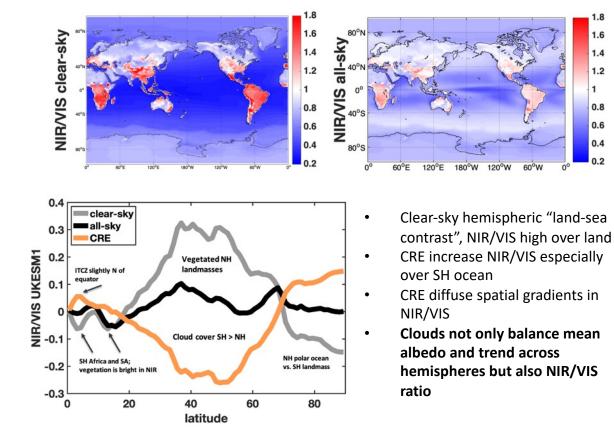
Cloud SW absorption is ~15 % of SW CRE (~47 Wm⁻²) and subject to change; example shows impact of cloud base height, shifting the ratio in favor of NIR **Change in cloud height alone does impact albedo!?**

Outlook: Libera SW sub-band measurement

- Three idealized process examples indicate NIR is particularly sensitive to changes in climate relevant properties
- Set expectations for NIR/VIS of different processes and scene types
 - Calculations and ML
 - Set of ancillary information
- Improved knowledge of SW absorption processes and change therein can help understand bigger picture mysteries... or unveil more mystery...

UKESM1 is symmetric in albedo and NIR/VIS (20y pi control)

| All-sky | Glo | NH | SH |
|---------------|--------------------|------------|------------|
| TOT SW | 99.3 | 99.1 | 99.6 |
| NIR | 44.0 | 44.2 | 43.7 |
| VIS | 55.0 | 55.8 | 55.4 |
| NIR/VIS | 0.79 | 0.80 | 0.78 |
| | | | |
| Clear- sky | Glo | NH | SH |
| | Glo 54.4 | NH 57.4 | SH 51.4 |
| sky | | | |
| sky TOT SW | 54.4 | 57.4 | 51.4 |



