

# Measuring Methane on Mars

(from telescopes on Earth, from the surface of Mars, and from orbit around Mars)



**Paul Mahaffy**  
**NASA Goddard Space Flight Center**  
**September 18, 2019**

# Measuring Methane on Mars

(from telescopes on Earth, from the surface of Mars, and from orbit around Mars)

## The Martian METHANE MYSTERY

- Where does it come from?
- Why don't all the measurements agree?
- What does it all mean as we look for possible past or present life on Mars?



Paul Mahaffy  
NASA Goddard Space Flight Center  
September 18, 2019

# ARE WE ALONE IN THE UNIVERSE ?

## WHERE AND HOW DO WE SEARCH FOR LIFE ?

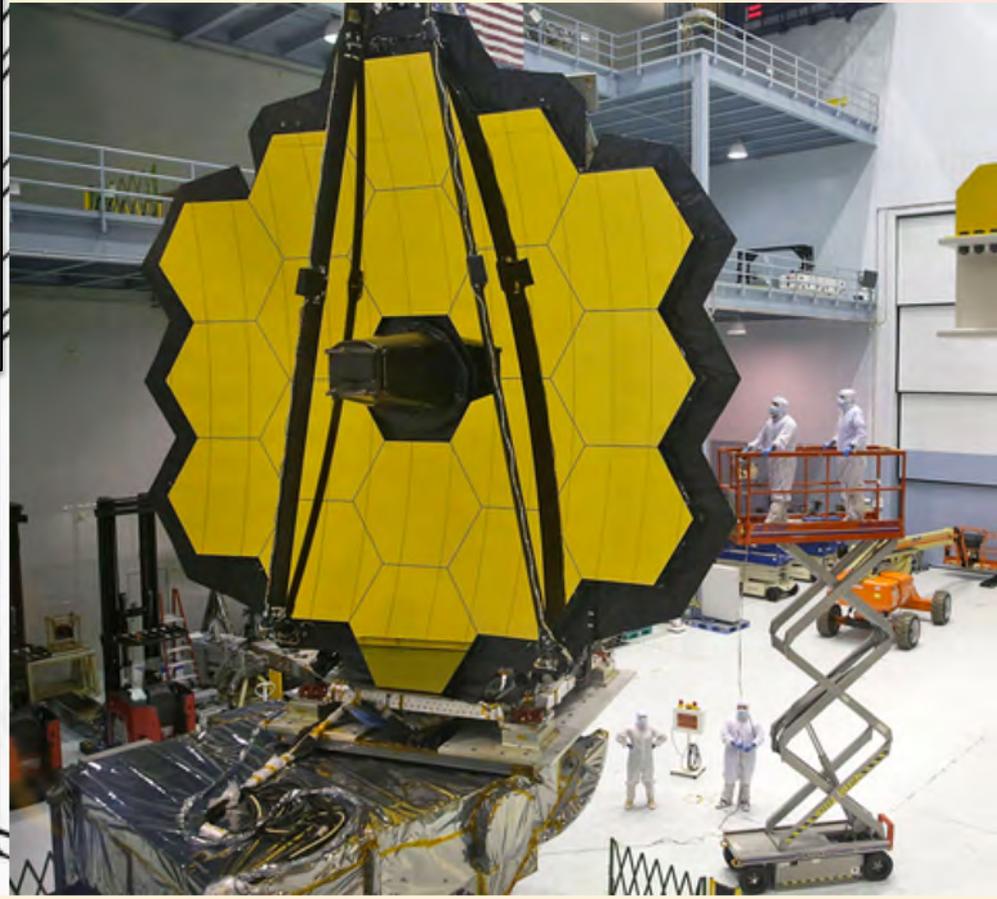
### NASA's Strategic Objectives

- Discover the Secrets of the Universe
- Search for Life Elsewhere
- Safeguard and Improve Life on Earth

Although Galileo could not detect methane with his 1610 telescope this measurement will be important for future exoplanet observations with big telescopes such as the James Webb and follow on space observatories.



Currently there are ~4,000 confirmed planets in thousands of systems, with hundreds of systems having more than one planet.



In our own solar system molecular biosignatures may reveal prebiotic chemistry or even the presence of present or past life in ocean worlds.

Measurements of methane in addition to complex organic molecules are an important element of this search



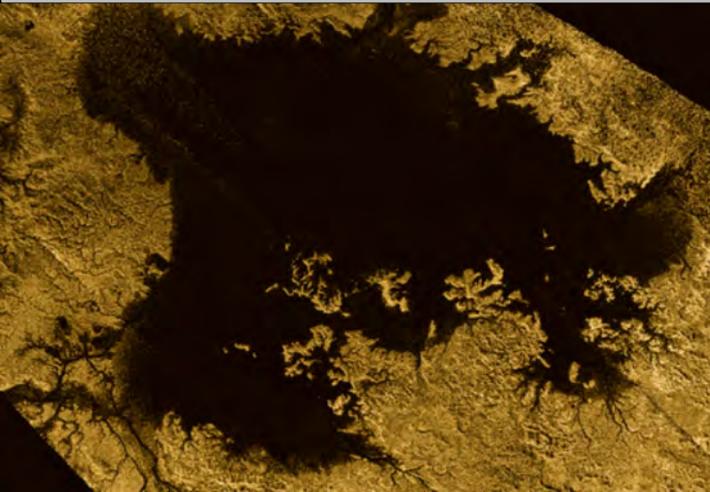
7/25/2018 Sub Glacial lake reported on Mars !

FUTURE MISSIONS TO OCEAN WORLDS  
Plumes or surface deposits at Europa or Enceladus whose oceans may contain microbial life



In our own solar system molecular biosignatures may reveal prebiotic chemistry or even the presence of present or past life in ocean worlds.

- Saturn's moon Titan has almost 5% methane in its atmosphere.
- This carbon building block together with nitrogen and hydrogen leads to a rich organic chemistry.
- But Titan is so cold that there is methane rain and methane lakes such as Ligia Mare shown below



## OCEAN WORLDS

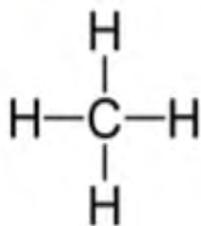
Titan's hydrocarbon-rich surface looking for chemical complexity –

This rotorcraft is being developed for NASA's latest New Frontier mission

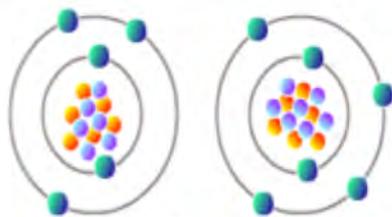
# Measurements of simple molecules, their isotopes, and complex molecules all address the grand challenge to find extraterrestrial life !

Reasonable Doubt

(or processed beyond attribution)



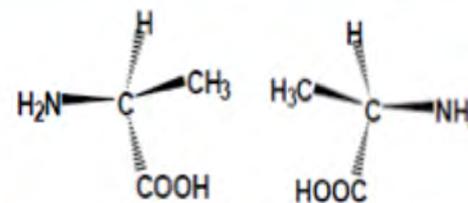
Food



Isotopes



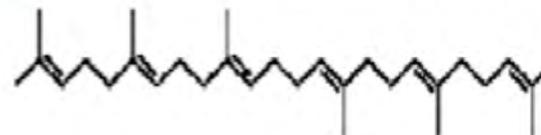
Cell-like Structures



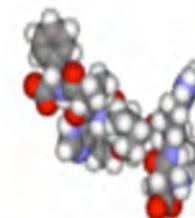
L-alanine

D-alanine

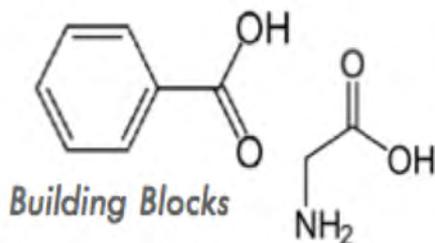
Enantiomeric Excess



Structural Preferences



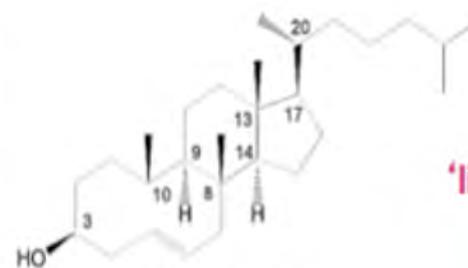
Informational Polymers



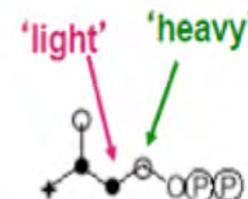
Building Blocks



Biofabrics



Structural Selectivity



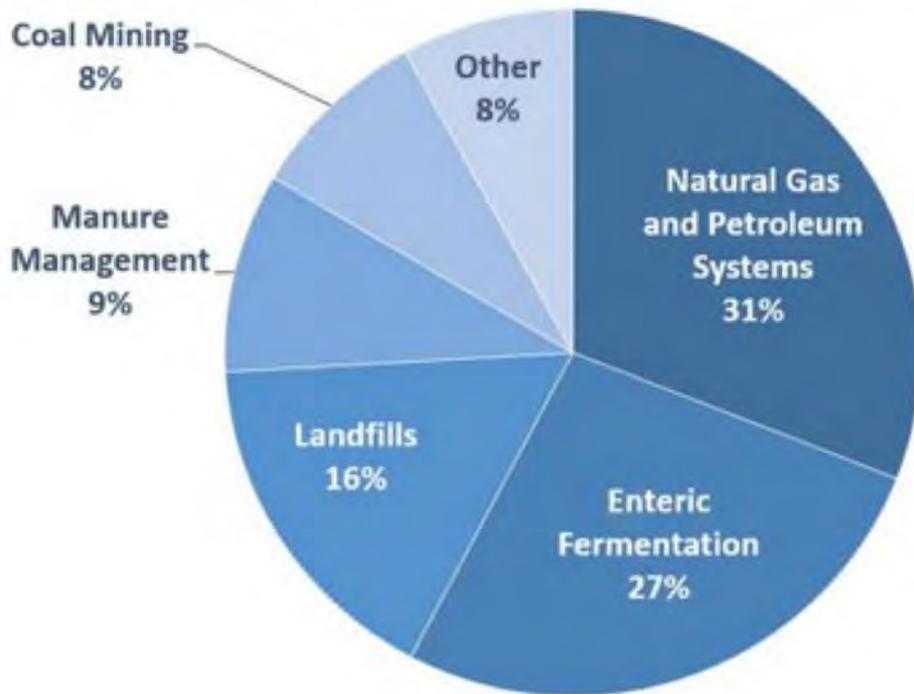
Isotopic ordering

✓ Multiple life signatures provide high confidence

✓ Abiotic signatures provide important context

Most methane in Earth's atmosphere comes from human activity  
Natural sources include volcanoes, methane clathrates, and permafrost

### 2017 U.S. Methane Emissions, By Source



Note: All emission estimates from the [Inventory of U.S. Greenhouse Gas Emissions and Sinks: 1990-2017](#).

**Lifetime in Atmosphere:** 12 years  
**Global Warming Potential (100-year):** 25<sup>1</sup>



[en.wikipedia.org/wiki/Enteric\\_fermentation](https://en.wikipedia.org/wiki/Enteric_fermentation)

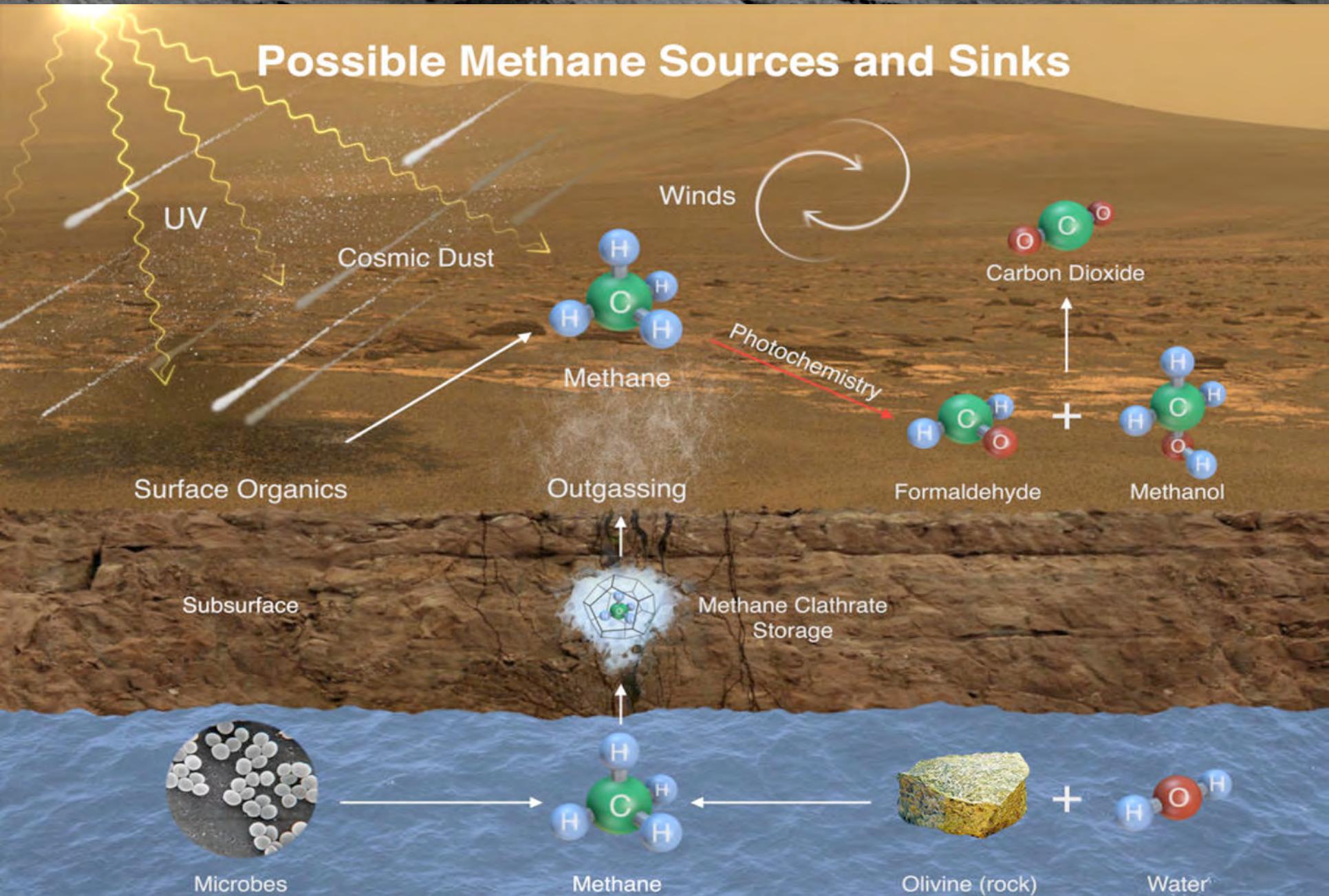
Ruminant animals are those that have a [rumen](#). A rumen is a multichambered stomach found almost exclusively among some [artiodactyl](#) mammals, such as [cattle](#), [deer](#), and [camels](#), enabling them to eat [cellulose](#)-enhanced tough plants and grains that [monogastric](#) (i.e., "single-chambered stomach") animals, such as [humans](#), [dogs](#), and [cats](#), cannot digest.

Enteric fermentation occurs when methane (CH<sub>4</sub>) is produced in the rumen as microbial fermentation takes place. Over 200 species of microorganisms are present in the rumen, although only about 10% of these play an important role in digestion. Most of the CH<sub>4</sub> byproduct is [belched](#) by the animal, however, a small percentage of CH<sub>4</sub> is also produced in the [large intestine](#) and passed out as flatulence.



BUT methane by itself on Mars is not a biosignature

## Possible Methane Sources and Sinks



# The search for martian methane from telescopes on Earth and from ESA's Mars Express Planetary Fourier Spectrometer

**Michael Mumma et al.** from 2004 observations

- 10's of parts per billion enhanced in regions of changing topography and over Valles Marineris

**Vladimir Kranopolski et al.**, from another telescope

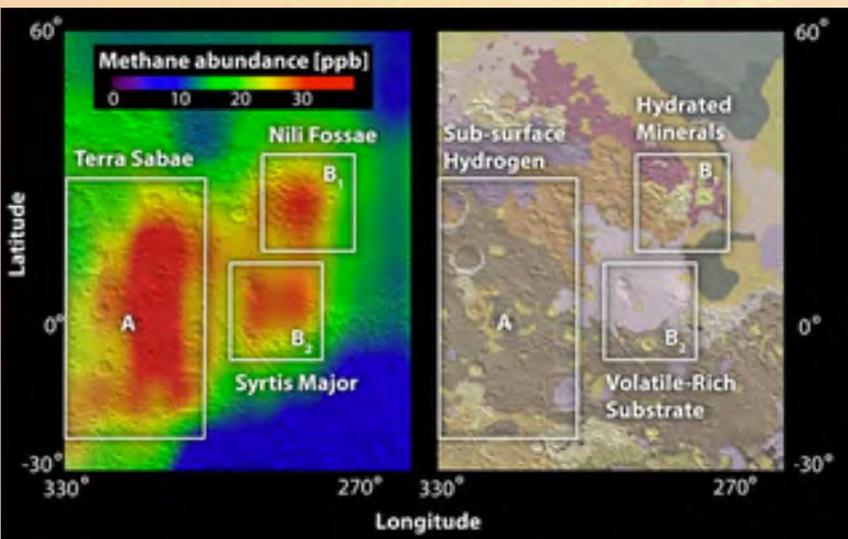
- 10 ppb suggesting 270 tons of  $\text{CH}_4$  produced per year with an atmospheric lifetime of 340 years

**Geronimo Villanueva et al.** from later observations

- Upper limit of 7 ppb

**Vittorio Formisano et al.** from Mars Express

- 15 ppb secured by averaging more than 15,000 spectra
- Suggested source was northern cap summer release



