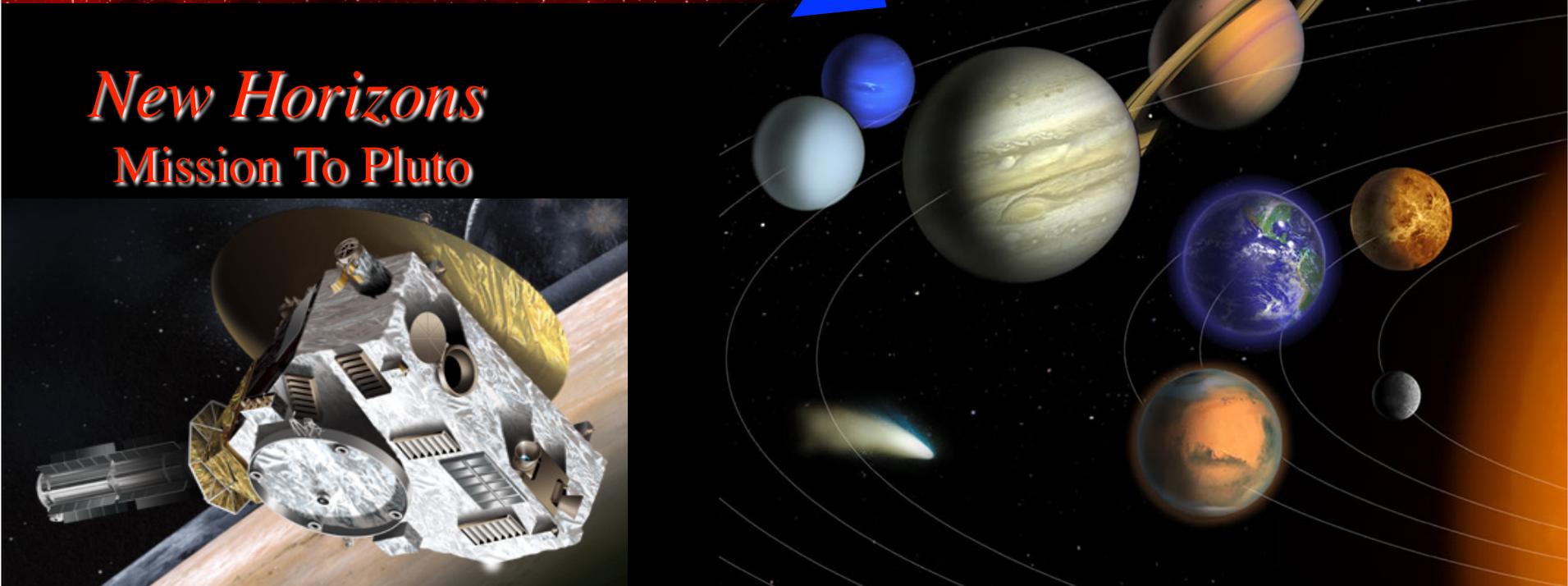
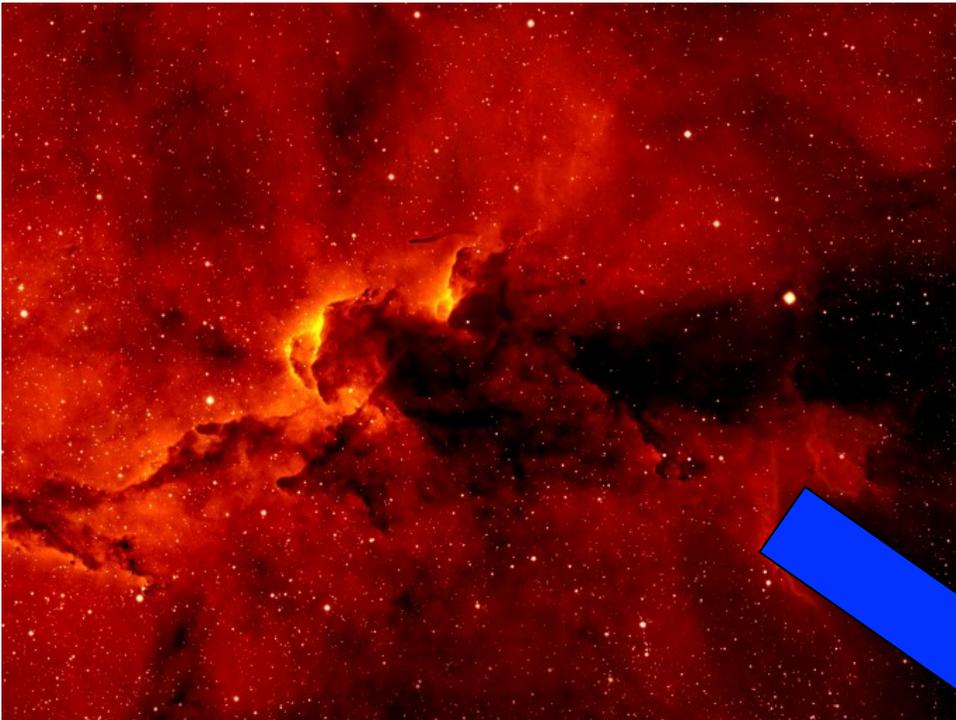
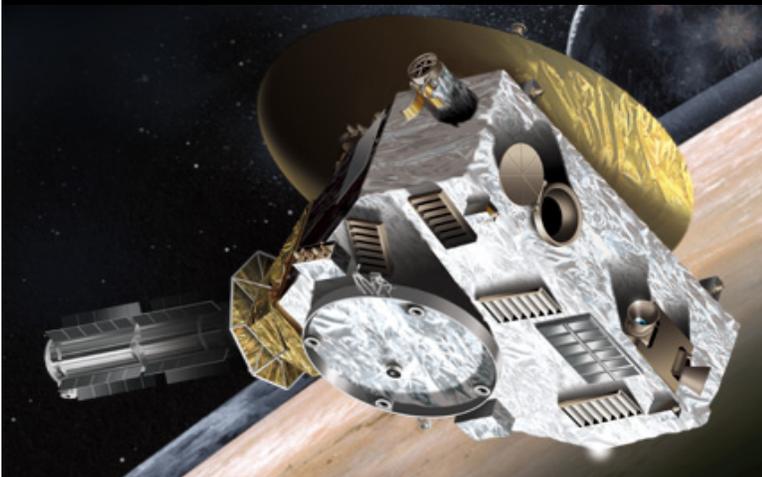