1.8

1.6

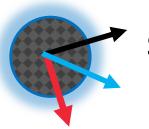
1.4

1.2

0.8

0.6 0.4

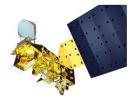
0.2



Space Balls

Direct measurement of the net radiative flux (EEI) at TOA

EEI = Global mean Net radiative flux at TOA



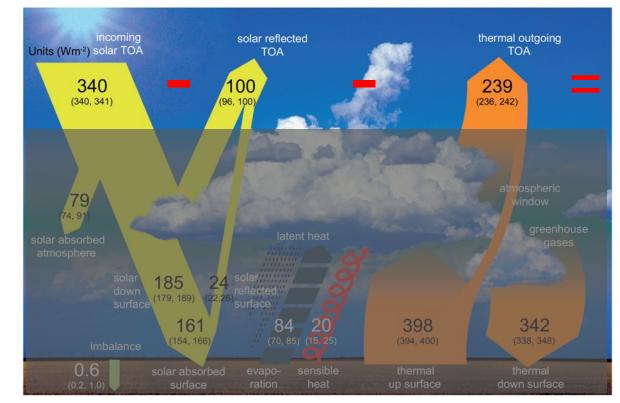
CERES EBAF & Libera

Clouds and Earth's Radiant Energy System

SORCE & TSIS

Solar Radiation and Climate Experiment

Total and Spectral Solar Irradiance Sensor

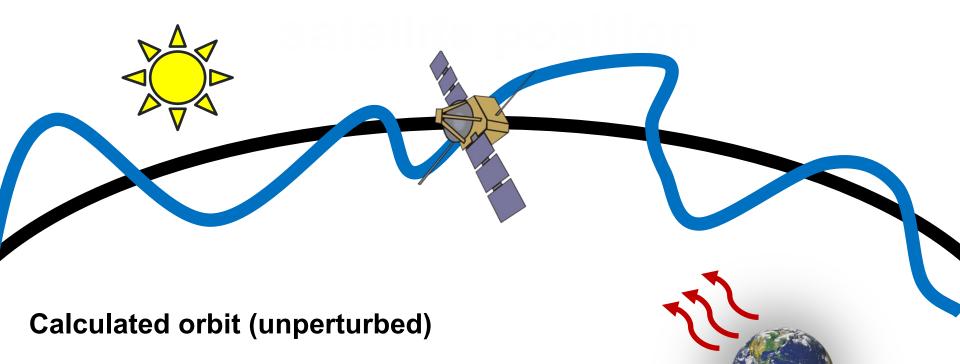


Updated from IPCC AR5 / Wild et al. 2013, 2015 Climate Dynamics

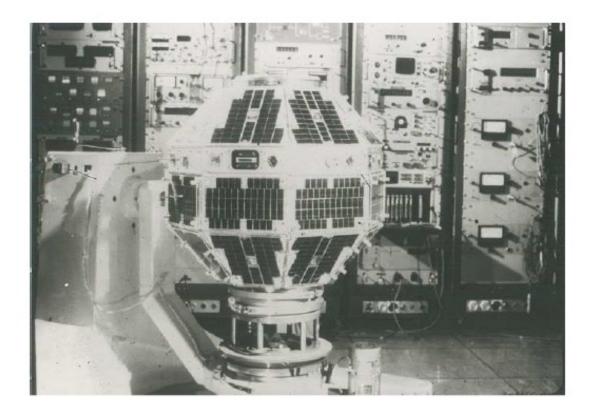
~0.7 Wm⁻² (Johnson et al., 2016)

Spoiler: this number does not come from TOA radiation data, but from heat uptake assessments!

Radiation pressure affects satellite position



Real orbit (perturbed)

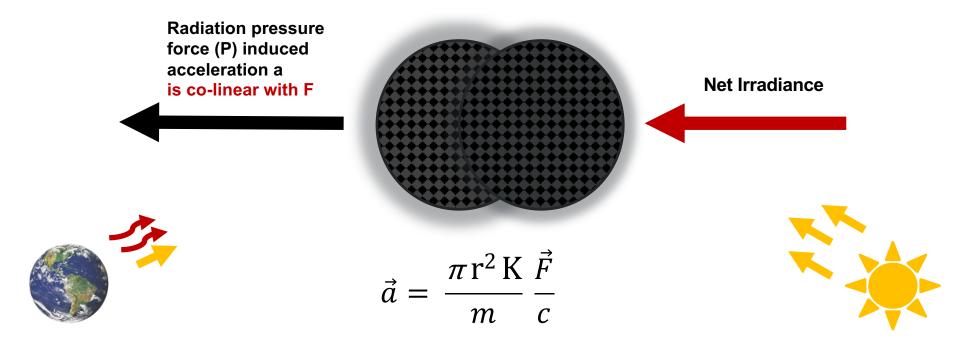


CASTOR D5B Satellite

CNES, 1975

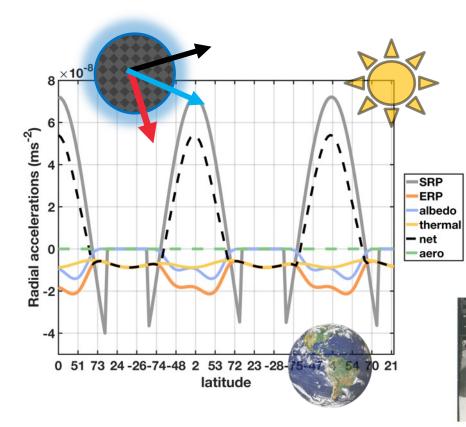
radius: 0.8 m, mass: 76 kg

Simple relationship between net irradiance and induced acceleration if space craft is spherical and perfectly absorbing across Earth and Solar radiation



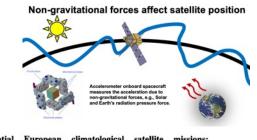
Space Balls Update (~2-yr JPL R&TD project)

Earth's Energy Imbalance via radiation pressure accelerations



Is a high-accuracy measurement of Earth's Energy Imbalance (EEI) feasible via radiation pressure accelerations experienced in orbit? Objectives:

