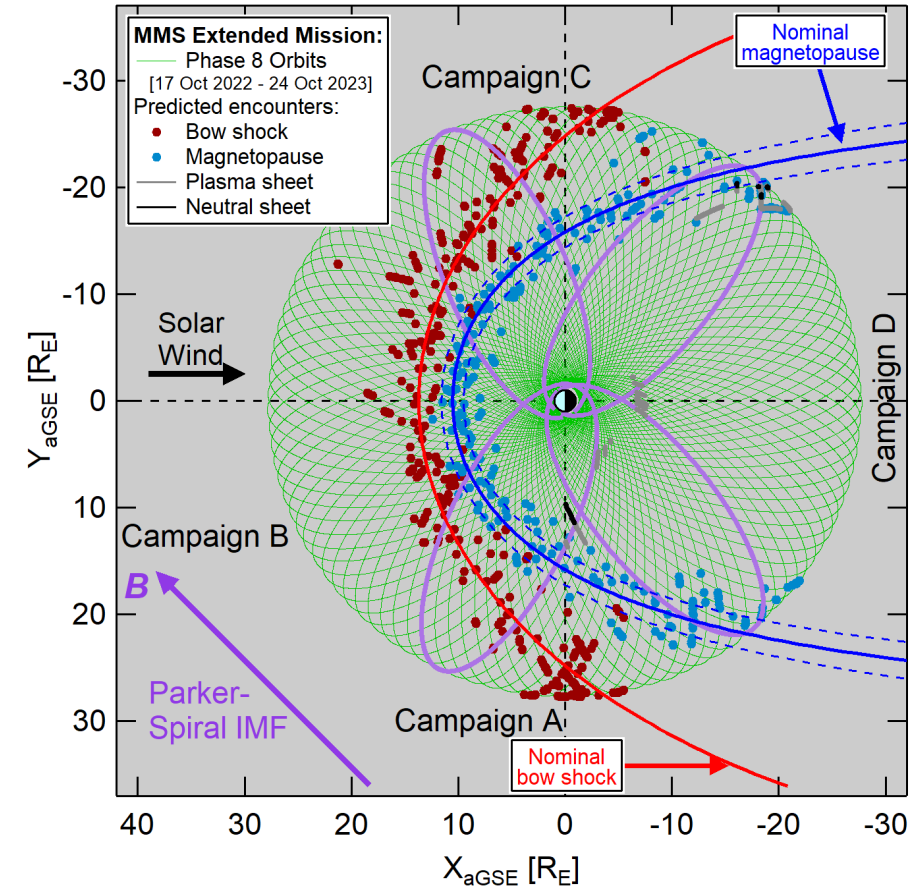


MMS-ARASE Conjunction observations during Phase 8d

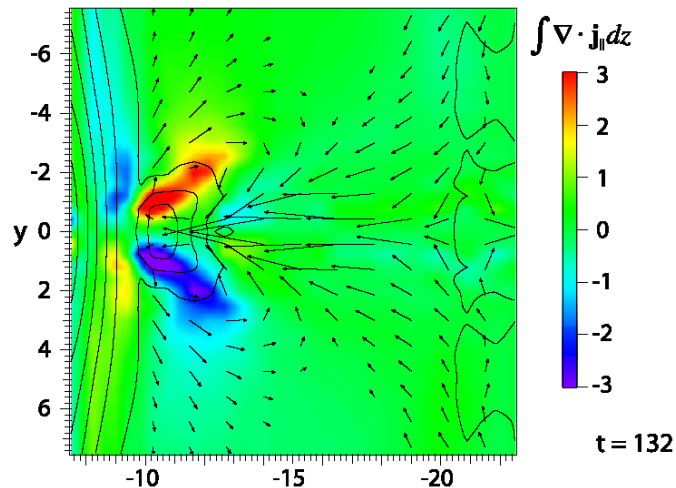
R. Nakamura, Y. Miyoshi, N. Kitamura

2023 Summer-Autumn

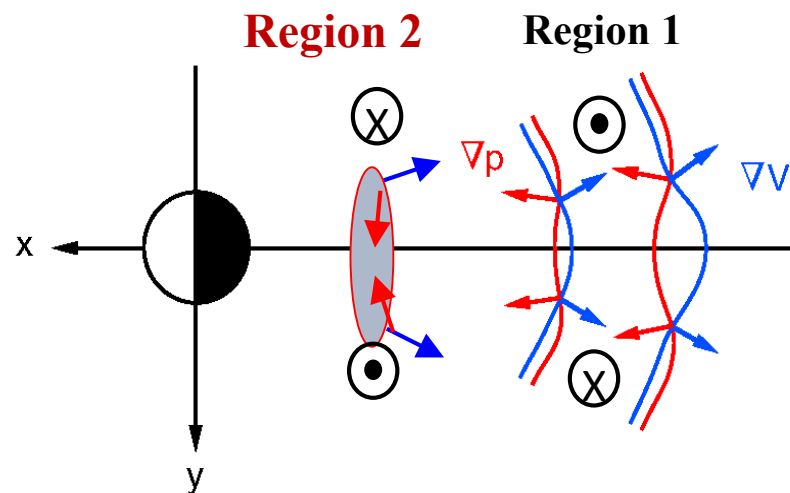
- MMS current sheet crossings take place in inner magnetosphere during Phase 8D (apogee in lobe)
- >12 inner magnetospheric (inside GEO) conjunction with ARASE spacecraft
- Most of them are >3 missions conjunctions with MMS, ARASE, THEMIS or Cluster, and some with foot point in North-American sector
- Opportunities for night-side sciences: flow braking, FAC, injection, instabilities/waves around the transition region, wave-particle interaction related to pulsating aurora



Parallel current and Pressure gradient in Flow Braking region



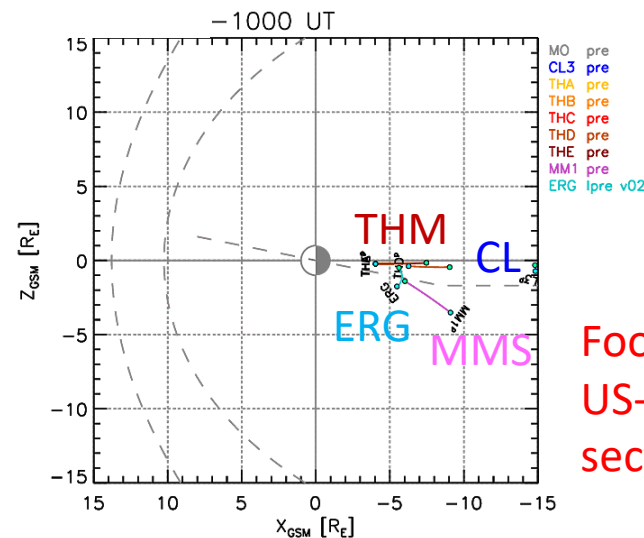
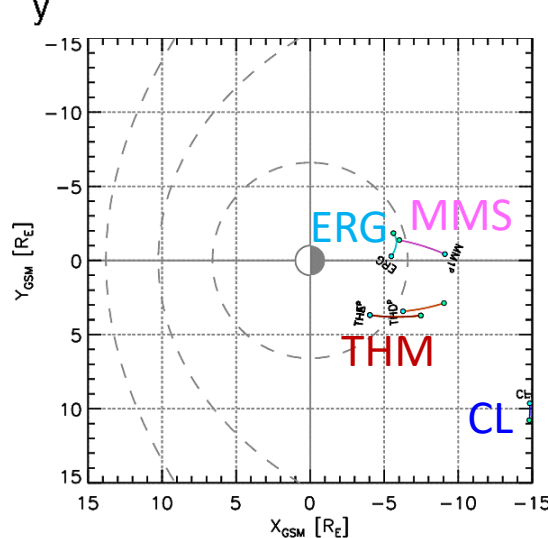
(adapted from Birn et al., 2011)



$$\int_{equator}^{ionosphere} \frac{j_{\parallel}}{B} ds = -\frac{\mathbf{B}}{B_{eq}^2} \cdot \nabla p_{eq} \times \nabla V$$