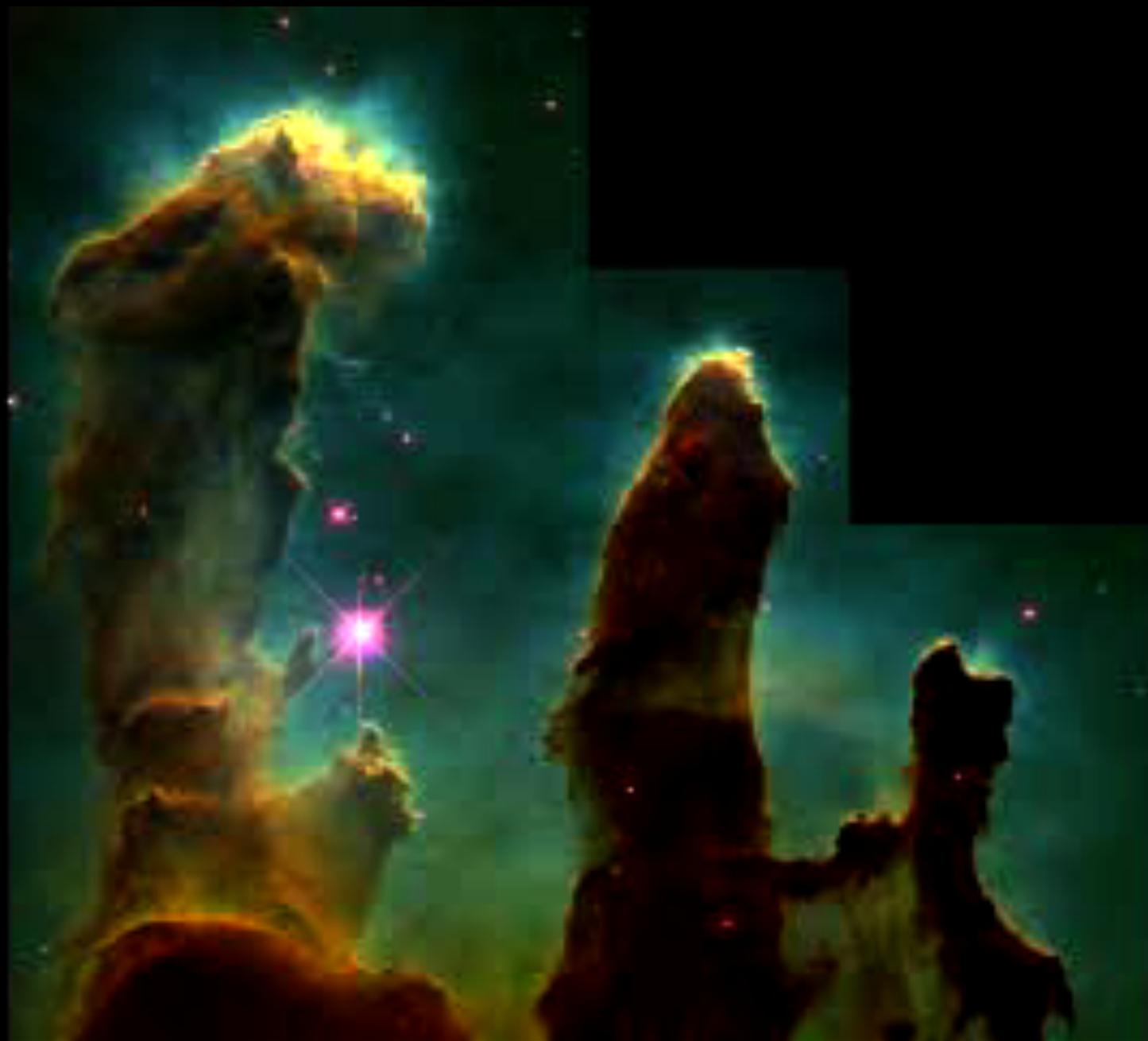
Fran Bagenal

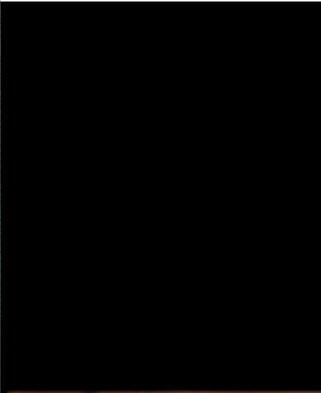
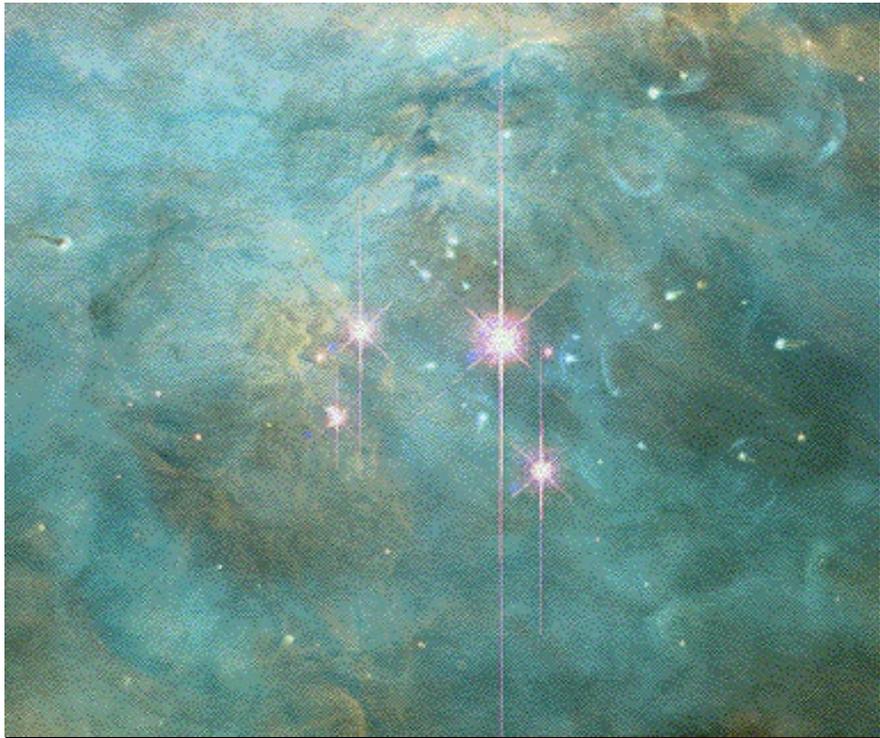
Solar System
Formation



New Horizons
Mission To Pluto





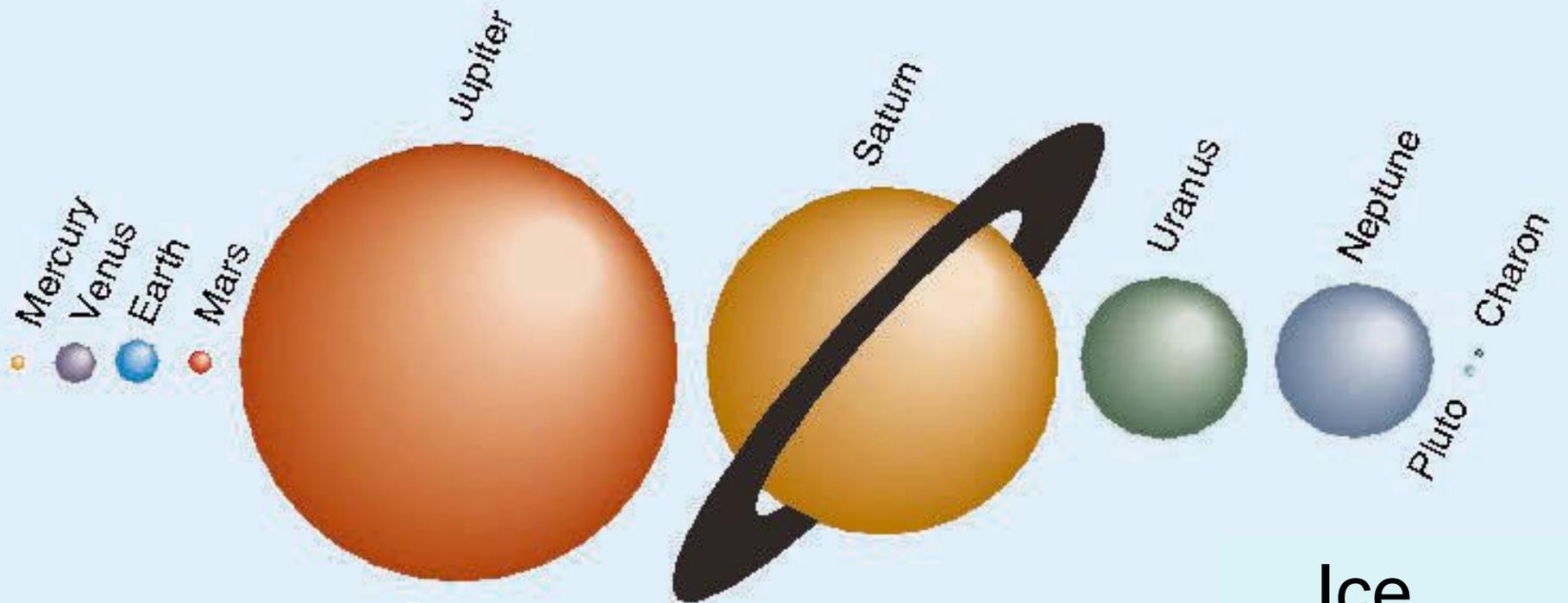




**Protoplanetary Disks in the Orion Nebula
Hubble Space Telescope • WFPC2**

NASA, J. Bally (University of Colorado), H. Throop (SWRI), and C.R. O'Dell (Vanderbilt University)
STScI-PRC01-13

