

GOES-16 and 17 SUVI Performance Analysis and Comparison with Other Instruments

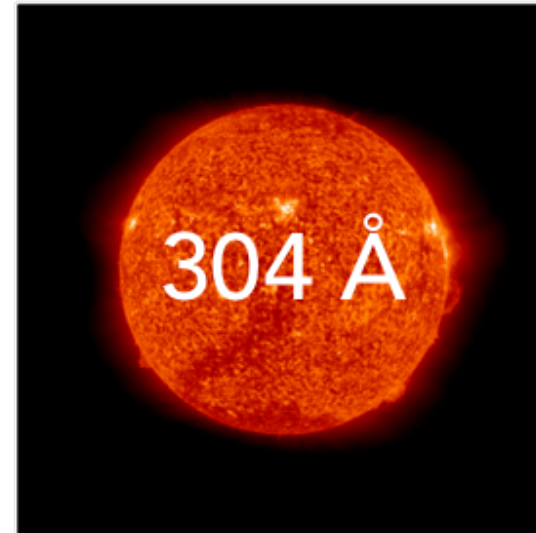
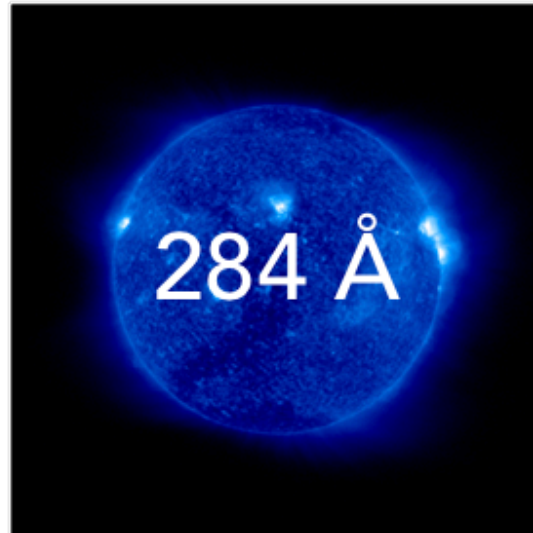
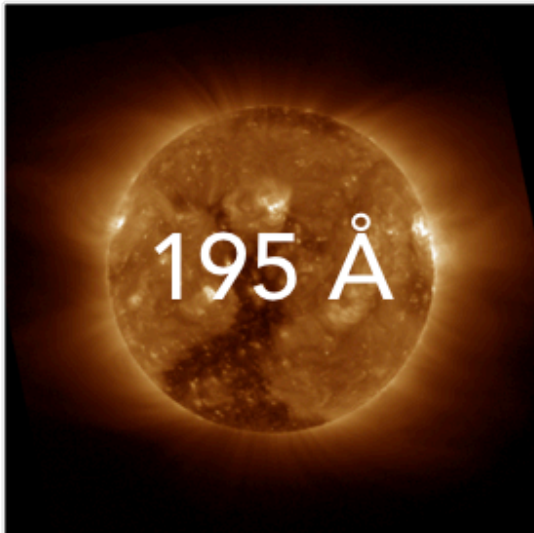
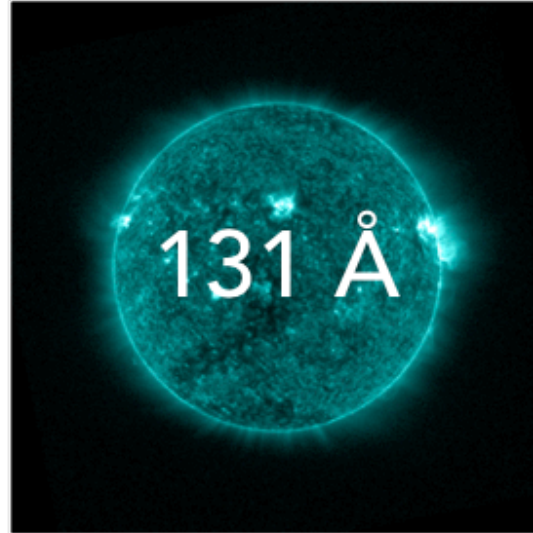
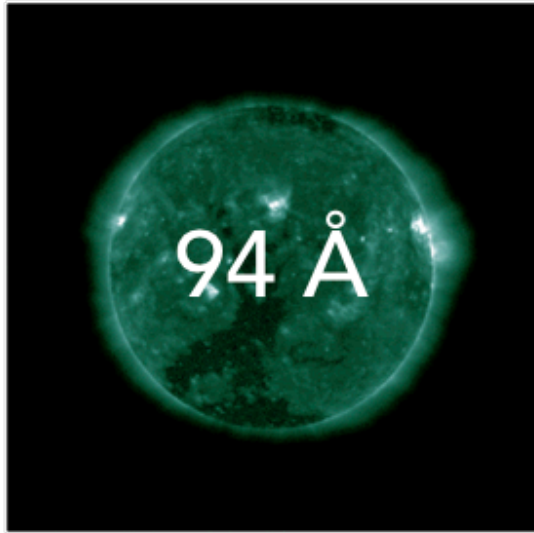
Courtney Peck, Daniel Seaton, and Jonathan Darnel

Cooperative Institute for Research in Environmental Sciences (CIRES)

University of Colorado - Boulder



The Solar UltraViolet Imager (SUVI) is an EUV imager with 6 channels



53 x 53 arcmin field-of-view

1280 x 1280 CCD

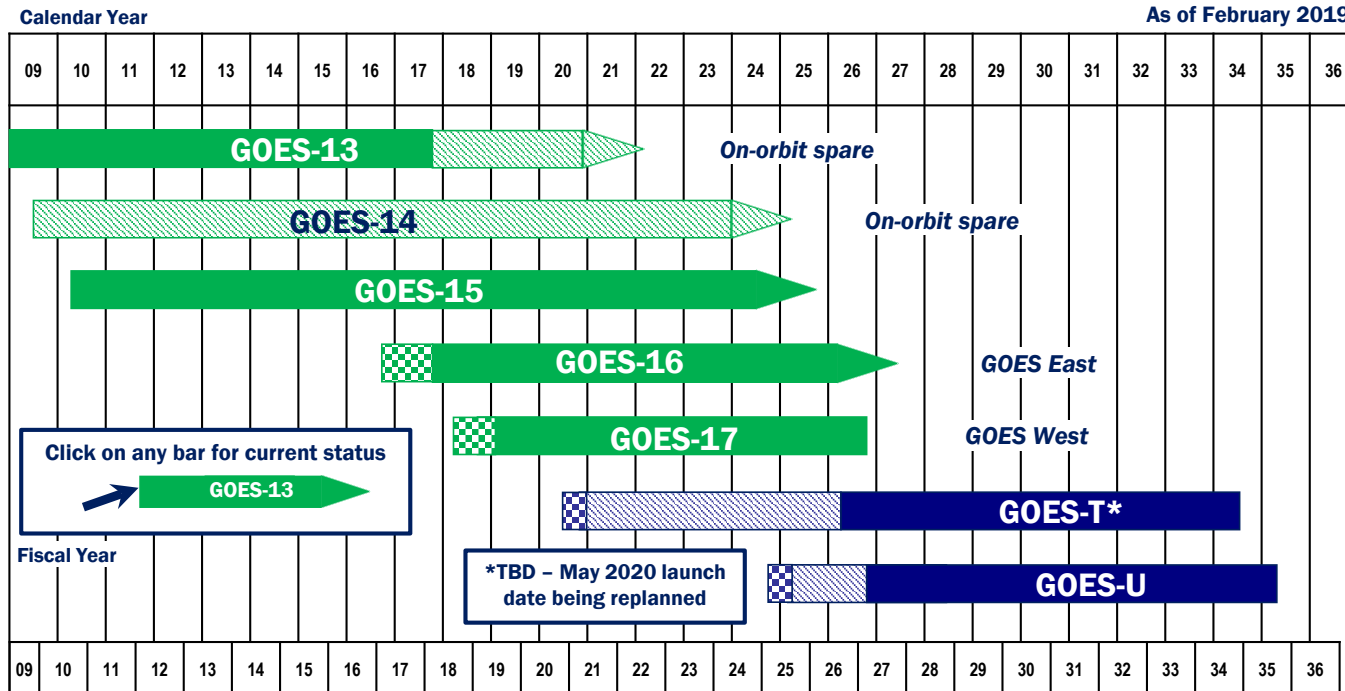
2.5 arcsec/pixel resolution

10 s operational cadence

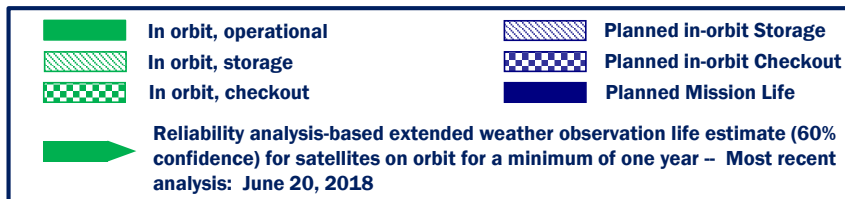
Planned lifetime is almost 20 years between 4 (nearly) identical instruments



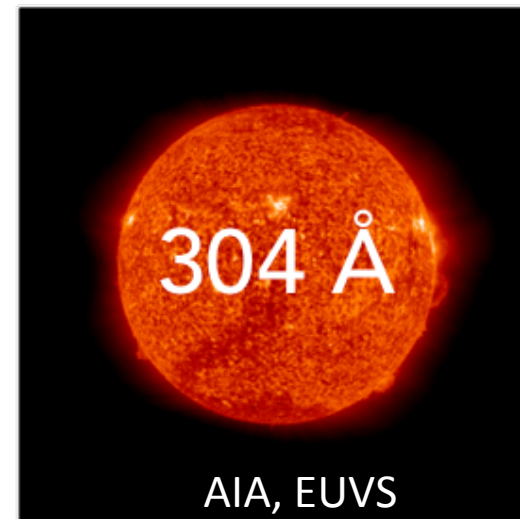
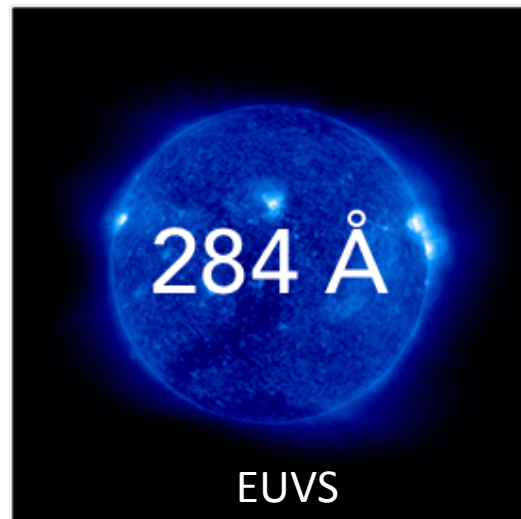
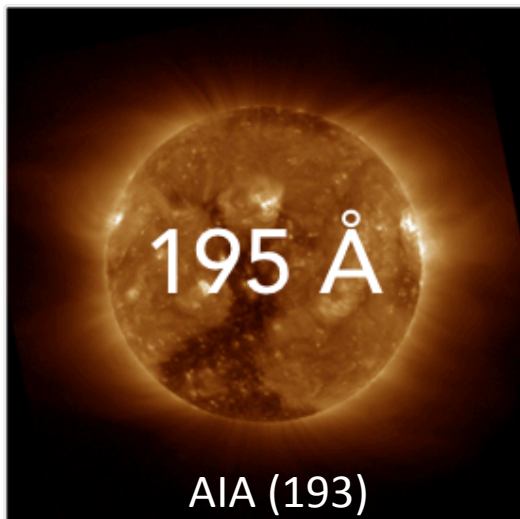
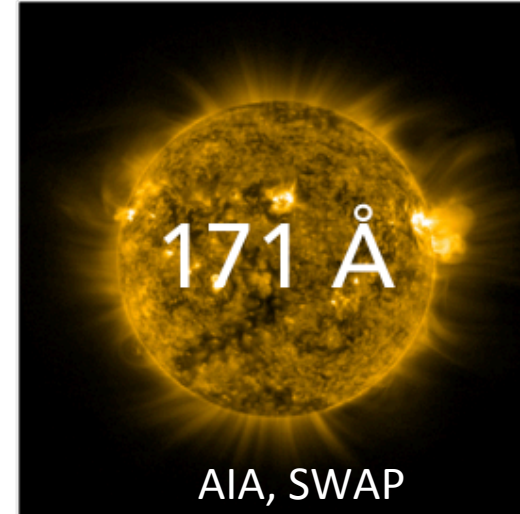
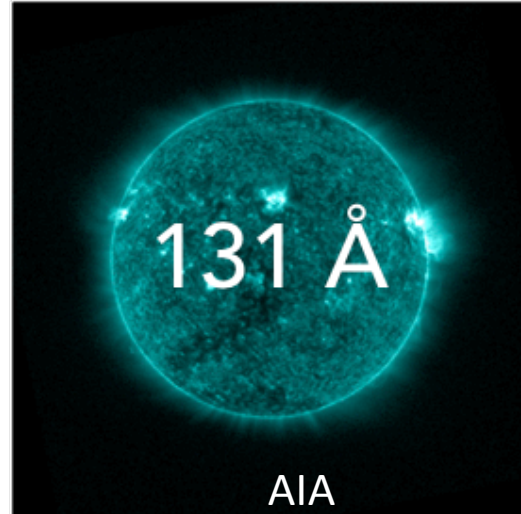
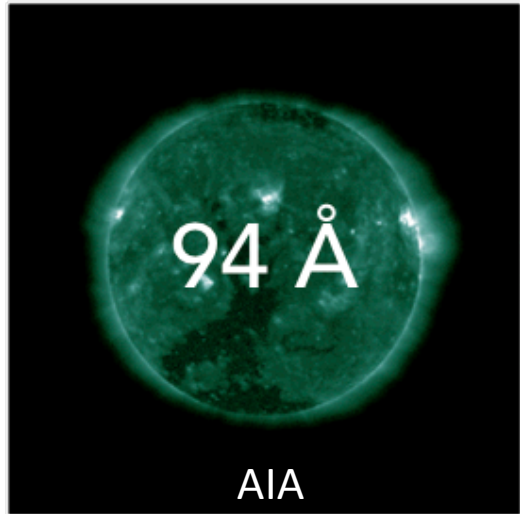
NOAA Geostationary Satellite Programs Continuity of Weather Observations



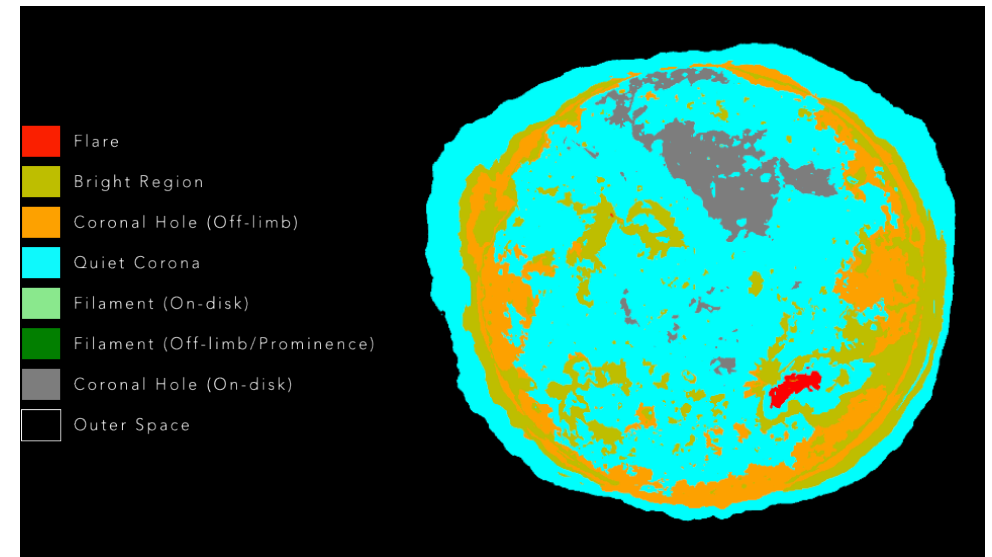
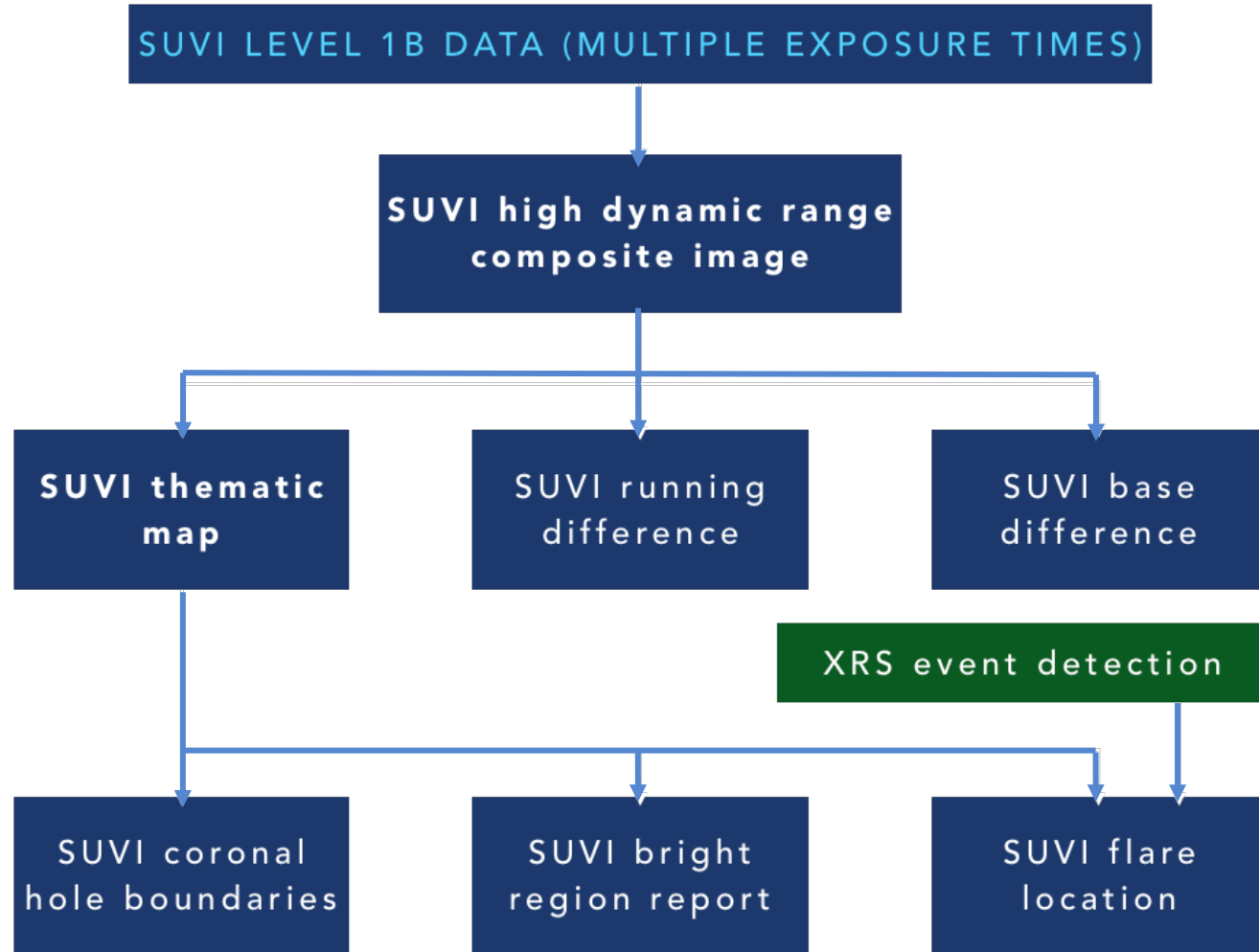
Approved: Assistant Administrator for Satellite and Information Services