# These observations immediately raised some great questions

*Mischna et al.*

*LeFevre and Forget*

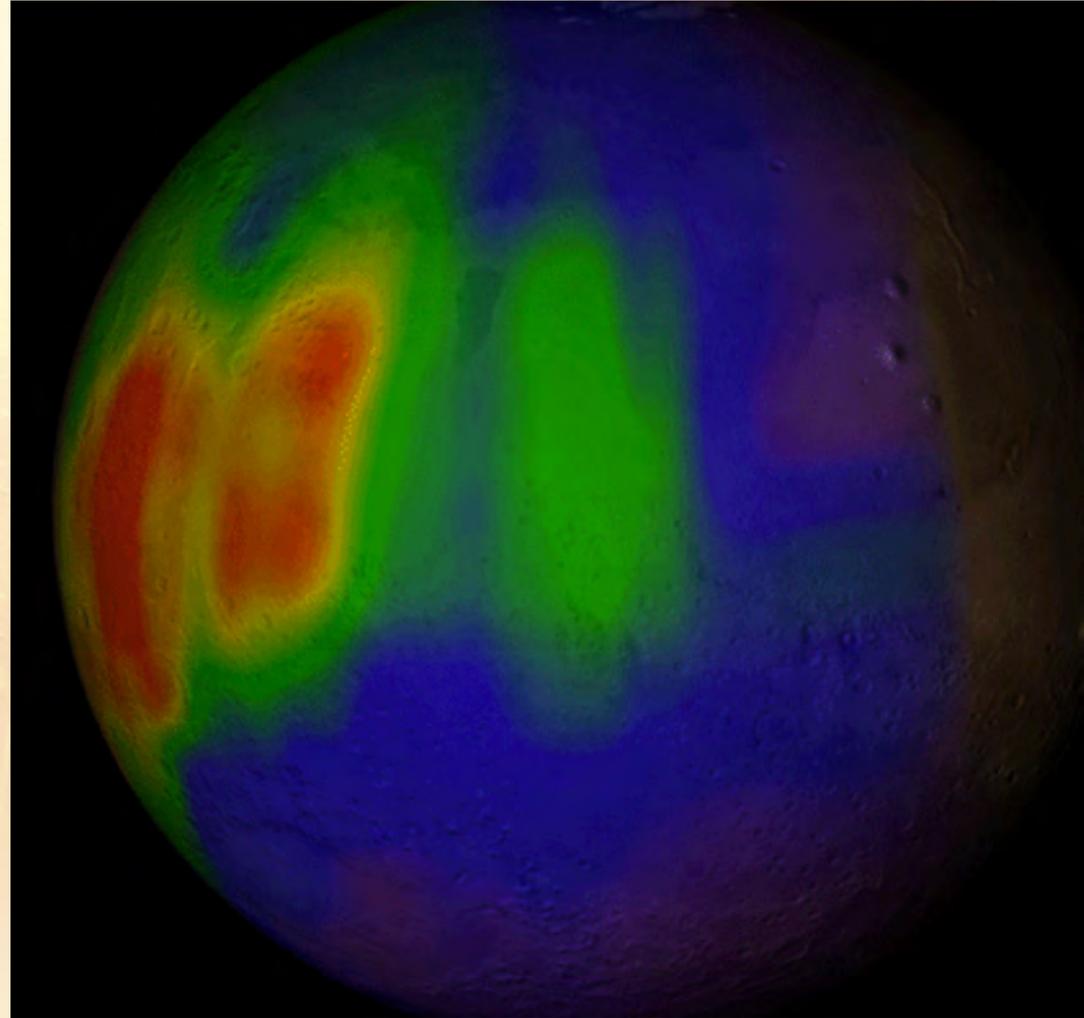
**How could you have these gradients if the methane lifetime is really greater than 300 years?**

**Would not a persistent plume require a release over months and a faster decay than photochemistry would allow?**

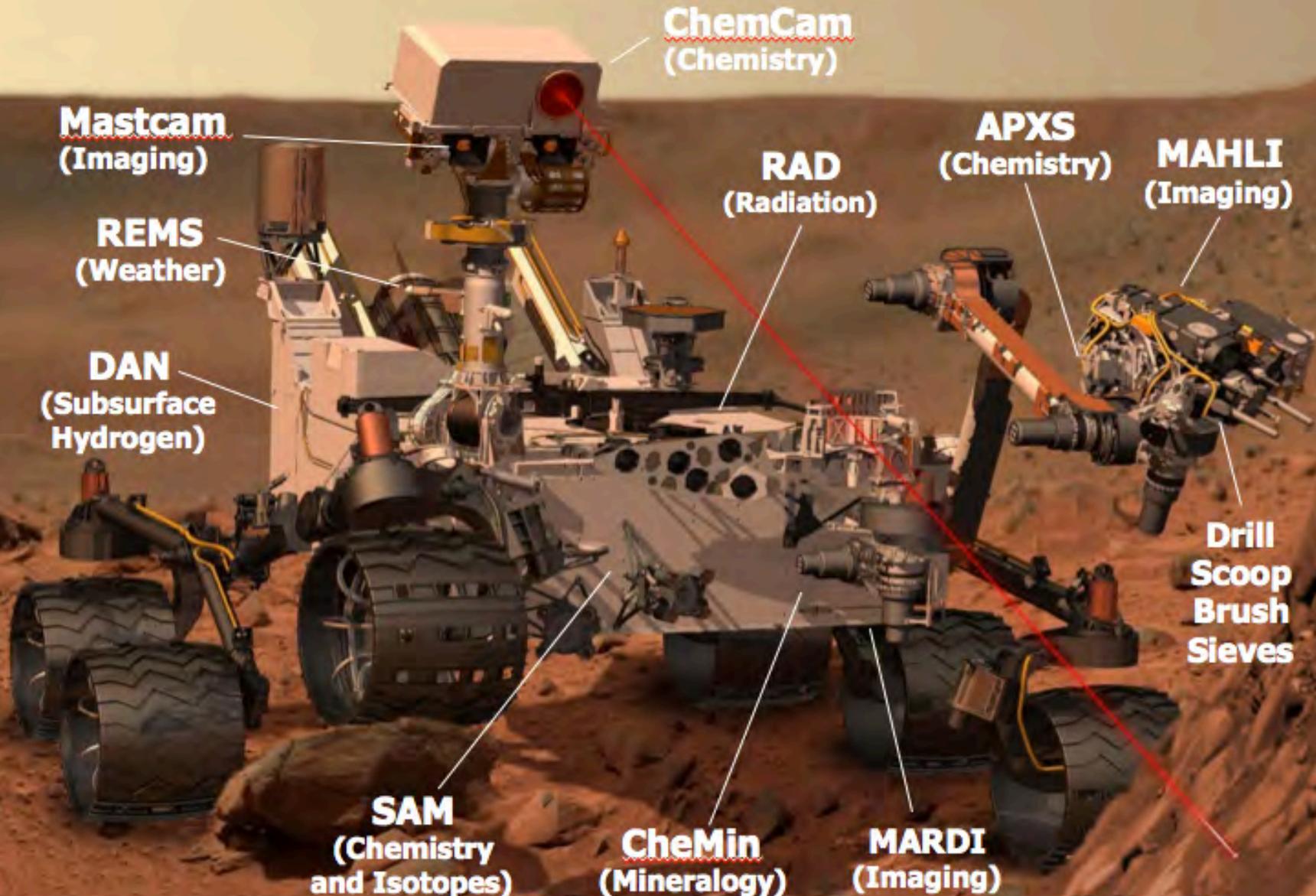
*Zahnle et al.*

**Was the terrestrial methane properly accounted for?**

**If chemical oxidation what was the source of the oxygen since over thousands of years the atmospheric oxygen would be consumed?**



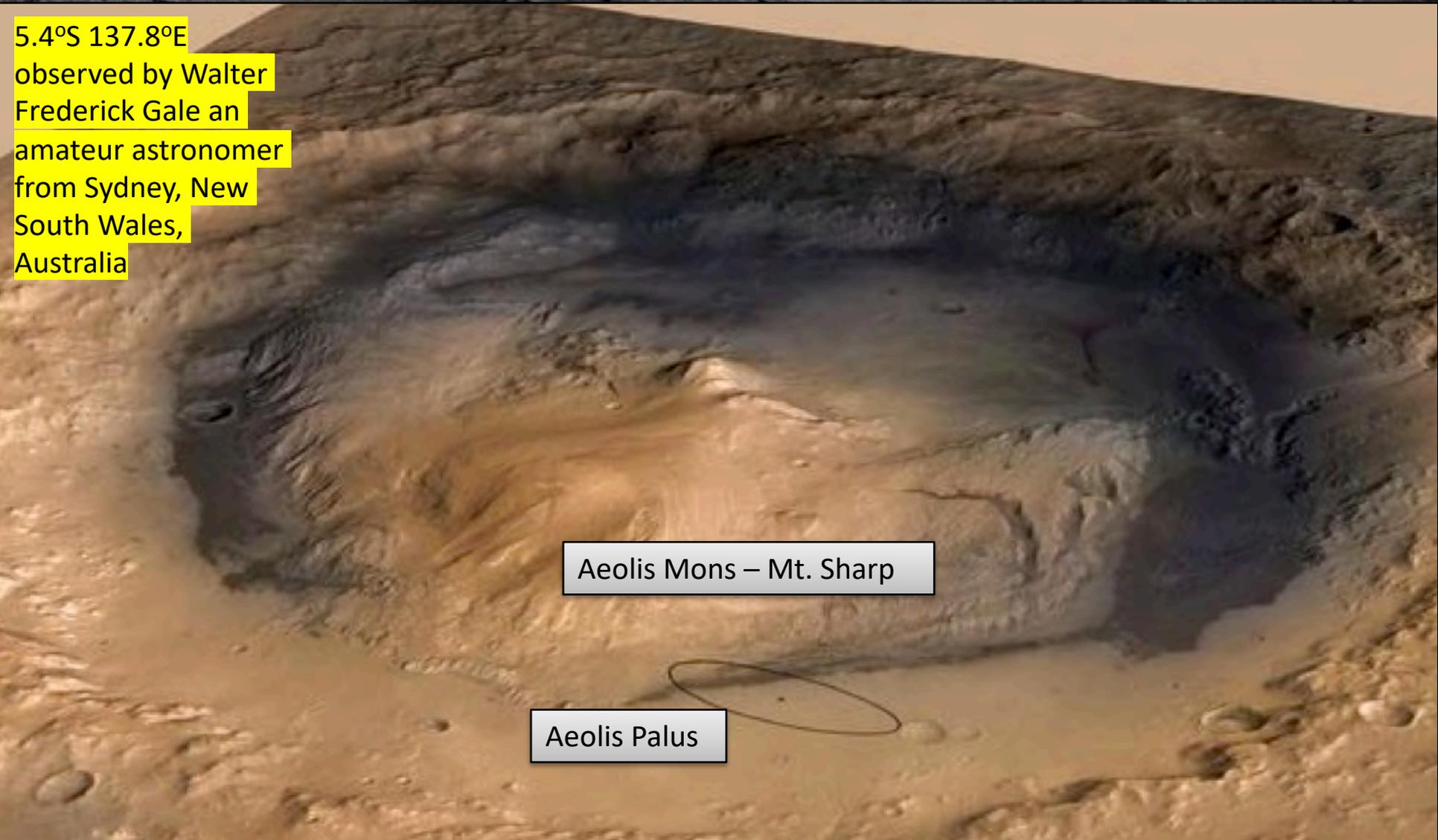
Curiosity's primary scientific goal is to explore and quantitatively assess a local region on Mars' surface as a potential habitat for life, past or present.



# CAN ORGANICS FROM A CRATER FORMED MORE THAN BE PRESERVED ON MARS ???

5.4°S 137.8°E

observed by Walter  
Frederick Gale an  
amateur astronomer  
from Sydney, New  
South Wales,  
Australia



Aeolis Mons – Mt. Sharp

Aeolis Palus

Gale Crater is a 150 km diameter crater with a central 5 km high mound

Curiosity's primary scientific goal is to explore and quantitatively assess a local region on Mars' surface as a potential habitat for life, past or present.

**Methane is only one of the measurements that give insight into chemical processes related to possible life on early Mars when surface water was more abundant**

(1) Mineralogy and geology

→ **what is it telling us about the ancient crater environment ?**

(2) Isotopic composition of light elements

→ **is isotopic fractionation from biological or physical/chemical processes ?**

(3) Age of formation of rocks

→ **how long have aqueous conditions persisted on Mars ?**

(4) Cosmic radiation of surface materials

→ **are near surface organics preserved from cosmic radiation ?**

(5) Clays, perchlorates, sulfates, and hydrated minerals

→ **do chemical environments enable preservation or destruction of organics ?**

(6) Organic compounds extracted from soils and rocks

→ **what have we learned so far ?**

(7) Atmospheric composition and methane

→ **what does methane have to do with present or past life?**

# Sample Analysis at Mars (SAM) Instrument Suite

**SAM suite instruments and major subsystems:**

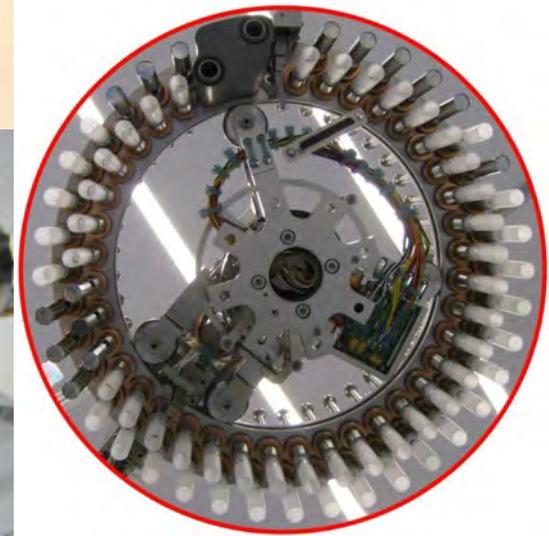
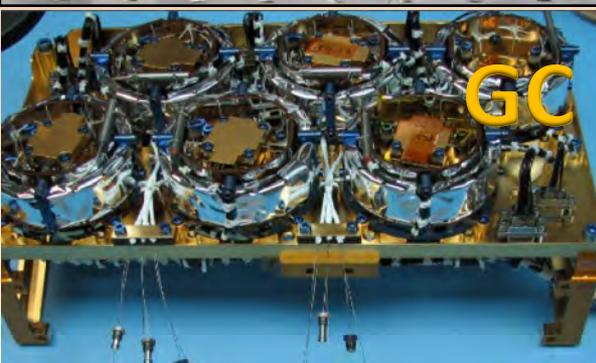
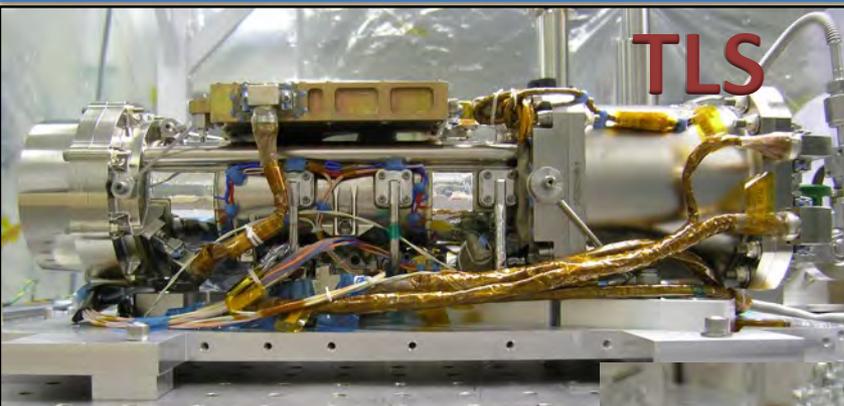
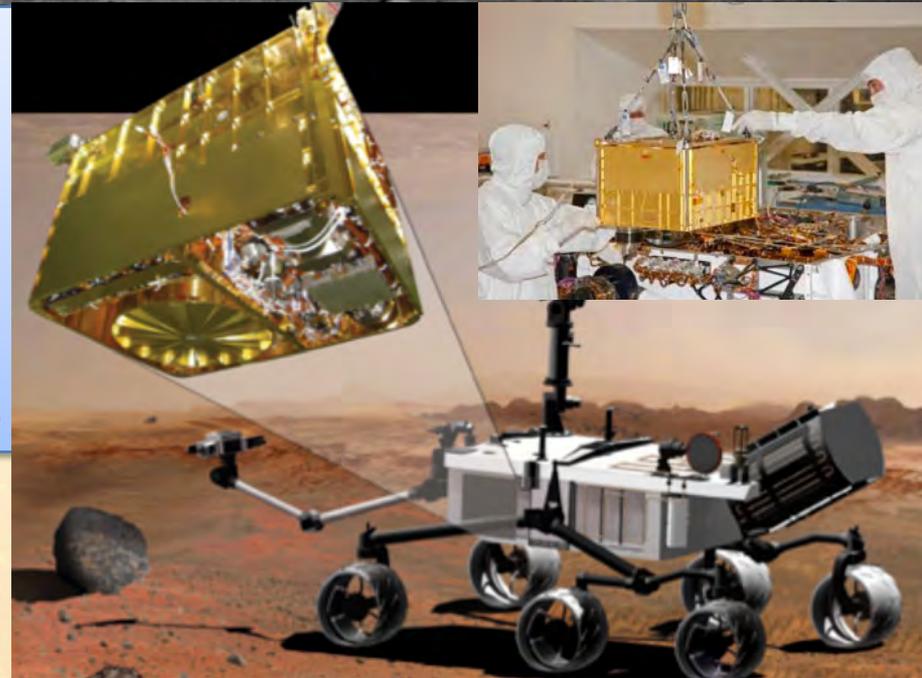
**Quadrupole Mass Spectrometer-NASA GSFC**

**Gas Chromatograph-University of Paris, France**

**Tunable Laser Spectrometer-JPL**

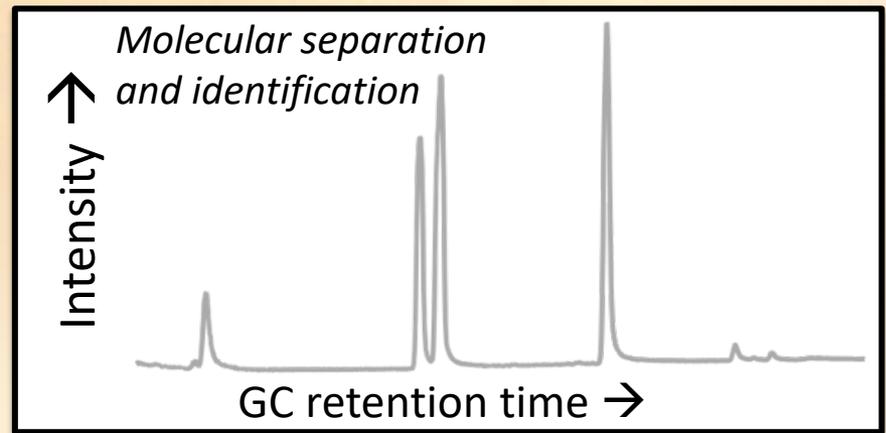
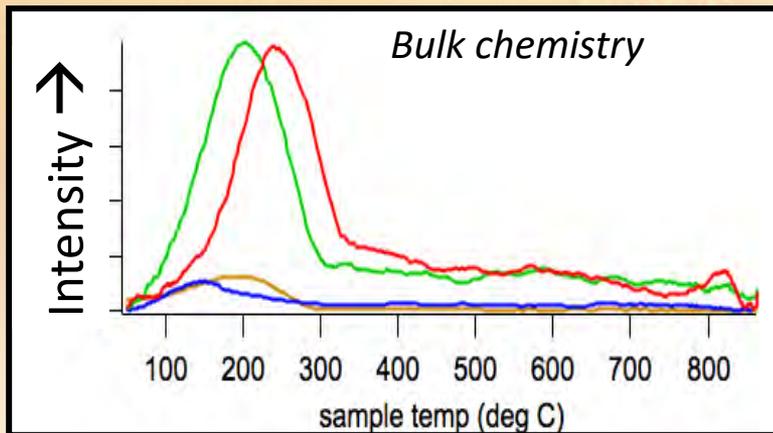
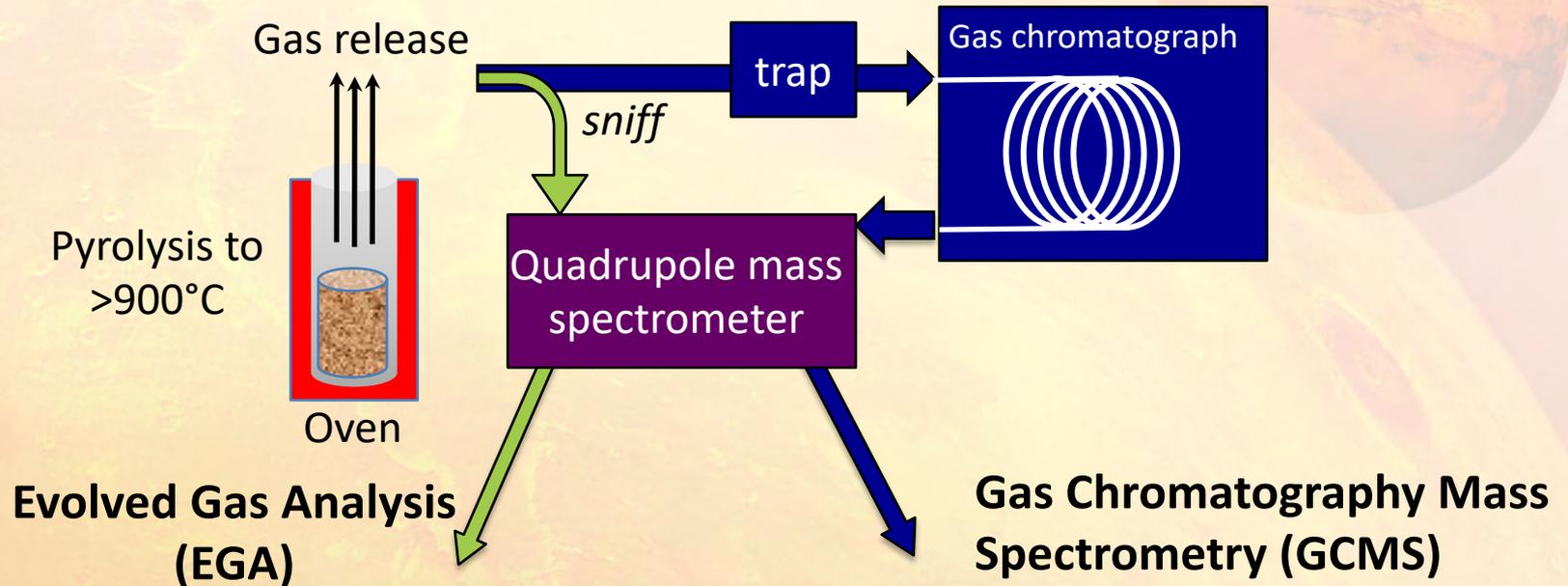
**Gas Processing System-NASA GSFC**