$$V = \int_{eq}^{io} \frac{ds}{B} \text{ flux tube volume}$$

- Multi-point gradient measurements
- FAC

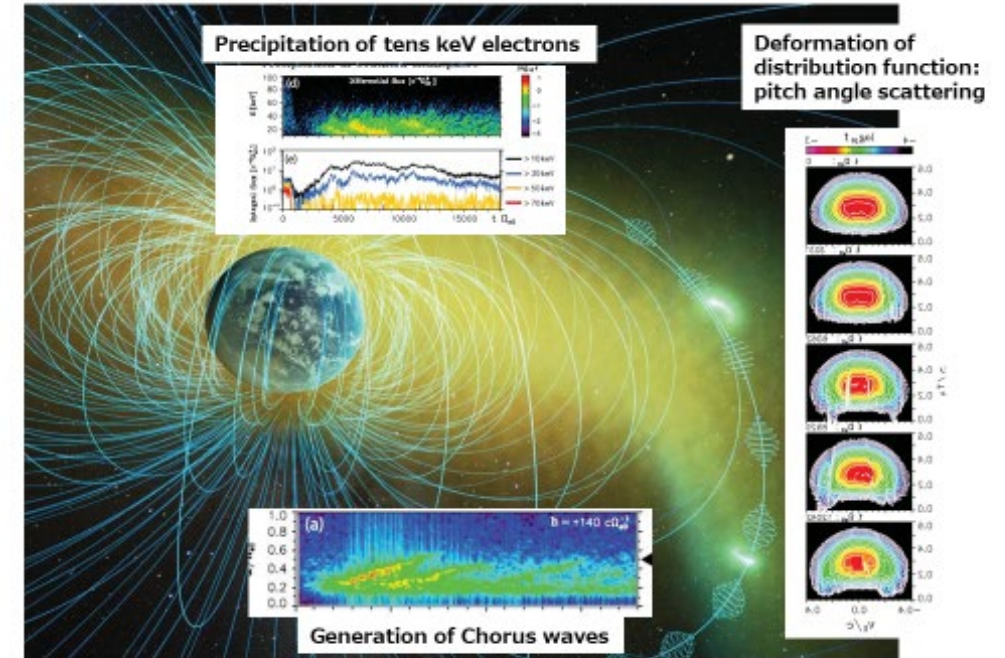


Footprint:
US-Canada
sector

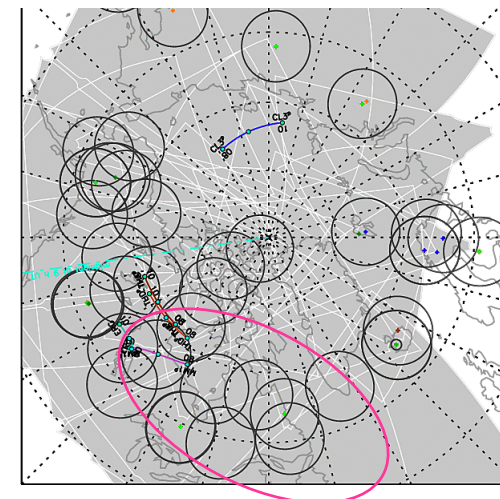
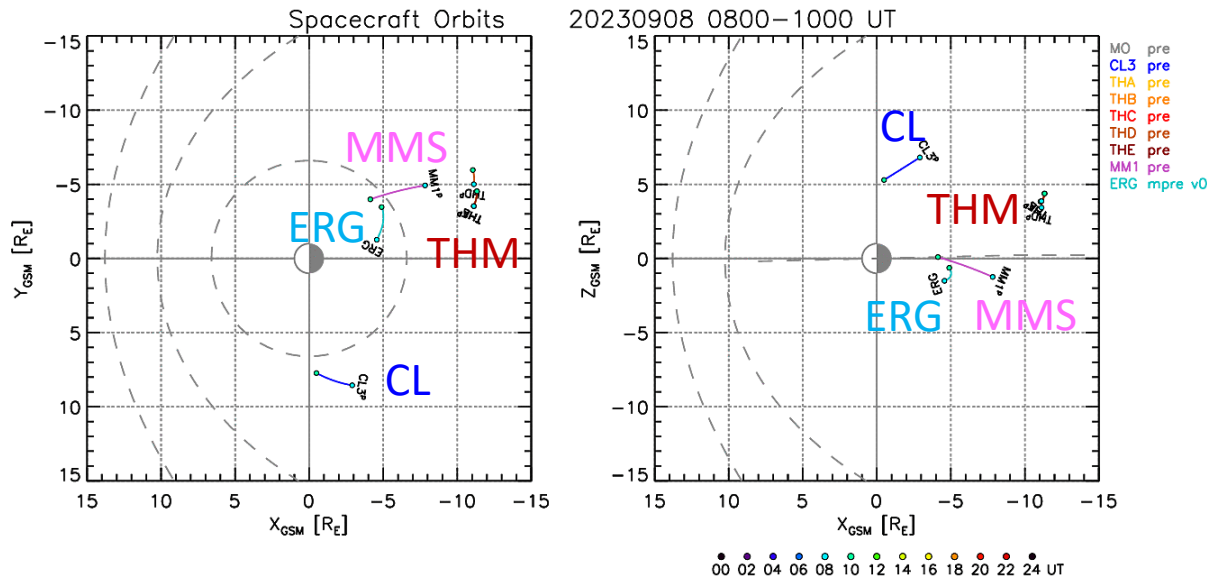
Pitch angle scattering by wave-particle interactions

Pulsating aurora

- Monitoring all ingredients: injection, density irregularities, wave-particle interaction, auroral precipitation



Hikishima+, GRL, 2010



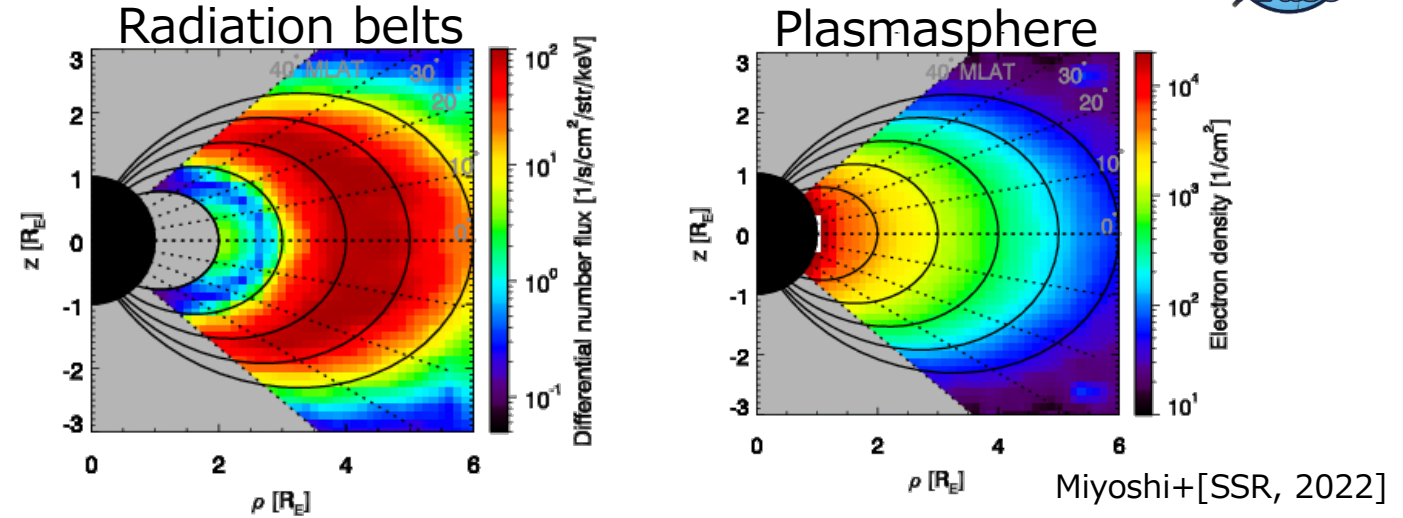
Foot-point in northern american auroral stations

Arase/ERG



Orbit:

- Apogee: 32,000 km
- Perigee: 460 km
- Orbital period: 9.4 hours
- Inclination: 31 deg
- Spin period: 8 sec



- Particles:** Ion (w/ mass discrimination): 10 eV/q – 180 keV/q (3D distribution function)
Electrons: ~ 100 eV - ~ 5 MeV (3D distribution function)
- Fields/waves:** Electric fields: DC – 10 MHz
Magnetic fields: DC – 100 kHz
(DC-256 Hz: waveform, -64 kHz: wave burst)

WPIA (direct measurements of phase between electrons/ions and waves) has been operated.