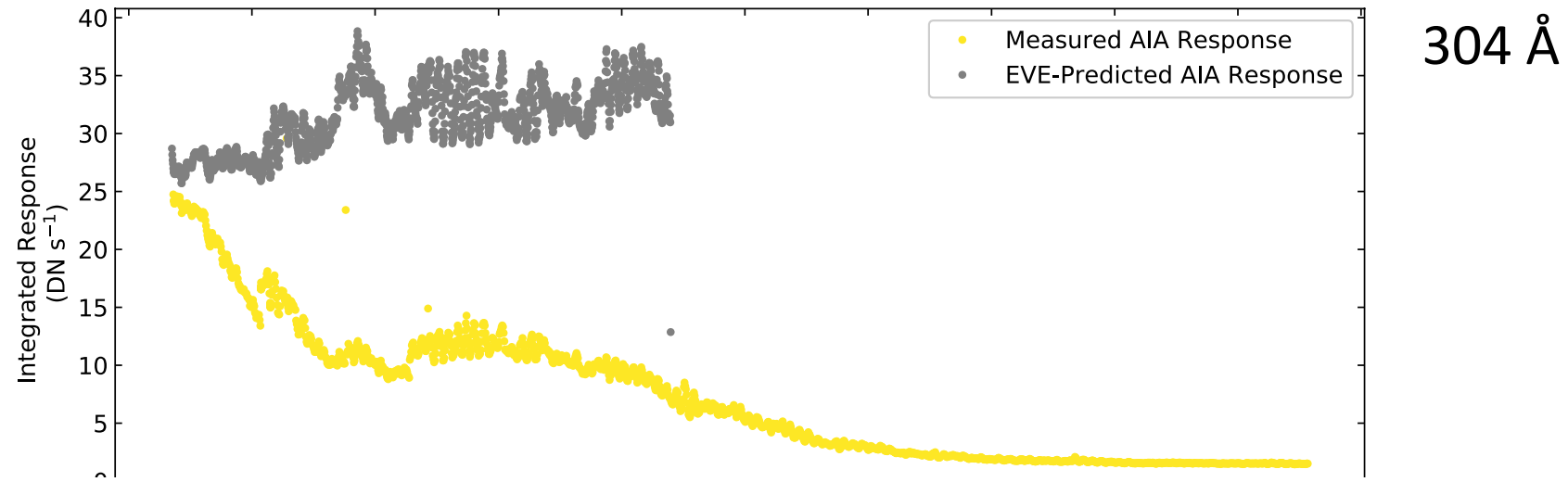
SUVI passbands closely match those of AIA and SWAP imagers, as well as two EUVS channels



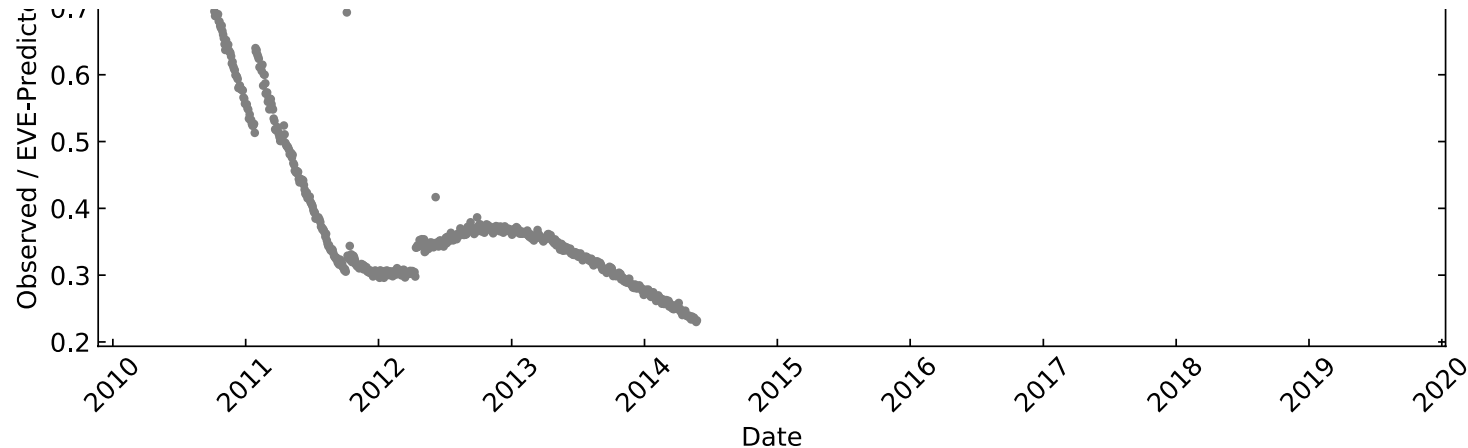
SUVI requires image degradation calibration to maintain accurate measurements for both science and operational products



AIA degradation could be calibrated with EVE until MEGS-A failure in 2014

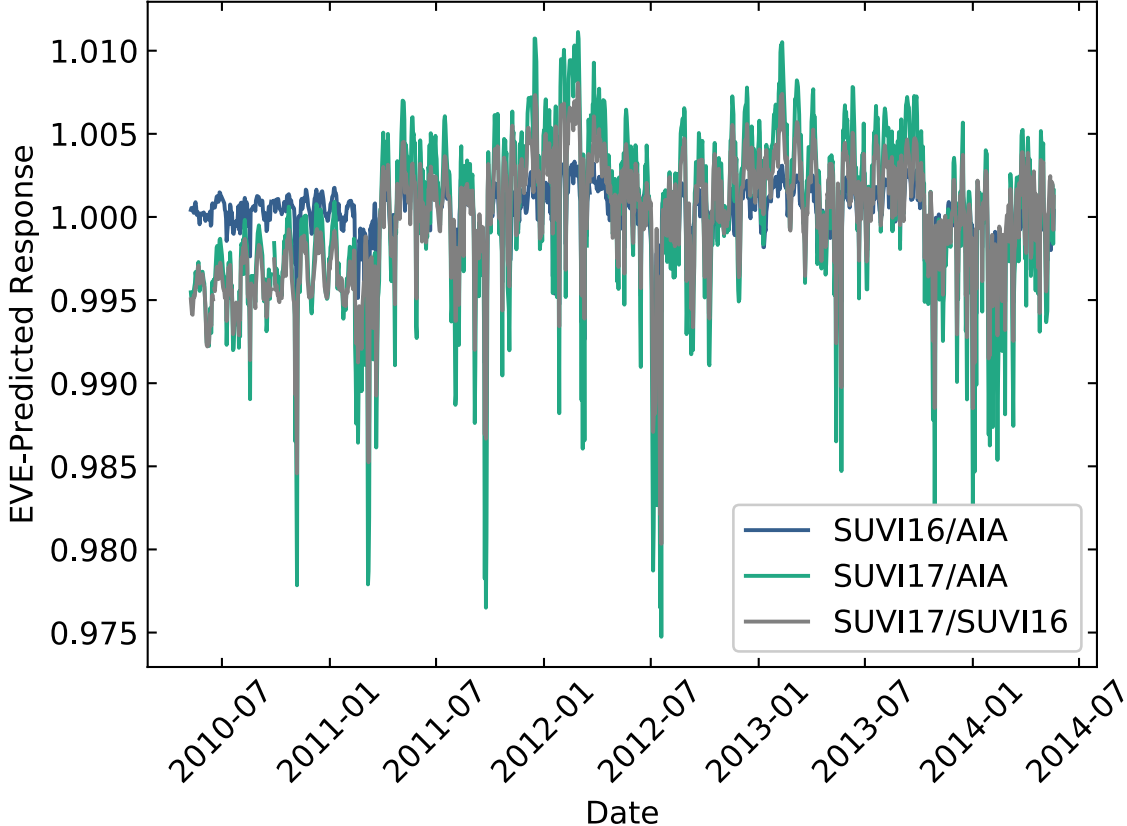
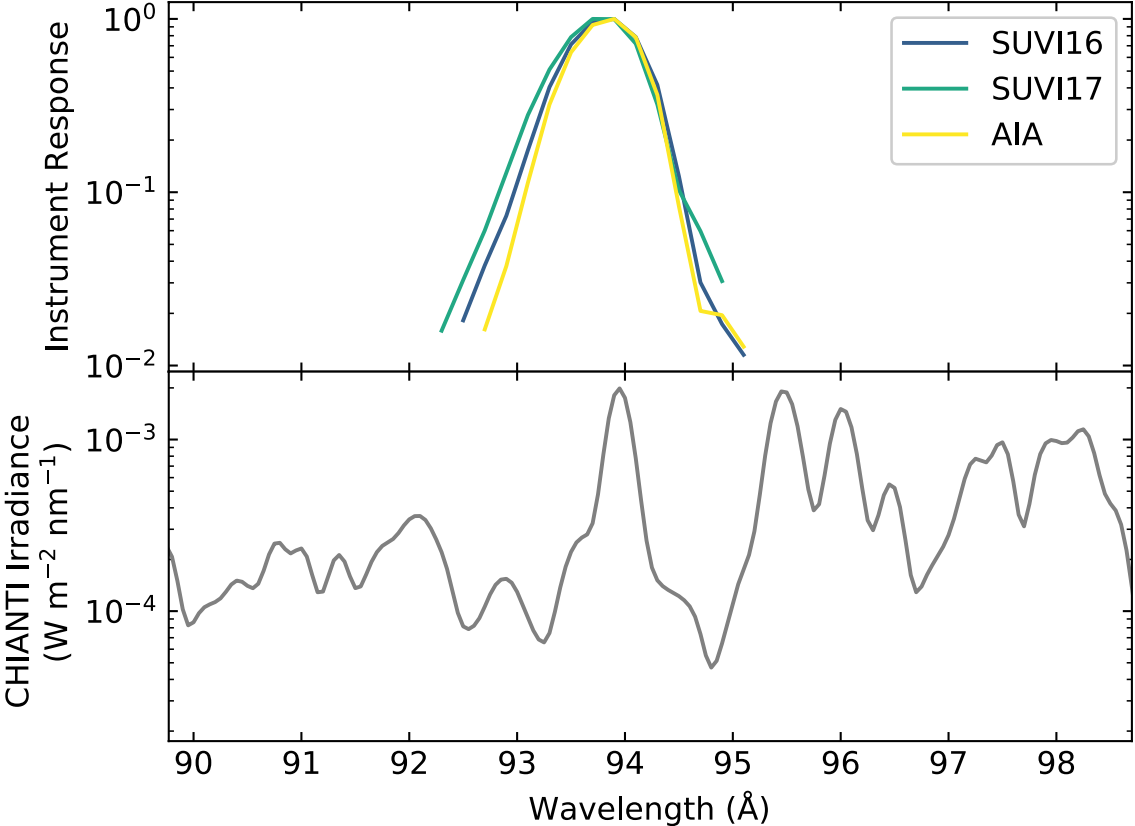


Our goal is to leverage the multi-instrument aspect of SUVI with other EUV instruments to understand and correct image degradation



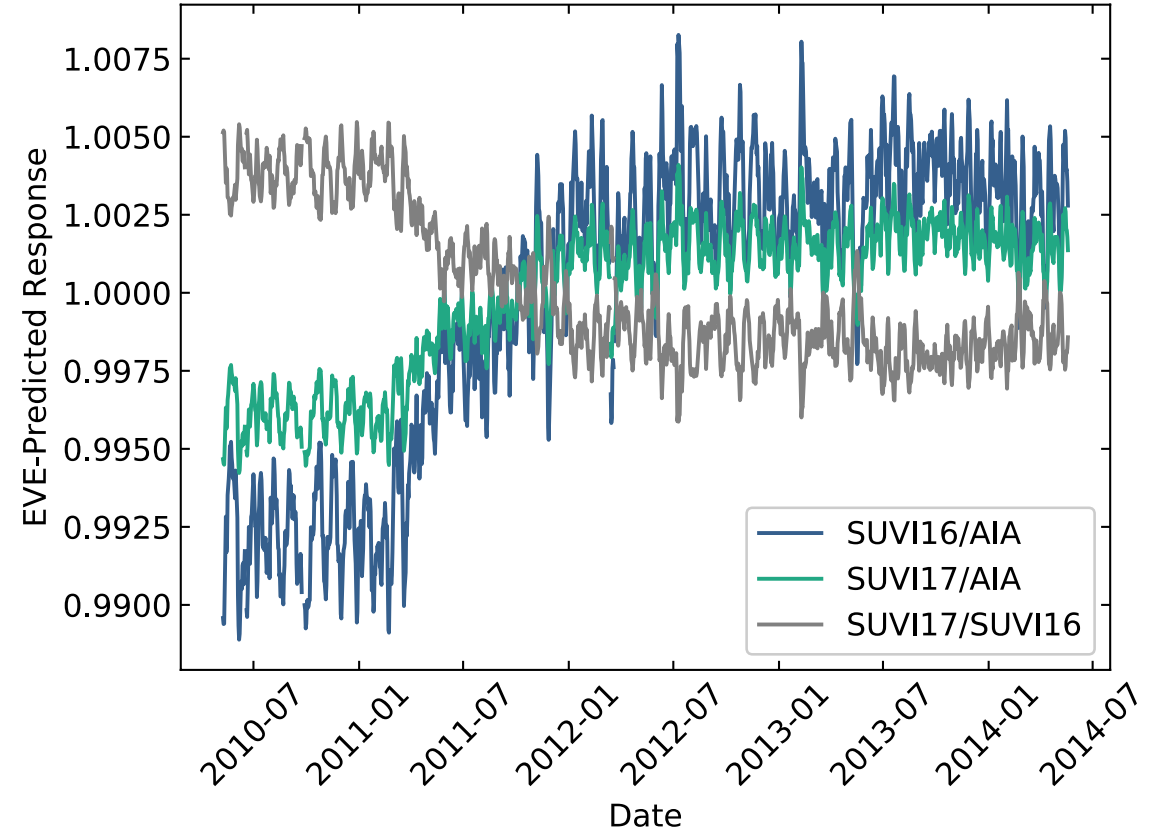
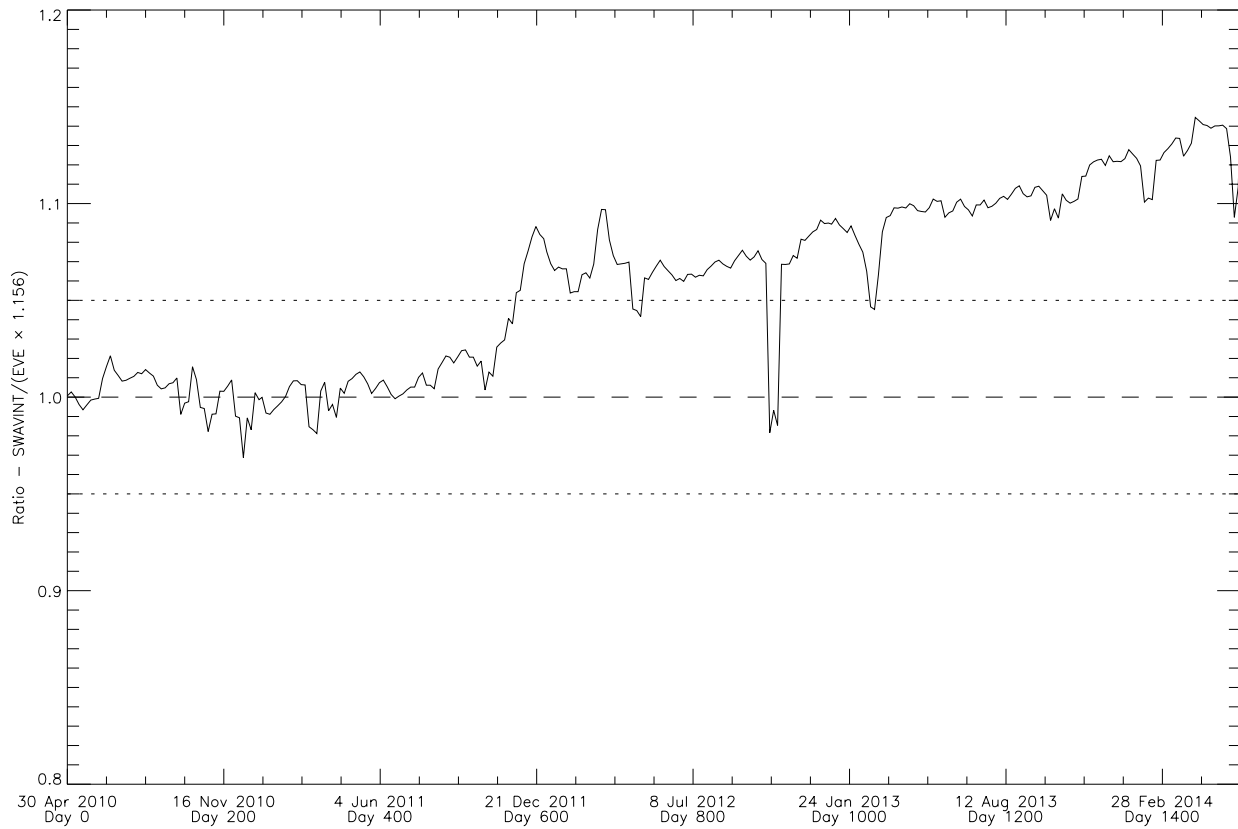
SUVI and AIA response function comparison