**Sample Manipulation System-Honeybee Robotics**

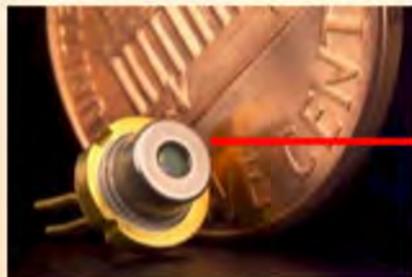




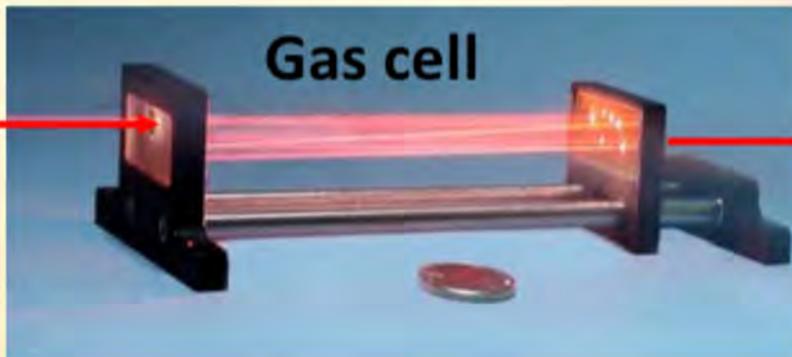
# SAM Detects Volatiles Released from Sample



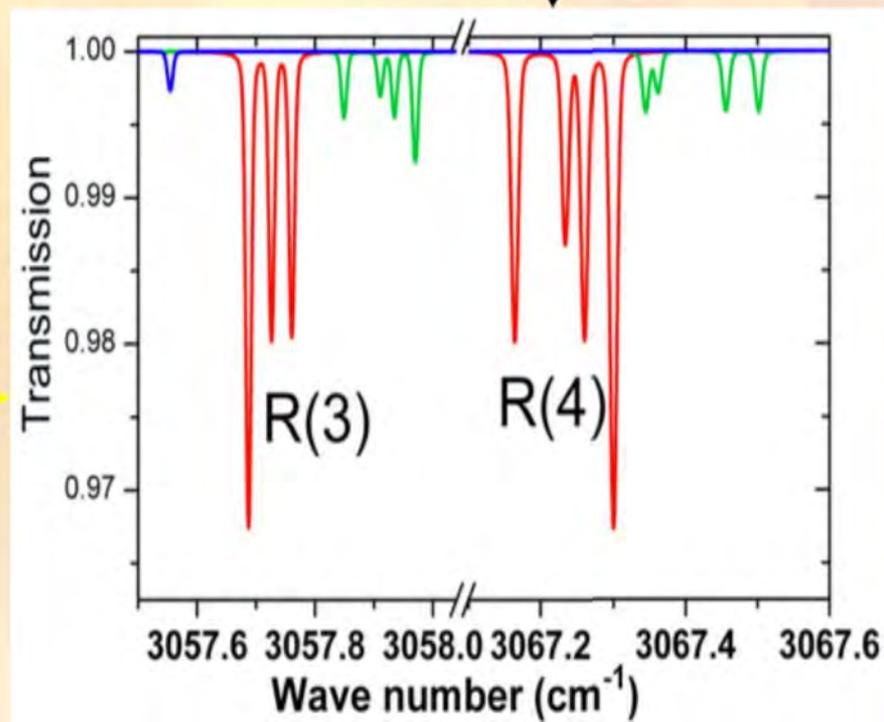
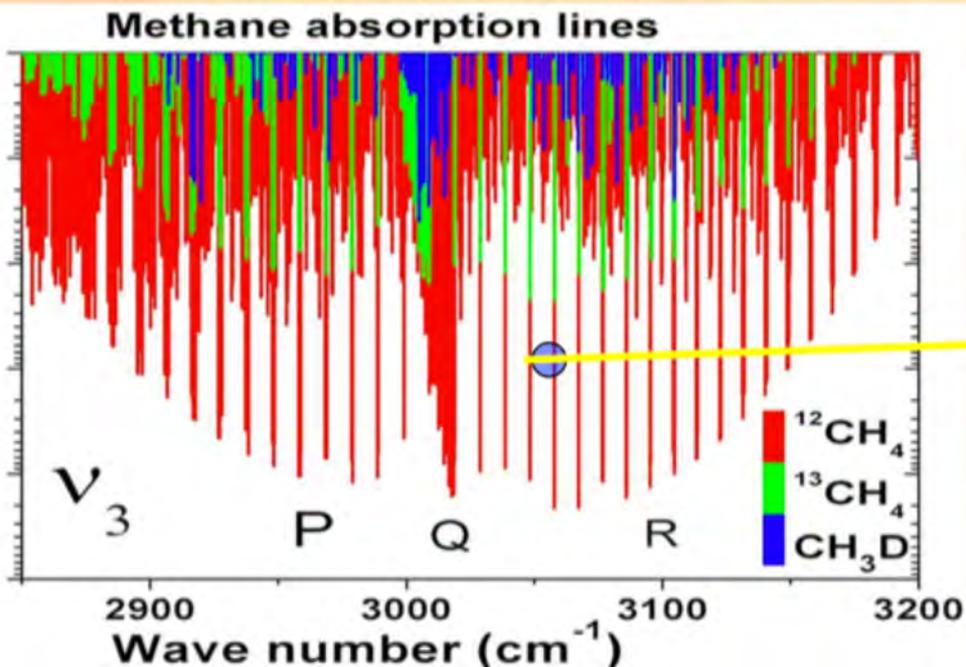
# SAM's Tunable Laser Spectrometer (TLS) methane detector

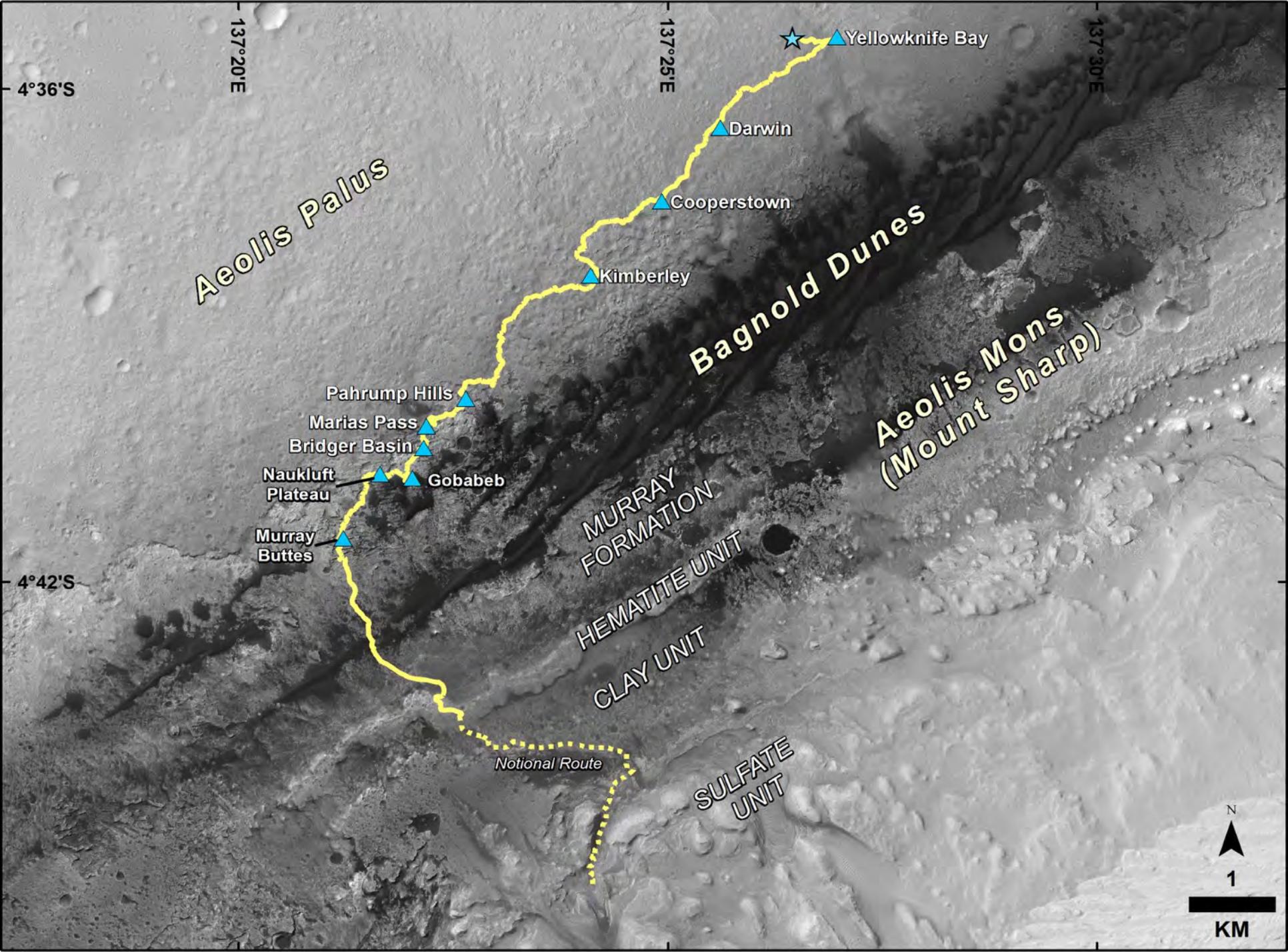


Tunable laser 1-12  $\mu\text{m}$



Detector







John Klein  
Sol 182



Cumberland  
Sol 279



Windjana  
Sol 621



Confidence Hills  
Sol 759



Mojave  
Sol 882



Telegraph Peak  
Sol 908



Buckskin  
Sol 1060



Big Sky  
Sol 1119



Greenhorn  
Sol 1137



Lubango  
Sol 1320



Okoruso  
Sol 1332



Oudam  
Sol 1361



Marimba  
Sol 1422



Quela  
Sol 11464



Sebina  
Sol 1495

MAHLI images credit  
NASA/MSSS/ JPL  
arranged by  
Lakdawalia &  
Sorenson

Curiosity's primary scientific goal is to explore and quantitatively assess a local region on Mars' surface as a potential habitat for life, past or present.

**Measurements that give insight into chemical processes related to possible life on early Mars when surface water was more abundant**

(1) Mineralogy and geology

→ **what is it telling us about the ancient crater environment ?**

(2) Isotopic composition of light elements

→ **is isotopic fractionation from biological or physical/chemical processes ?**

(3) Age of formation of rocks

→ **how long have aqueous conditions persisted on Mars ?**

(4) Cosmic radiation of surface materials

→ **are near surface organics preserved from cosmic radiation ?**

(5) Clays, perchlorates, sulfates, and hydrated minerals

→ **do chemical environments enable preservation or destruction of organics ?**

(6) Organic compounds extracted from soils and rocks

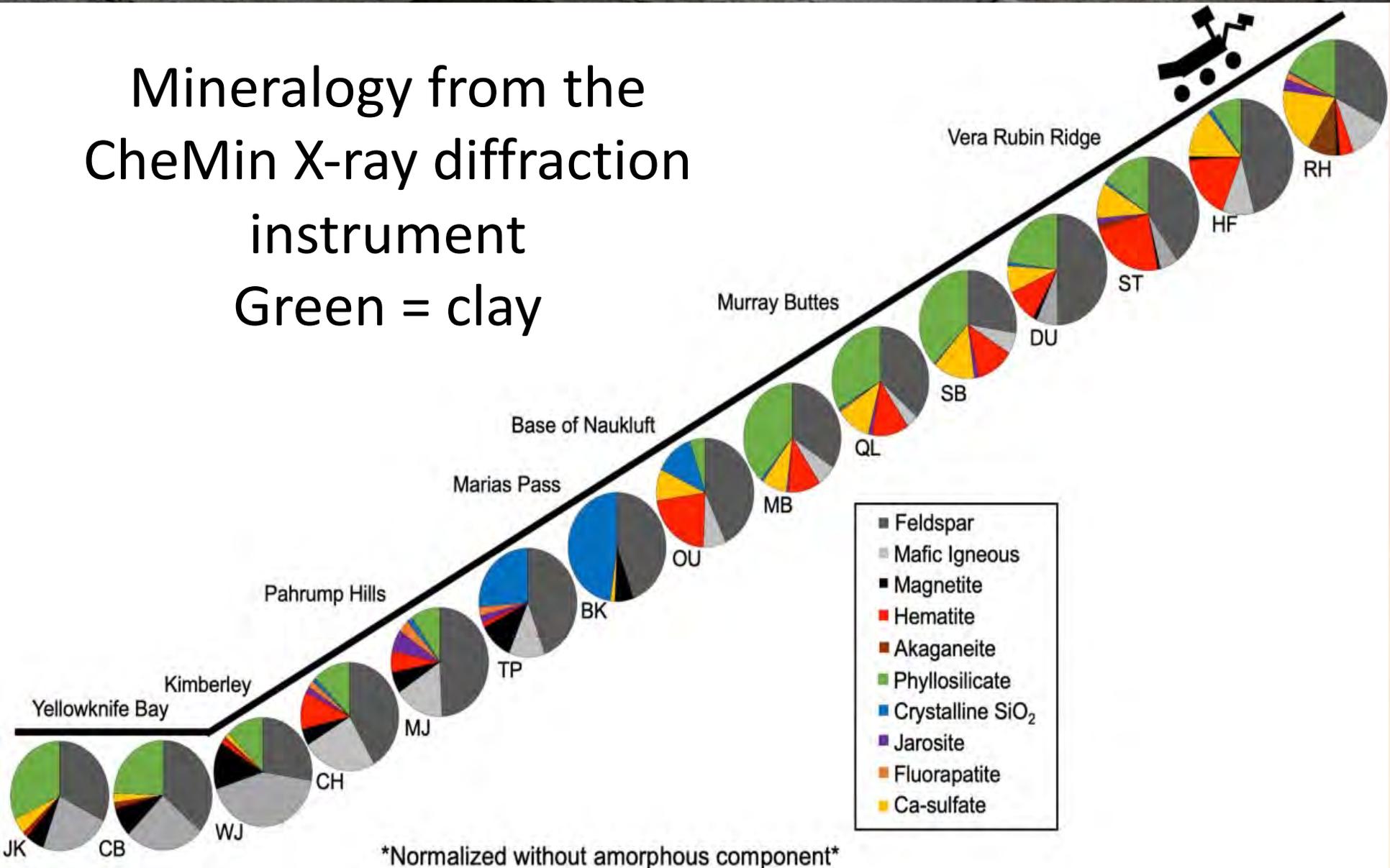
→ **what have we learned so far ?**

(7) Atmospheric composition and methane

→ **what does methane have to do with present or past life?**

Changes in mineralogy with elevation → differences in sediment source regions, rock-forming processes, and subsequent alteration.