Data & Analysis tool:

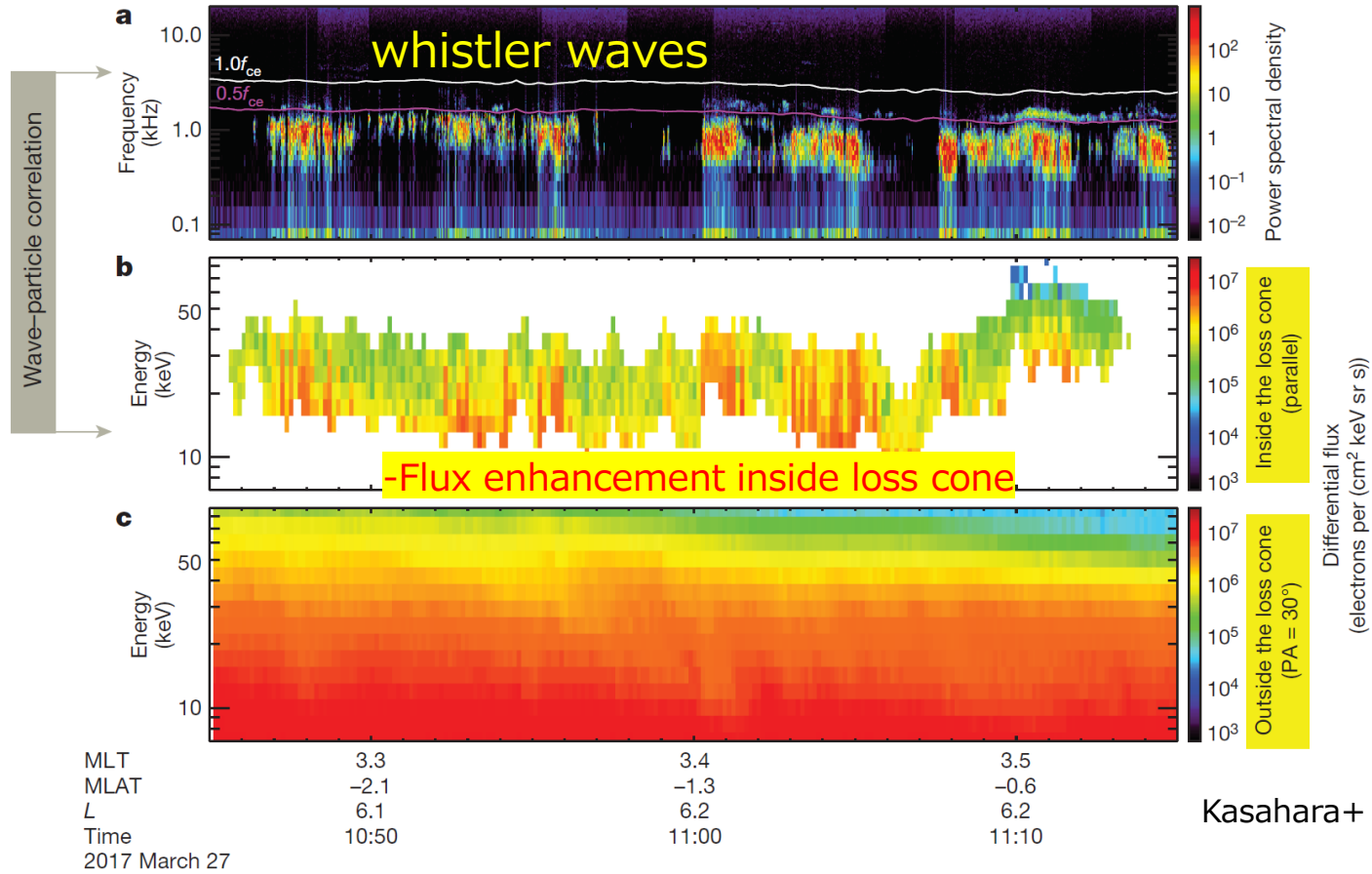
Data are available at ERG Science Center (<https://ergsc.isee.nagoya-u.ac.jp/index.shtml.ja>)
SPEDAS/PySPEDAS are available for the Arase data and ERG-ground based data (ASI, magnetometer, high-speed camera data etc).

Example of observations by Arase

VLF waves:

Electrons:
Inside loss cone

Electrons:
Outside loss cone

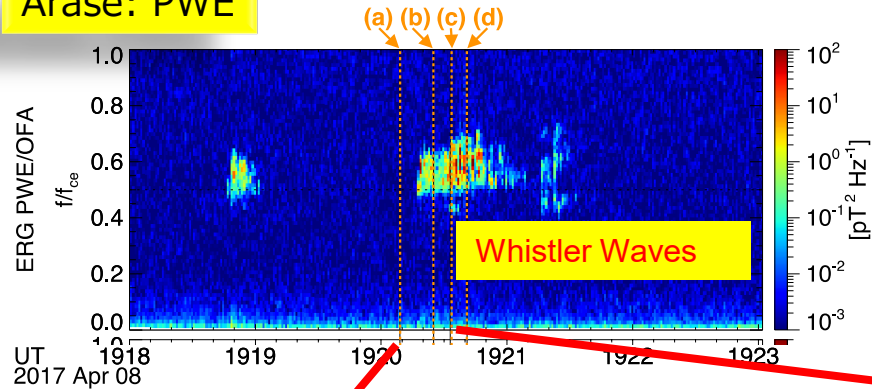


Kasahara+ [Nature, 2018]

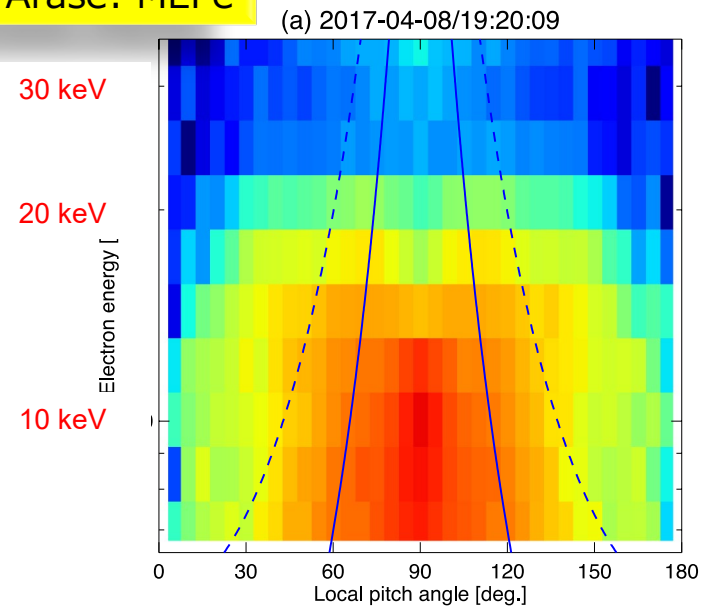
The Arase satellite can sometimes observe electrons inside loss cone ($\sim 100 \text{ eV} - 90 \text{ keV}$), and discuss how much electrons actually precipitate into the thermosphere through the pitch angle scattering.

Example of observations by Arase

Arase: PWE

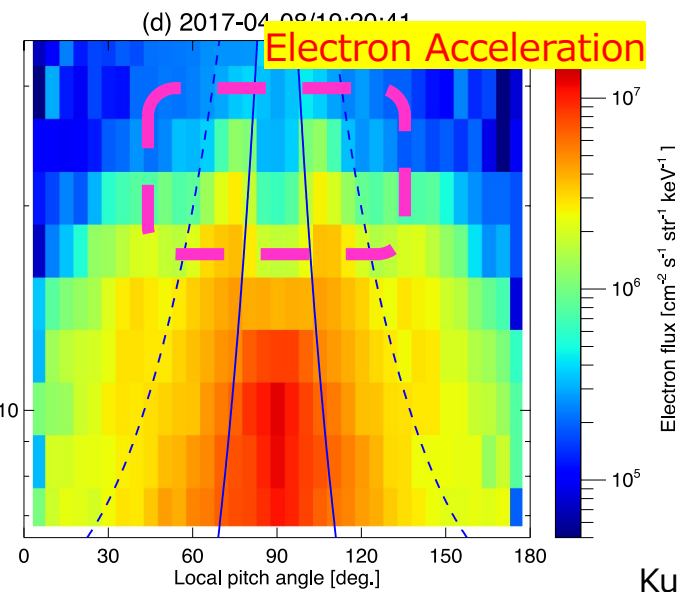


Arase: MEPe



(a) 2017-04-08/19:20:09

(40 seconds)