94 Å



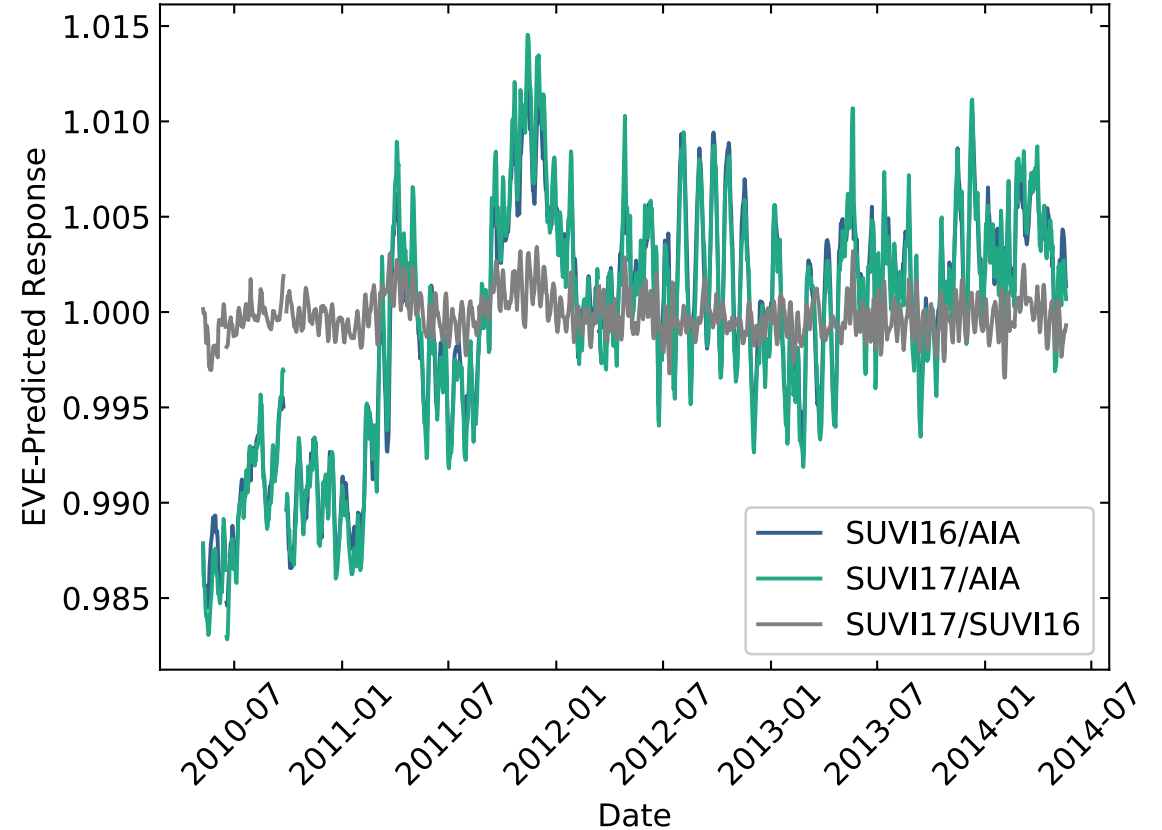
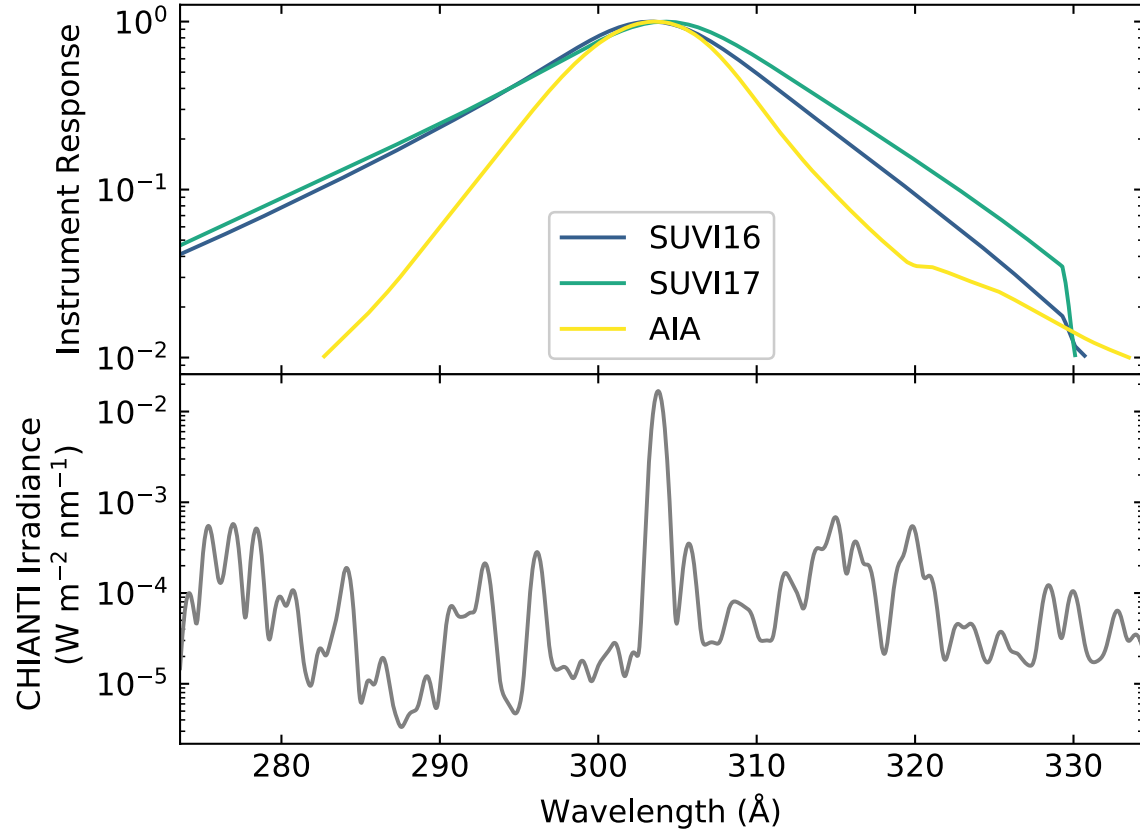
SUVI and AIA response function comparison

171 Å



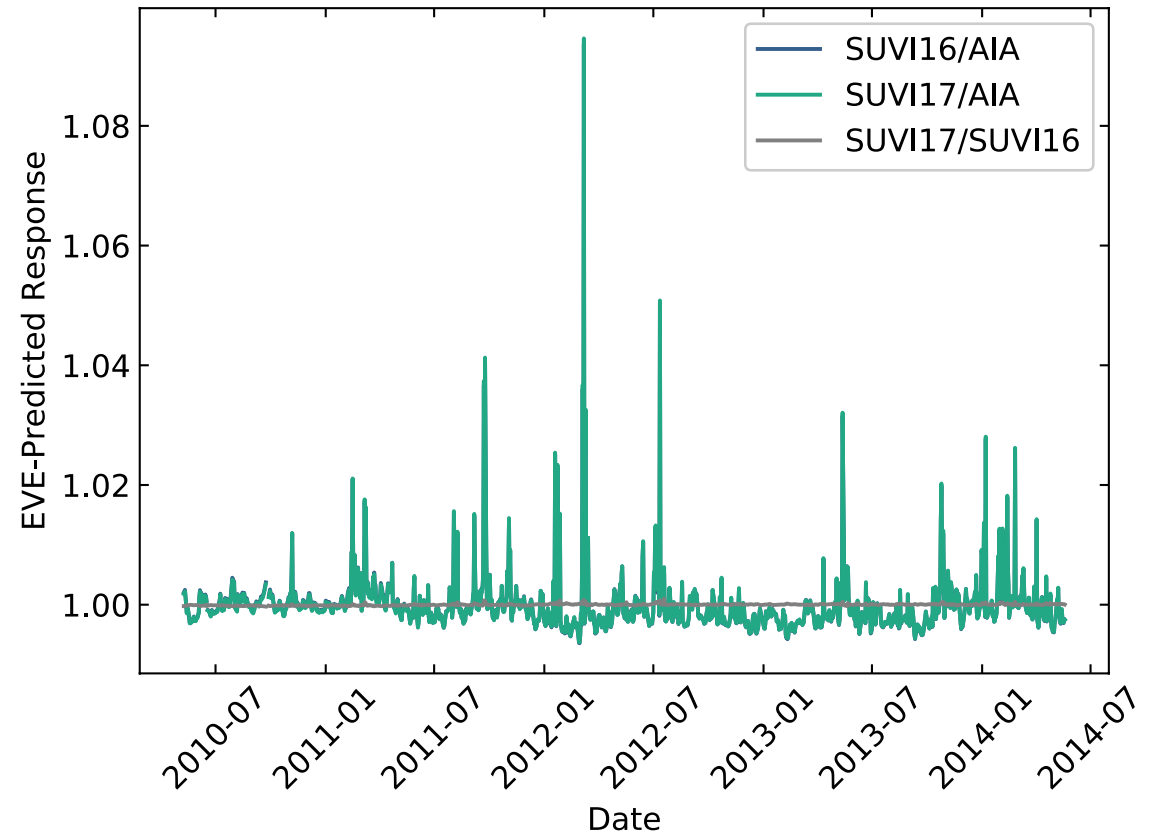
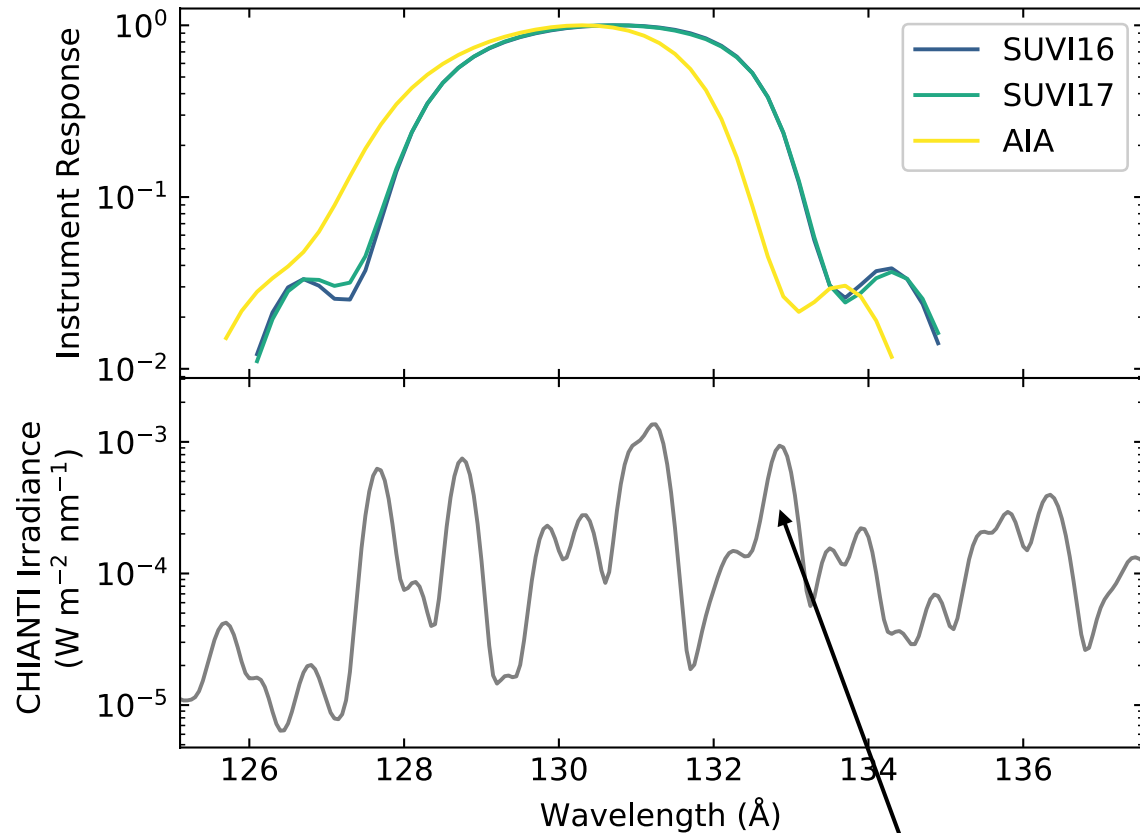
SUVI and AIA response function comparison