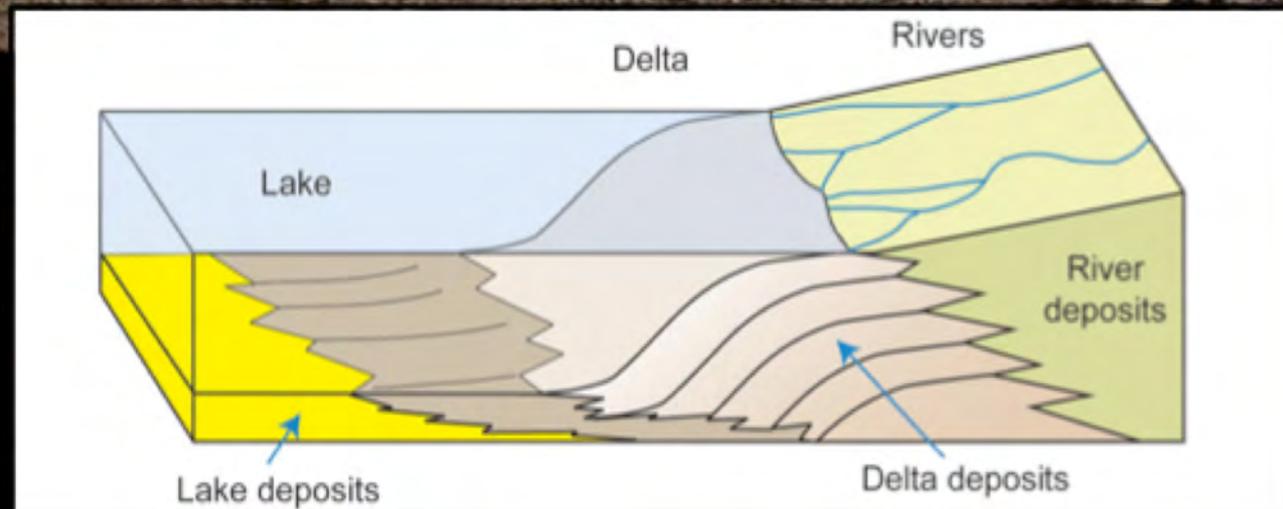
# Mineralogy from the CheMin X-ray diffraction instrument Green = clay



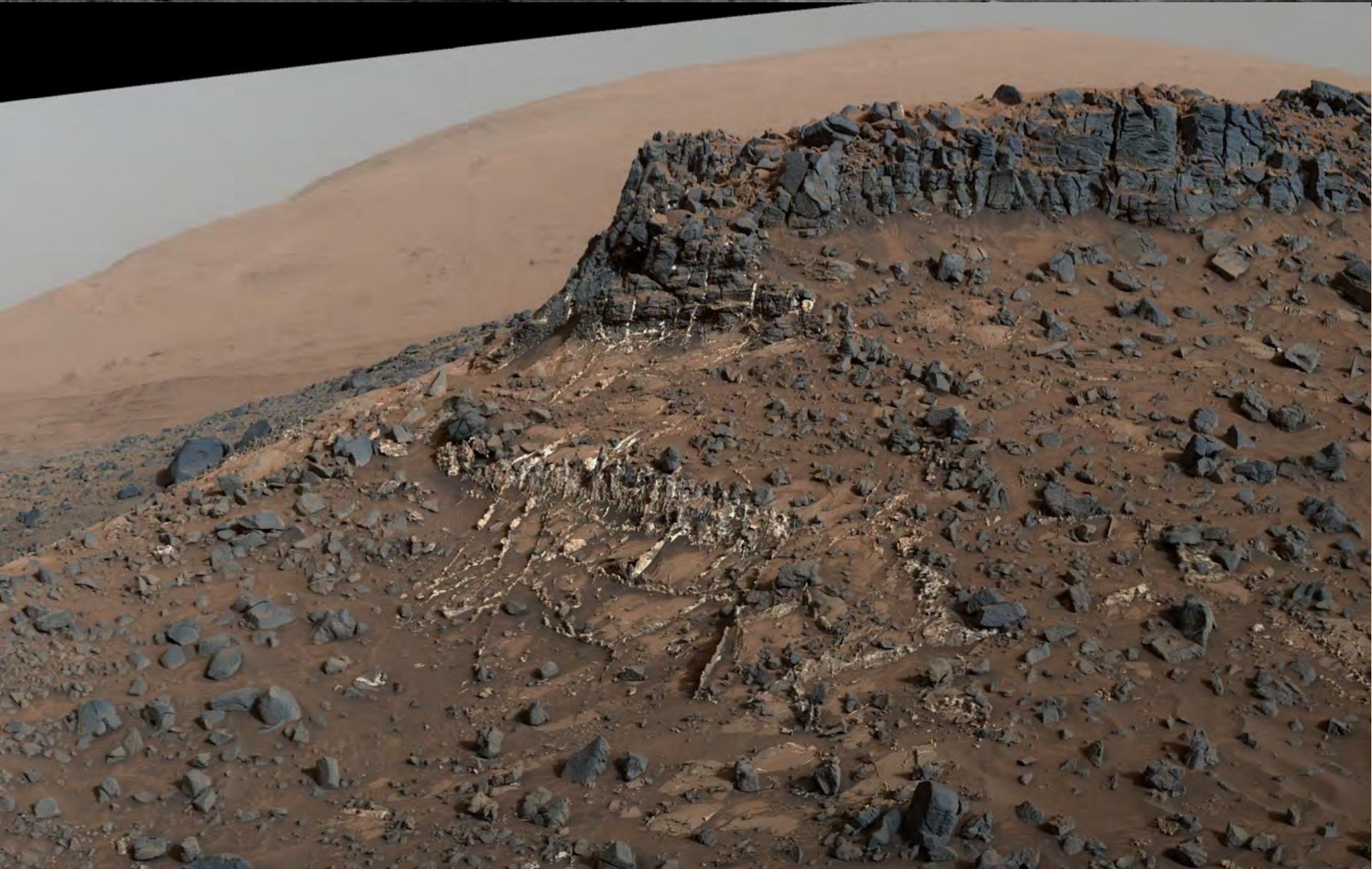
Sandstone beds on Gale Crater's plains indicate water-driven transport of sediment, building lower Mount Sharp from lake deposits



Possible lake deposits at the base of Mount Sharp



Mineral ( $\text{CaSO}_4$  & B-rich) veins of fill fractures within the lake-formed mudstone  $\rightarrow$  multiple generations of interaction with liquid water



Curiosity's primary scientific goal is to explore and quantitatively assess a local region on Mars' surface as a potential habitat for life, past or present.

**Measurements that give insight into chemical processes related to possible life on early Mars when surface water was more abundant**

(1) Mineralogy and geology

→ what is it telling us about the ancient crater environment ?

(2) Isotopic composition of light elements

→ is isotopic fractionation from biological or physical/chemical processes ?

(3) Age of formation of rocks

→ how long have aqueous conditions persisted on Mars ?

(4) Cosmic radiation of surface materials

→ are near surface organics preserved from cosmic radiation ?

(5) Clays, perchlorates, sulfates, and hydrated minerals

→ do chemical environments enable preservation or destruction of organics ?

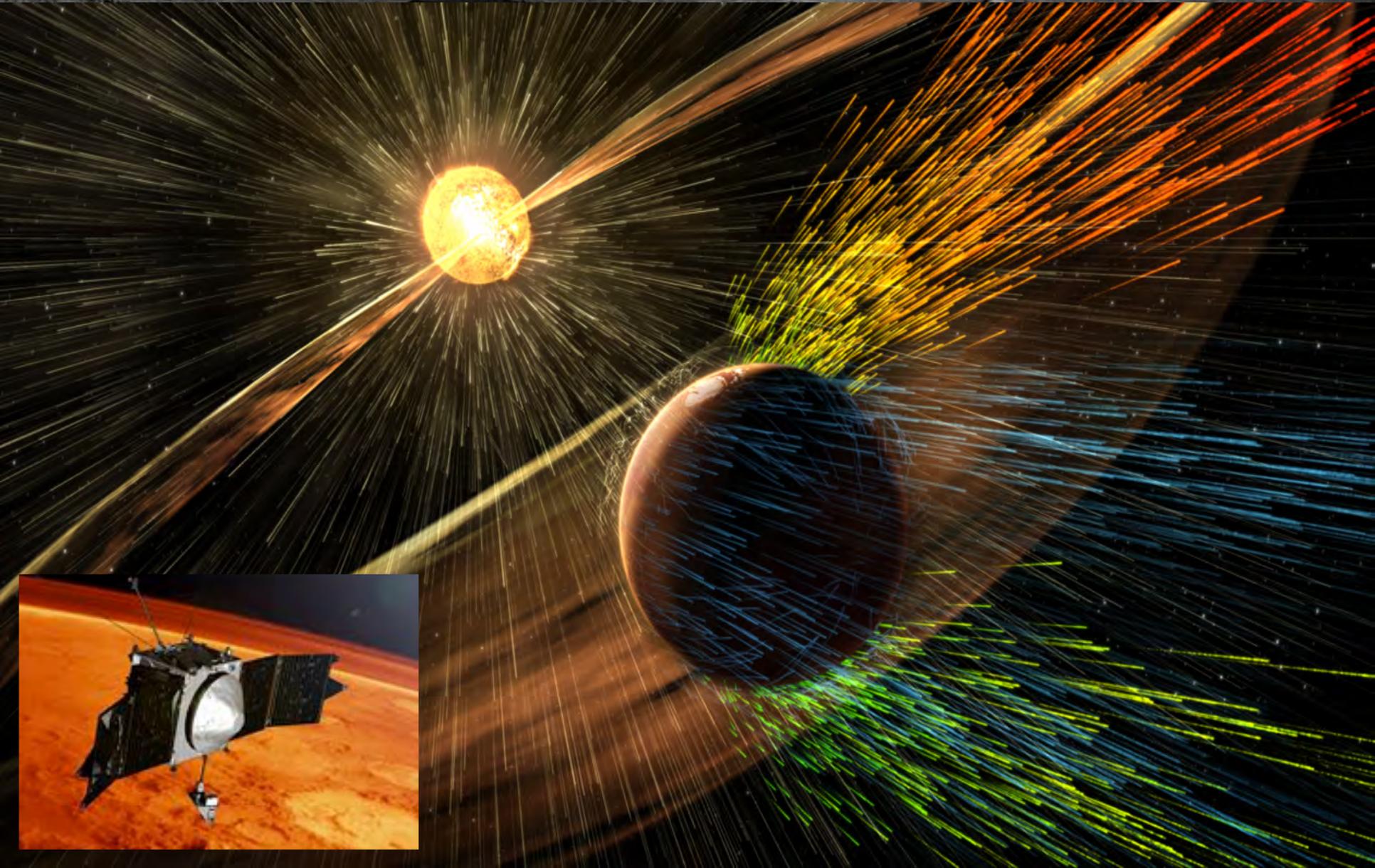
(6) Organic compounds extracted from soils and rocks

→ what have we learned so far ?

(7) Atmospheric composition and methane

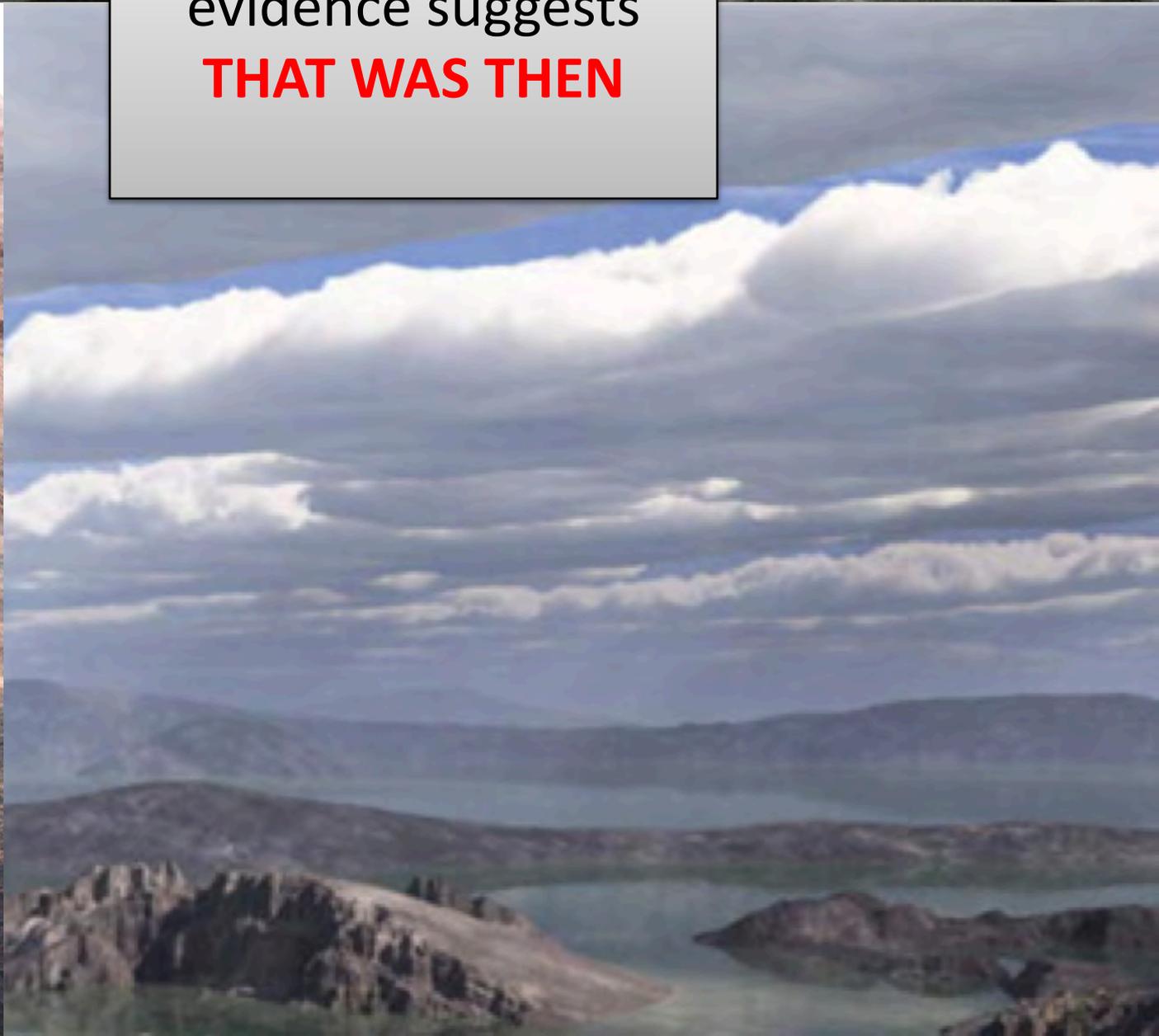
→ what does methane have to do with present or past life?

MSL and MAVEN exploring martian atmospheric loss over time. Atmospheric H, C, O, N, Ar, & Xe isotopes dominated by atmospheric escape processes



**This is now**

D/H and geological  
evidence suggests  
**THAT WAS THEN**



P.R. Mahaffy and 26  
CoAuthors, The imprint of  
atmospheric evolution in  
the D/H of Hesperian clay  
minerals on Mars, Science  
23, 412-414 (2015).

Curiosity's primary scientific goal is to explore and quantitatively assess a local region on Mars' surface as a potential habitat for life, past or present.

**Measurements that give insight into chemical processes related to possible life on early Mars when surface water was more abundant**

(1) Mineralogy and geology

→ what is it telling us about the ancient crater environment ?

(2) Isotopic composition of light elements

→ is isotopic fractionation from biological or physical/chemical processes ?

(3) Age of formation of rocks

→ how long have aqueous conditions persisted on Mars ?

(4) Cosmic radiation of surface materials

→ are near surface organics preserved from cosmic radiation ?

(5) Clays, perchlorates, sulfates, and hydrated minerals

→ do chemical environments enable preservation or destruction of organics ?

(6) Organic compounds extracted from soils and rocks

→ what have we learned so far ?

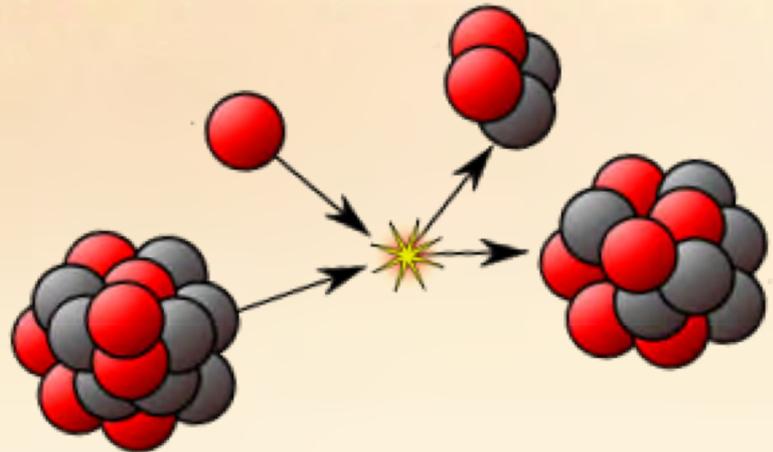
(7) Atmospheric composition and methane

→ what does methane have to do with present or past life?

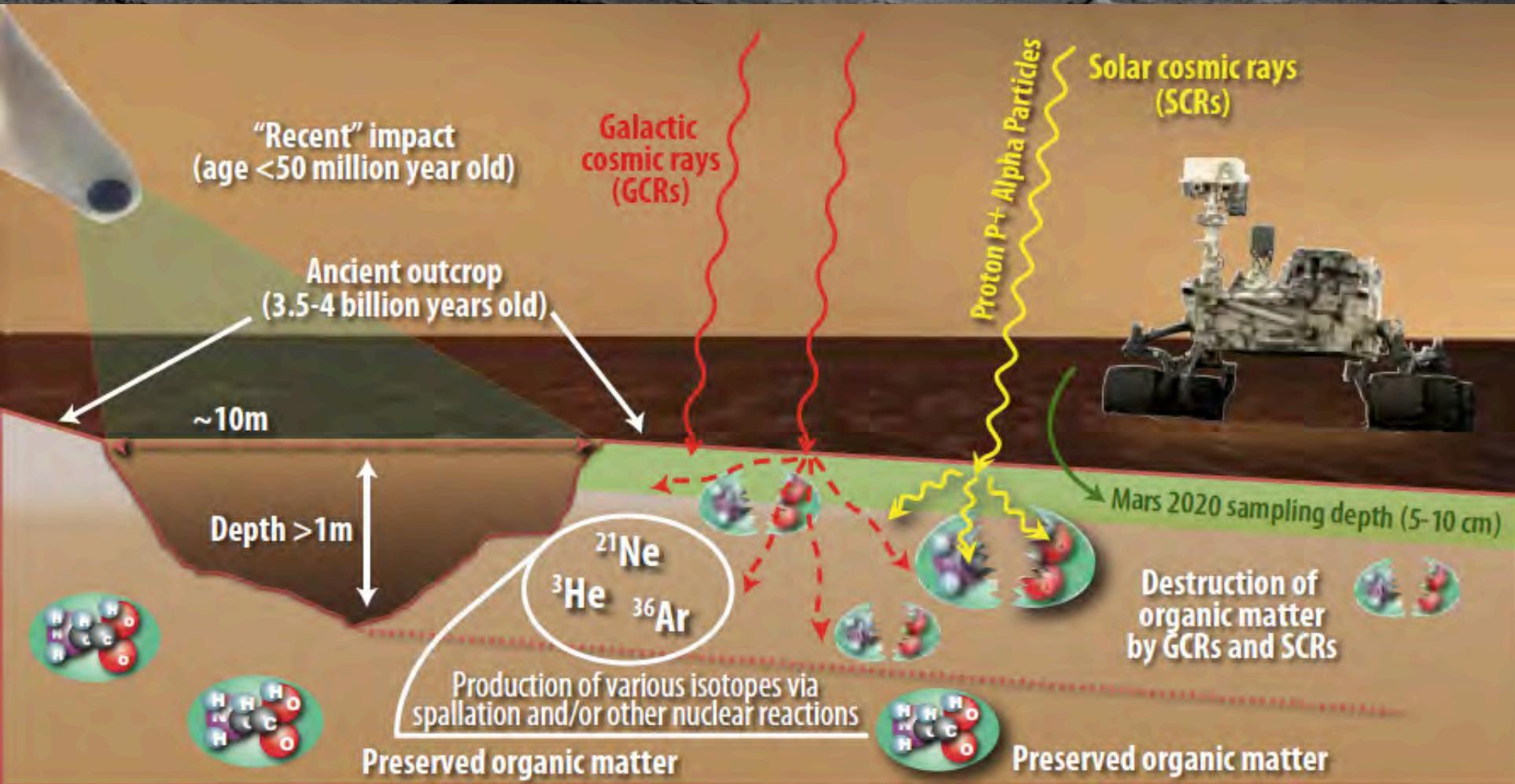
# COSMIC RADIATION

- $^3\text{He}$ ,  $^{21}\text{Ne}$ ,  $^{36}\text{Ar}$  in Martian rocks can be produced only by Cosmic Rays (protons, alpha particles, etc.)
- CRs bombard Martian surface and cause nuclear reactions in the top few meters
- $^3\text{He}$ ,  $^{21}\text{Ne}$  – “spallation”;  $^{36}\text{Ar}$  – “neutron capture”
- The more  $^3\text{He}$ ,  $^{21}\text{Ne}$ ,  $^{36}\text{Ar}$  are in the rock the longer CR exposure of that rock had to be.