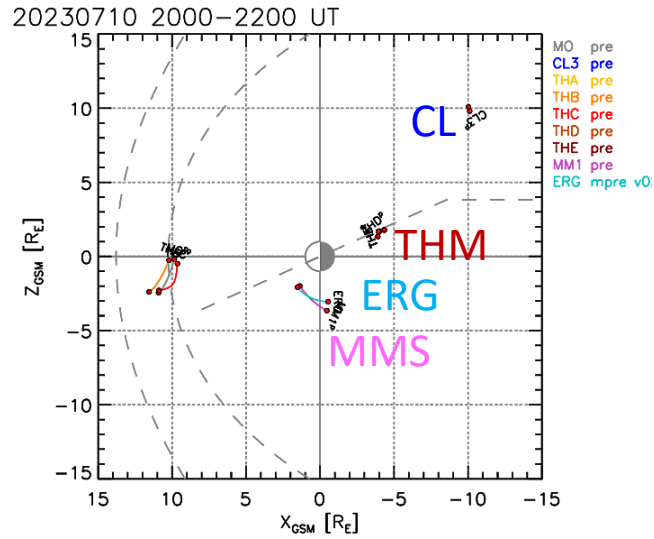
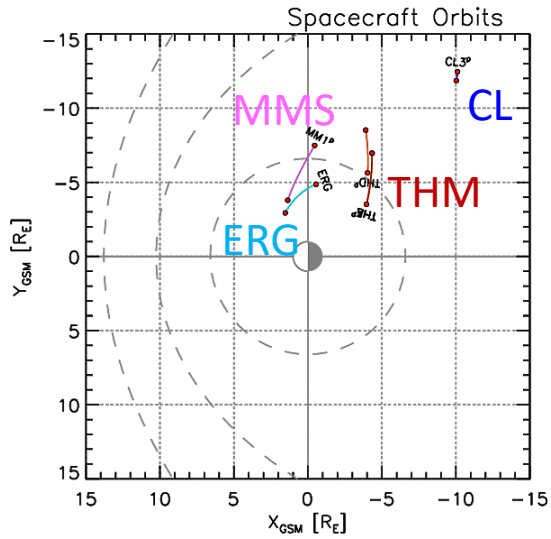


(d) 2017-04-08/19:20:41

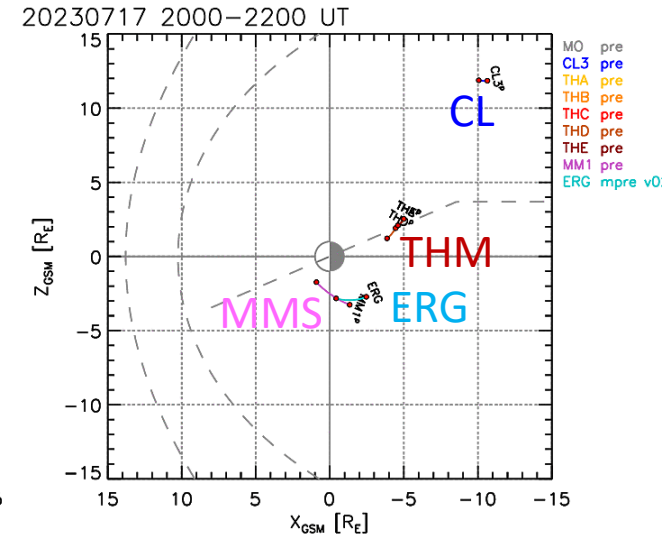
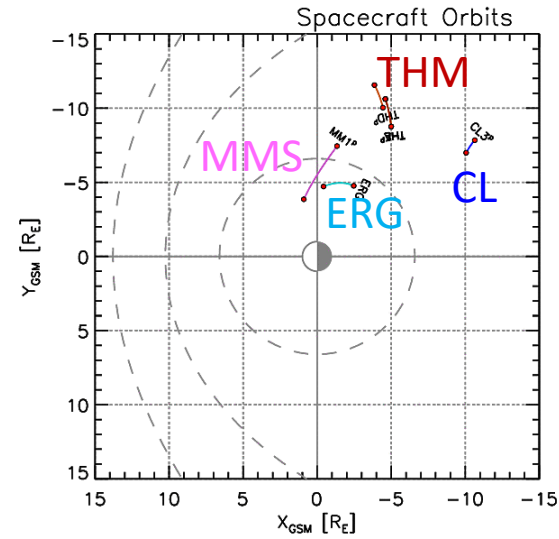
Kurita+ [JGR, 2018]

Arase observes 3-D distribution function every 8 sec (spin period) and found the deformation of the distribution function of electrons through chorus wave-particle interactions.

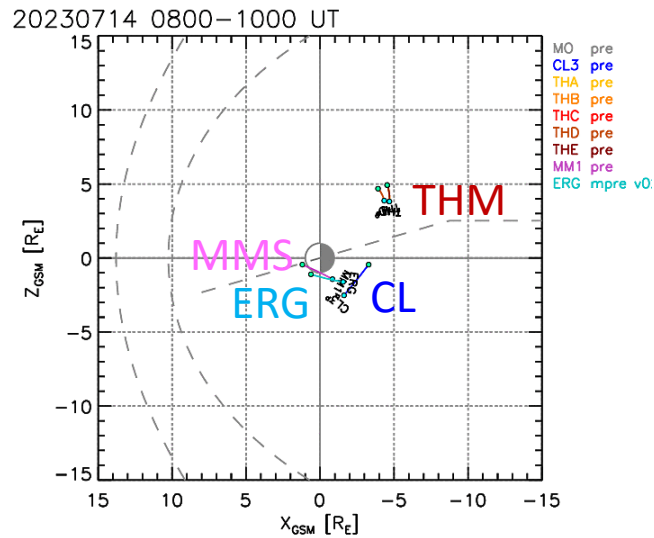
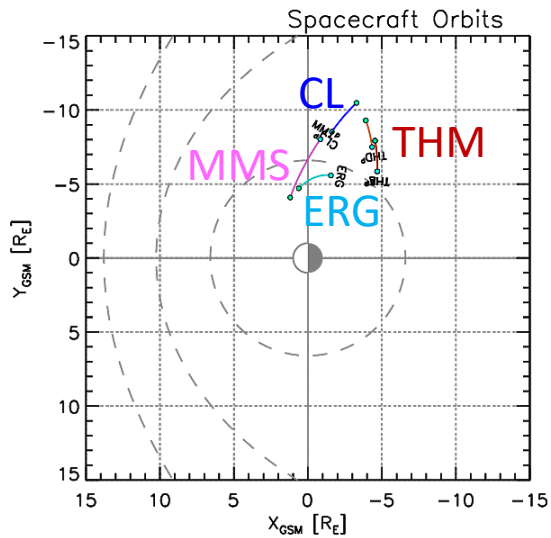
4 (Moderate) Conjunctions from Jul. 10 to Jul. 21



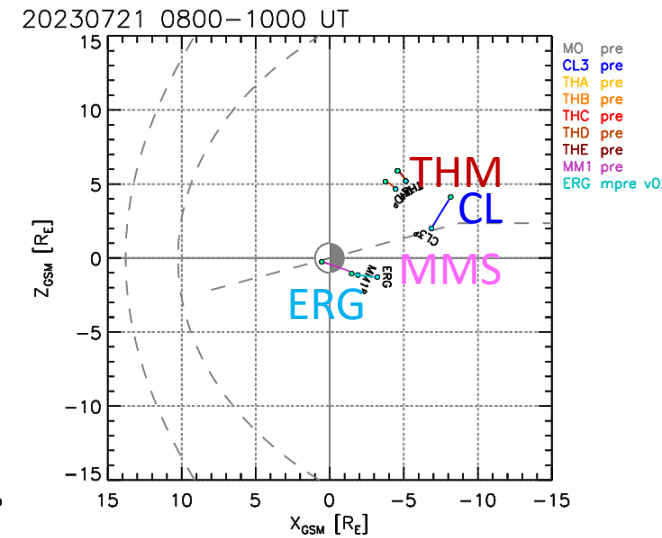
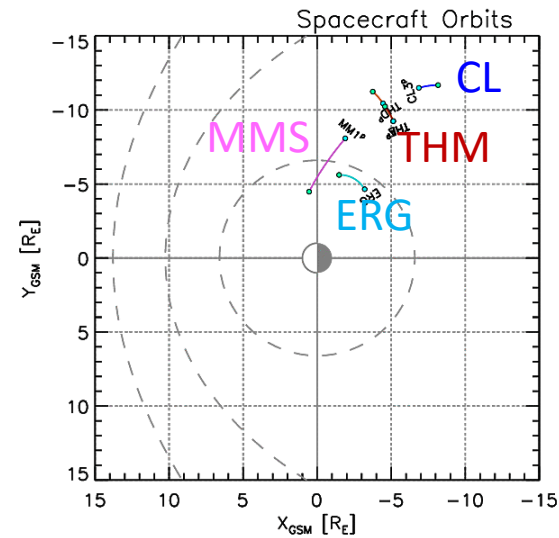
● 00 02 04 06 08 10 12 14 16 18 20 22 24 UT