The Planets at a Glance



Small
Inner
Rocky
Planets

Giant

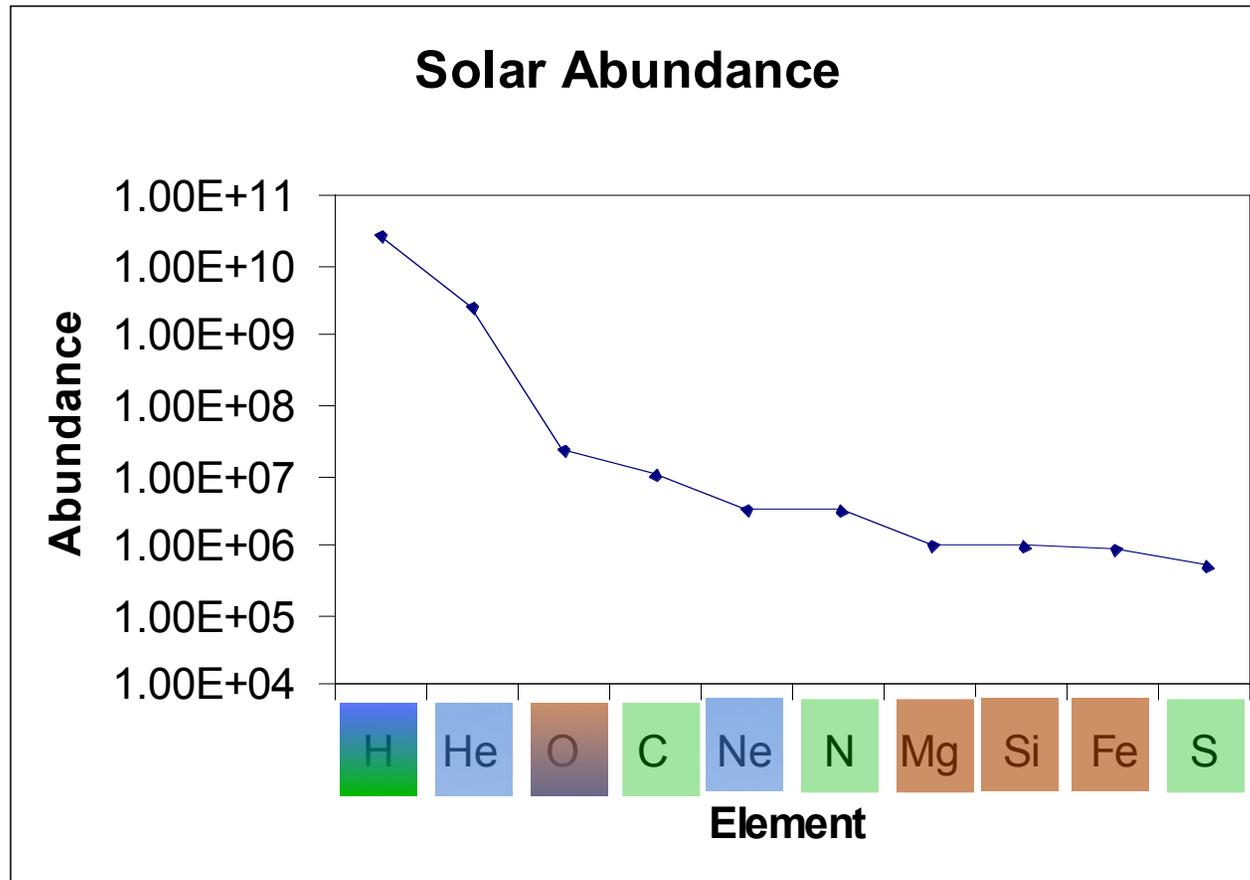
Outer

Gas

Planets

Ice
Dwarf
Planets

The Universe's Top Ten

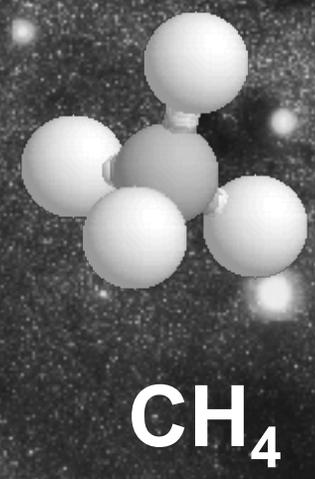
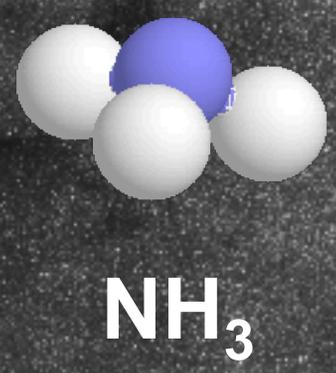
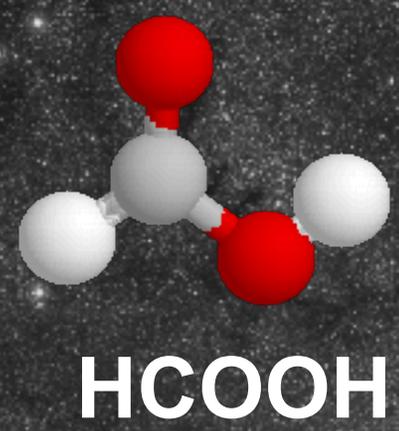
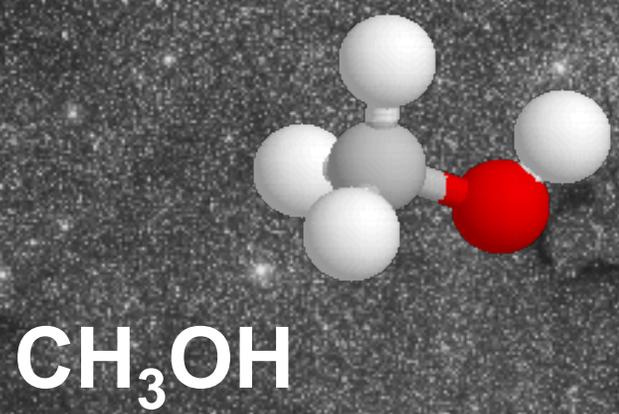
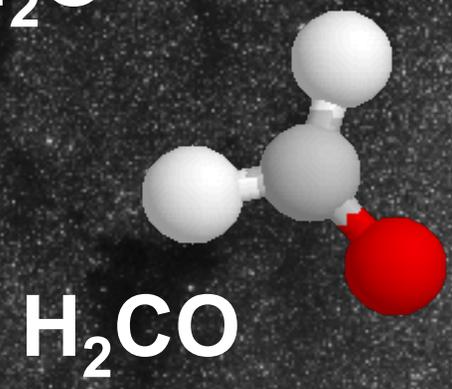
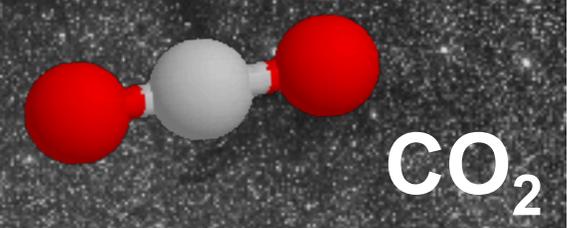
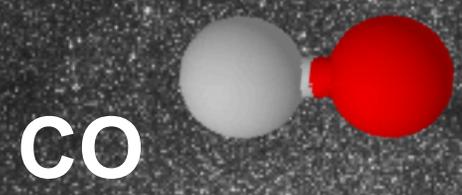
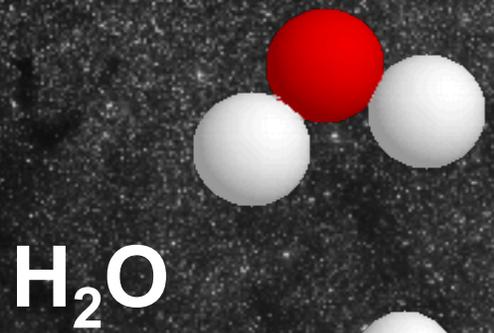