Even relatively short exposure to cosmic rays can transform most of the organic compounds



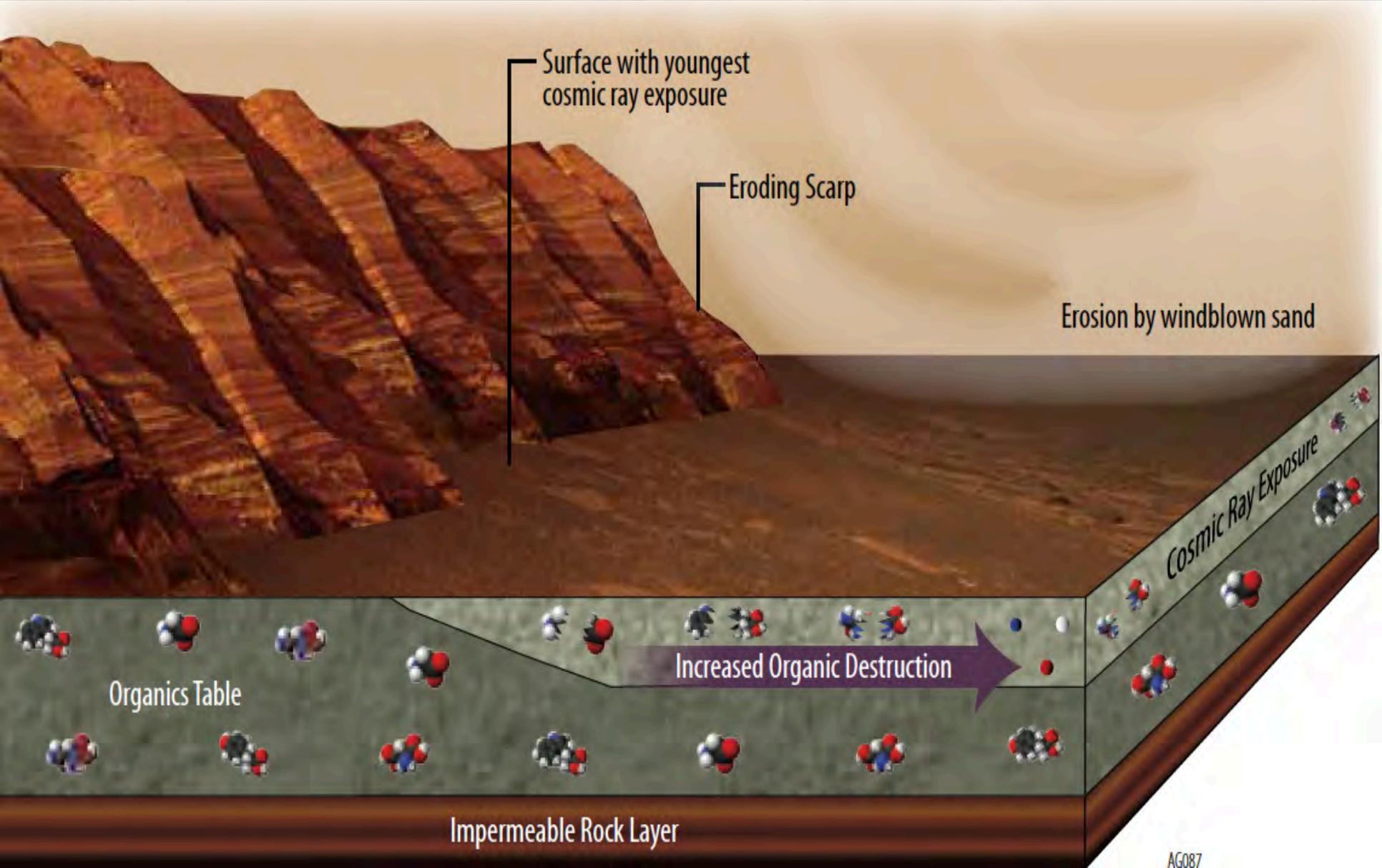
# COSMIC RAY EXPOSURE AGE - but this radiation can also transform organic compounds



Nuclear spallation -  ${}^3\text{He}$  from many elements;  ${}^{21}\text{Ne}$  - Mg, Al, Si;  ${}^{36}\text{Ar}$  - Ca, K  
Neutron capture -  ${}^{36}\text{Ar}$  from  ${}^{35}\text{Cl}$ ,  ${}^{80}\text{Kr}$  from Br;  ${}^{128}\text{Xe}$  from I

**INDEPENDENT EXPOSURE AGE RESULTS (Ma) =  ${}^3\text{He}$  72 $\pm$ 15,  ${}^{21}\text{Ne}$  84 $\pm$ 28,  ${}^{36}\text{Ar}$  78 $\pm$ 24**

# SAM noble gas measurements are pointing toward the best locations to search for minimally transformed organic compounds



# First *in situ* rock formation age experiment on another planet

Sample Mass	0.135 ± 0.018 g	
K-Ar System	±1σ	
K <sub>2</sub> O (wt %)	0.50	0.08
<sup>40</sup> Ar (nmol/g)	11.95	1.71
K-Ar Age (Ga)	4.21	0.35

**Sheepbed Mudstone sample**  
(sediment mixture)

**Sediment sources**

## K-Ar Result

**Gale Crater region:  
4.2 (± 0.4) billion years old**

**Crater density  
age estimates:  
3.6 - 4.1 billion years**

**Conclusion – rock sediment that washed down from the rim of Gale Crater was formed ~4.2 billion years ago – this consistent with crater densities**

A 2-step K/Ar chronology measurement on a drilled sample containing jarosite gave an age of  $2.12 \pm 0.36$  billion years for this mineral



→ fluvial activity in Gale crater long after fluvial activity had been supposed to cease !!

Curiosity's primary scientific goal is to explore and quantitatively assess a local region on Mars' surface as a potential habitat for life, past or present.

Measurements that give insight into chemical processes related to possible life on early Mars when surface water was more abundant

(1) Mineralogy and geology

→ what is it telling us about the ancient crater environment ?

(2) Isotopic composition of light elements

→ is isotopic fractionation from biological or physical/chemical processes ?

(3) Age of formation of rocks

→ how long have aqueous conditions persisted on Mars ?

(4) Cosmic radiation of surface materials

→ are near surface organics preserved from cosmic radiation ?

(5) Clays, perchlorates, sulfates, and hydrated minerals

→ do chemical environments enable preservation or destruction of organics ?

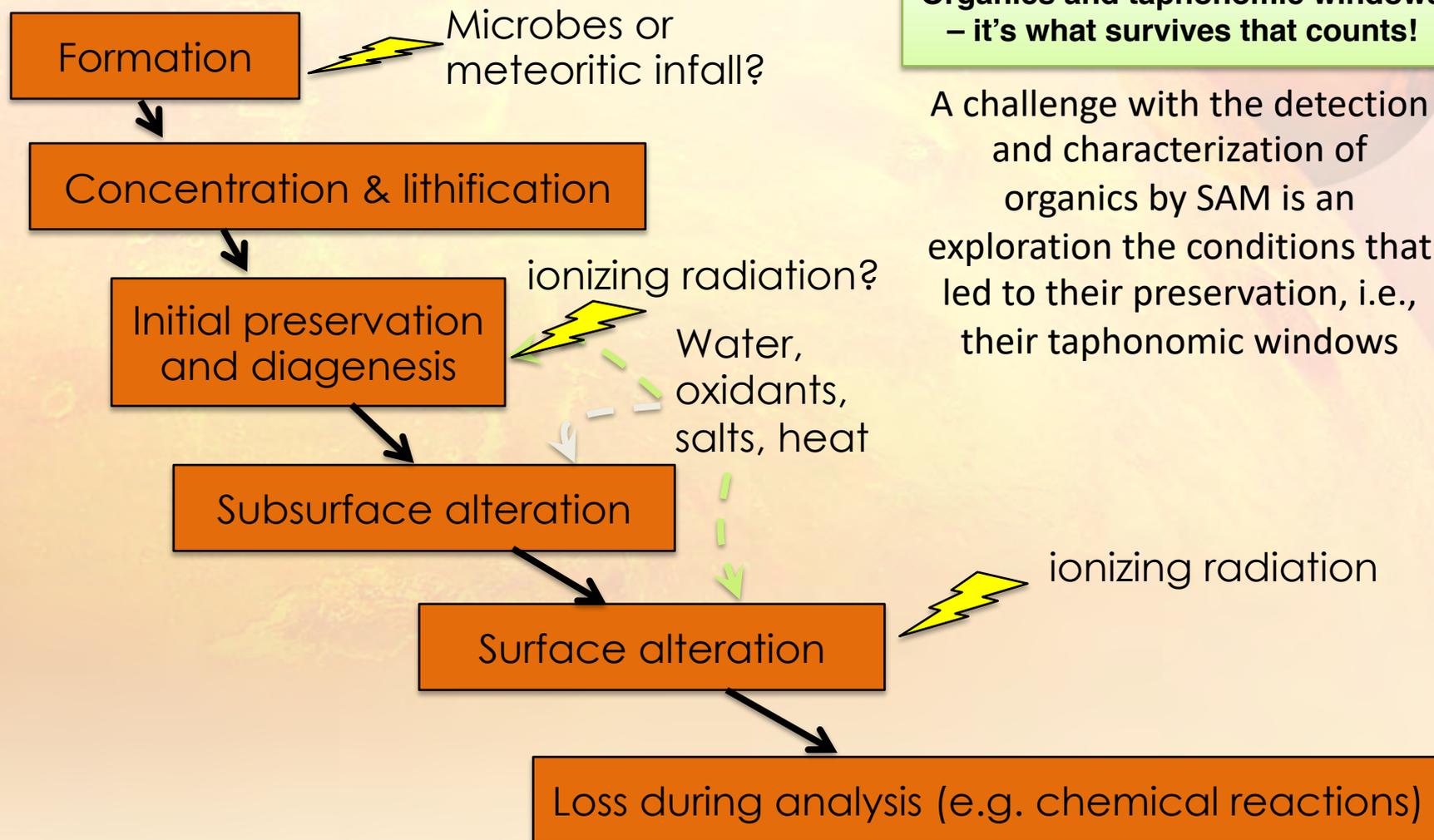
(6) Organic compounds extracted from soils and rocks

→ what have we learned so far ?

(7) Atmospheric composition and methane

→ what does methane have to do with present or past life?

Curiosity's primary scientific goal is to explore and quantitatively assess a local region on Mars' surface as a potential habitat for life, past or present.



# From SAM Evolved Gas (EGA) Experiment → noble gases, organic & inorganic compounds & many isotopes

## Evolved gases

**H<sub>2</sub>O** – from hydrated minerals, adsorbed water, or clays

**SO<sub>2</sub> & H<sub>2</sub>S** - from sulfates & sulfides

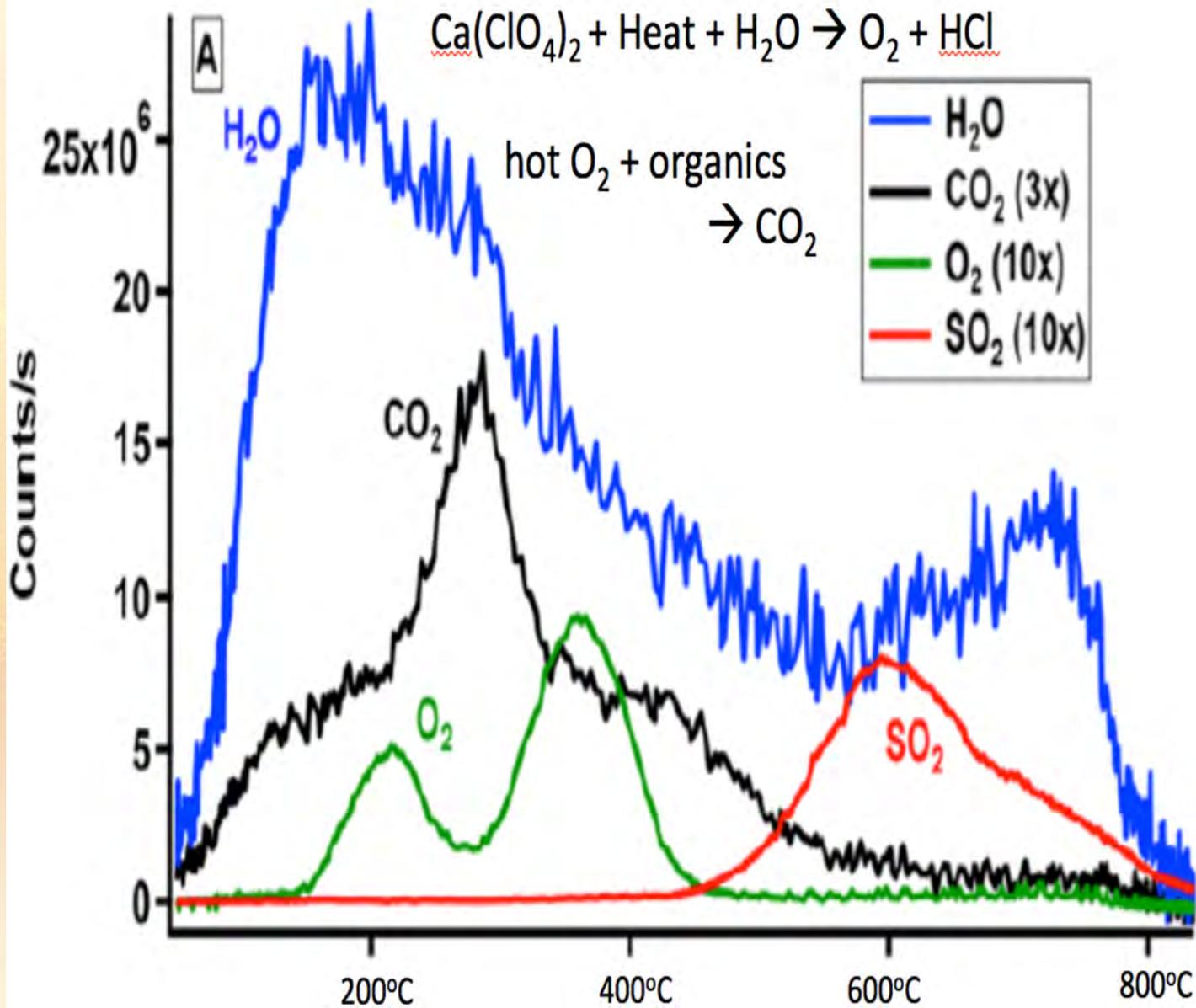
**O<sub>2</sub> & HCl** - from perchlorates or oxychlorine compounds

**NO** - from nitrates

**He, Ar, Ne** – rock formation & exposure age determinations

**CO<sub>2</sub> & CO** – from carbonates or combusted organics

**Organics (incl. CH<sub>4</sub>)**

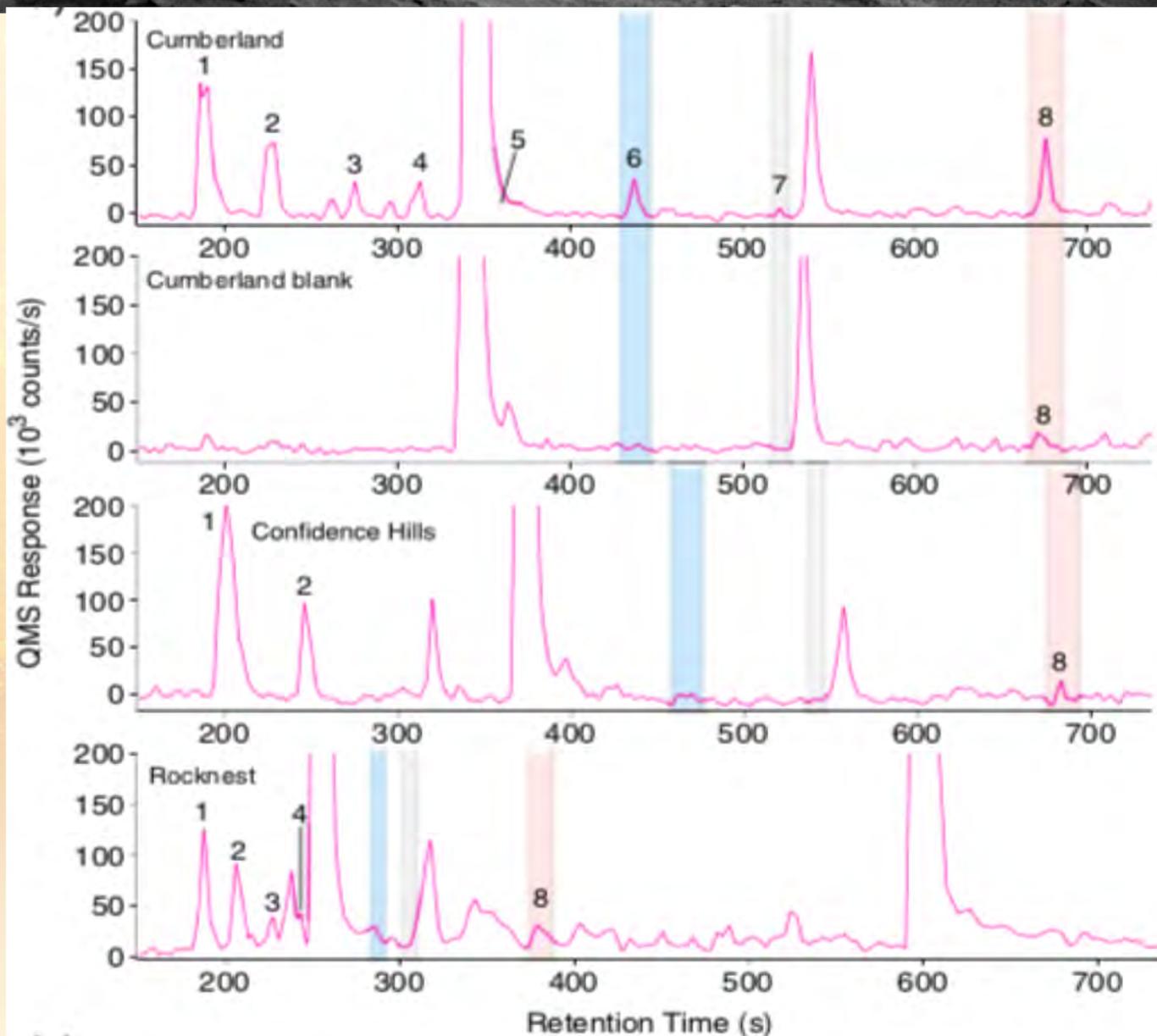


# SAM chromatograms and SAM GCMS identification of chlorohydrocarbons validated with retention times on SAM TB

- 1: chloromethane**
- 2: dichloromethane**
- 3: trichloromethane,**
- 4: carbon tetrachloride**
- 5: 1,2-dichloroethane**
- 6: 1,2-dichloropropane**
- 7: 1,2-dichlorobutane**
- 8: chlorobenzene**