304 Å



SUVI and AIA response function comparison

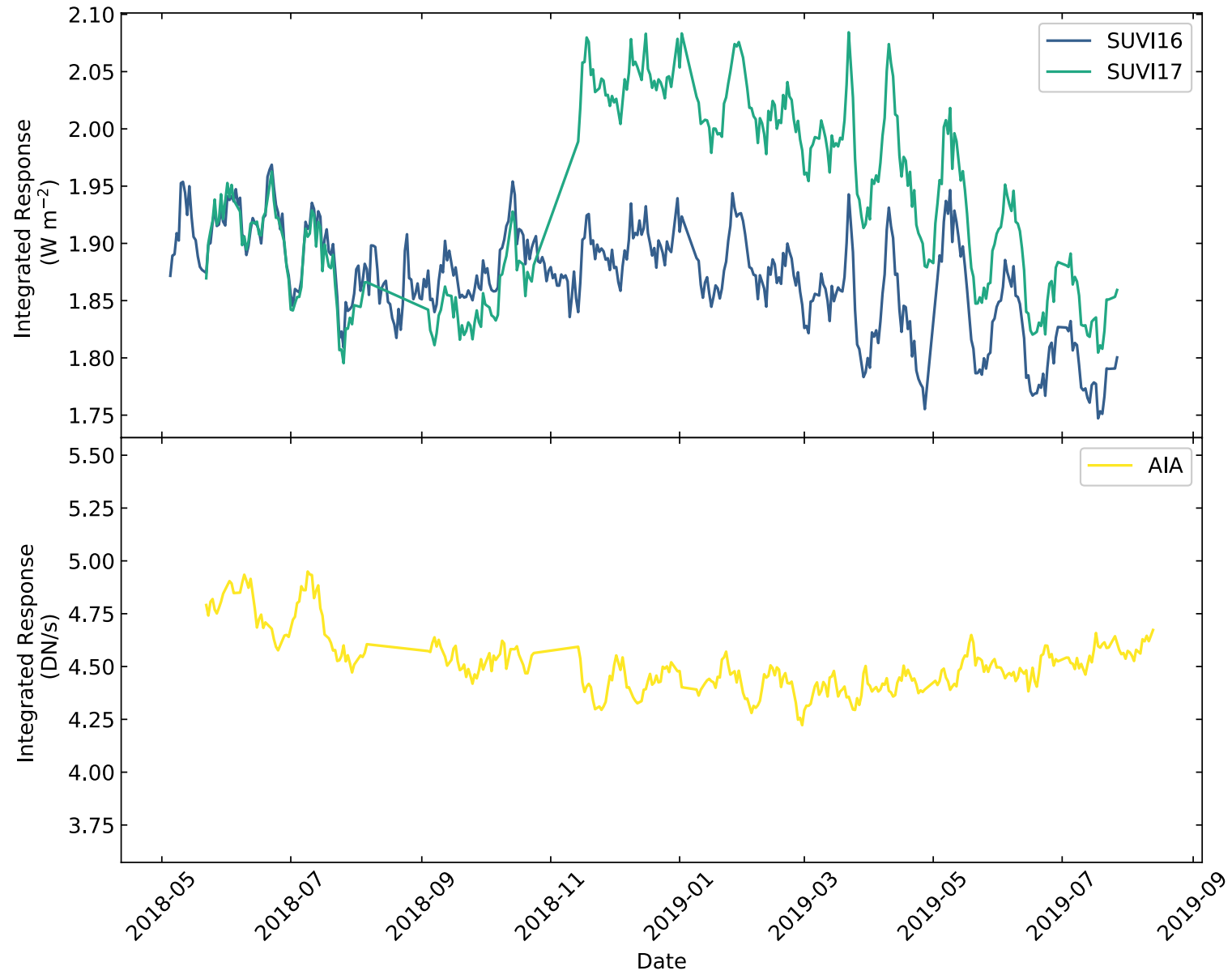
131 Å



Fe XXII flare line

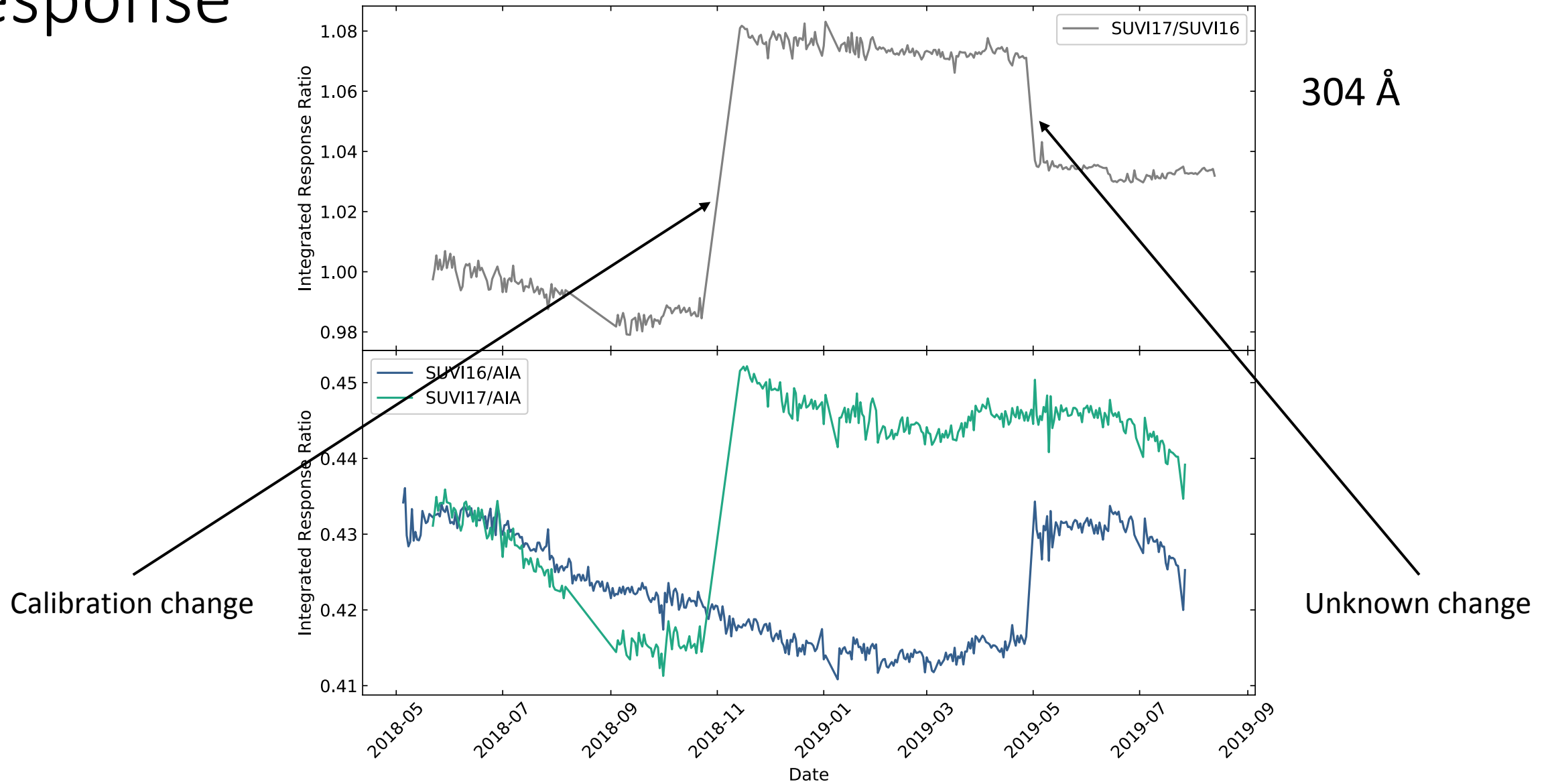
SUVI 16/17 and AIA passbands largely agree (with exception of 131 Å)