Noble gases

Water Avg. solid => 50/50 rock/ice

Rocks

Life building blocks



Interstellar medium – ISM – space between stars

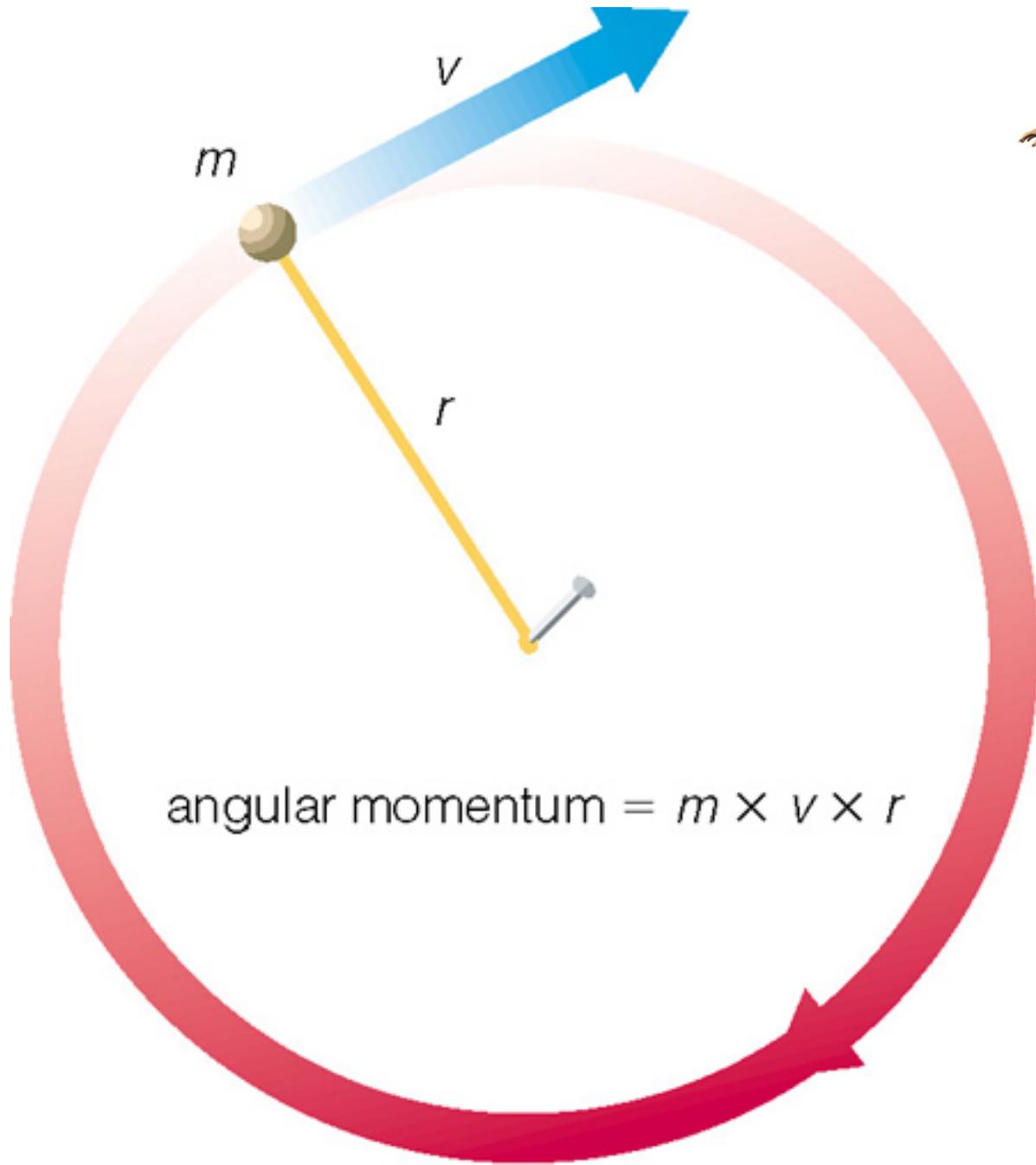
Collapse of the Solar Nebula



Collapse

Spin-up

Form Disk



Conservation of
Angular
Momentum

$MVR = \text{Constant}$

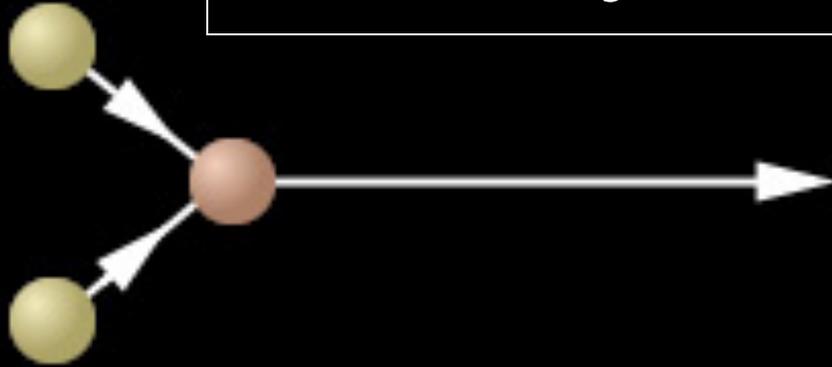
Where did the angular momentum come from???

Small random motions averaging out to a tiny bulk motion

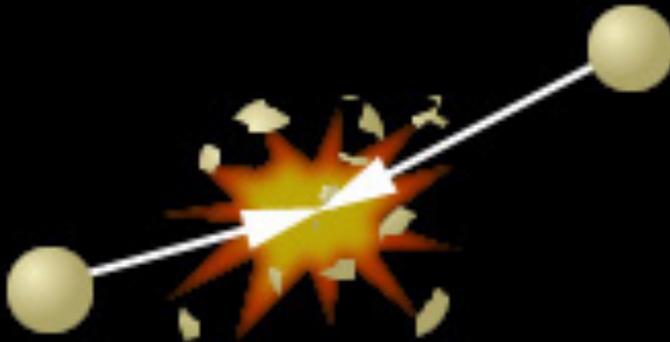
- this bulk motion is then “amplified” (due to conservation of angular momentum) as the cloud collapses

Demonstration: Take a flat-bottomed bowl (preferably glass) and fill it to ~2" deep with water. Shake a little pepper into the water. Stab the water with a stick/pencil - not stir. Then wait. After about a minute or two you should see a slow bulk motion develop, rotating one direction or the other.

Why a Disk?

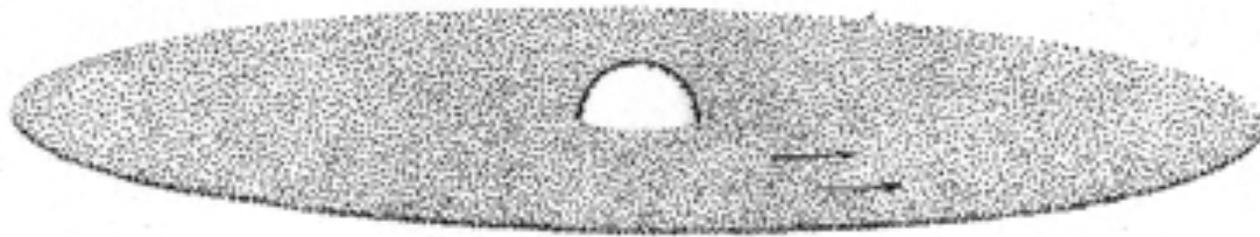


Oblique collisions ▶ regular orbits

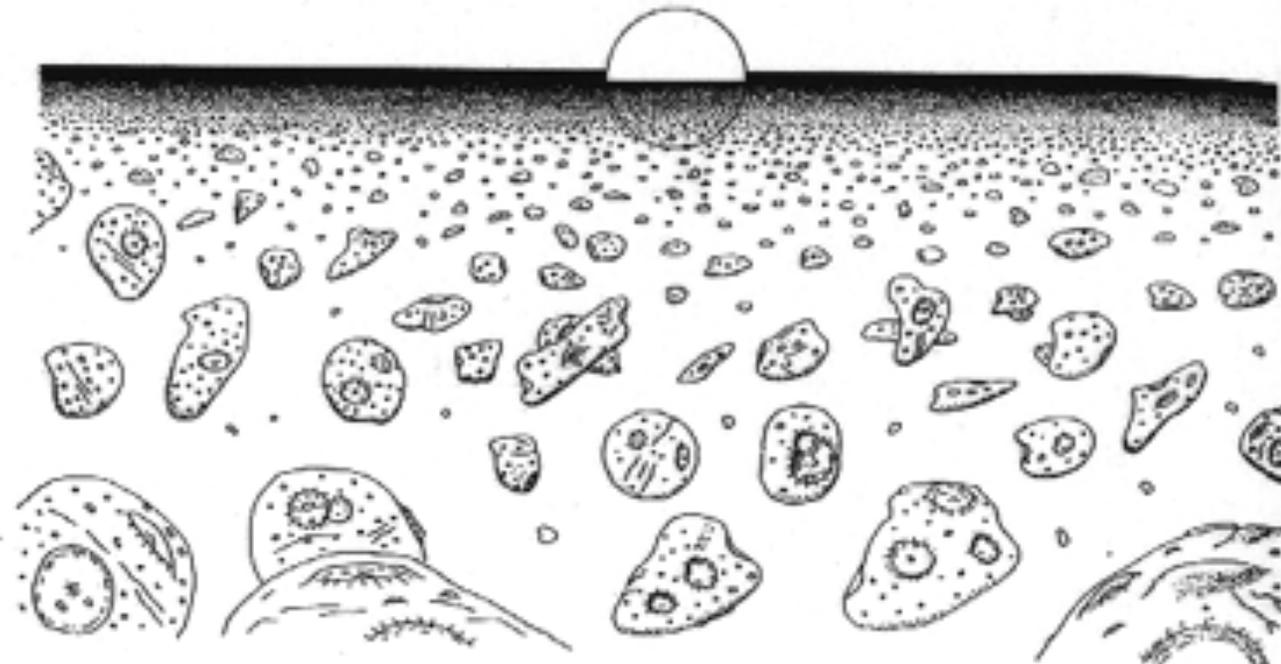


Head-on collisions ▶ smaller object

- As the cloud collapses (due to gravity) the gases, dust and stuff orbit the central mass.
- On the timescale of an orbit, gravity still balances the centrifugal force.
- The disk is not formed by being “flung out into a disk”.
- Nor does gravity of the disk “pull the material down onto a disk”.
- These are common misconceptions.