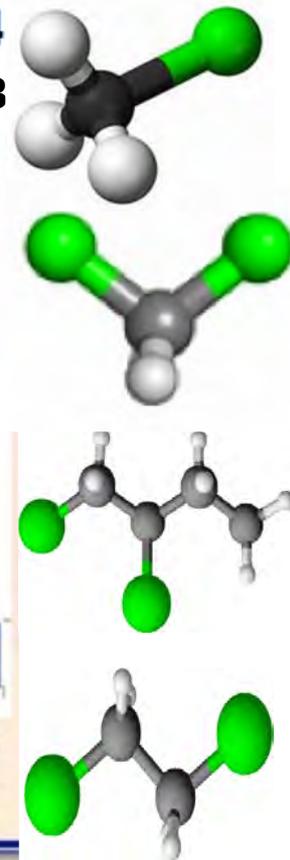
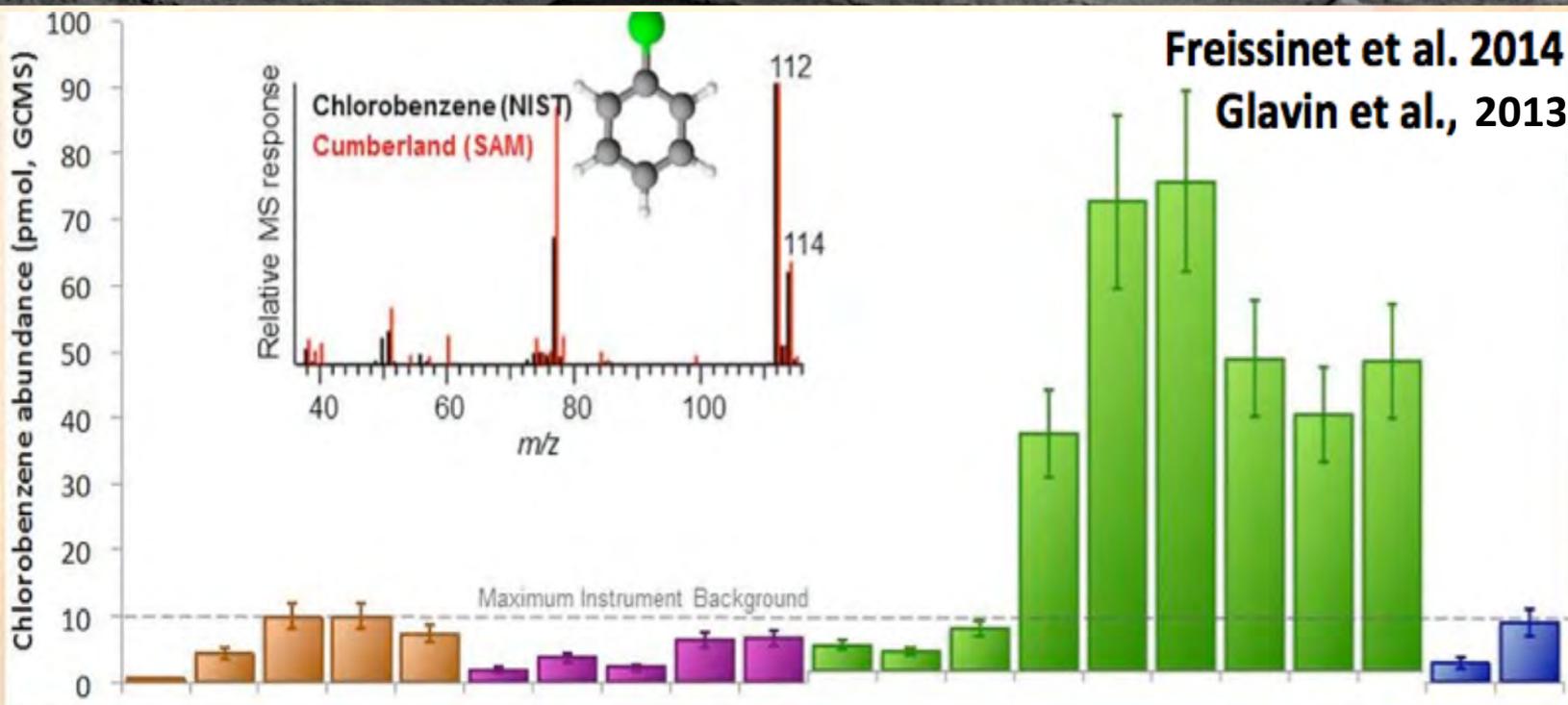
*Reconstructed chromatograms*

*m/z 52 × 2 +*  
*m/z 84 +*  
*m/z 83 × 8 +*  
*m/z 117 × 35 +*  
*m/z 63 × 8 +*  
*m/z 90 × 10 +*  
*m/z 112 × 7*



# SAM has a special focus on carbon compounds

## SAM GCMS – chlorobenzene and dichlorinated alkanes

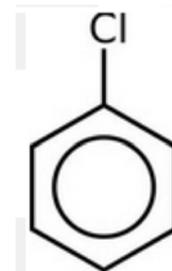


ROCKNEST

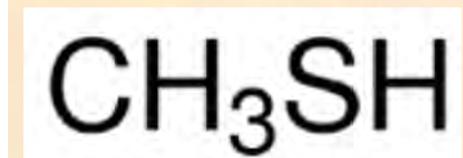
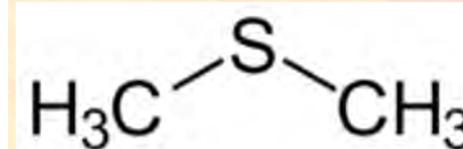
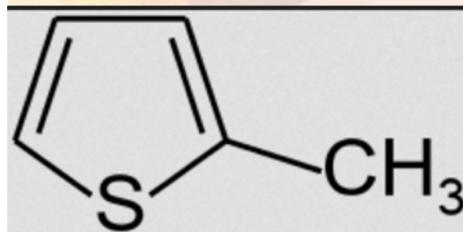
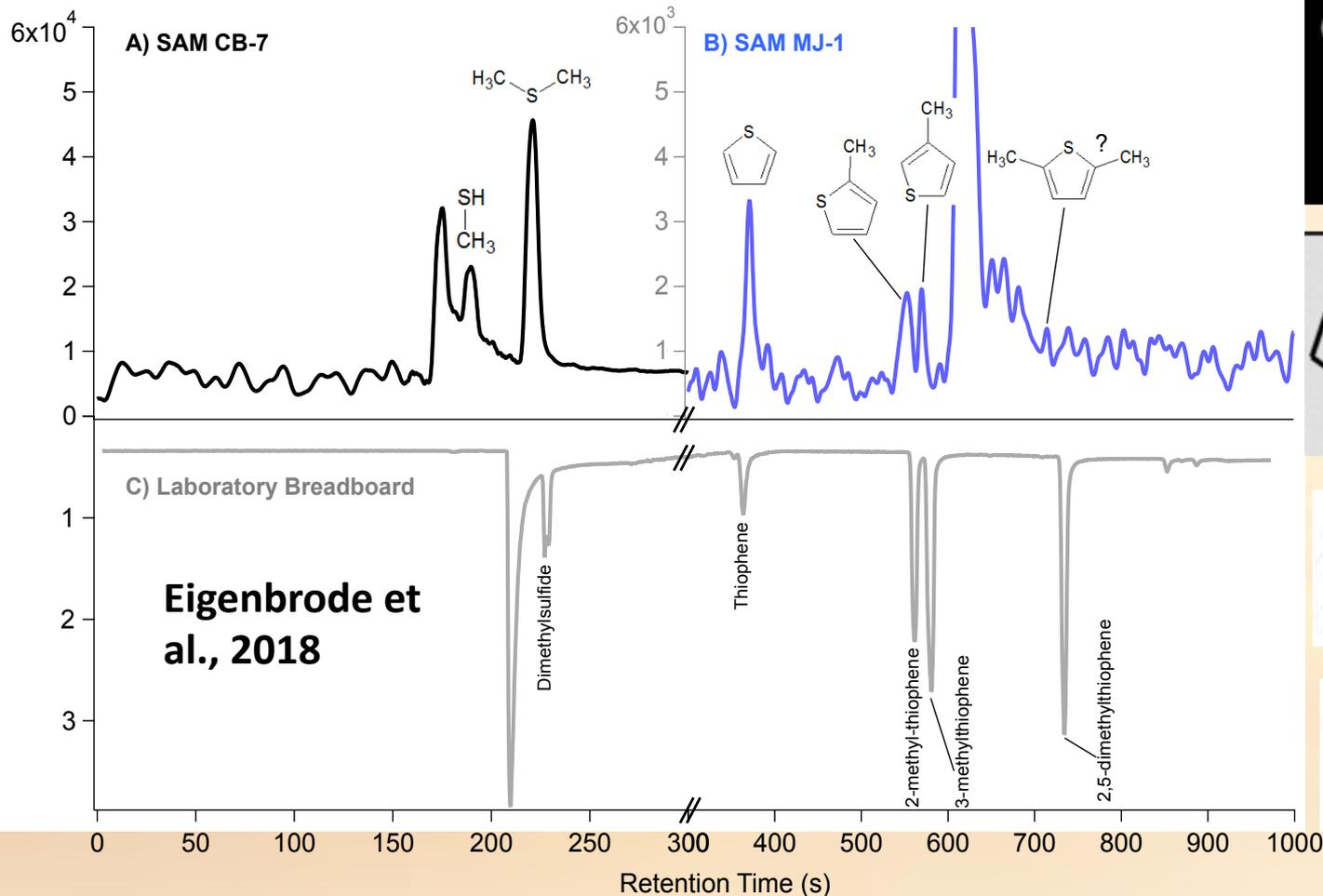
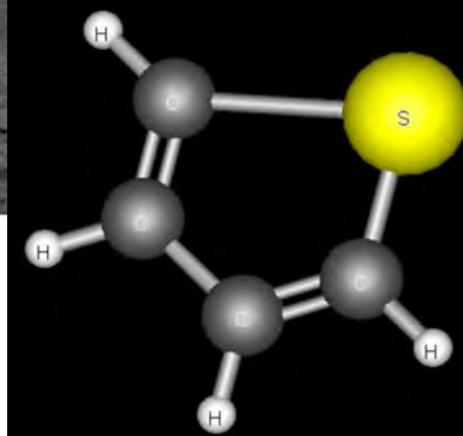
JOHN KLEIN

CUMBERLAND

CONFIDENCE HILLS

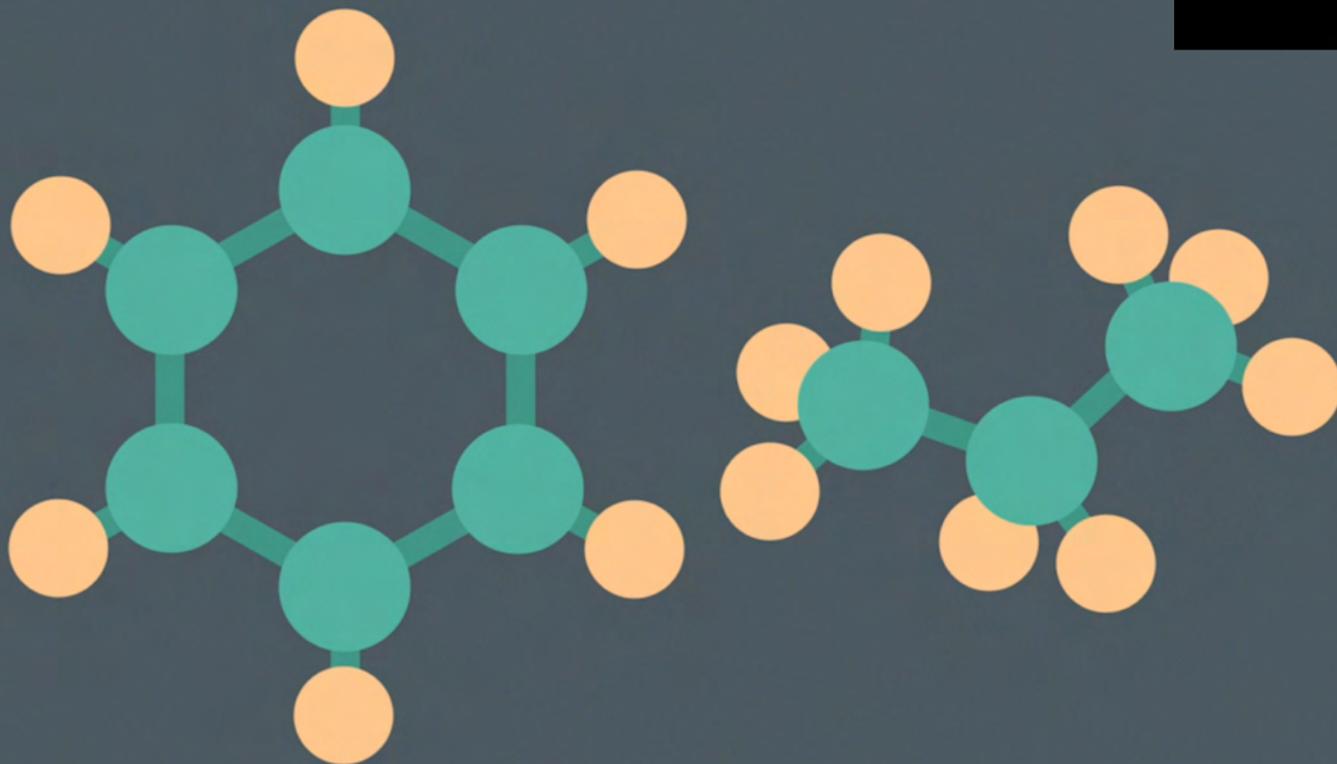
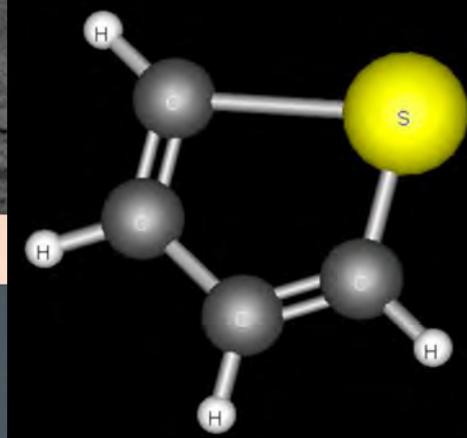


# S-containing compounds without Cl !

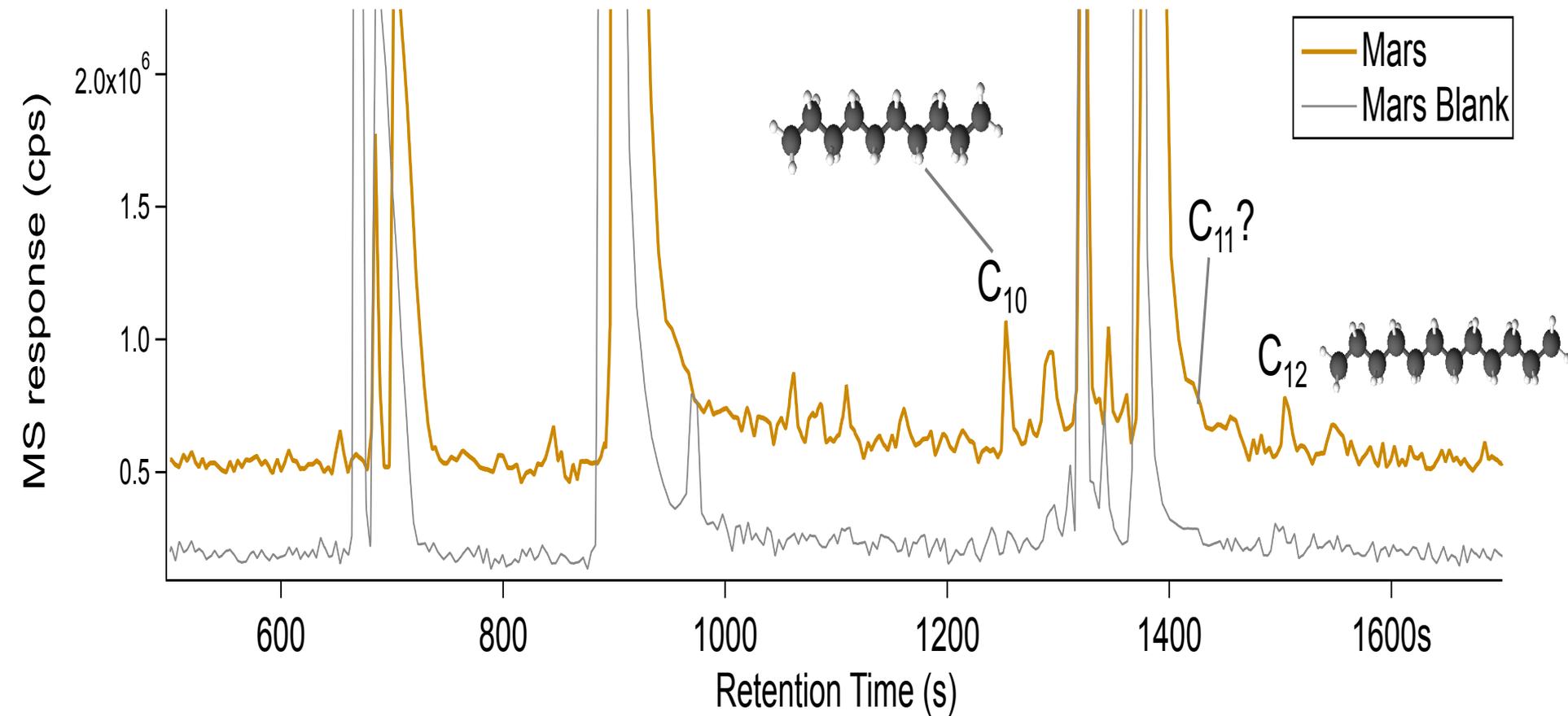


- ✓ Thiophene, methyl and di-methyl thiophene, dimethylsulfide, methanethiol
- ✓ Seen in CB high temperature cut & Mojave (MJ) wide temperature cut.
- ✓ MJ in Murray formation at Parhump Hills - lower mount outcrop at Gale.