● 00 02 04 06 08 10 12 14 16 18 20 22 24 UT

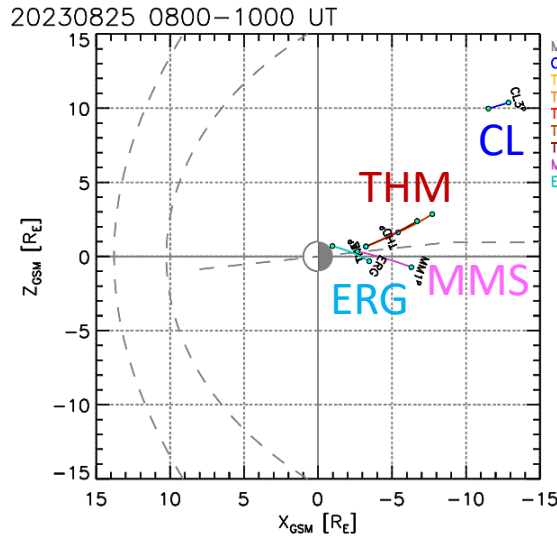
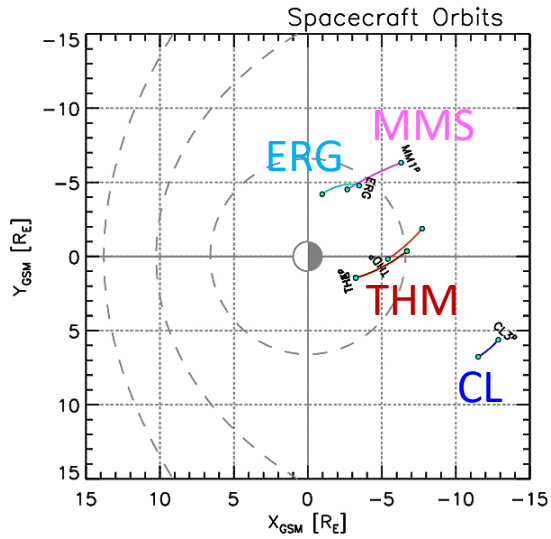


● 00 02 04 06 08 10 12 14 16 18 20 22 24 UT



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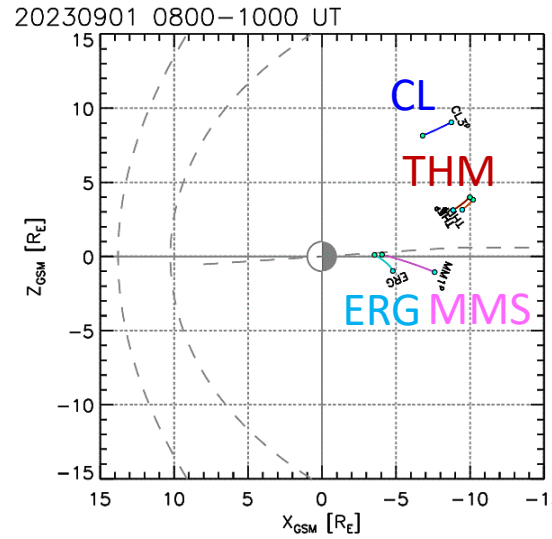
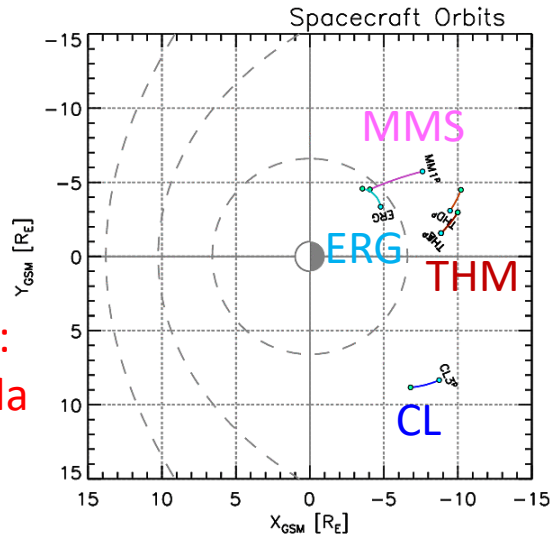
5 Conjunctions from Aug. 25 to Sep. 8



MO pre
CL3 pre
THA pre
THB pre
THC pre
THD pre
THE pre
MM1 pre
ERG lpre v02

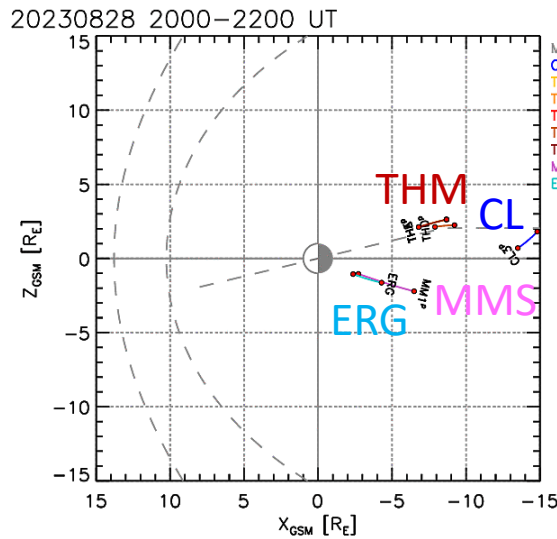
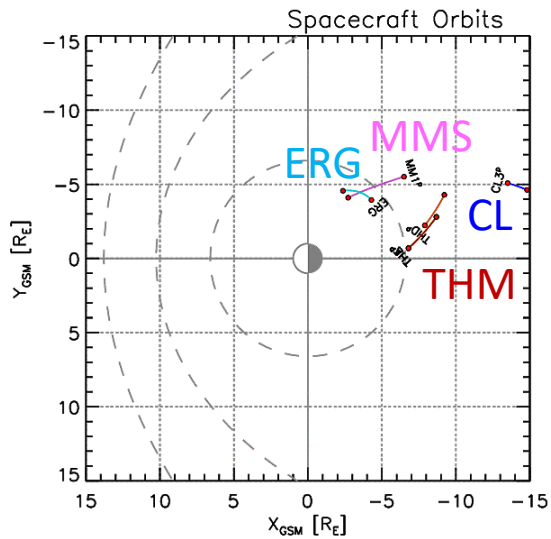
Footprint:
US-Canada
sector

● 00 02 04 06 08 10 12 14 16 18 20 22 24 UT



MO pre
CL3 pre
THA pre
THB pre
THC pre
THD pre
THE pre
MM1 pre
ERG lpre v02

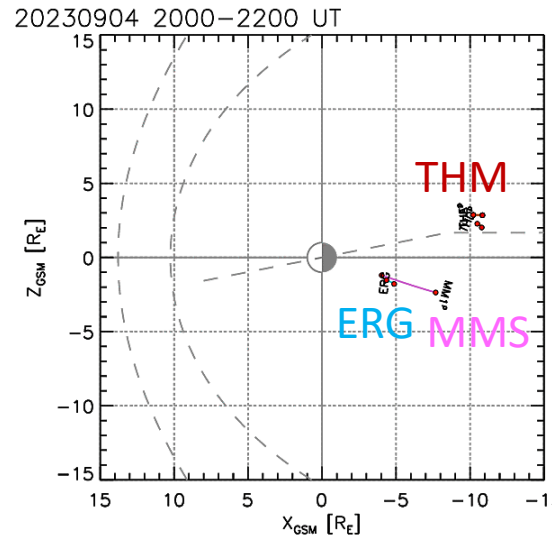
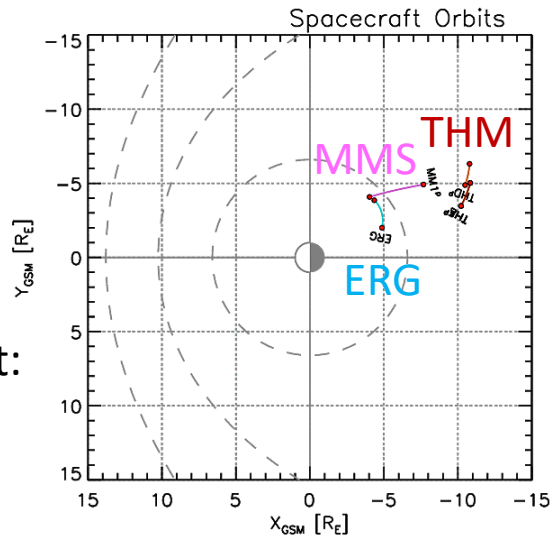
● 00 02 04 06 08 10 12 14 16 18 20 22 24 UT