- 1. Build SB simulation environment using mission design software Monte
- 2. Enhance fidelity of force and shape models
- 3. Study measurement errors due to S/C and orbit characteristics, and confounding forces
- 4. Explore different sampling strategies



Potential European climatological satellite missions: SEOCS and BIRAMIS

G. DUCHOSSOIS Department of Application Programmes, Directorate of Planning and Future Programmes, Eu Space Agency, 8 rue Mario Nikis, 7738, Paris Cedex 15, France

(Received 19 December 1979)

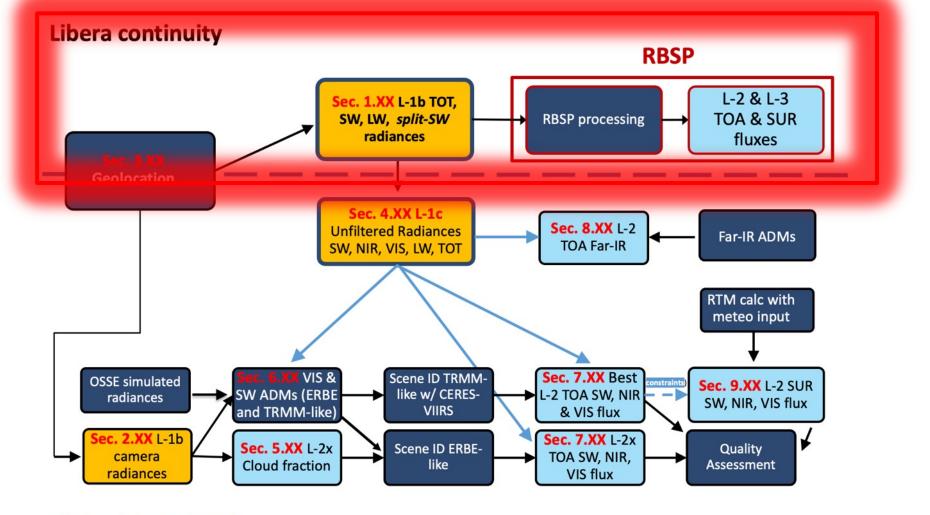


CASTOR D5B Satellite

Theoretical Comparison Between Radiometric and Radiation

Pressure Measurements for Determination of the Earth's Radiation Budget T.H. Vonder Haar and E.A. Smith

Thank you!