*Go with the flow or crash to oblivion
- Extreme Conformism!*



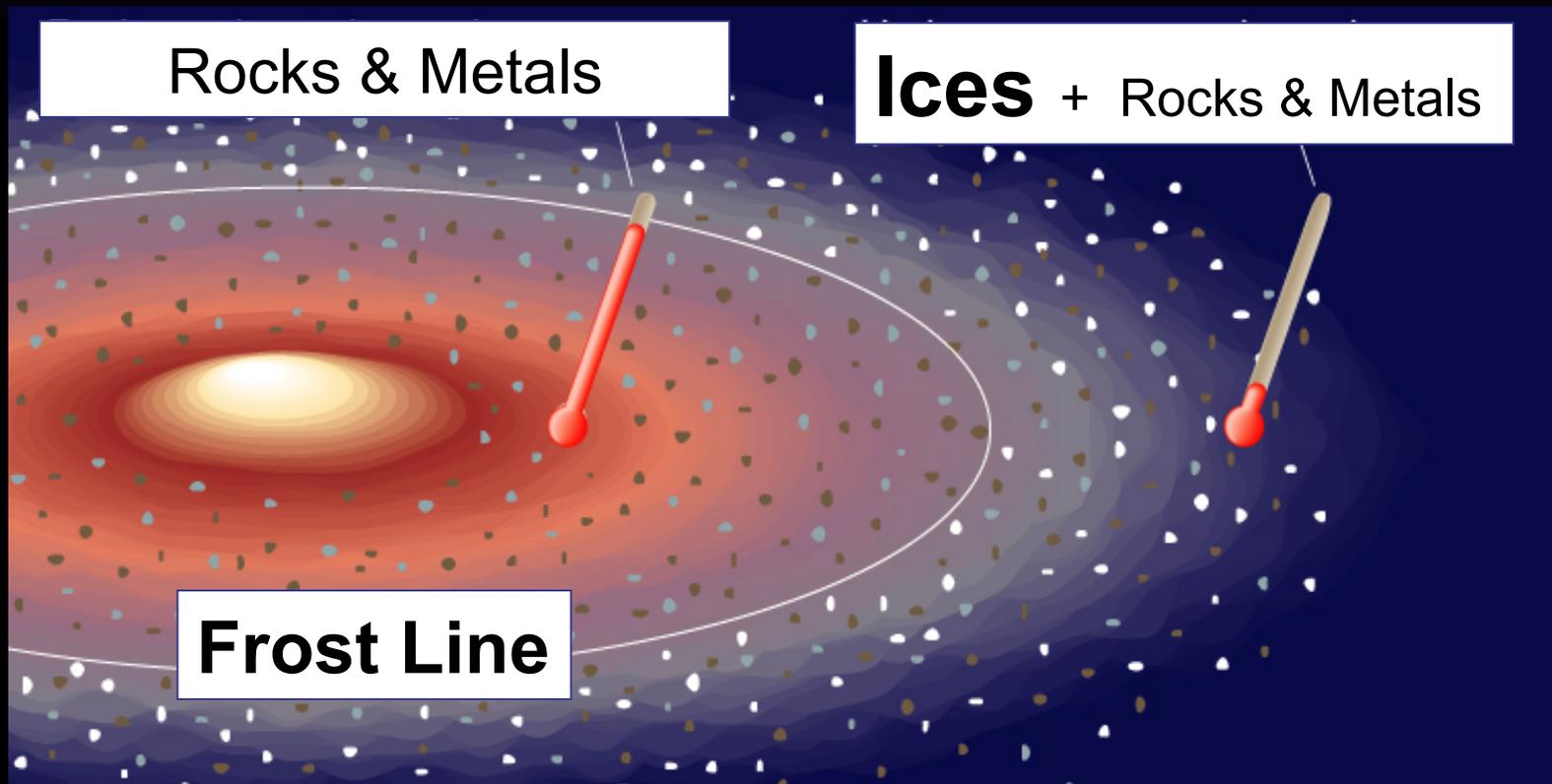
Collapse of the Solar Nebula



Rocks & Metals

Ices + Rocks & Metals

Frost Line



Summary of Solar System Formation - Part 1

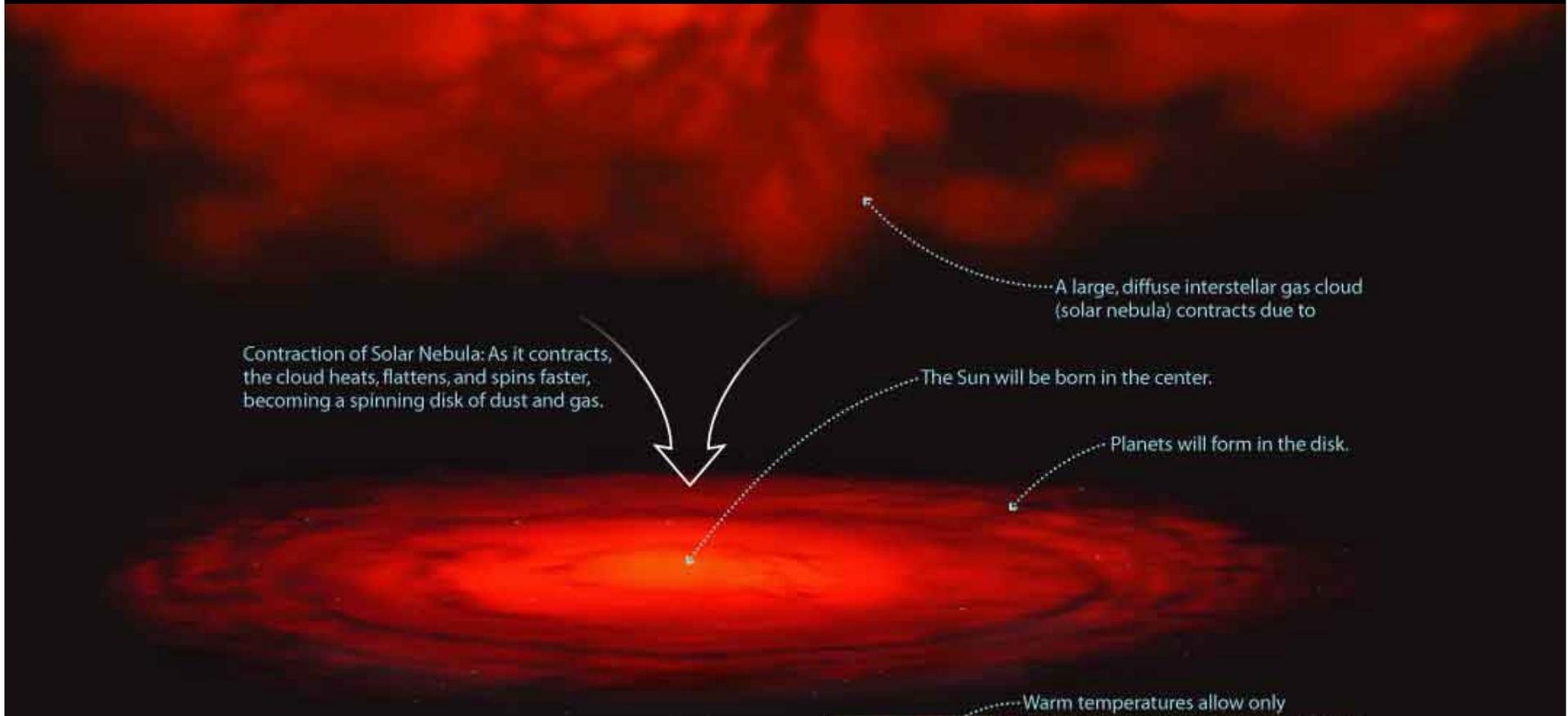
Contraction of Solar Nebula: As it contracts, the cloud heats, flattens, and spins faster, becoming a spinning disk of dust and gas.