MO pre
CL3 pre
THA pre
THB pre
THC pre
THD pre
THE pre
MM1 pre
ERG lpre v02

Footprint:
Russia
sector

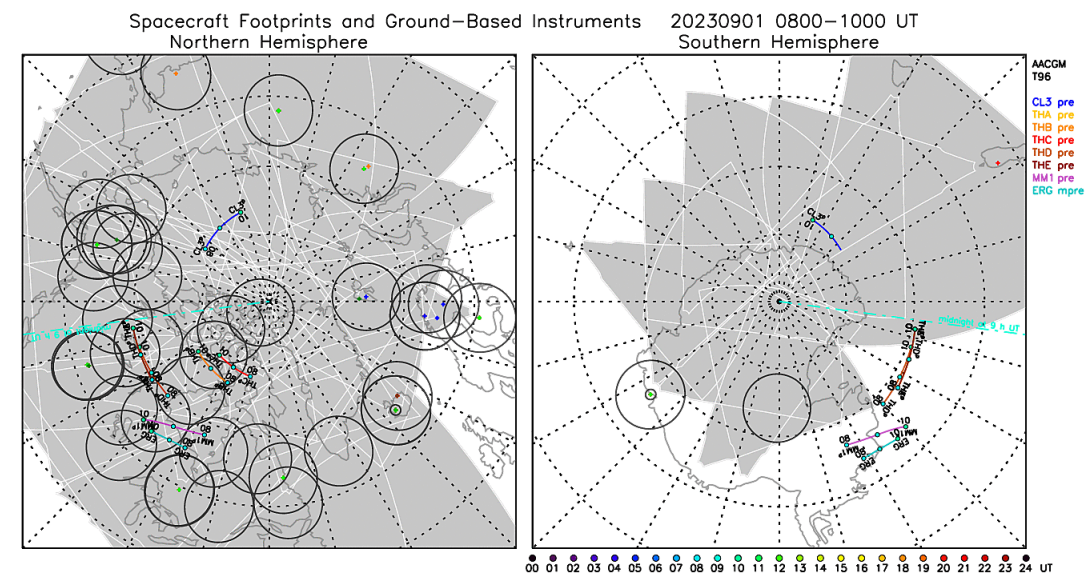
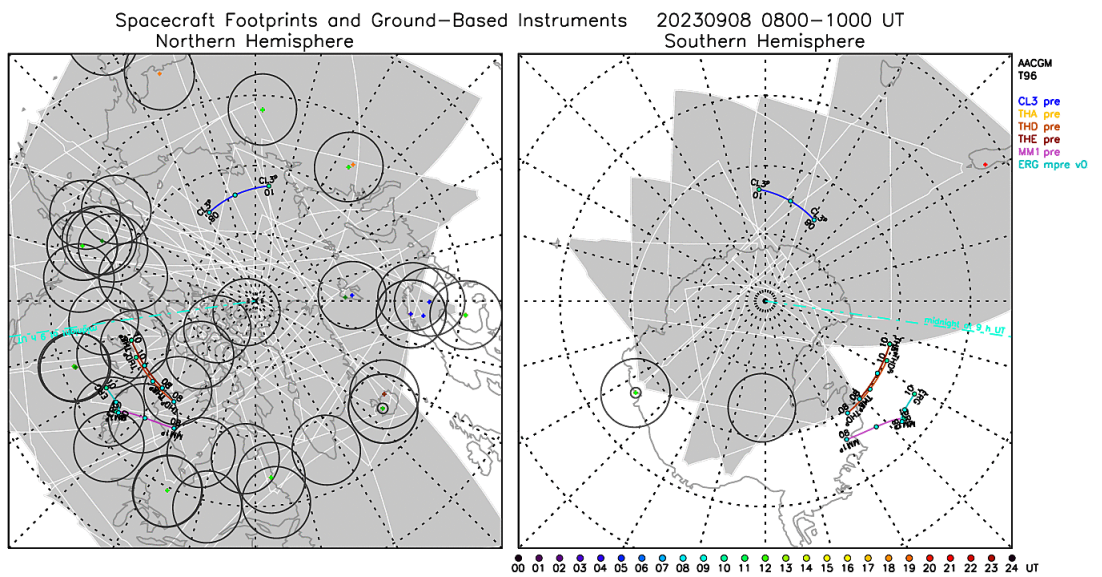
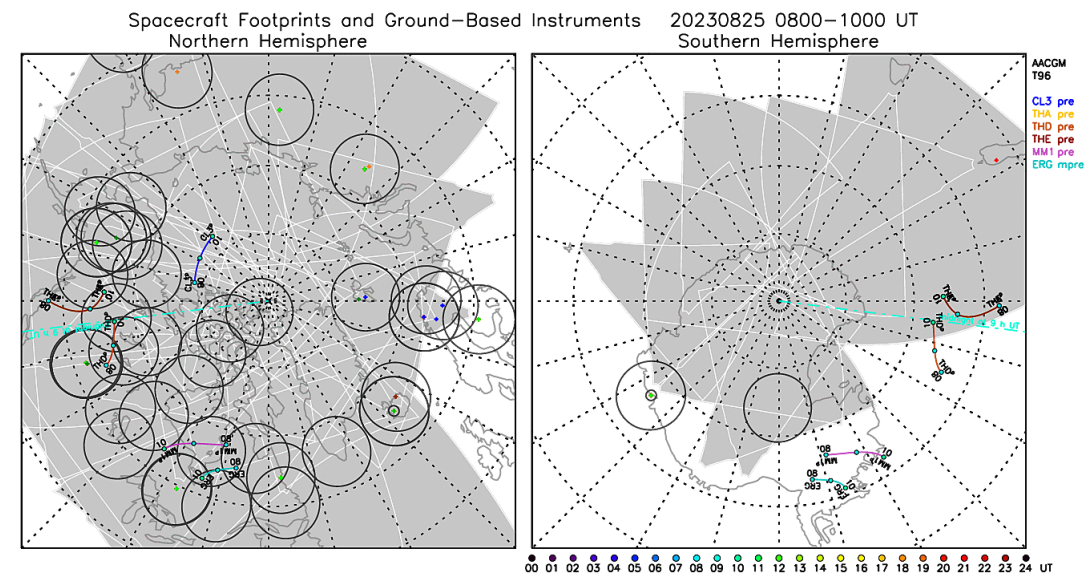
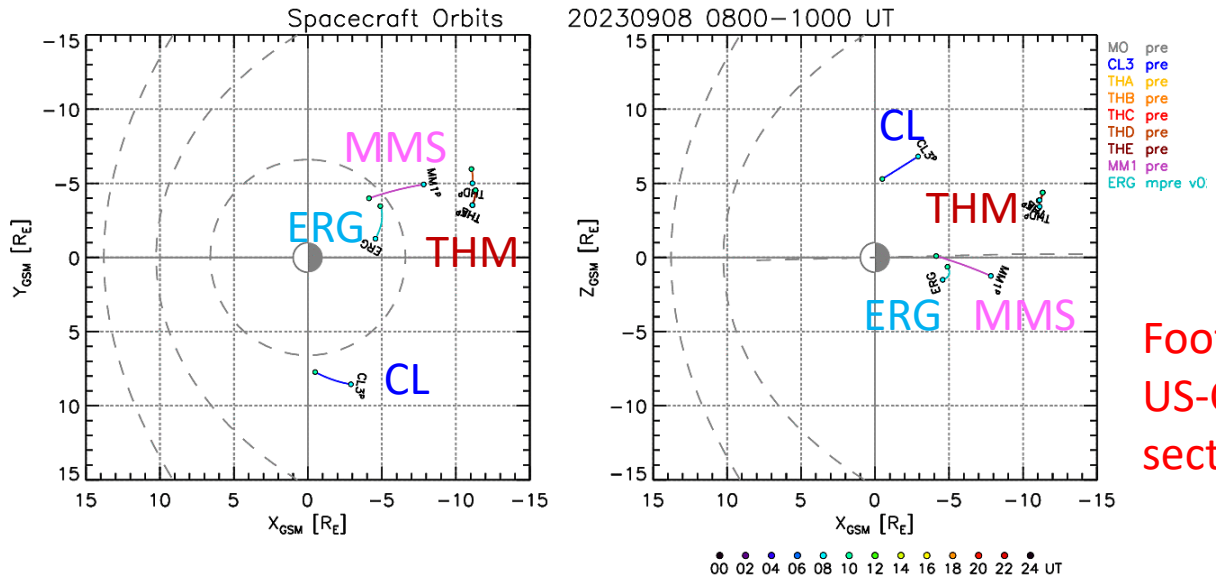
● 00 02 04 06 08 10 12 14 16 18 20 22 24 UT