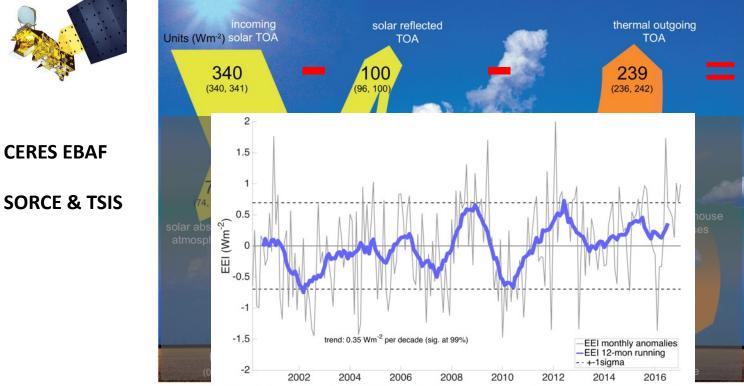


Libera beyond L-1b

EEI = radiative forcings – radiative feedbacks



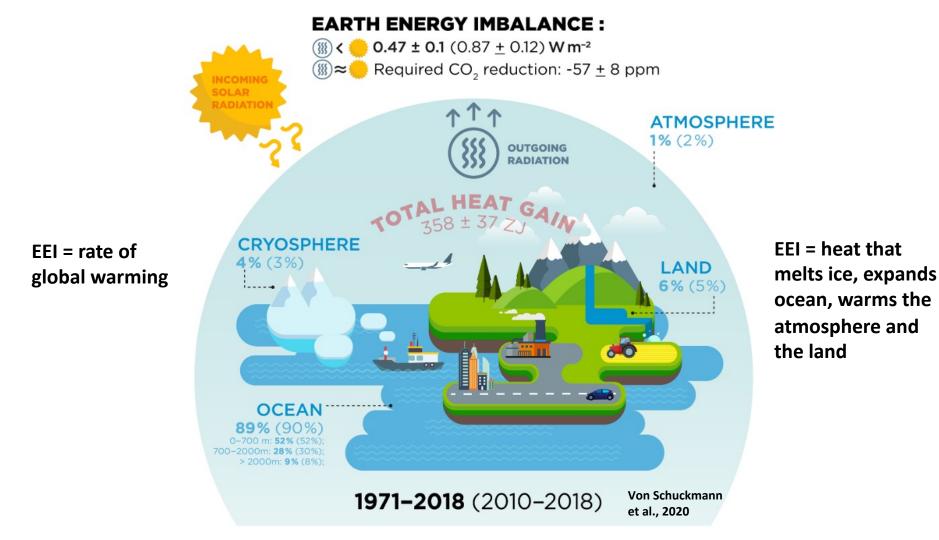
CERES EBAF



~0.7 Wm-2

global long-term mean 2005-2015

EEI variability reflects natural and anthropogenic radiative effects on multiple time scales



Geophysical Research Letters*

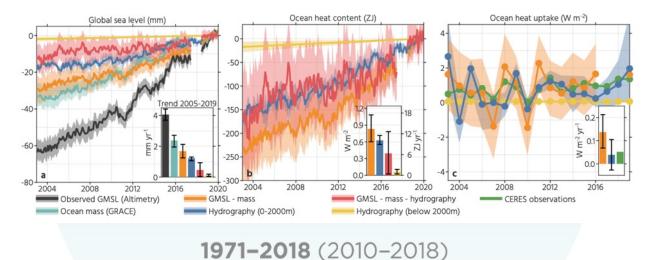
Research Letter

RE

Earth's Energy Imbalance From the Ocean Perspective (2005–2019)

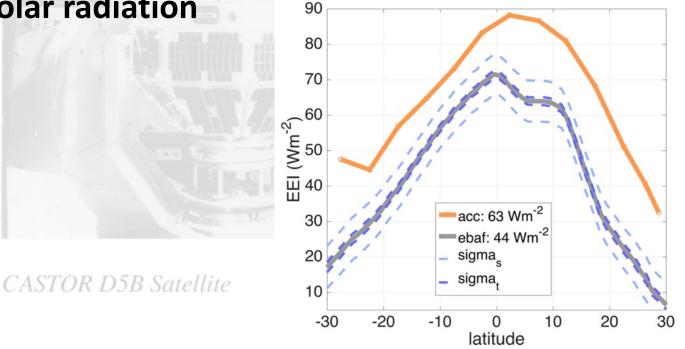
M. Z. Hakuba 🔀, T. Frederikse, F. W. Landerer

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CACTUS analysis delivered

- Zonal and seasonal variability of net radiative flux
- Longwave cloud effects and land-sea contrasts
- Incoming solar radiation

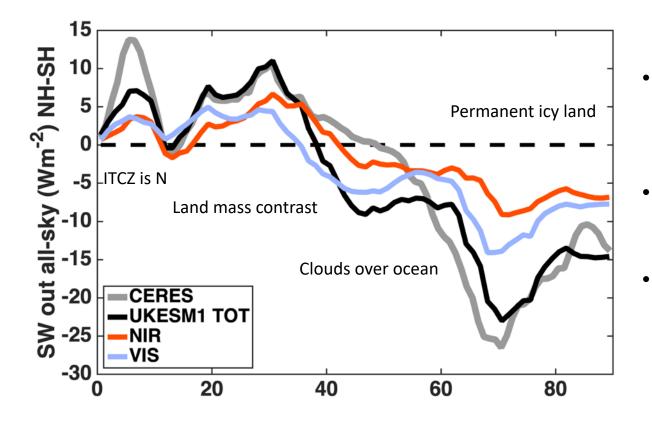


UKESM1 SW, NIR, and VIS fluxes

- UK Earth system modeling project
- Successor to HadGEM2-ES
- "The complexity of coupling between the ocean, land, and atmosphere physical climate and biogeochemical cycles in UKESM1 is unprecedented for an Earth system model." (Sellar et al., 2019)
- Limits of spectral intervals (wavelengths in m.)
 - Band Lower limit Upper limit
 - 1 2.0000000E-07 3.2000000E-07
 - 2 3.2000000E-07 5.05000000E-07
 - 3 5.05000000E-07 6.90000000E-07
 - 4 6.9000000E-07 1.19000000E-06
 - 5 1.19000000E-06 2.38000000E-06
 - 6 2.38000000E-06 1.00000000E-05
- Integrate bands 2-3 for visible, and bands 4-5 for near-IR