A large, diffuse interstellar gas cloud (solar nebula) contracts due to

The Sun will be born in the center.

Planets will form in the disk.

Warm temperatures allow only



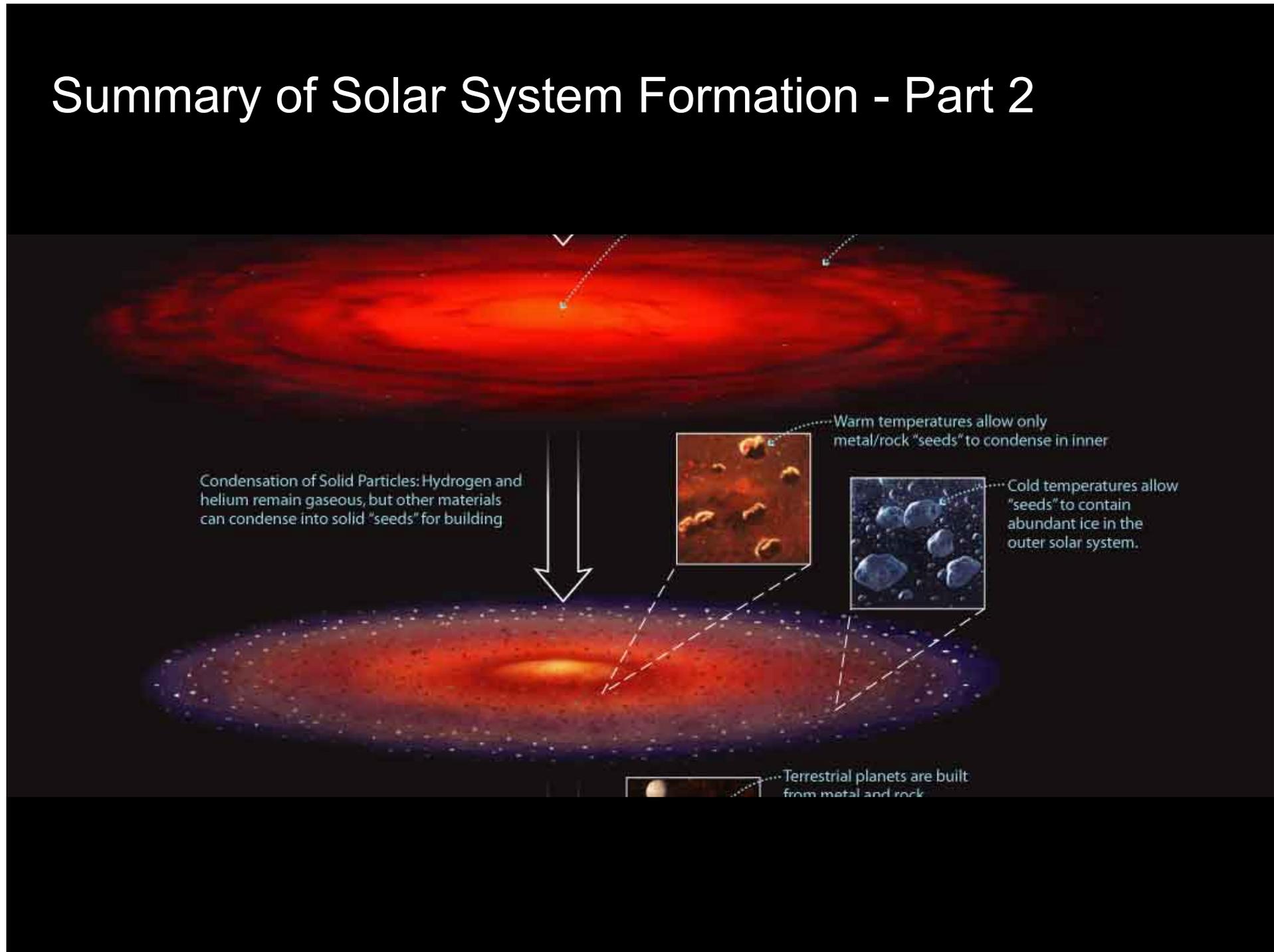
Summary of Solar System Formation - Part 2

Condensation of Solid Particles: Hydrogen and helium remain gaseous, but other materials can condense into solid "seeds" for building

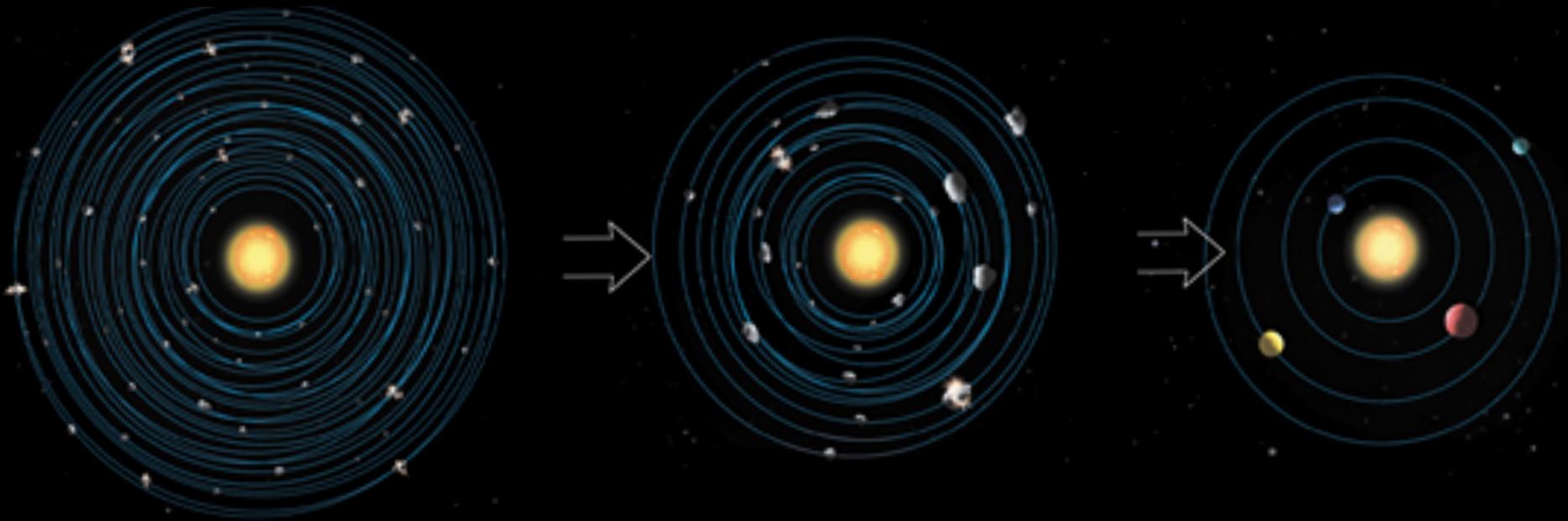
Warm temperatures allow only metal/rock "seeds" to condense in inner

Cold temperatures allow "seeds" to contain abundant ice in the outer solar system.