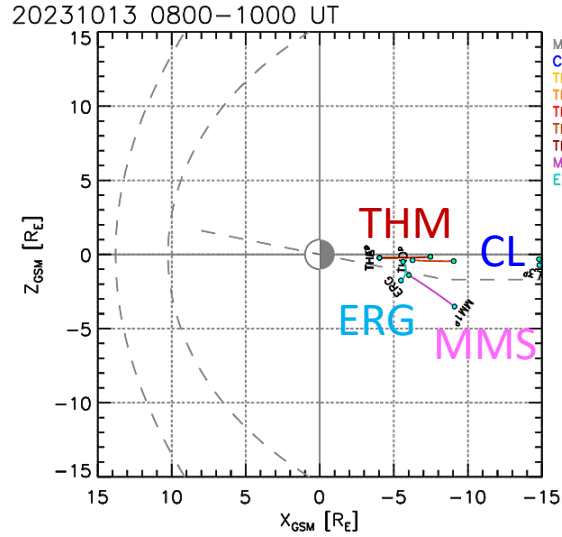
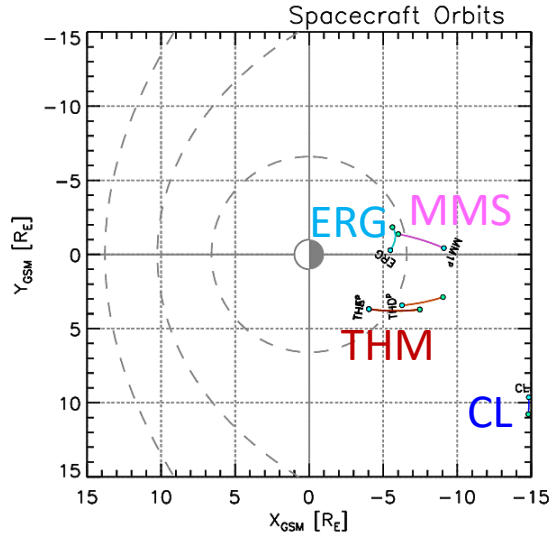
MO pre
CL3 pre
THA pre
THB pre
THC pre
THD pre
THE pre
MM1 pre
ERG mpre v0

● 00 02 04 06 08 10 12 14 16 18 20 22 24 UT

5 Conjunctions from Aug. 25 to Sep. 8



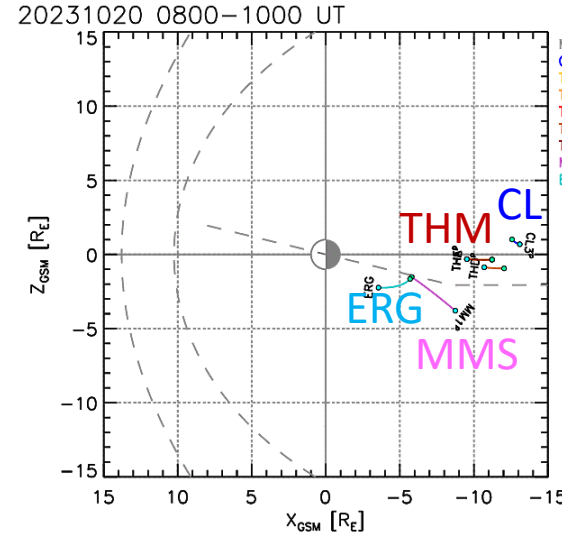
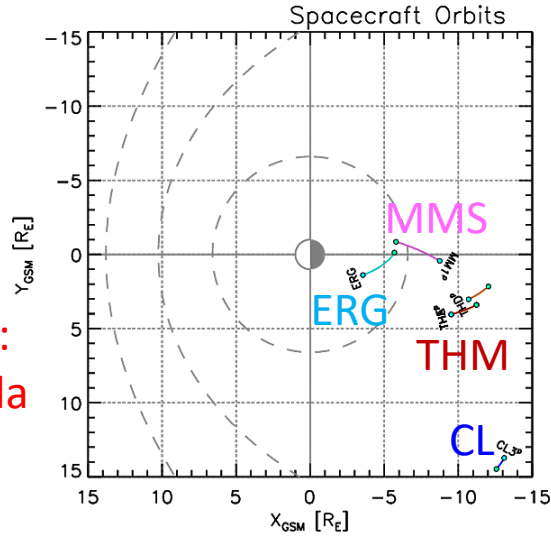
>3 Conjunctions from Oct. 13



MO pre
CL3 pre
THA pre
THB pre
THC pre
THD pre
THE pre
MM1 pre
ERG lpre v02

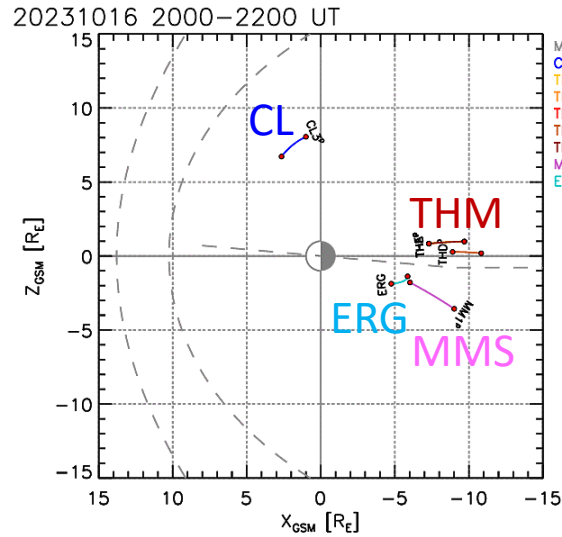
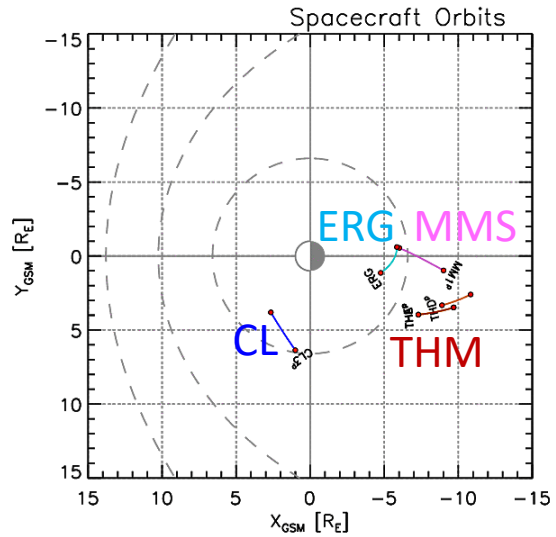
Footprint:
US-Canada
sector

00 02 04 06 08 10 12 14 16 18 20 22 24 UT



MO pre
CL3 pre
THA pre
THB pre
THC pre
THD pre
THE pre
MM1 pre
ERG lpre v02

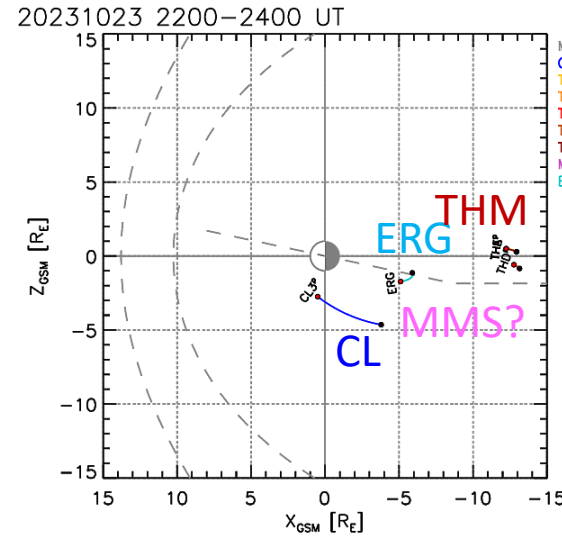
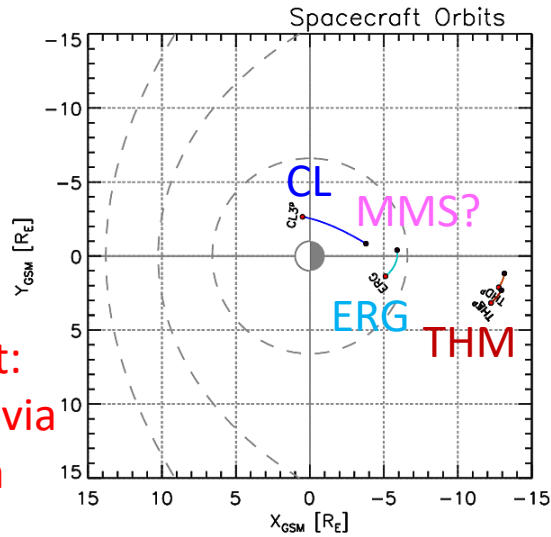
00 02 04 06 08 10 12 14 16 18 20 22 24 UT