# S-containing compounds without Cl and evidence of macromolecules !



# Long chain alkanes from a Cumberland “doggy bag” sample!



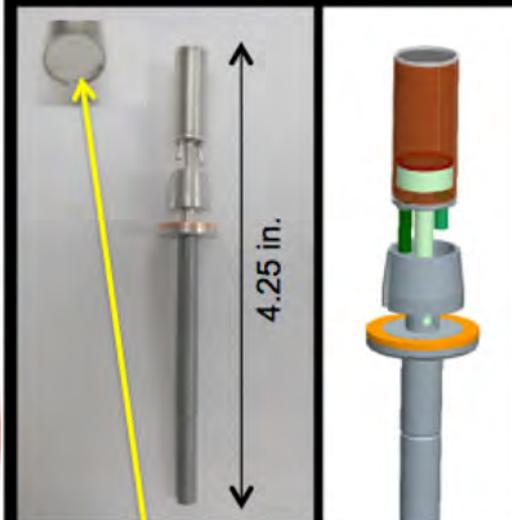
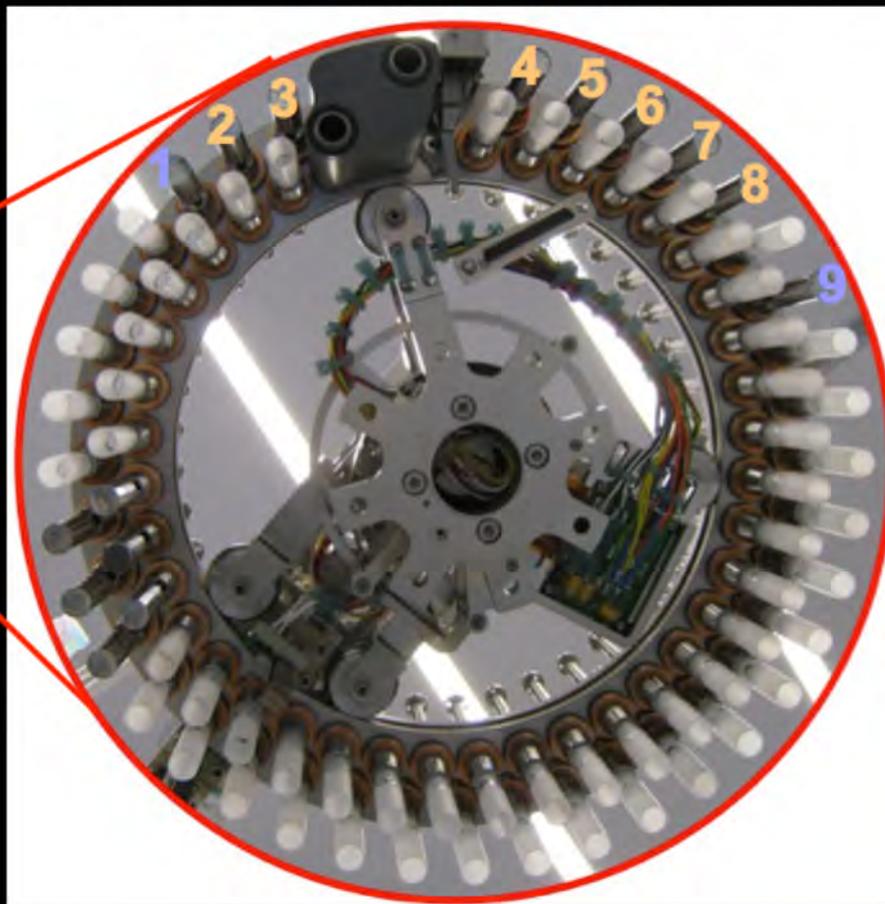
Are decane and dodecane decomposition products of fatty acids?

# Wet chemistry cups – OD = use of residual vapor in the SMS and long exposures to sample to enable MTBSTFA Rx

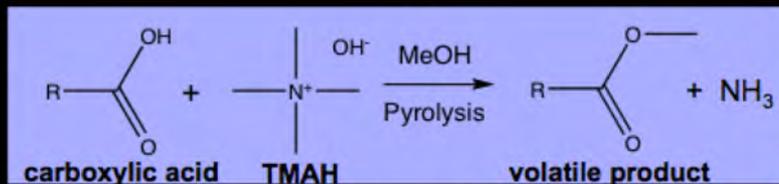
Nine cups for low temperature extraction targeting less volatile organic compounds



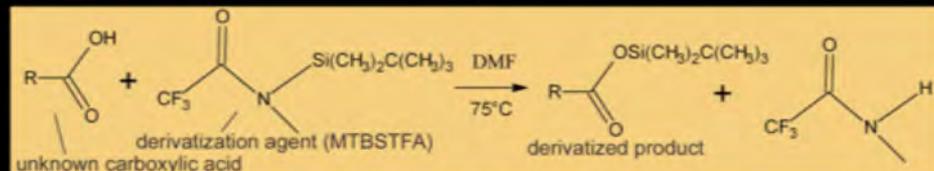
Several unidentified heavy organics showing up in OD runs



Foil cap designed for puncture using pin – Mars sample dropped into solvent filled cup through inlet tube



**TMAH: Tetramethylammonium hydroxide**



**MTBSTFA: *N*-(*tert*-butyldimethylsilyl)-*N*-methyltrifluoroacetamide**

## PERSPECTIVES

By Inge Loes ten Kate



PLANETARY SCIENCE

# *Organic molecules on Mars*

Data from the Curiosity rover provide evidence for organic molecules in ancient martian rocks and in the atmosphere

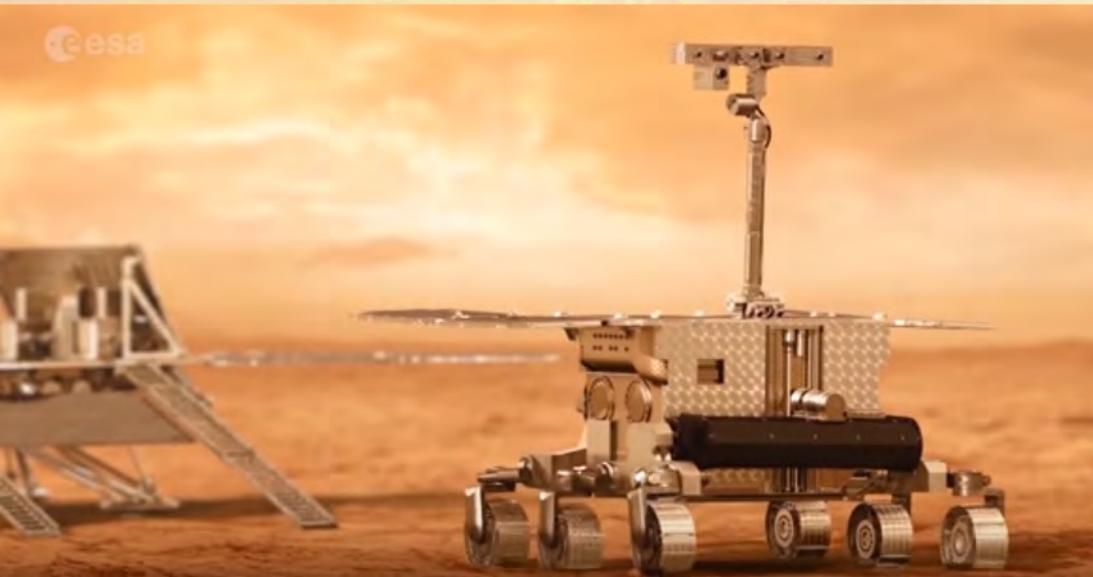
The detection of organic molecules and methane on Mars has far-ranging implications in light of potential past life on Mars. Curiosity has shown that Gale crater was habitable around 3.5 billion years ago (15), with conditions comparable to those on the early Earth, where life evolved around that time. The question of whether life might have originated or existed on Mars is a lot more opportune now that we know that organic molecules were present on its surface at that time.

# Planetary goals for mass spectrometry are moving toward direct detection of molecular biosignatures

Example -The MOMA investigation is the first direct search for martian life since the Viking landers in 1976. MOMA-MS just delivered to ESA from Goddard.

## **MOMA-MS advanced capabilities and development notes**

- Laser ablation mass spectrometry will be a planetary first
- Tandem mass spectrometry enables more robust analysis of complex molecular structures and provides a new capability for future missions
- GCMS with derivatization (DMF-DMA) for enantiomeric separations
- Significant contribution to international collaborations in the search for life



Curiosity's primary scientific goal is to explore and quantitatively assess a local region on Mars' surface as a potential habitat for life, past or present.

**Measurements that give insight into chemical processes related to possible life on early Mars when surface water was more abundant**

(1) Mineralogy and geology

→ what is it telling us about the ancient crater environment ?

(2) Isotopic composition of light elements

→ is isotopic fractionation from biological or physical/chemical processes ?

(3) Age of formation of rocks

→ how long have aqueous conditions persisted on Mars ?

(4) Cosmic radiation of surface materials

→ are near surface organics preserved from cosmic radiation ?

(5) Clays, perchlorates, sulfates, and hydrated minerals

→ do chemical environments enable preservation or destruction of organics ?

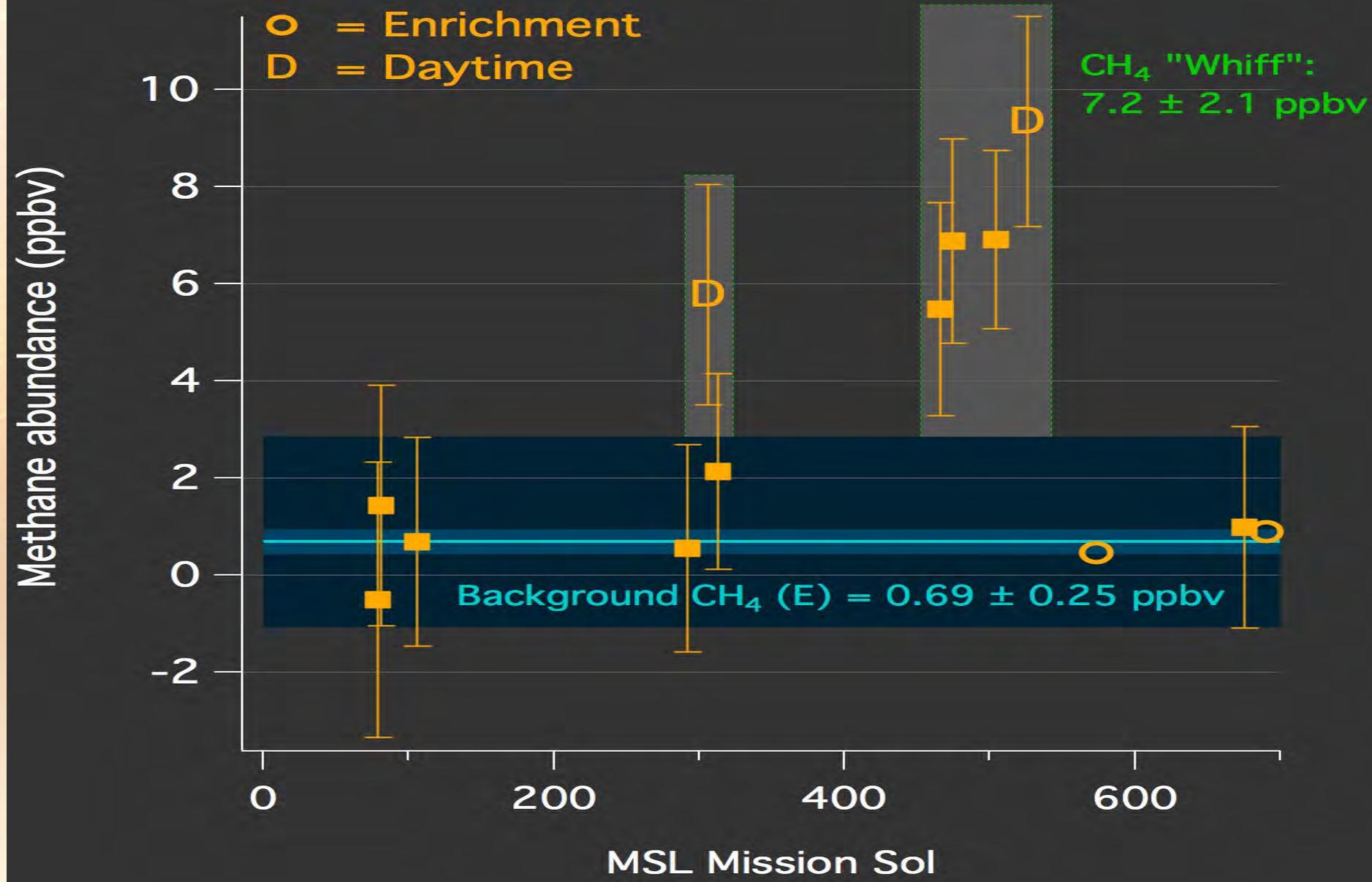
(6) Organic compounds extracted from soils and rocks

→ what have we learned so far ?

(7) Atmospheric composition and methane

→ what does methane have to do with present or past life?

# Methane "spikes" early in the mission



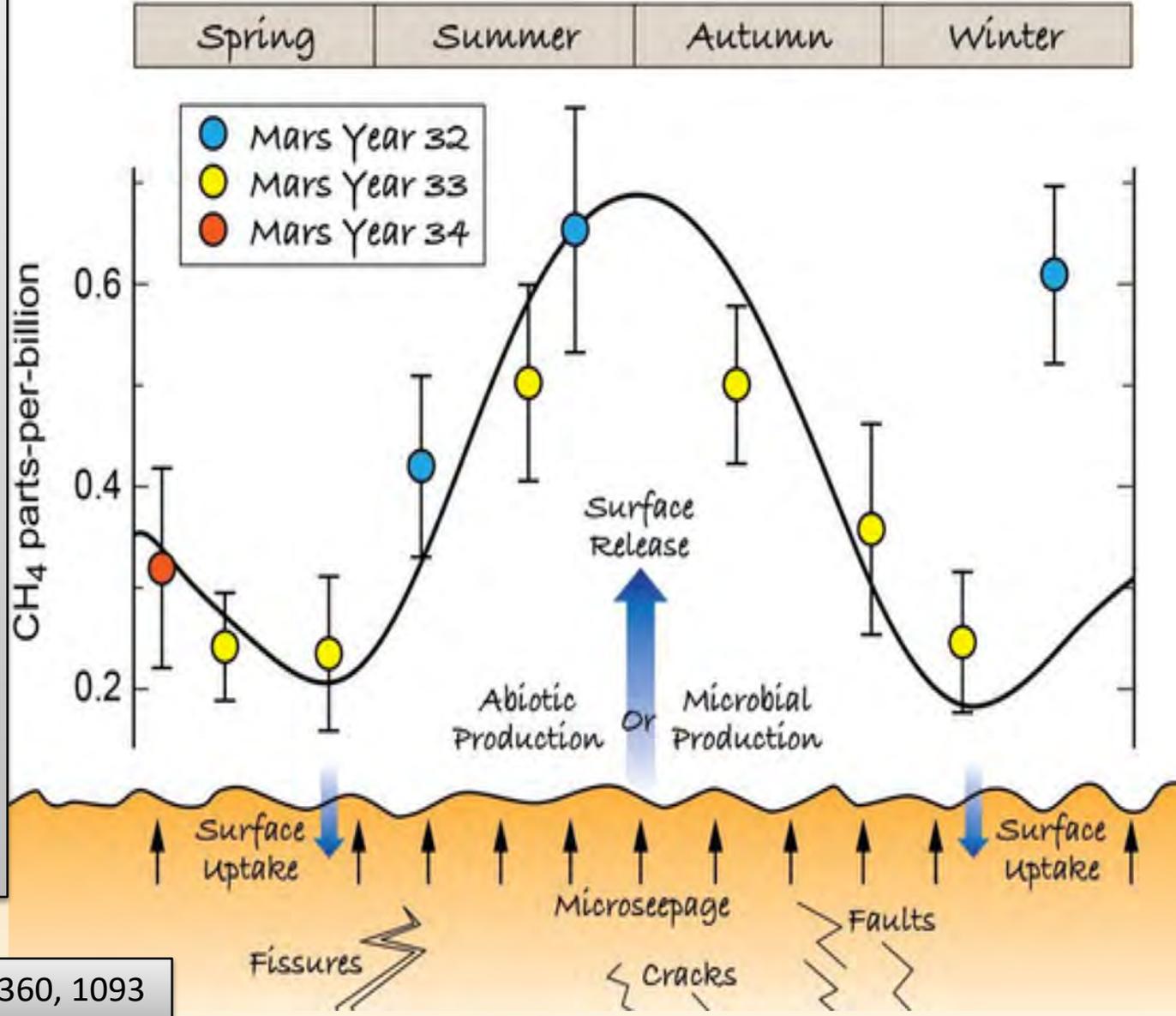
**SAM's commanding language allows us great flexibility in generation of new sequences on Mars as we make discoveries – we keep a duplicate SAM operational at NASA Goddard to test these scripts**



# Atmospheric composition using both QMS & TLS

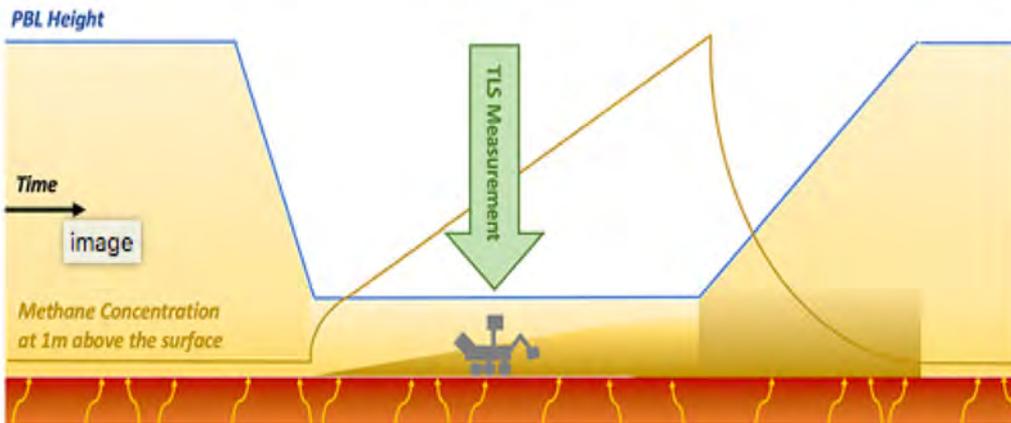
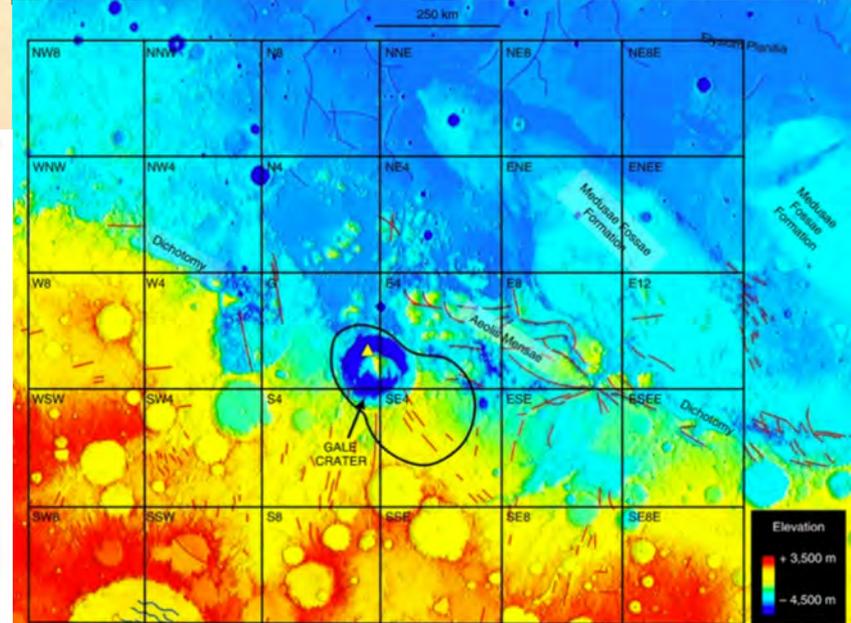
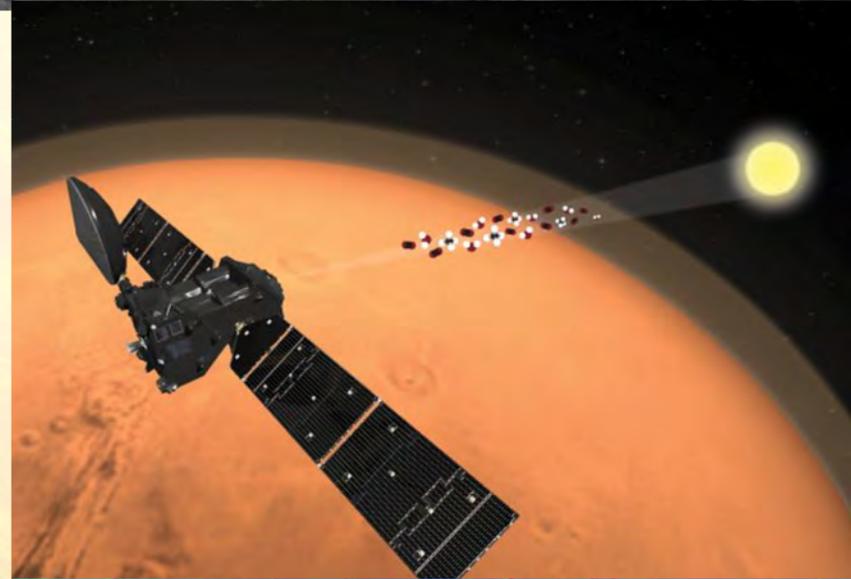
- Seasonal variations in mixing ratios of major atmospheric gases CO<sub>2</sub>, N<sub>2</sub>, Ar, O<sub>2</sub>, and CO
- "Spikes" in methane mixing ratio to 5-10 ppn
- With enrichment sub-ppb levels and seasonal trends.