Comparison of SUVI 16 and 17 and AIA integrated response

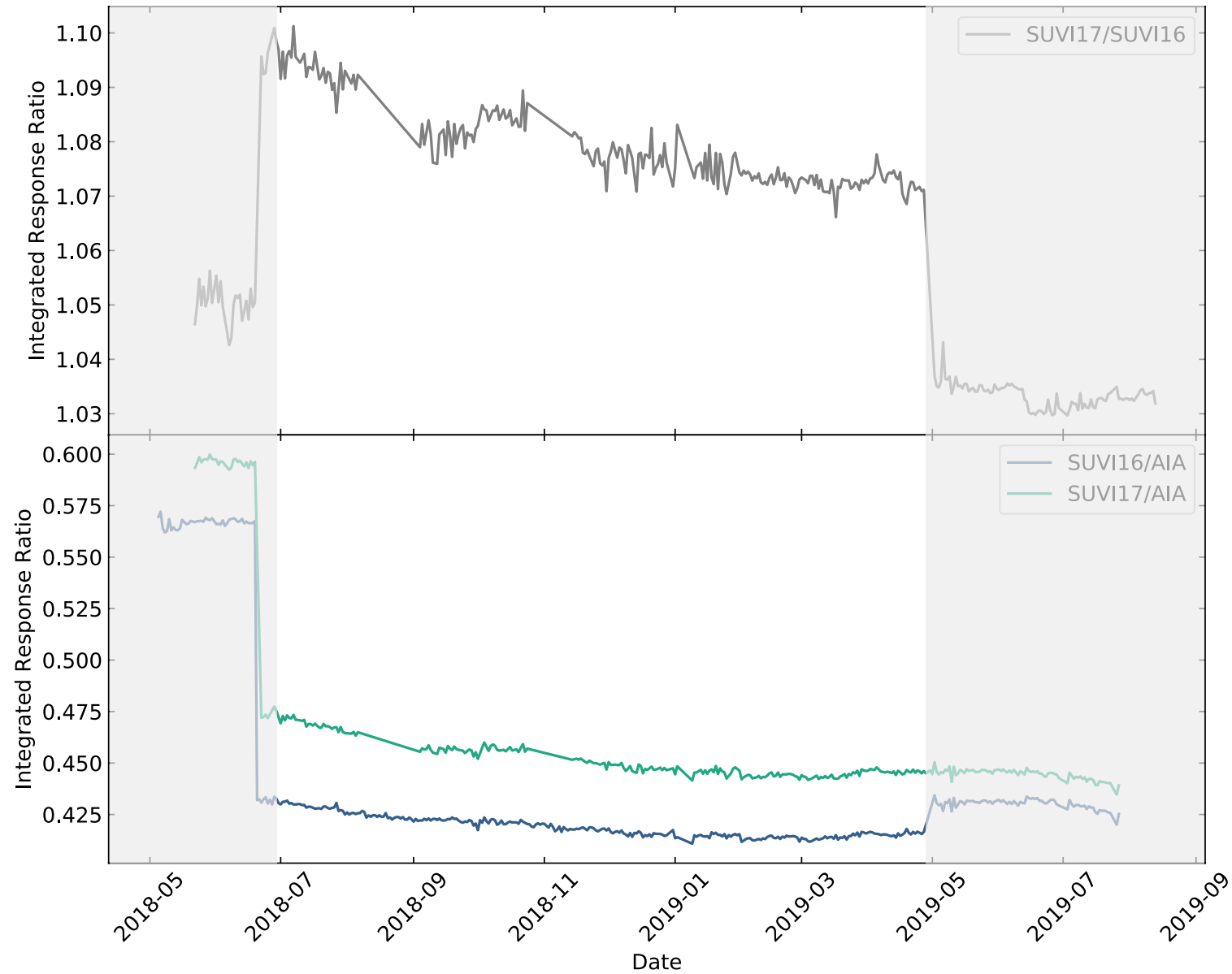


304 Å

Comparison of SUVI 16 and 17 and AIA integrated response

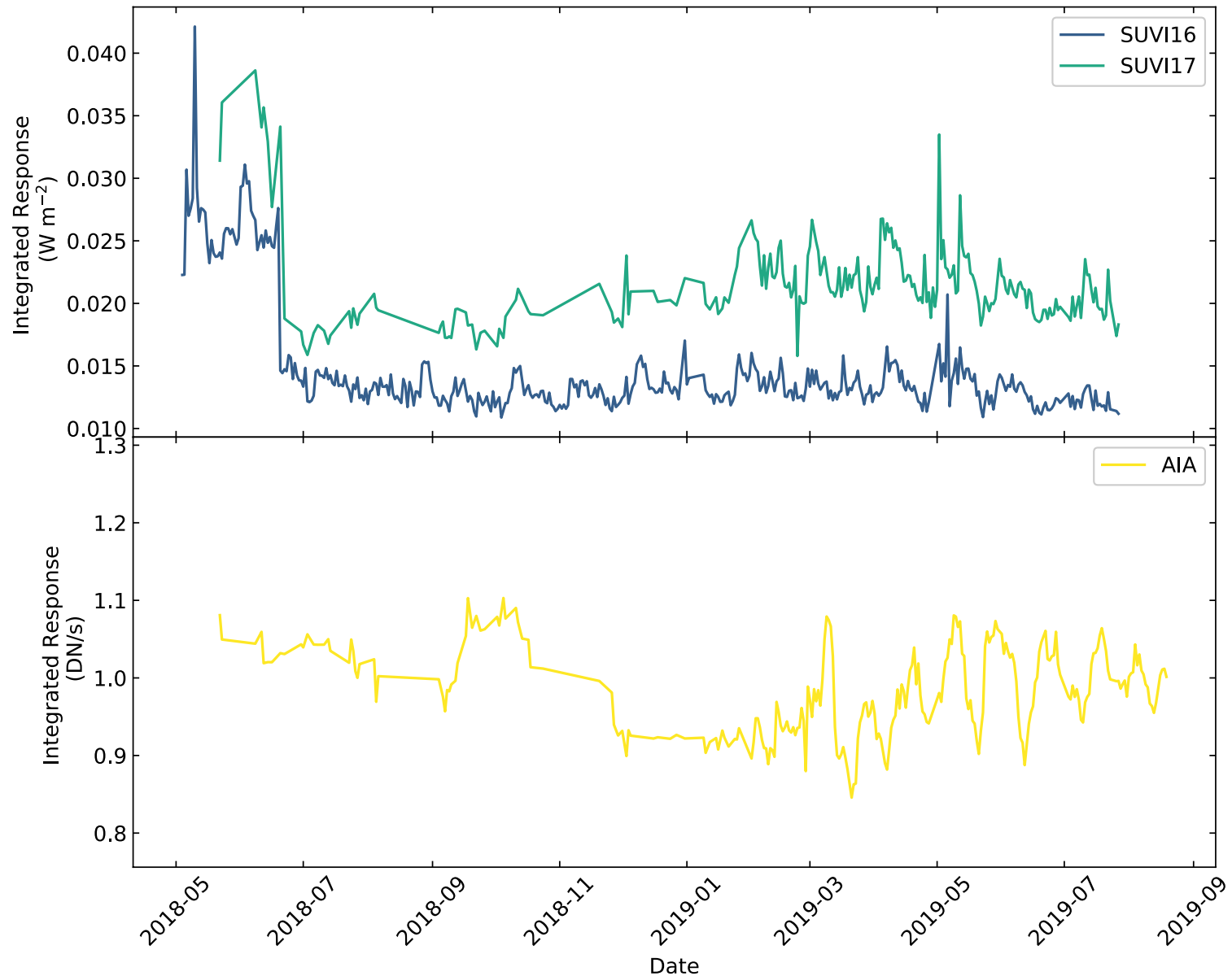


Comparison of SUVI 16 and 17 and AIA integrated response



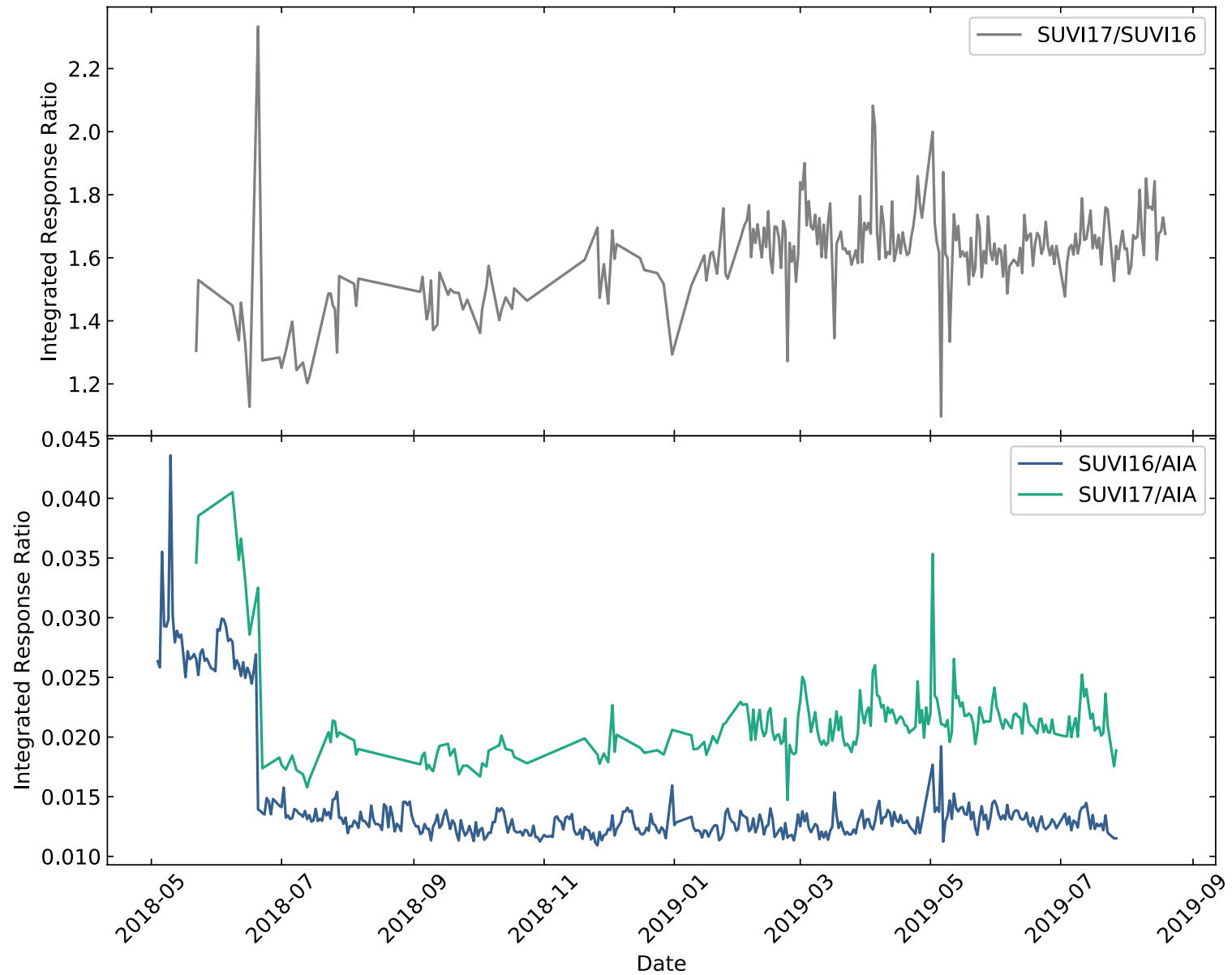
304 Å

Comparison of SUVI 16 and 17 and AIA integrated response

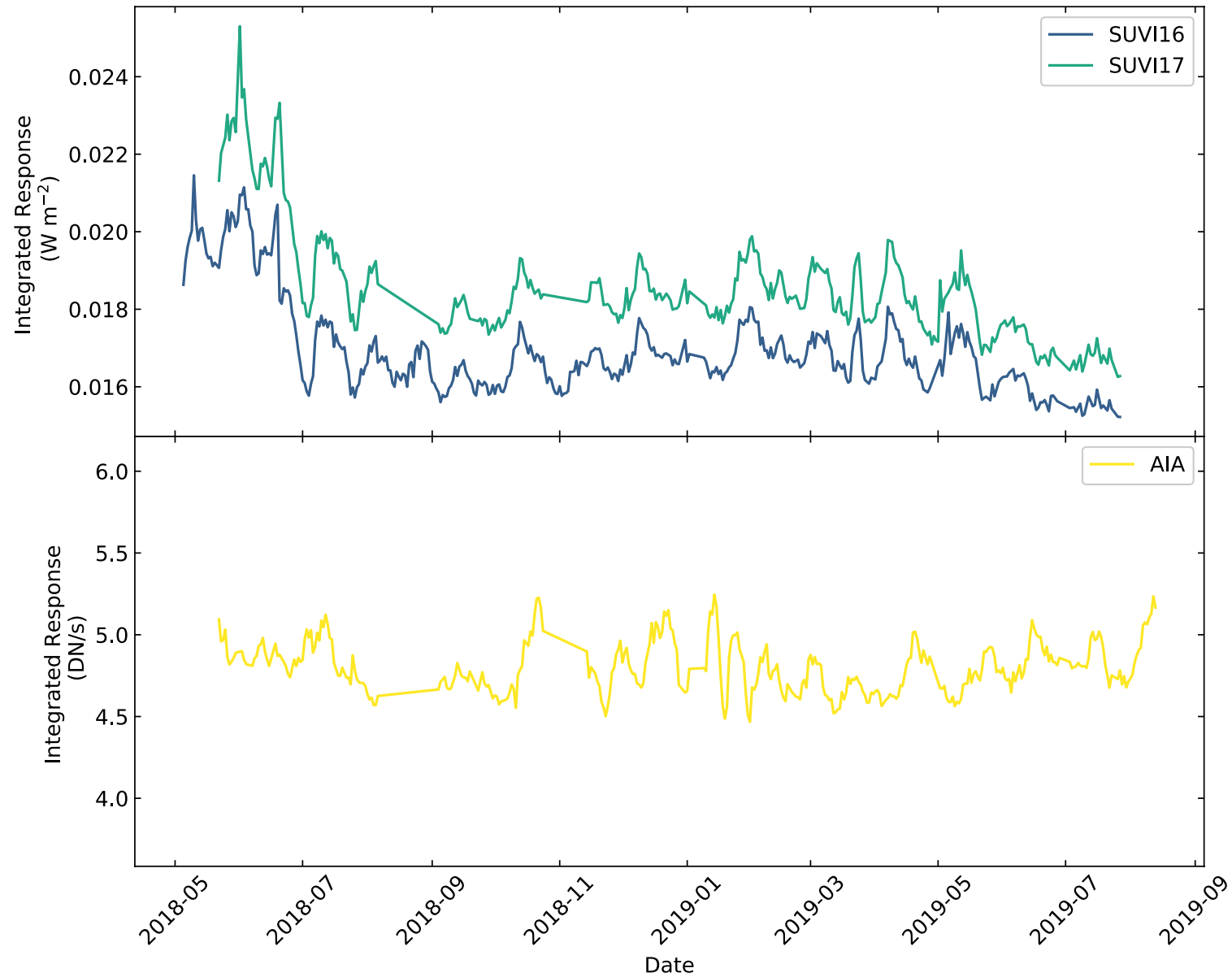


94 Å

Comparison of SUVI 16 and 17 and AIA integrated response

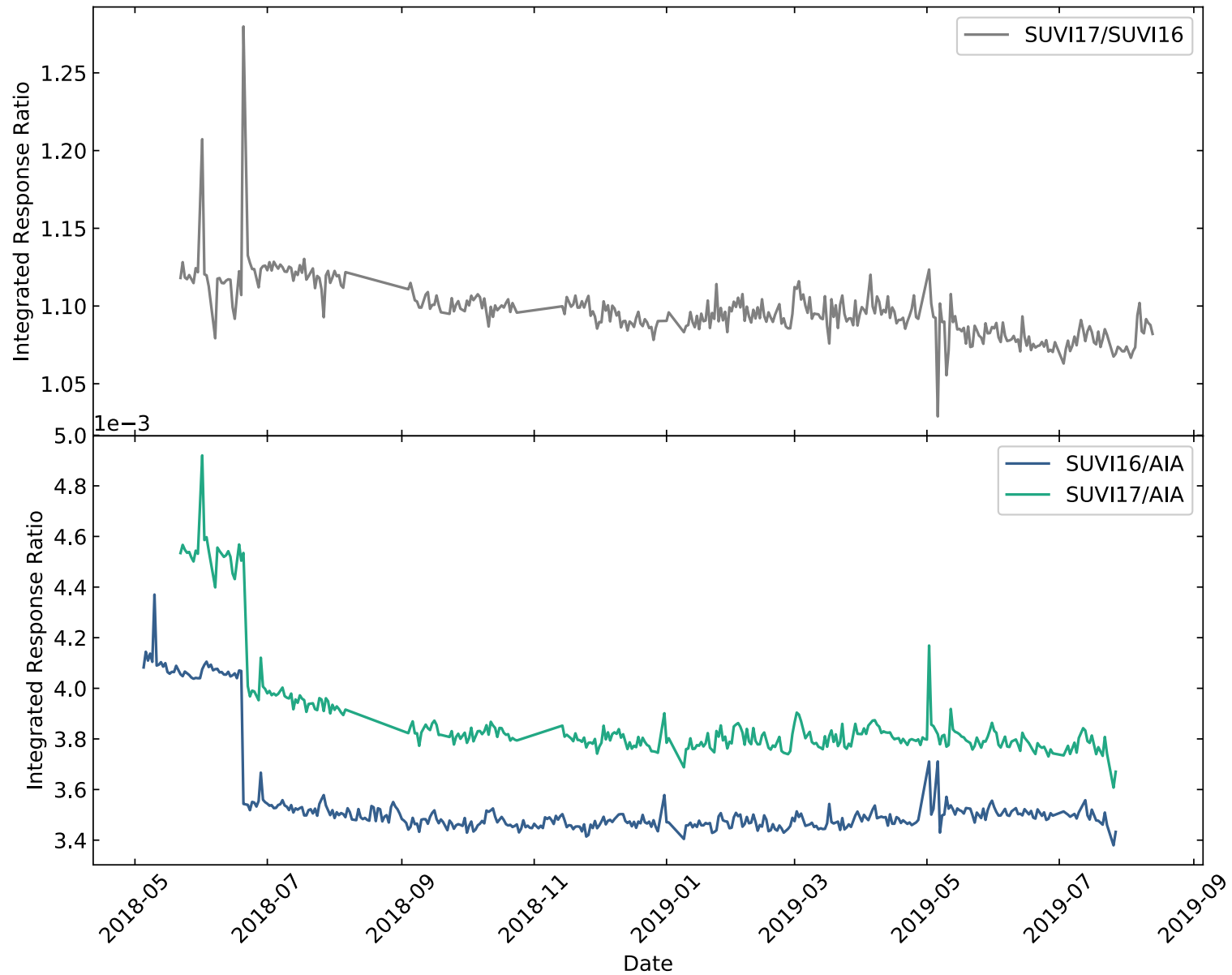


Comparison of SUVI 16 and 17 and AIA integrated response



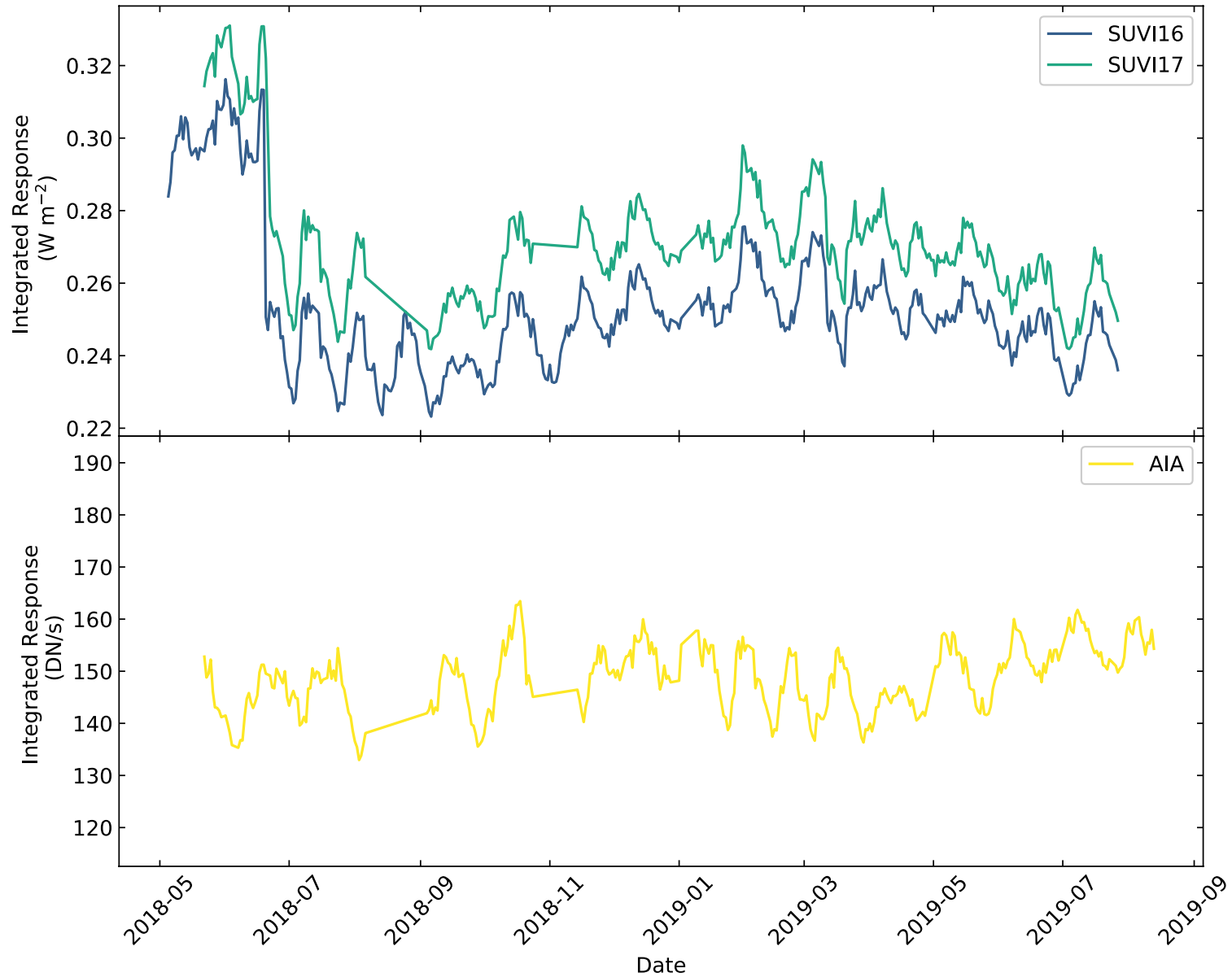
131 Å

Comparison of SUVI 16 and 17 and AIA integrated response



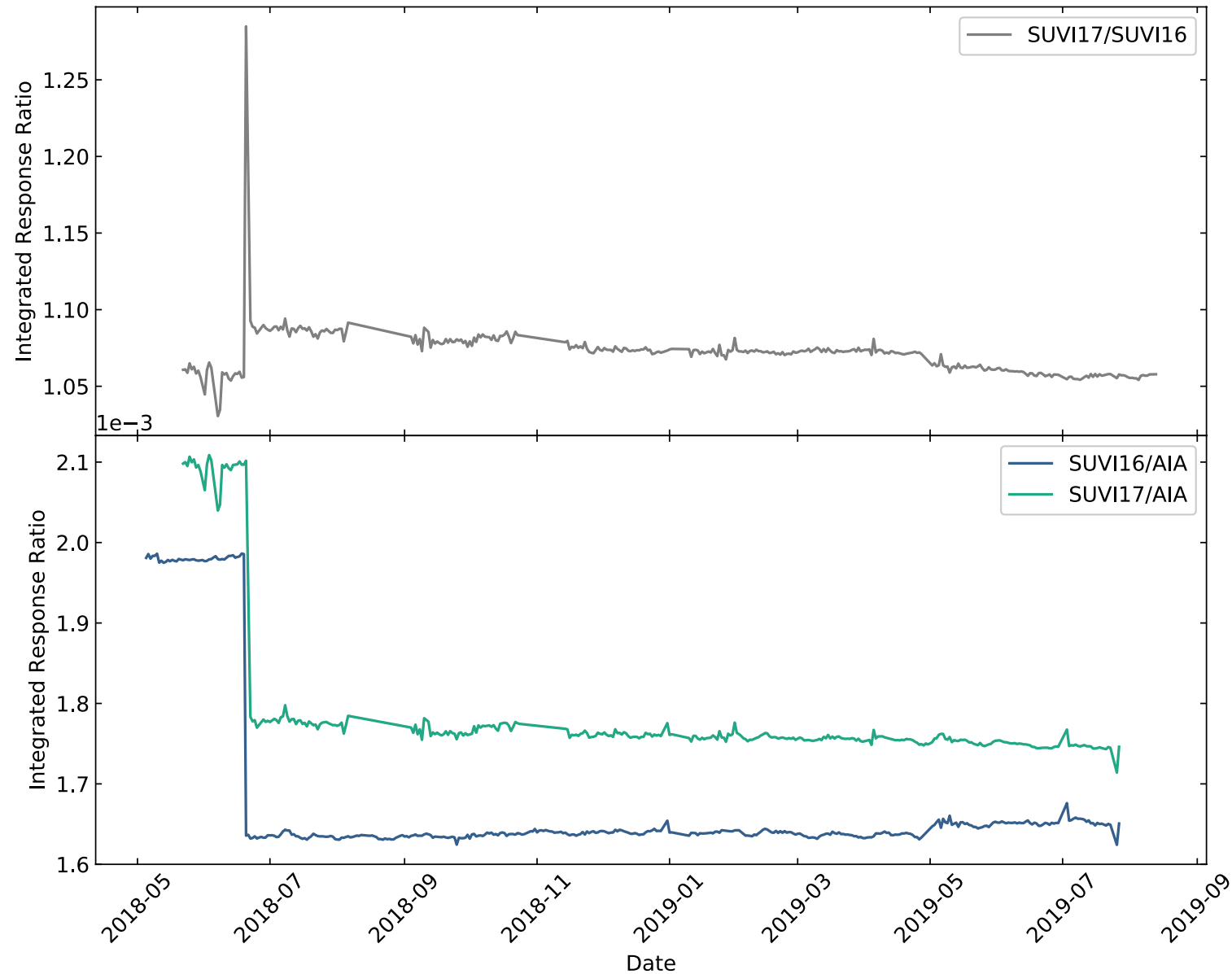
131 Å

Comparison of SUVI 16 and 17 and AIA integrated response



171 Å

Comparison of SUVI 16 and 17 and AIA integrated response



171 Å

Fitting degradation trend between SUVI and AIA
Can we back out the degradation coefficient?

$$\frac{T_1}{T_2} = \frac{T_{i,1} e^{-\tau_1(t-t_{i,1})}}{T_{i,2} e^{-\tau_2(t-t_{i,2})}} \quad \Delta t = t - t_{i,1}$$

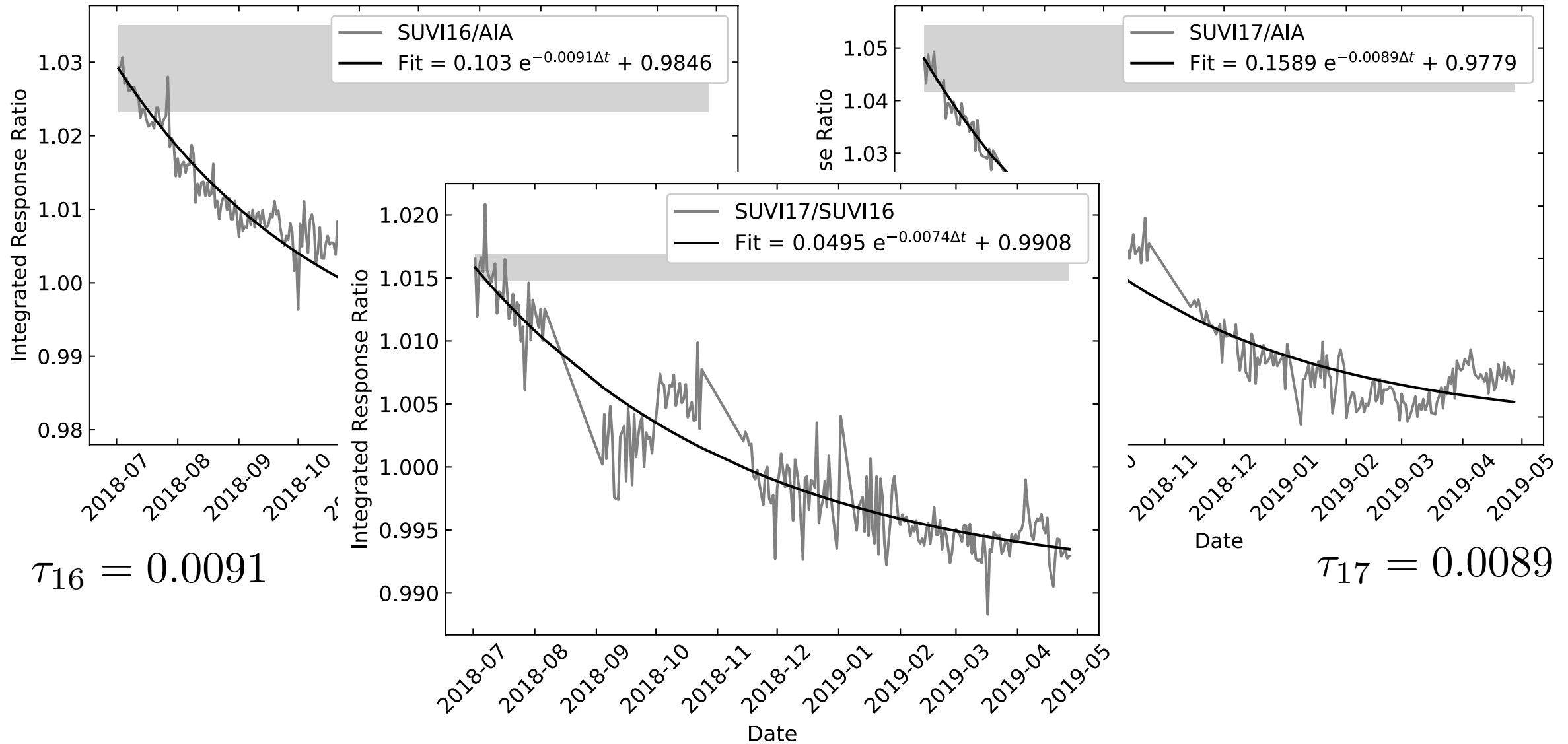
$$= \frac{T_{i,1} e^{-\tau_1(\Delta t)}}{T_{i,2} e^{-\tau_2(\Delta t + t_{i,1} - t_{i,2})}}$$

$$\frac{T_1}{T_2} = C e^{-(\tau_1 - \tau_2)\Delta t}$$

Fitting degradation trend to the data (SUVI 16 to AIA)

Can we back out the degradation coefficient?