Terrestrial planets are built from metal and rock

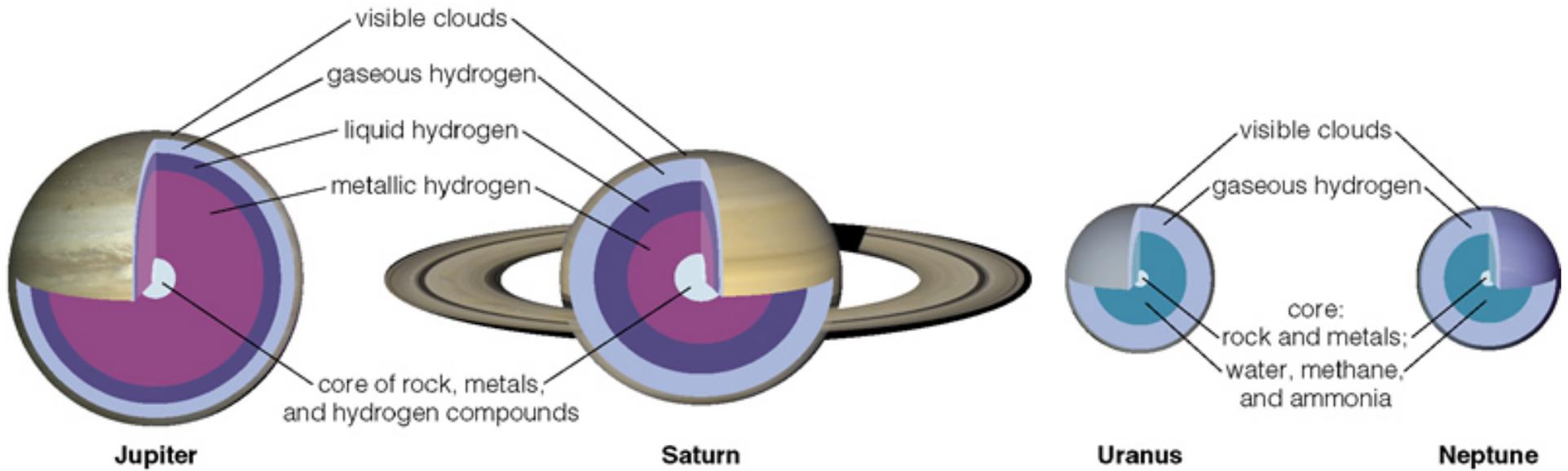




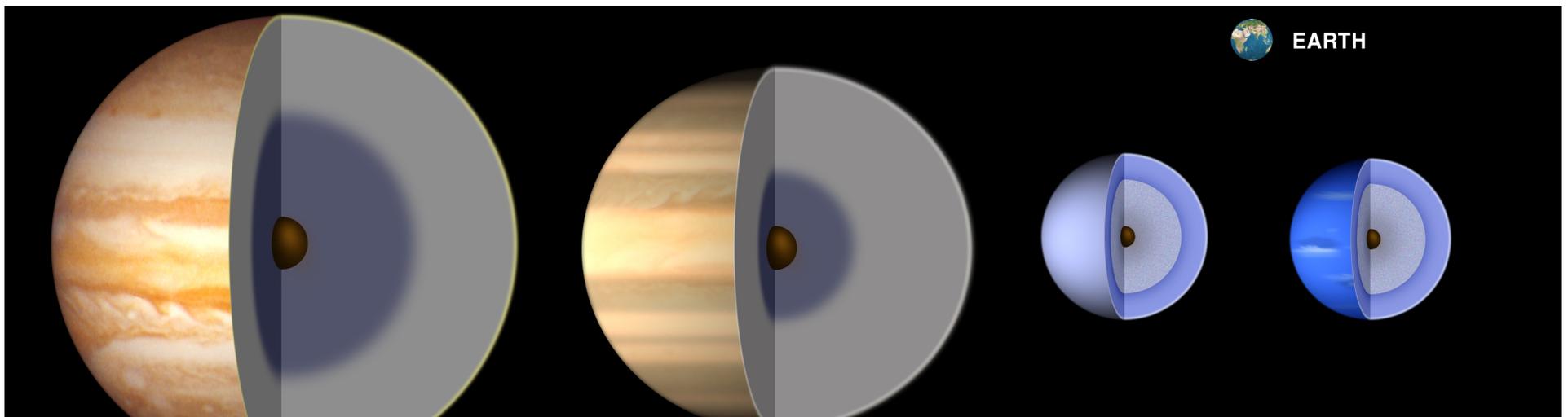


Planetesimals

- ***Oligarchic growth*** - the bigger get bigger
- REALLY big planetesimals (~20 x Earth) strong gravity pulls in hydrogen - the most abundant gas - and become GIANT.



Hydrogen envelopes over cores of rock, metals and **Water**, **Ammonia**, **Methane**



Why 3 Types of Planets?

1. Cosmic Abundance of Elements

2. Temperature Colder Farther from Sun

∅ Abundant ices condense beyond frost line

∅ Snowballs -> bigger snowballs

∅ Giant snowballs have enough gravity to hold H - most abundant element - > giant gas planets - *Jupiter, Saturn, Uranus, Neptune*

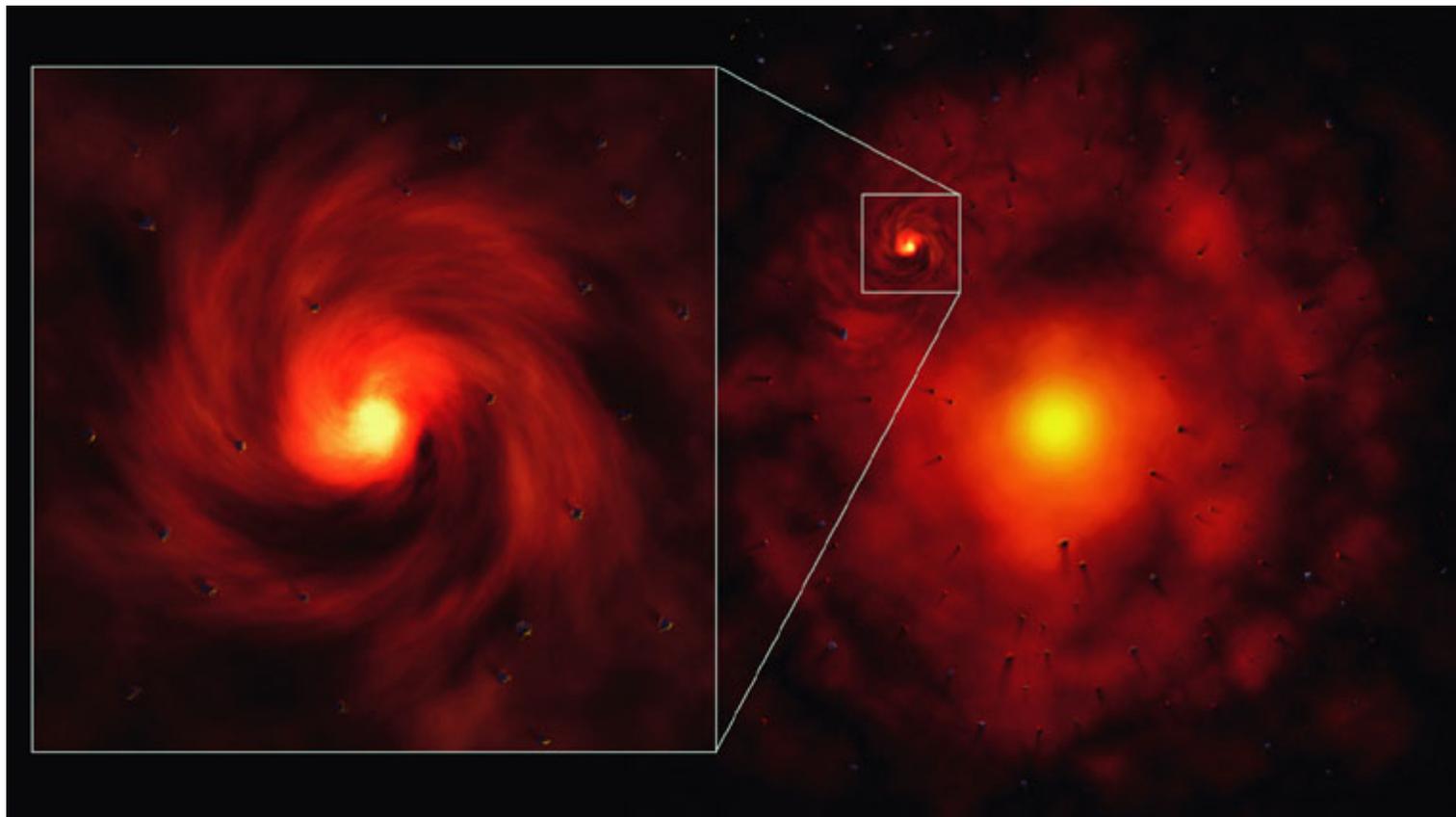
∅ Small amounts of rock & metal-> terrestrial planets - *Mercury, Venus, Earth, Mars*

∅ Ice dwarfs, comets, asteroids = leftovers

- Pluto, Xena, and ?

Building the Planets

- Each jovian planet formed its own “miniature” solar nebula.
- Moons formed out of the disk.