MO pre
CL3 pre
THA pre
THB pre
THC pre
THD pre
THE pre
MM1 pre
ERG lpre v02

Footprint:
Scandinavia
or Syowa
sector

00 02 04 06 08 10 12 14 16 18 20 22 24 UT

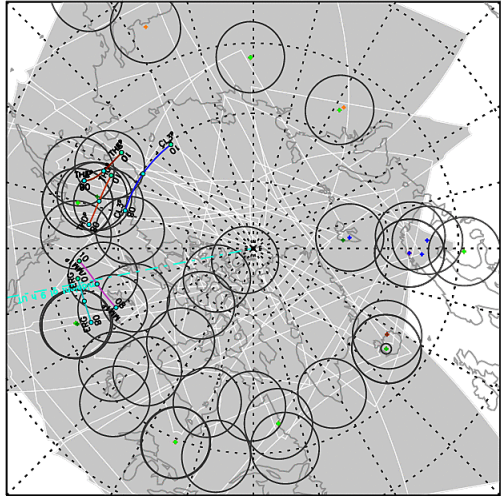


MO pre
CL3 pre
THA pre
THB pre
THC pre
THD pre
THE pre
MM1 pre
ERG lpre v02

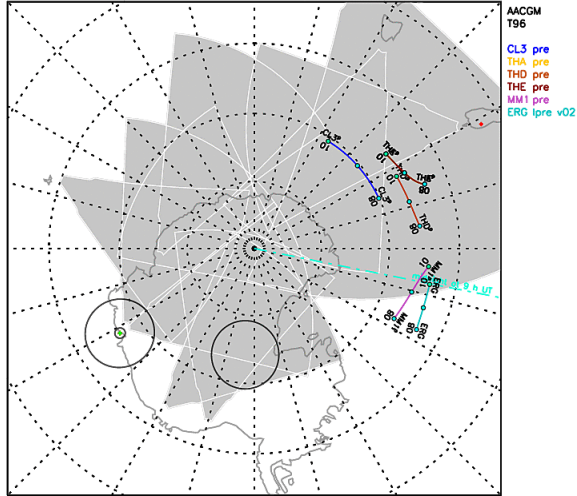
00 02 04 06 08 10 12 14 16 18 20 22 24 UT

>3 Conjunctions from Oct. 13

Spacecraft Footprints and Ground-Based Instruments 20231013 0800–1000 UT
Northern Hemisphere

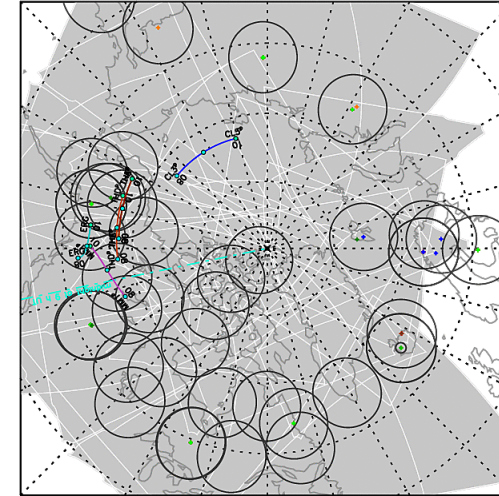


20231013 0800–1000 UT
Southern Hemisphere

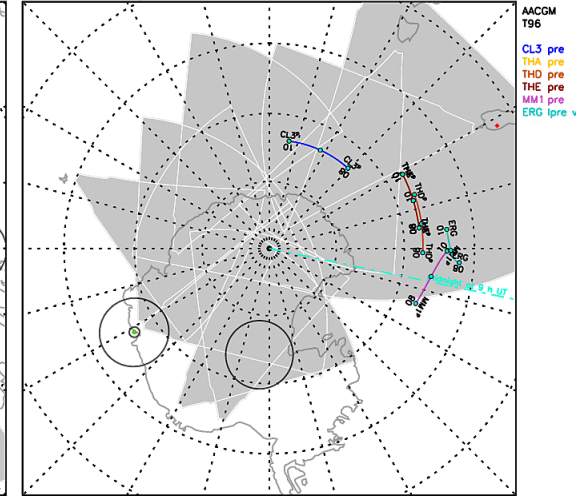


Footprint:
US-Canada
sector

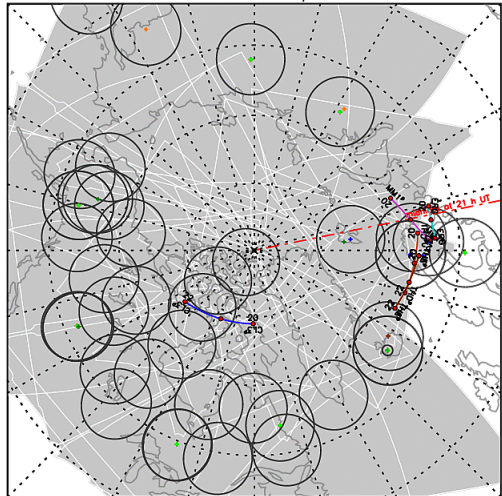
Spacecraft Footprints and Ground-Based Instruments 20231020 0800–1000 UT
Northern Hemisphere



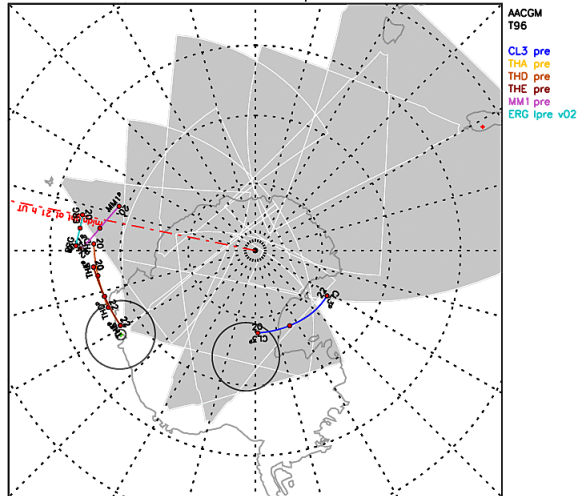
20231020 0800–1000 UT
Southern Hemisphere



Spacecraft Footprints and Ground-Based Instruments 20231016 2000–2200 UT
Northern Hemisphere

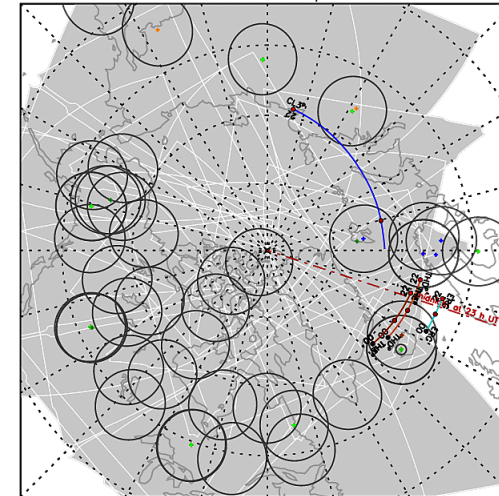


20231016 2000–2200 UT
Southern Hemisphere

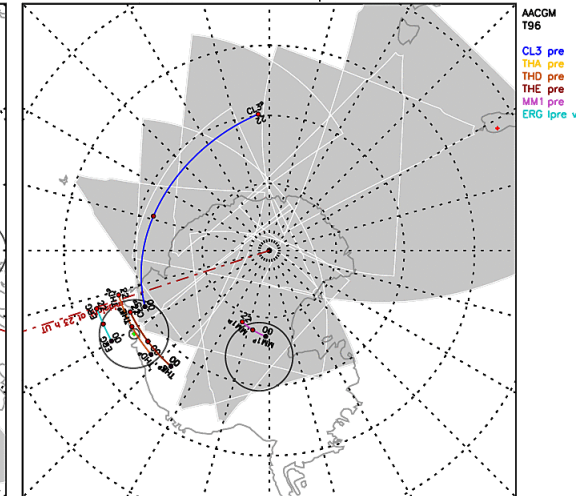


Footprint:
Scandinavia
or Showa
sector

Spacecraft Footprints and Ground-Based Instruments 20231023 2200–2400 UT
Northern Hemisphere



20231023 2200–2400 UT
Southern Hemisphere



Discussion points

- Conjunction events include $L < 10$
- Is it possible to extend SROI region toward inner magnetosphere ?
- Interest from ground-based community ?