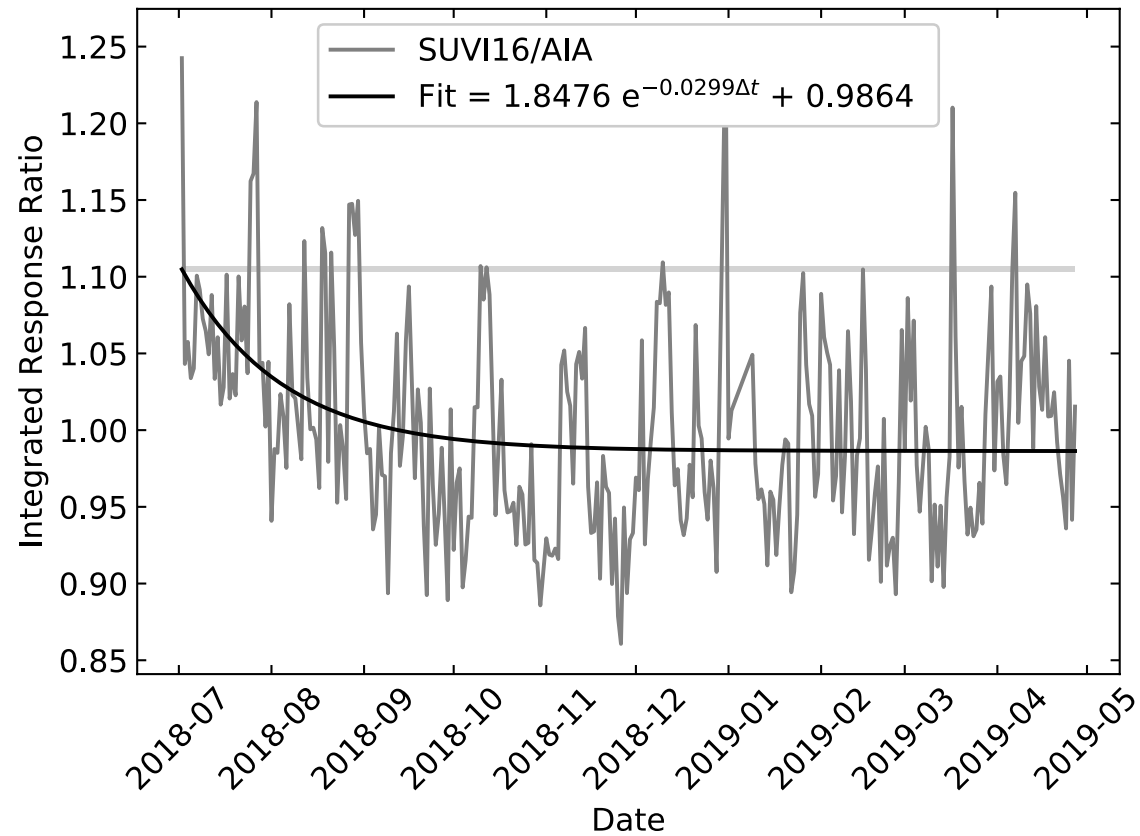
304 Å



Fitting degradation trend to the data (SUVI 16 to AIA)

Can we back out the degradation coefficient?

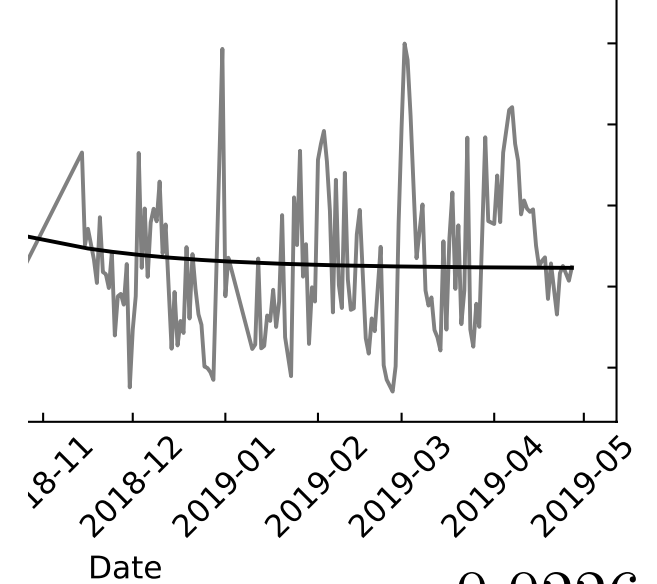
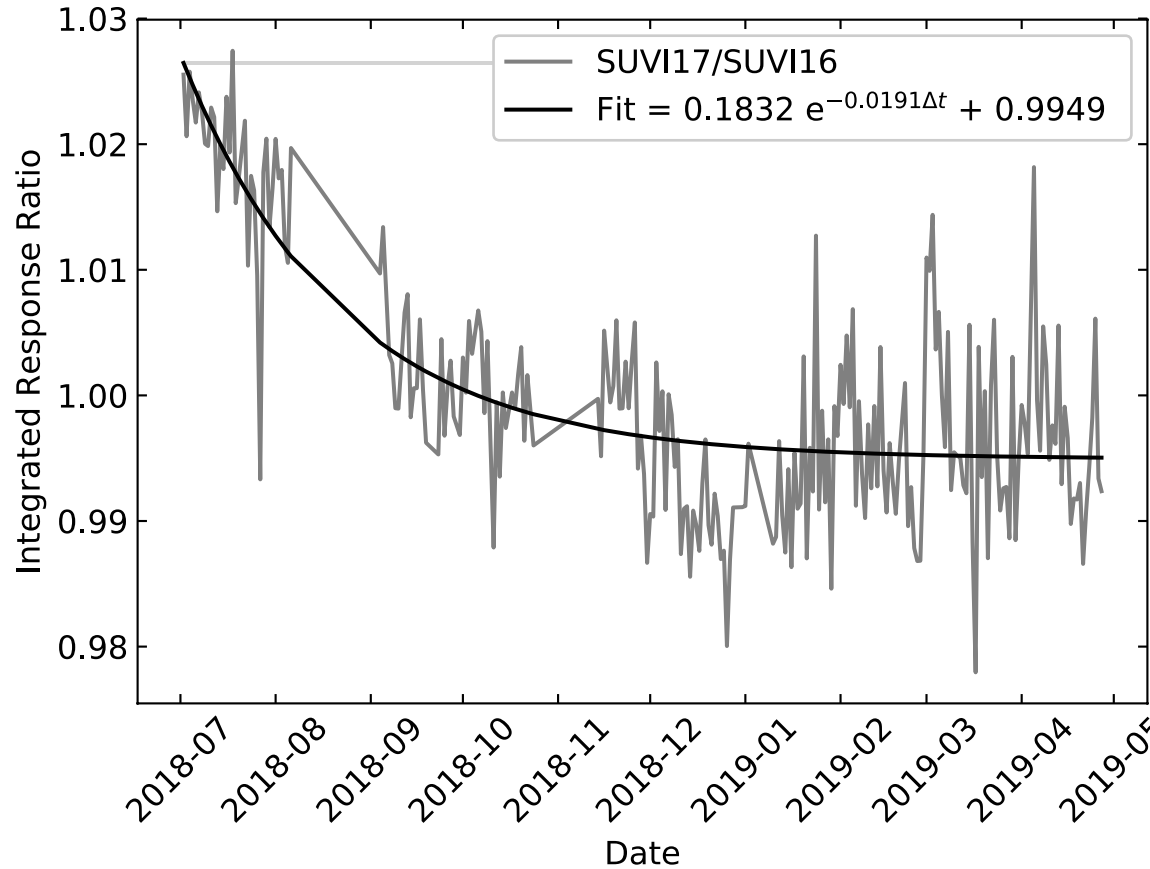
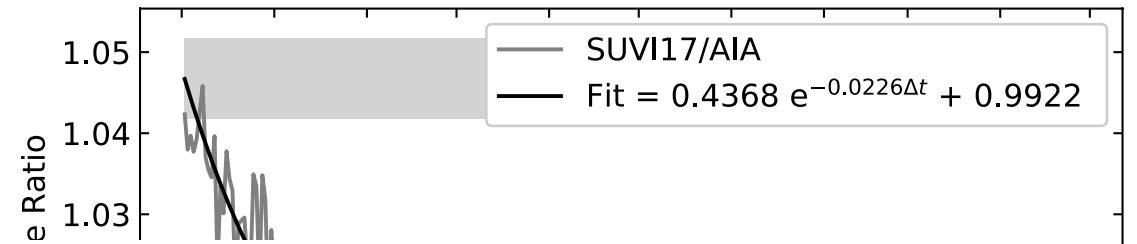
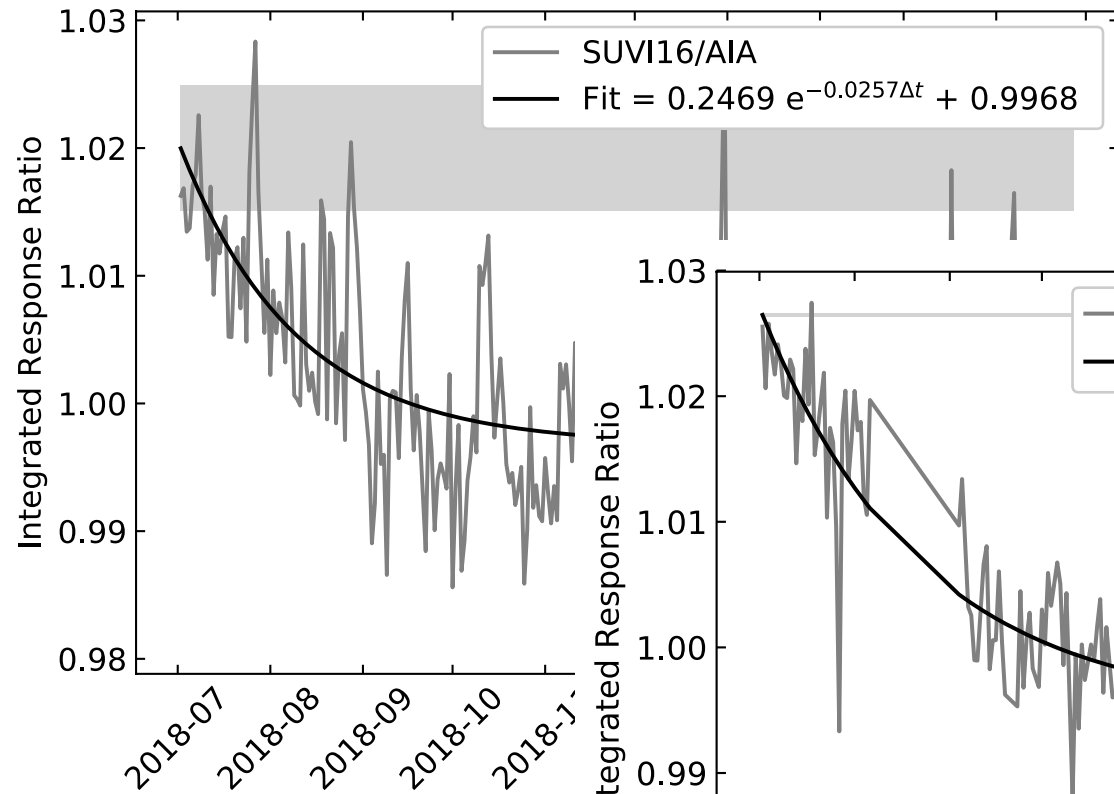
94 Å



Fitting degradation trend to the data (SUVI 16 to AIA)

Can we back out the degradation coefficient?