Curiosity Discovers Seasonal Cycle in Mars Methane

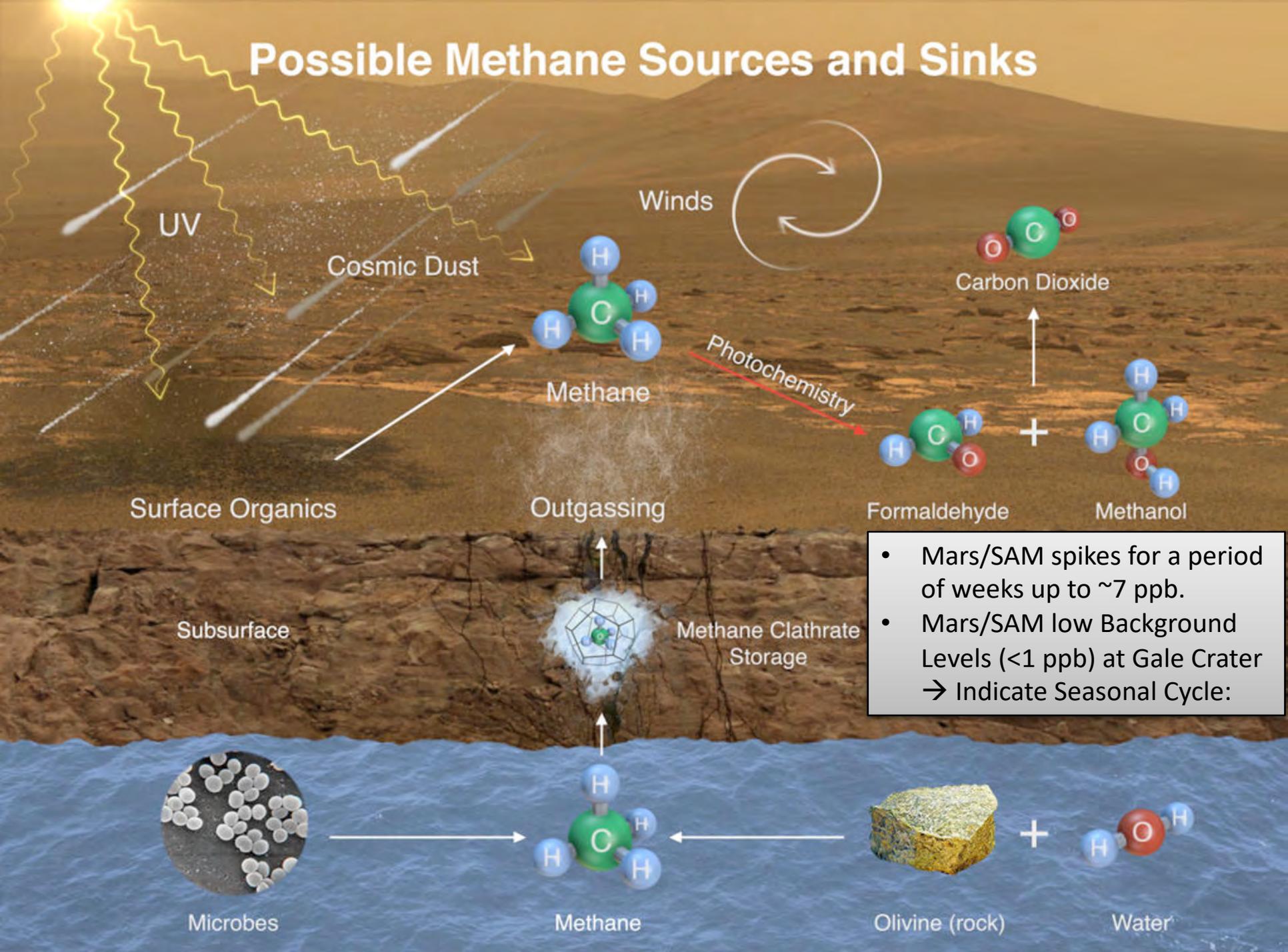


# The methane mystery persists - 2019 developments

- 1) ESA's Trace Gas Orbiter gets to work and finds very low upper limits only of methane
- 2) ESA's Mars Express PFS instrument reports finding methane near Gale crater at the time of a previous Curiosity TLS methane spike
- 3) A model of methane seep in Gale crater (Moore et al.) predicts methane accumulation at night near the surface
- 4) Curiosity conducts an experiment to test the Moore's theory but instead finds the biggest methane spike yet on the surface ~20 ppb



# Possible Methane Sources and Sinks

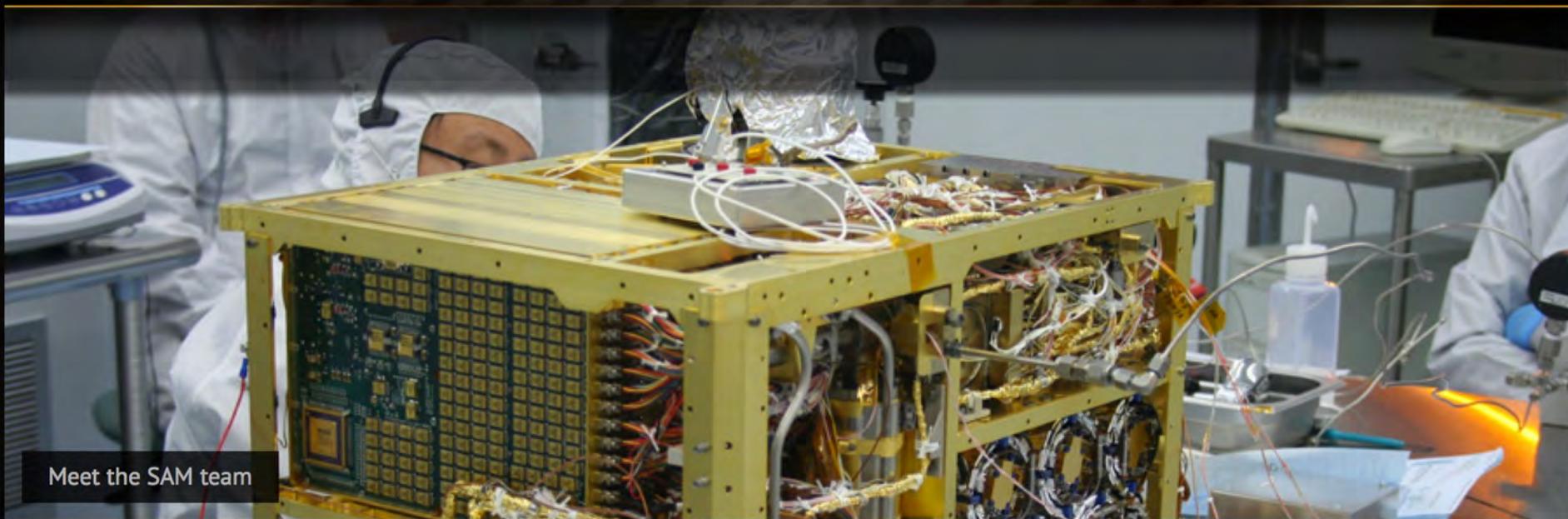


- Mars/SAM spikes for a period of weeks up to ~7 ppb.
- Mars/SAM low Background Levels (<1 ppb) at Gale Crater  
→ Indicate Seasonal Cycle:

<http://ssed.gsfc.nasa.gov/sam/>

**SAM: SAMPLE ANALYSIS AT MARS** 

**SAM: SAMPLE ANALYSIS AT MARS**  
ON THE ROVER CURIOSITY 



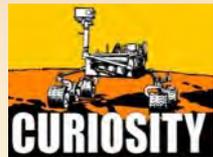
Meet the SAM team

<http://ssed.gsfc.nasa.gov/sam/>

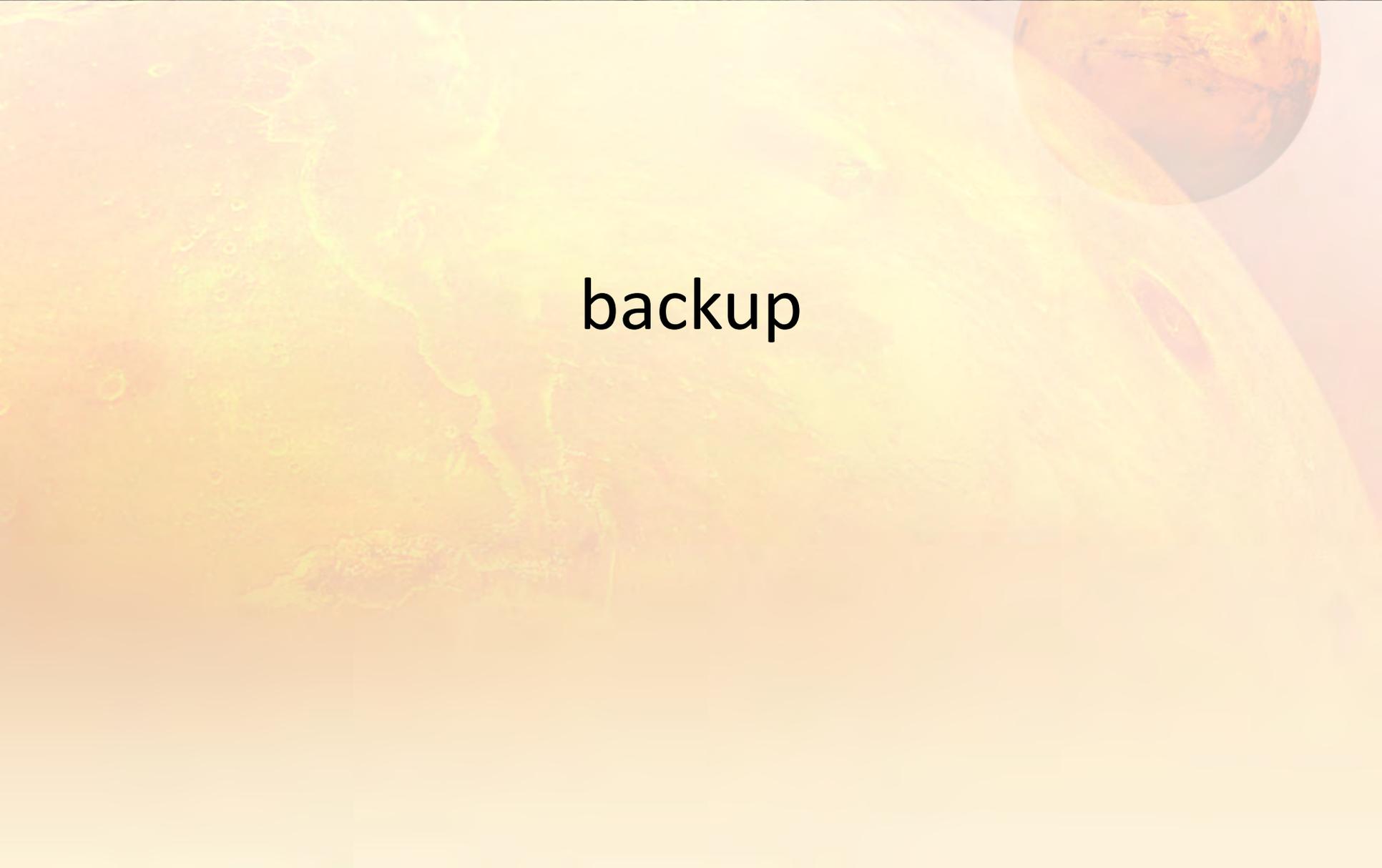


The Martians sent out a patrol  
To see what was doing one sol.  
They got a big thrill  
To see over a hill  
Curiosity taking a stroll

Curiosity limerick by Dr. J.T. Nolan



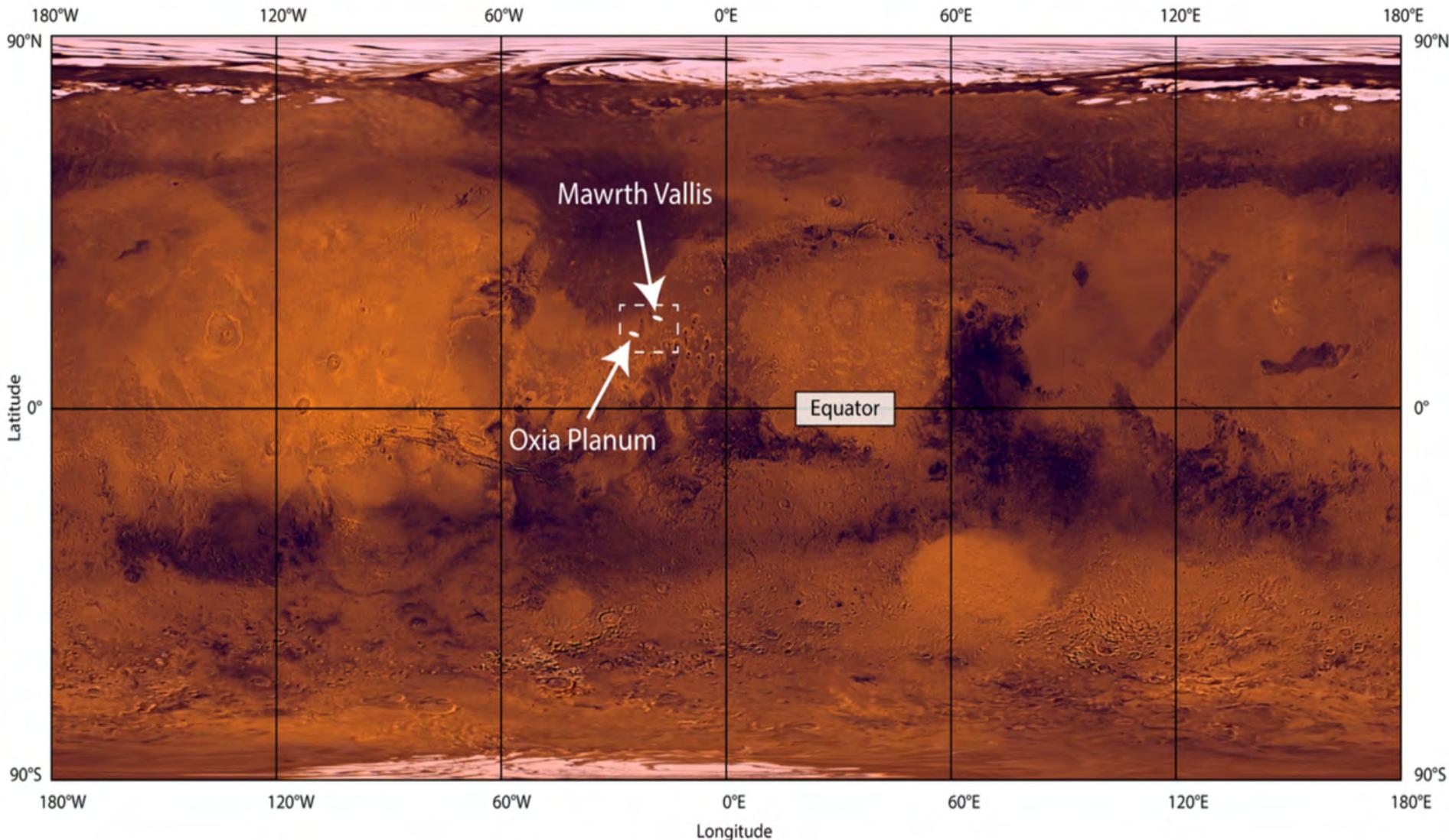
**Curiosity now in the clay unit will continue to study the  
Layers, Canyons, and Buttes of Mount Sharp**



backup

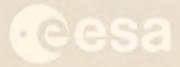


# Ancient Noachian site Oxia Planum is the selected landing site for the Rosalind Franklin rover

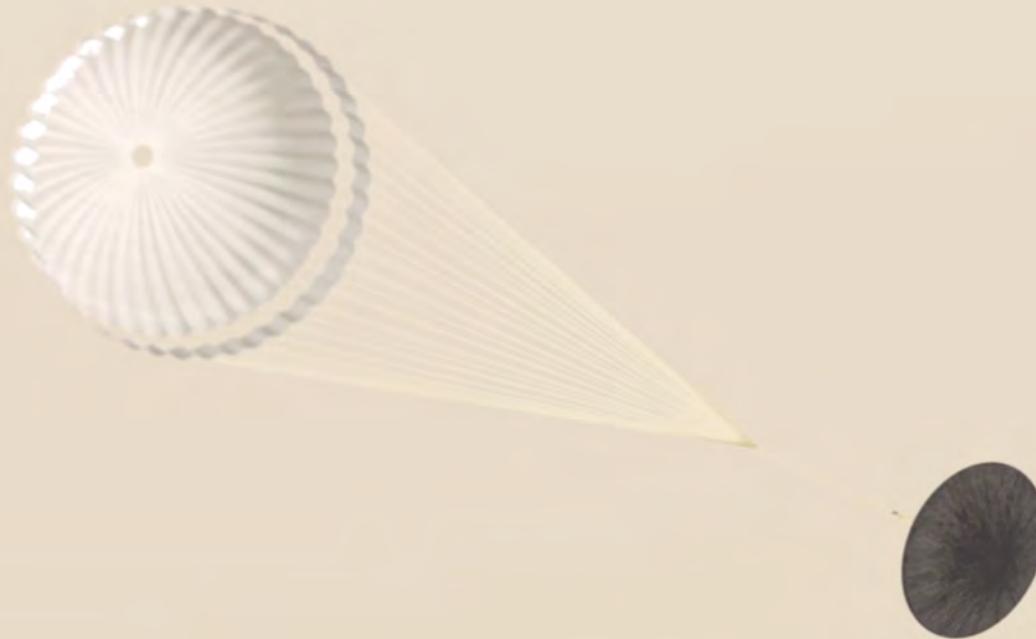


Clay rich material recently exposed giving a window into very early Mars

# An entirely different world preserved from ancient times may be revealed in the martian subsurface



DRAFT MOMA\_Profile\_031318



00:00:00:00