EVE Rocket Analysis

A New Method to Analyze EVE Rocket Data

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EVE Diodes: (ESP, MEGS-P) 0.25 s Flight Data



Dark Corrected Channel Data



ESP and MEGS-P data. "Outage" at \sim 300 s is a stray-light check in ESP using a FS filter, it is ignored in the analysis. Soft X-ray and Lyman- α channels don't show atmospheric





- Isolate data close to apogee ($\sim \pm 40$ s).
- Average apogee data.
- Use NRL-MSIS to calculate residual atmospheric absorption.
- Calculate Air Mass Zero (AM0) response.
- Compare Rocket-EVE AM0 measurement with SDO-EVE measurement.





- Parabolic fit to WSMR radar altitude vs. time.
- Match radar altitude data to EVE data.
- Fit data vs. Altitude and to extrapolate to air mass zero (AM0).
- Compare Rocket-EVE AM0 measurement with SDO-EVE measurement.
- This method also provides the residual atmospheric absorption:

 $ABS_{ch} = DN_{ch}(AM0) - DN_{ch}(apogee)$

- I developed this using data from the 36.353 (9 Sept. 2021) flight.
- I have working radar data readers for all flights.
- I have not looked carefully at the data except for the 36.353, and 36.389 flights.
- I have problems reading data for a couple of early flights (36.258, 36.240).





The radar data file often arrives in different formats, but for this analysis, I am only using three fields:

TIME: The time after launch in seconds.

ALT: The altitude above sea level in meters.

ALT VEL: The vertical component of the velocity in m/s. I use this to determine when the rocket is in ballistic flight, It is sometimes not included, in which case I calculate it.



Radar Plots and fits







Fit Function



When the signal (DN_{ch}) shows atmospheric absorption during the flight I fit:

$$DN_{ch} = DN_{AM0} \times (1 - e^{-\frac{(z-z_0)}{H}})$$

 DN_{AM0} is the extrapolated AM0 signal

- z is the altitude in km
- z_0 is the extinction altitude in km
- ${\it H}\,$ is the scale height in km

If there is not enough extinction to fit the $1-e^{-\frac{Z}{H}}$ then a simple linear fit to the profile is used. In these cases extinction is not significant, so we are directly measuring the AMO.





• I used the SciPy fit function, which defaults to the Levenberg–Marquardt method. After trying to do the fit to altitude in meters (read directly from the radar data file) the fitter was very sensitive to the initial parameters, I changed to using km for the altitude, and the fitting performed much better.

def EXP_fit_fn(z, am0, h, z0):

return $am0^*(1.0 - numpy.exp(-(z-z0)/h))$

tparams, $tdcv = scipy.optimize.curve_fit(EXP_fit_fn, z/1000.0, dn, [p_0])$

- I fit the whole mission data as well as the up and down-legs individually.
- \cdot I use the diagonal of the covariance matrix (tdcv) to calculate the fit error.



ESP 19 nm Channel





ESP 19 nm channel fits. I fitted the up-leg, down-leg, and all data separately, and used the difference between the fits to calculate the uncertainty.



ESP 26 nm Channel



36353 ESP Signal vs. Altitude (26 nm channel)



ESP 26 nm channel fits.



ESP 30 nm Channel





ESP 30 nm channel fits.



ESP 36 nm Channel



36353 ESP Signal vs. Altitude (19 nm channel) 20.0 • Up Down ····· Fit up ····· Fit down ····· Fit all 15.0 AM0 Extrapolation: 12.00 ± 0.12 DN € 12.5 Jark (. <u>____</u> 0.0 150 200 250 300 350 Altitude (km)

ESP 36 nm channel fits.



ESP Quad Sum (Linear Fit)



36353 ESP Signal vs. Altitude (Quad Sum channel)



ESP Quad Sum linear fits



MEGS-P (Lyman- α) (Linear Fit)





ESP Quad Sum linear fits





Channel	AM0 Counts	Fit Errors ¹	Apogee Counts	Absorption
	(DN/.25s)	(%)	(DN/.25s)	(%)
E19	606.83	0.05	603.05	0.62
E26	137.99	0.06	136.52	1.06
E30	310.59	0.06	306.24	1.40
E36	11.94	0.83	11.58	3.07
QS ²	231.53	0.05		
MEGS-P ²	6.32	1.22		

¹These are calculated as the 1σ errors of the fit covariance AMO element ²There is not enough absorption to make an air-mass fit, so I use linear fit





Channel	AM0 Counts	Fit Errors ³	Apogee Counts	Absorption
	(DN/.25s)	(%)	(DN/.25s)	(%)
E19	1038.46	0.03	1022.98	1.49
E26	277.48	0.04	268.54	3.22
E30	478.22	0.03	458.41	4.14
E36	20.19	0.71	19.46	3.62
QS^4	775.84	0.08		
MEGS-P ²	12.48	0.49		

 $^3 These$ are calculated as the 1σ errors of the fit covariance AMO element $^4 There$ is not enough absorption to make an air-mass fit, so I use linear fit







Apogee Absorption vs. F_{10.7} Apogee Absorption vs. Wavelength

Data from flights: 36.389, 36.353, 36.336, 36.318, 36.290, 36.286, 36.275, 36.240 Still working on: 36.258, 36.240. 36.300 was lost





Advantages:

- The method seems to work well.
- It uses all [good] data.
- Averages over solar activity changes, and local air mass variations.
- Doesn't rely on external models e.g. NRL-MSIS to calculate AM0 results.
- Determined absorption could be used to 'correct' NRL-MSIS.

Disadvantages:

- \cdot Need to wait for WSMR Radar data (But NRL-MSIS needs average $F_{10.7}).$
- Probably can't use for all MEGS data.





I am not sure how easy it will be to apply to MEGS-A/B data at $40 \times$ slower cadence but Don has some promising results.

However, this method can be compared to the MEGS results, or used to scale the NRL-MSIS model.

The next step in this work are:

- Proceduralize the fit and plot routines.
 - Unfortunately, every flight has some slight differences in file formats, etc. so I need to hard-code some specifics for each flight.
 - Check data for all flights.
- \cdot Understand the $F_{10.7}$ results.
 - Calculate absorption to a standard altitude, not Apogee?
- Finish paper detailing this work.



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For deeper data access and posting you need a login, please let me know if you want access Andrew Jones and I will try and arrange it.





Backup Plots



Radar Detail







Apogee Altitude