131 Å



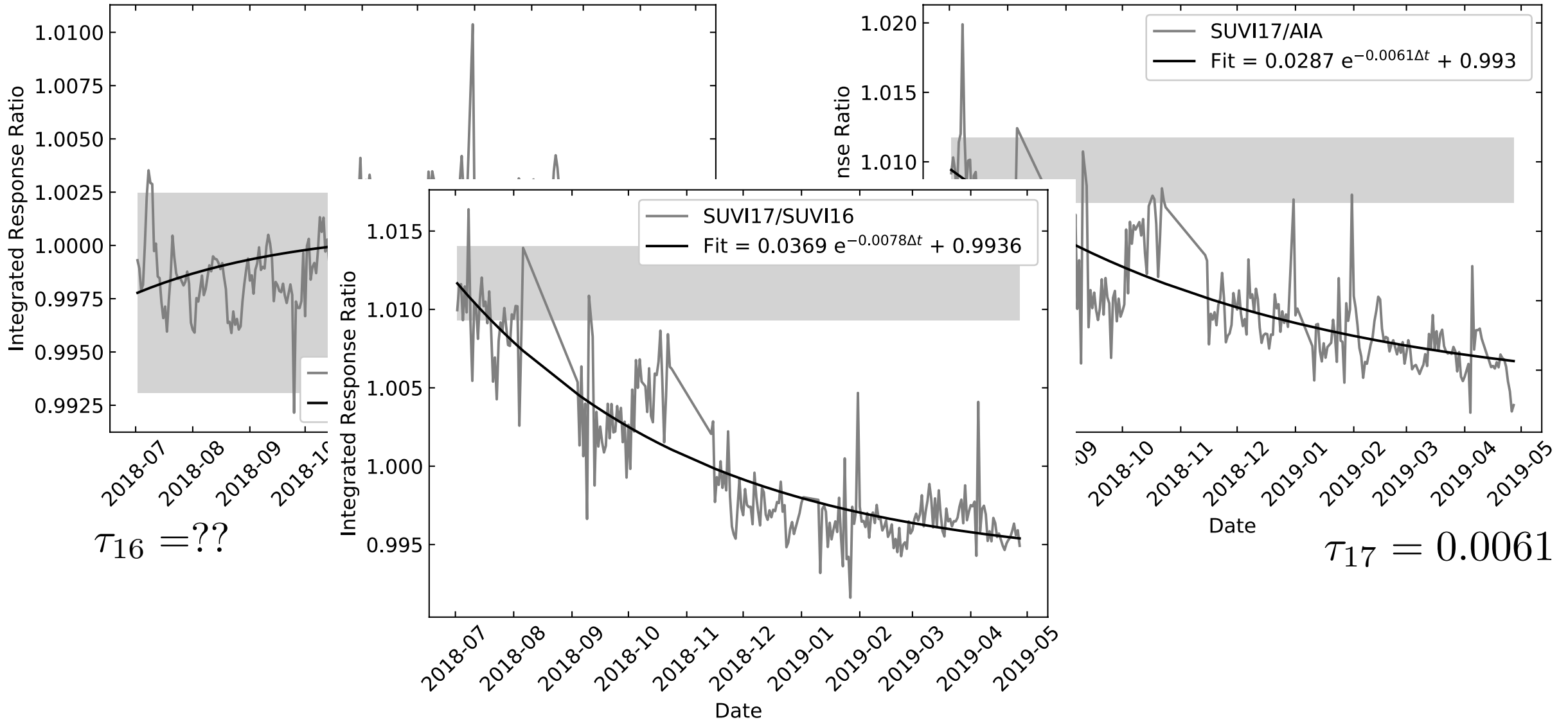
$$\tau_{16} = 0.0257$$

$$\tau_{17} = 0.0226$$

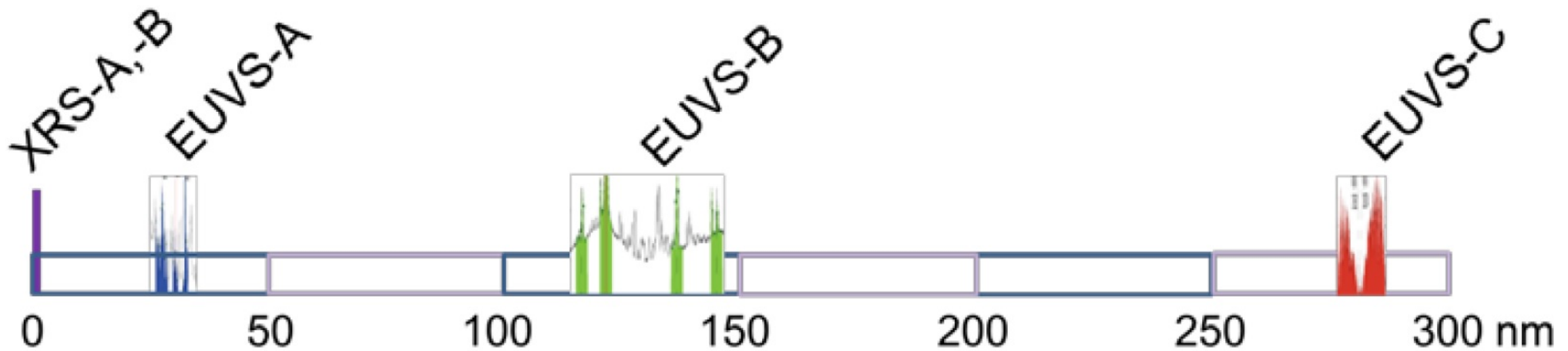
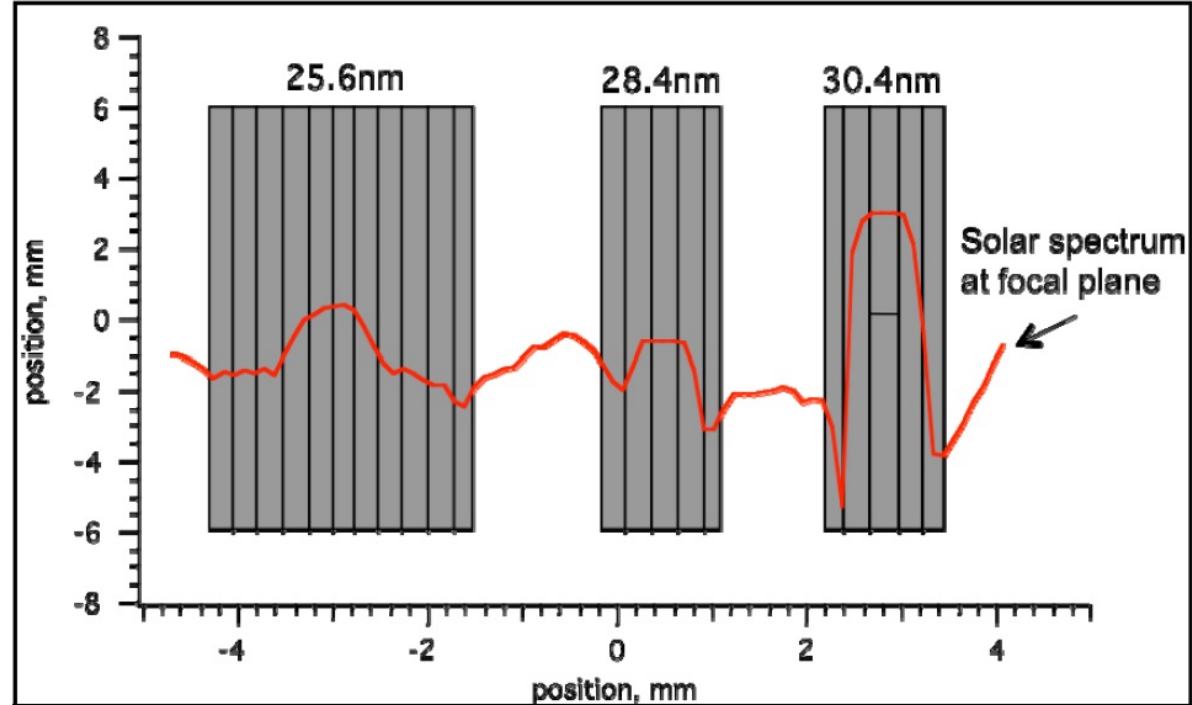
Fitting degradation trend to the data (SUVI 16 to AIA)

Can we back out the degradation coefficient?

171 Å

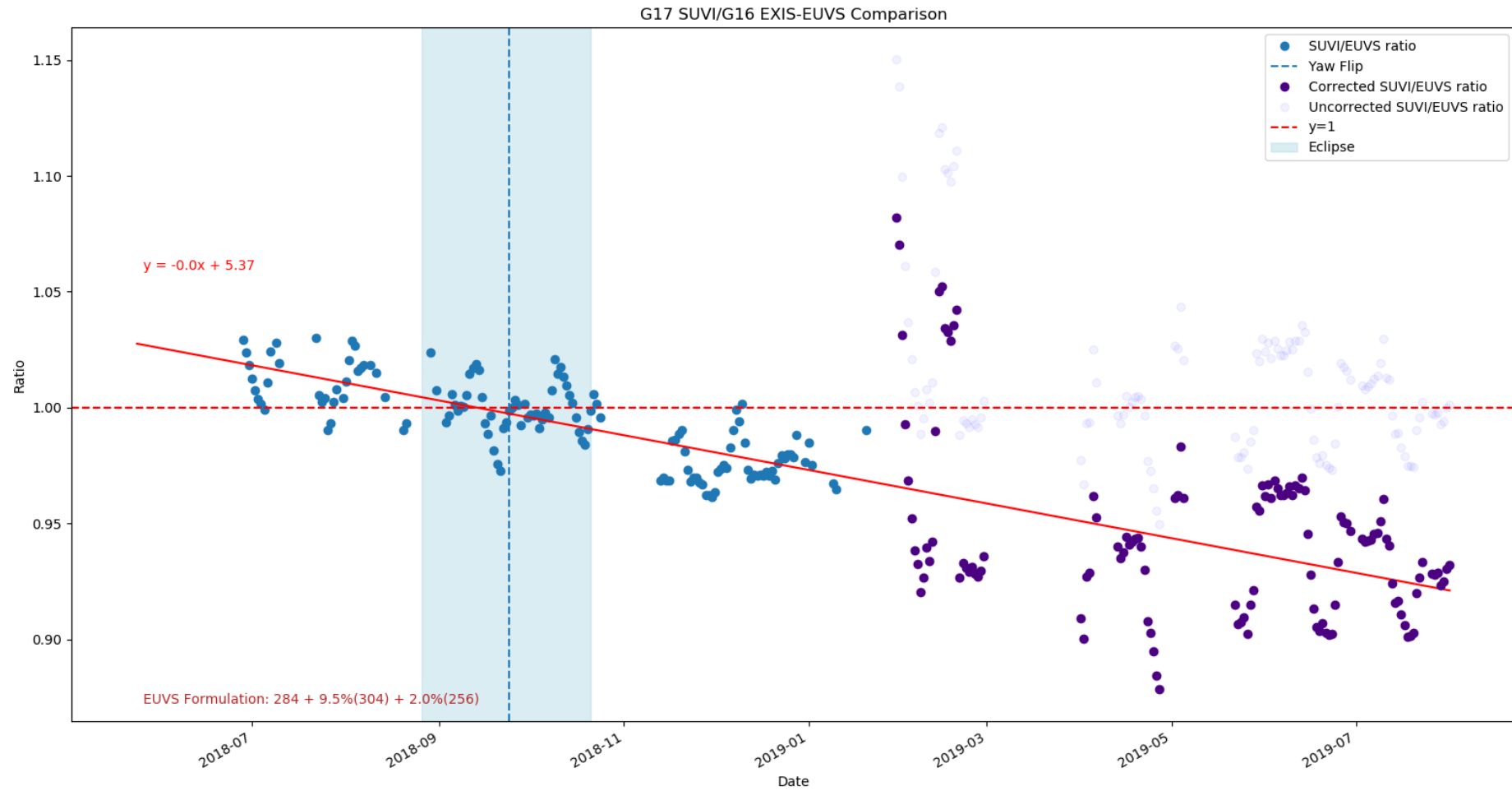


Comparison of SUVI 16 and 17 and EUVS



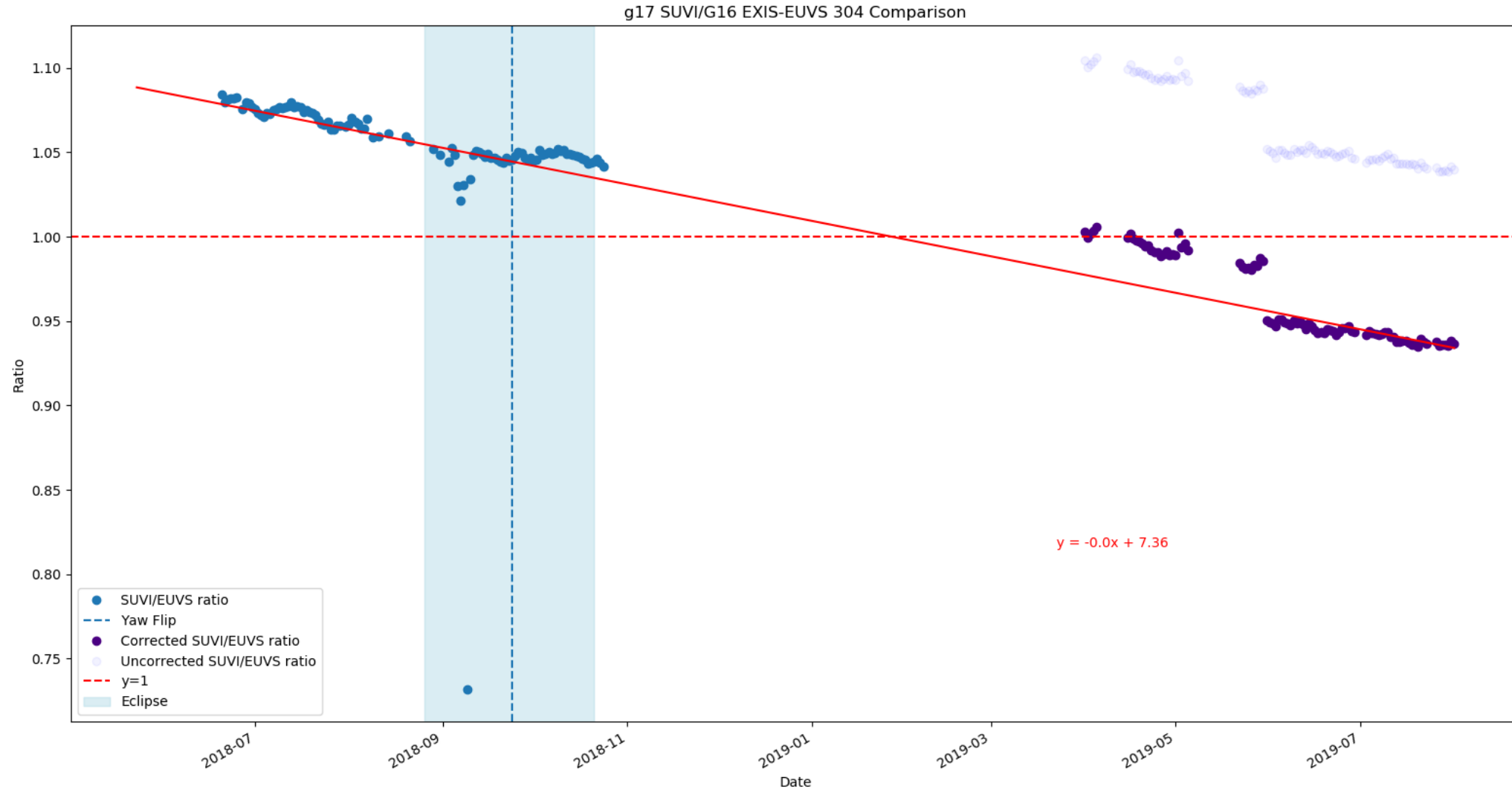
Comparison of SUVI 17 and EUVS

284 Å



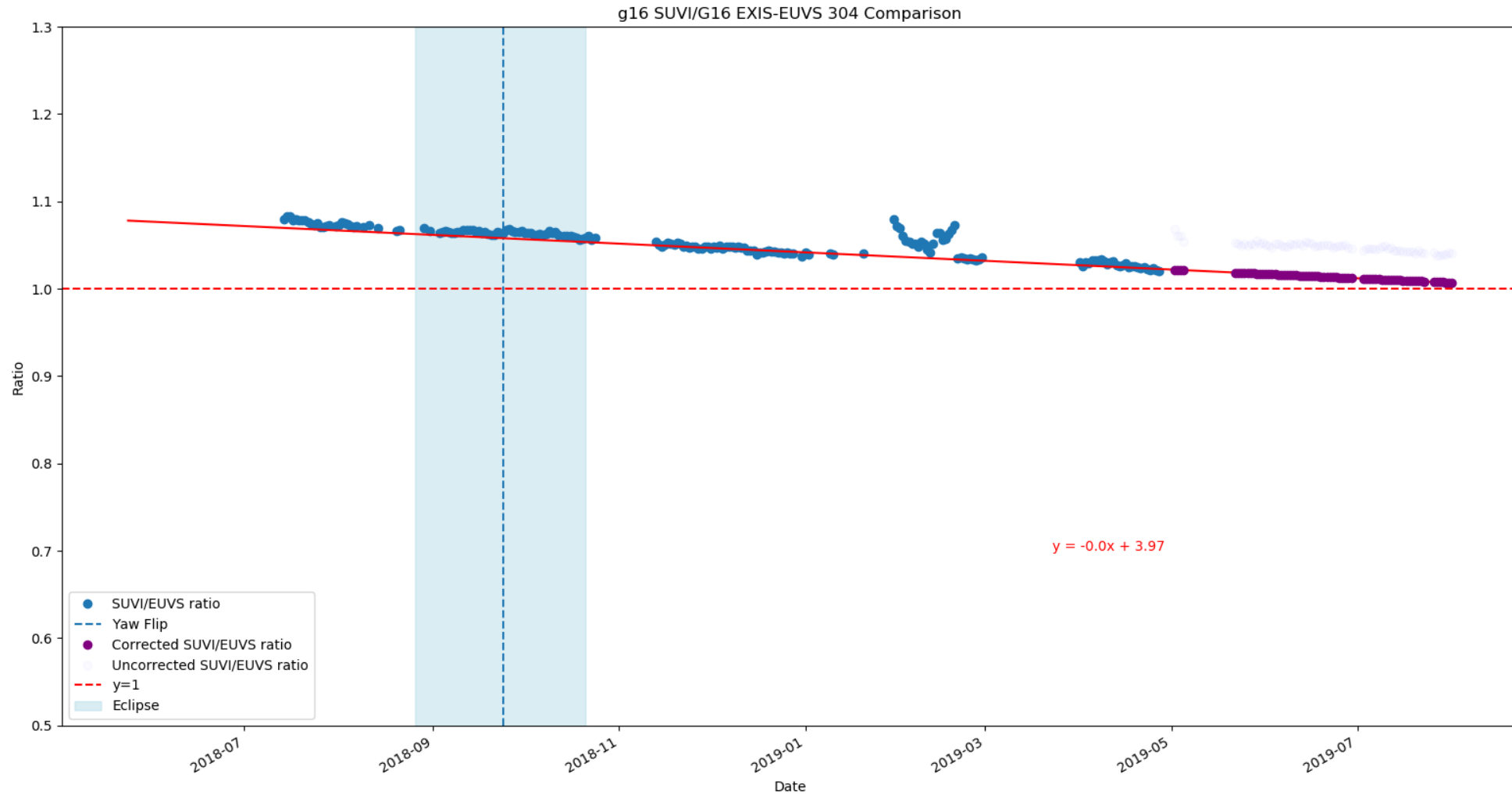
Comparison of SUVI 17 and EUVS

304 Å



Comparison of SUVI 16 and EUVS

304 Å



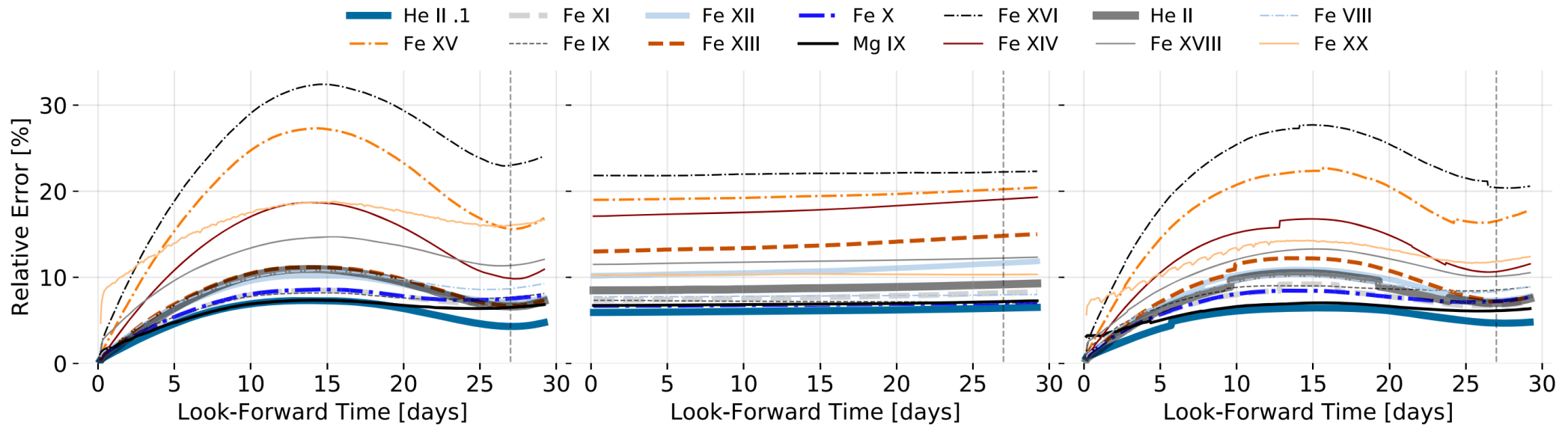
Comparison of SUVI 16 and 17 with EVE rocket underflights

June 17, 2018

	SUVI 16 L1b	Rocket conv SUVI 16	% Difference		SUVI 17 L1b	Rocket conv SUVI 17	% Difference
94 Å	1.80E+06	7.73E+05	132.86%		2.40E+06	5.91E+05	306.09%
131 Å	1.50E+07	1.44E+07	4.17%		1.57E+07	1.30E+07	20.77%
171 Å	2.80E+08	2.99E+08	-6.35%		2.80E+08	2.50E+08	12.00%
195 Å	1.51E+08	1.54E+08	-1.95%		1.77E+08	1.35E+08	31.11%
284 Å	3.40E+07	3.30E+07	3.03%		4.00E+07	3.60E+07	11.11%
304 Å	2.00E+08	1.61E+08	24.22%		2.00E+08	1.55E+08	29.03%

Future plans?