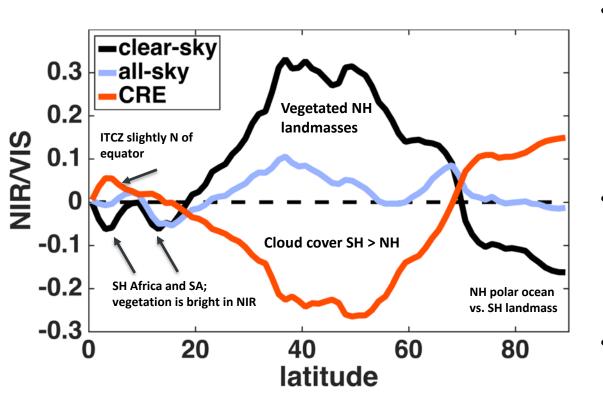
Model simulation: Alejandro Bodas-Salcedo

NH-SH differences per latitude



- NH is mostly brighter over 0-40 degree, but darker poleward
- Model agrees OK with CERES
- NIR & VIS zonal variability looks similar to total SW

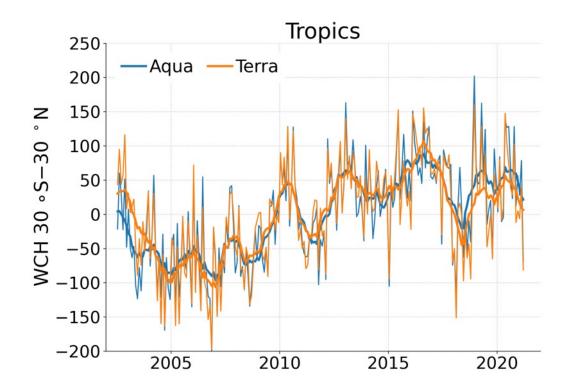
NH-SH differences NIR/VIS ratio per latitude



- Positive values: NIR/VIS ratio is larger on NH than on SH; especially true under clear-sky between 20-70 deg. (note: locally, SH Africa and SA have largest NIR/VIS)
- CRE balance the hemispheric NIR/VIS ratio zonally & mirror the Clearsky effects.
- But NIR/VIS ratio remains slightly larger on NH under all-sky conditions.

Tropical clouds are getting higher

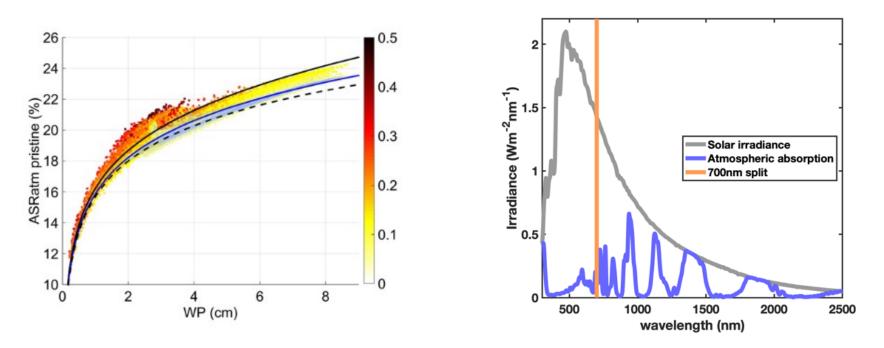
Richardson et al., 2022



Cloud height trend sensitivity: 330±60 m K⁻¹

This would imply doubling of NIR above-cloud absorption per K warming.

Spectral nature of shortwave radiation



- Water vapor saturation yields non-linear increase
- Underlying surface impacts relationship by several Wm⁻²