Machine learning



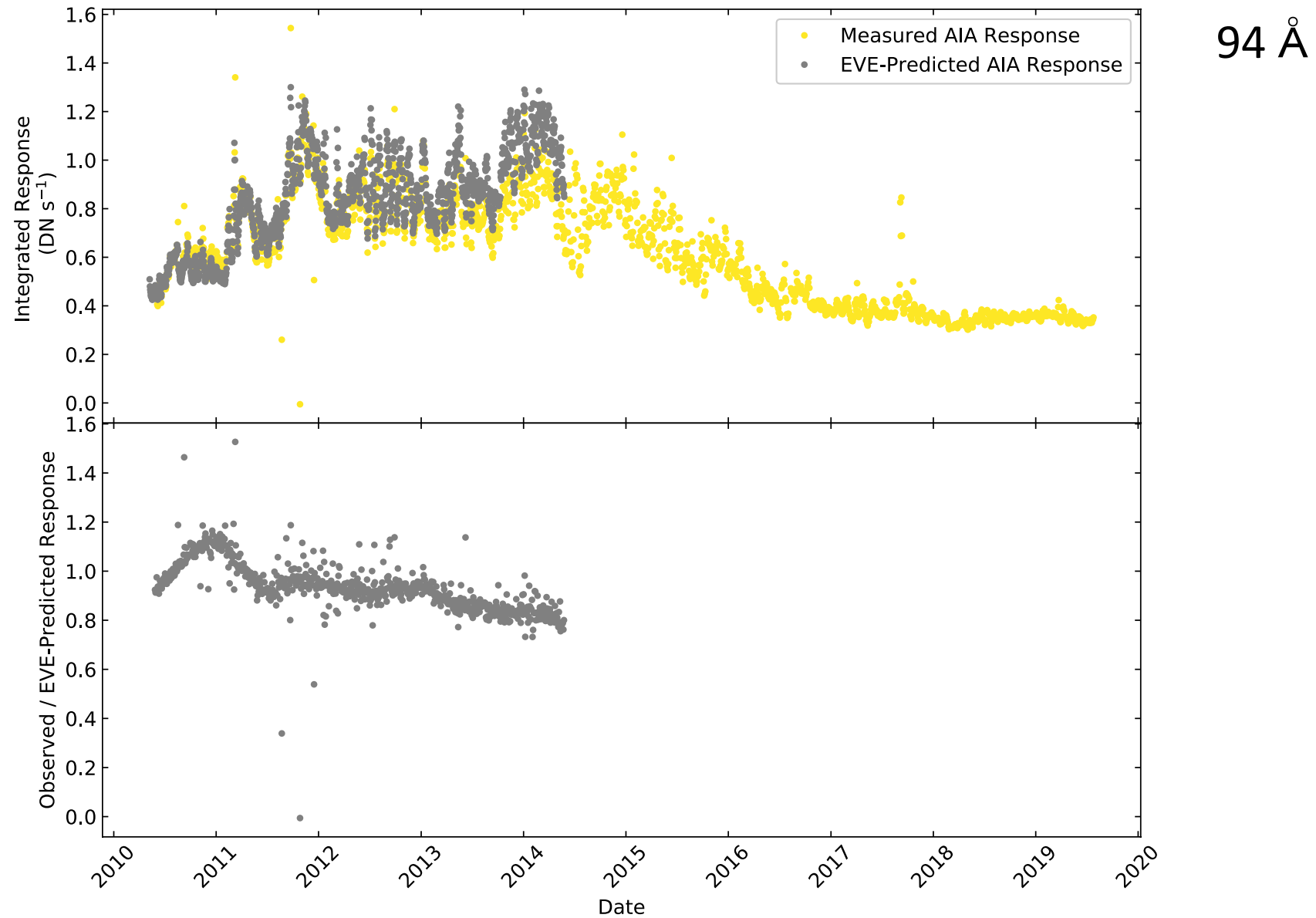
Galvez et al. 2019

Comparison between unsigned B flux and 304 Å

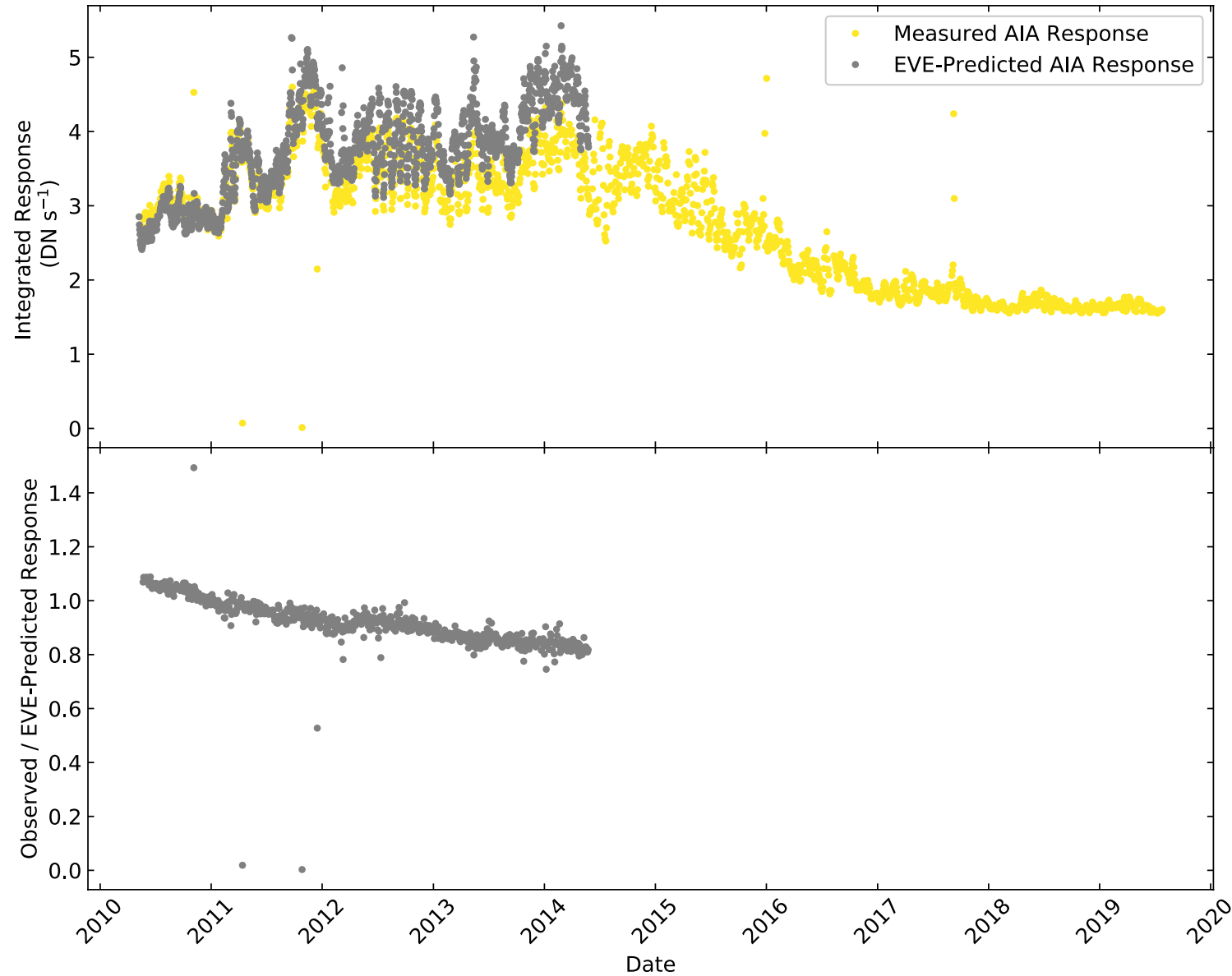
Reprocessing SUVI and EUVS data for more reliable comparisons

Extra Slides

AIA degradation and EVE integrated response comparison

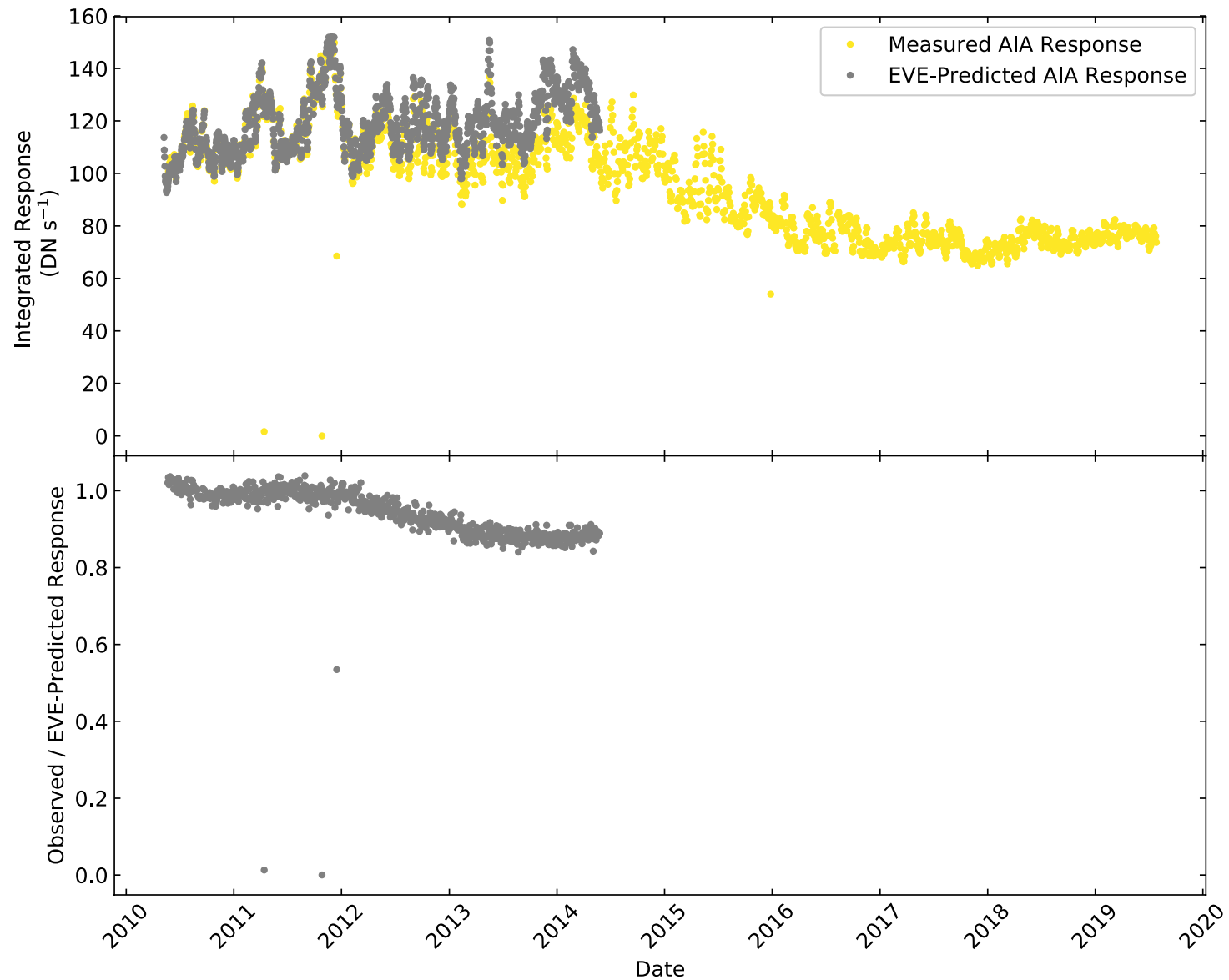


AIA degradation and EVE integrated response comparison

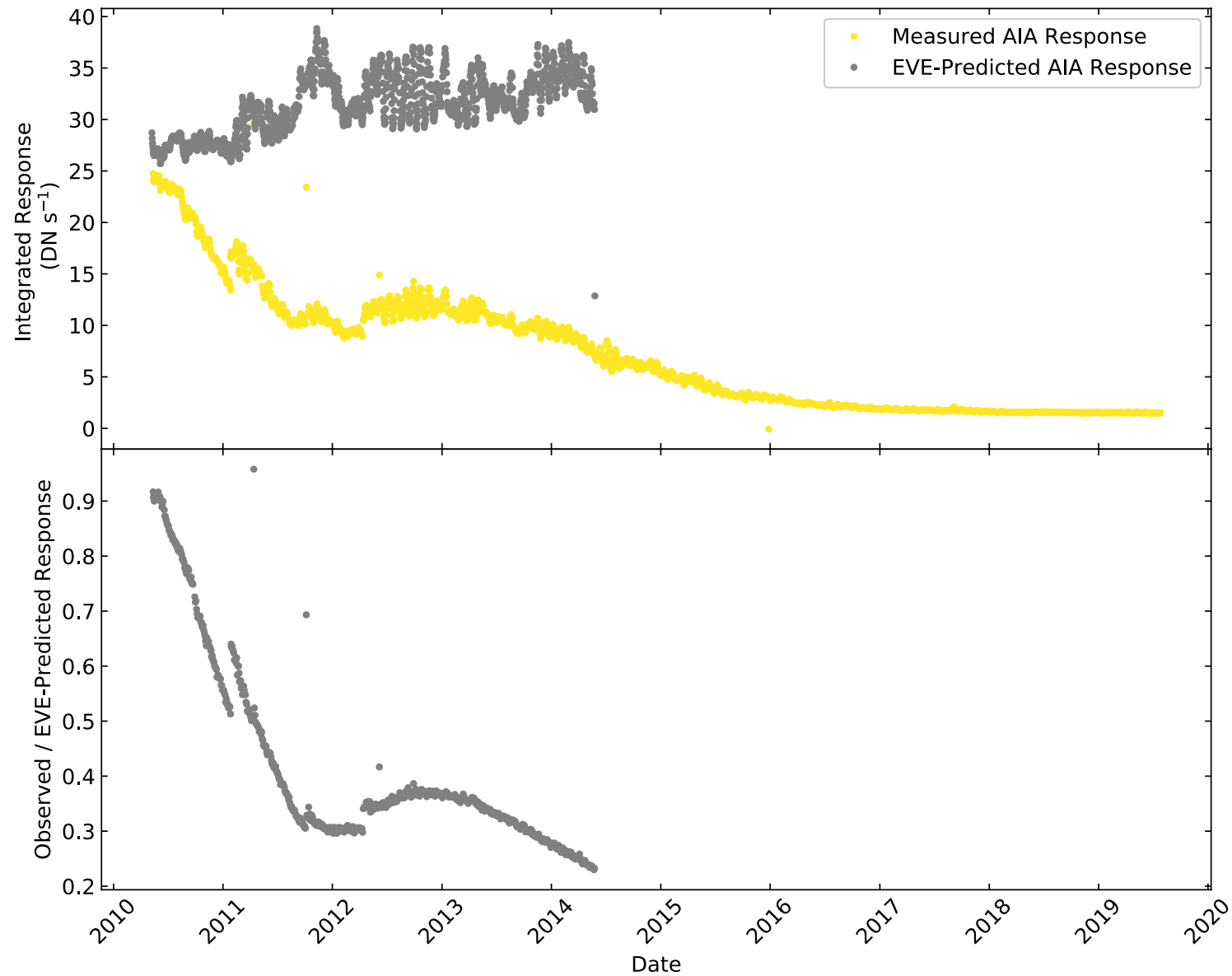


131 Å

AIA degradation and EVE integrated response comparison



AIA degradation and EVE integrated response comparison



SUVI 16 and 17 effective area change