Summary of Solar System Formation - Part 3

Condensation of Solid Particles: Hydrogen and helium remain gaseous, but other materials can condense into solid "seeds" for building

Warm temperatures allow only metal/rock "seeds" to condense in inner

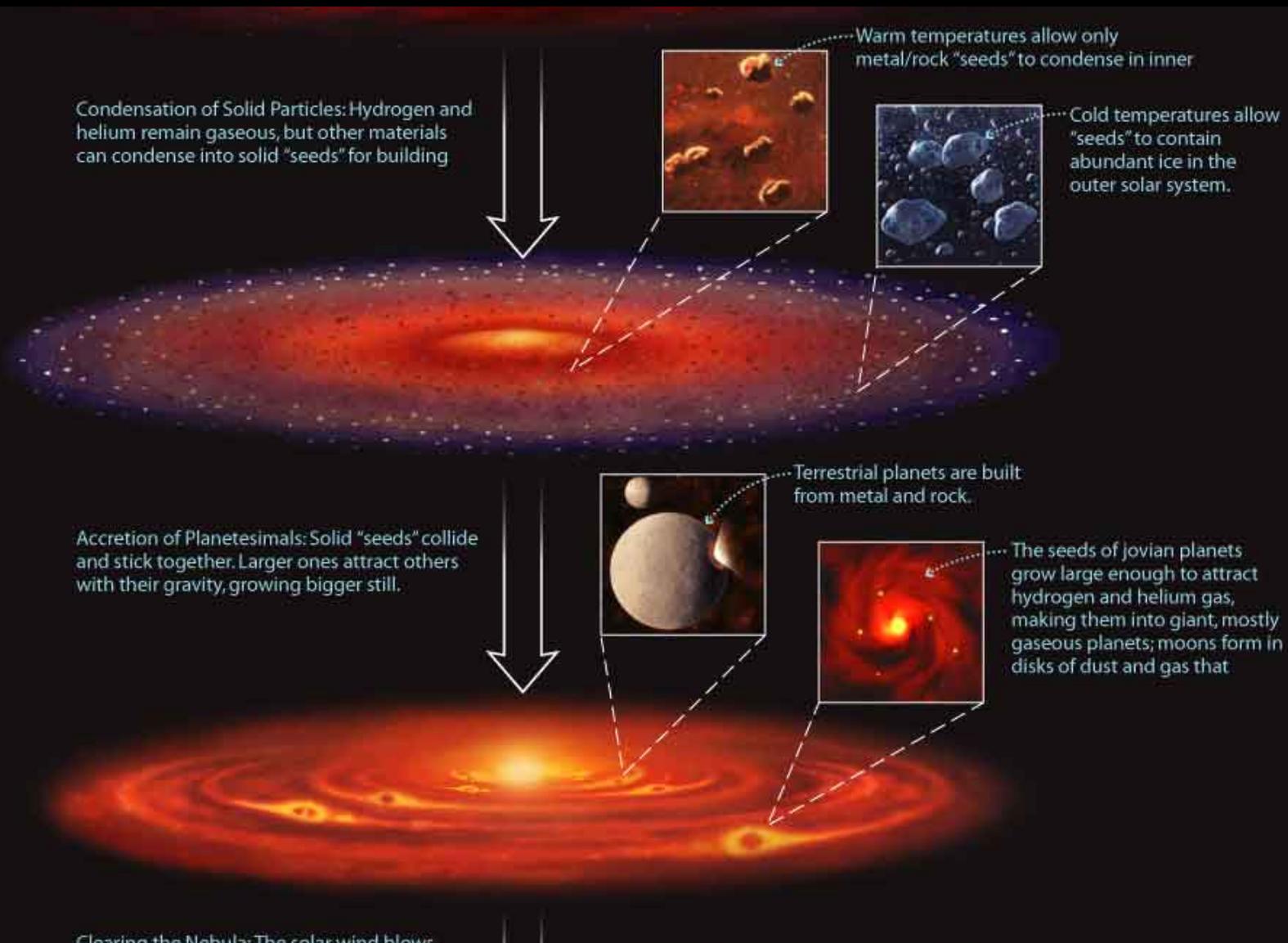
Cold temperatures allow "seeds" to contain abundant ice in the outer solar system.

Accretion of Planetesimals: Solid "seeds" collide and stick together. Larger ones attract others with their gravity, growing bigger still.

Terrestrial planets are built from metal and rock.

The seeds of jovian planets grow large enough to attract hydrogen and helium gas, making them into giant, mostly gaseous planets; moons form in disks of dust and gas that

Clearing the Nebula: The solar wind blows



But there's still (a) gas and (b)
junk between planets

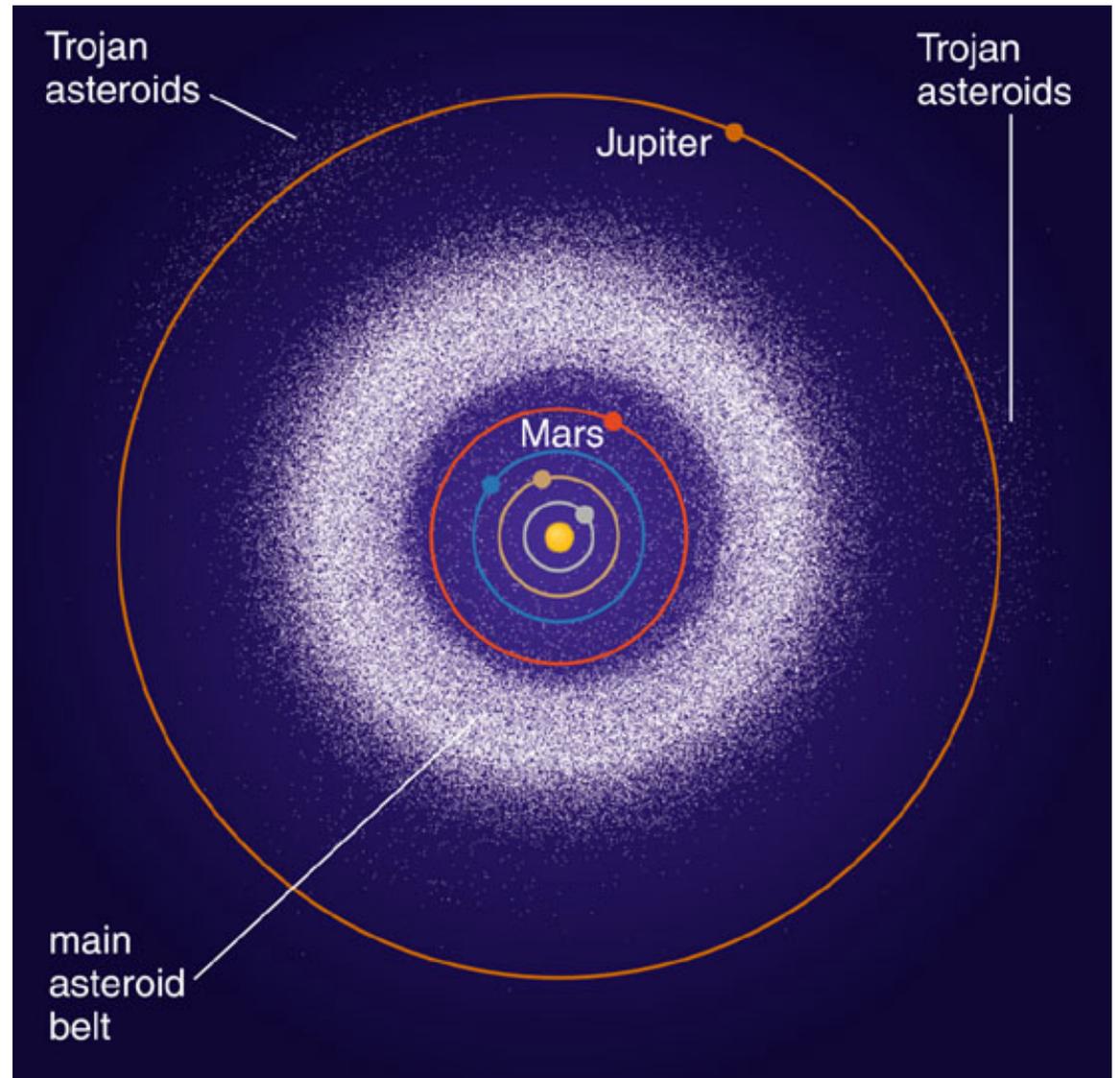
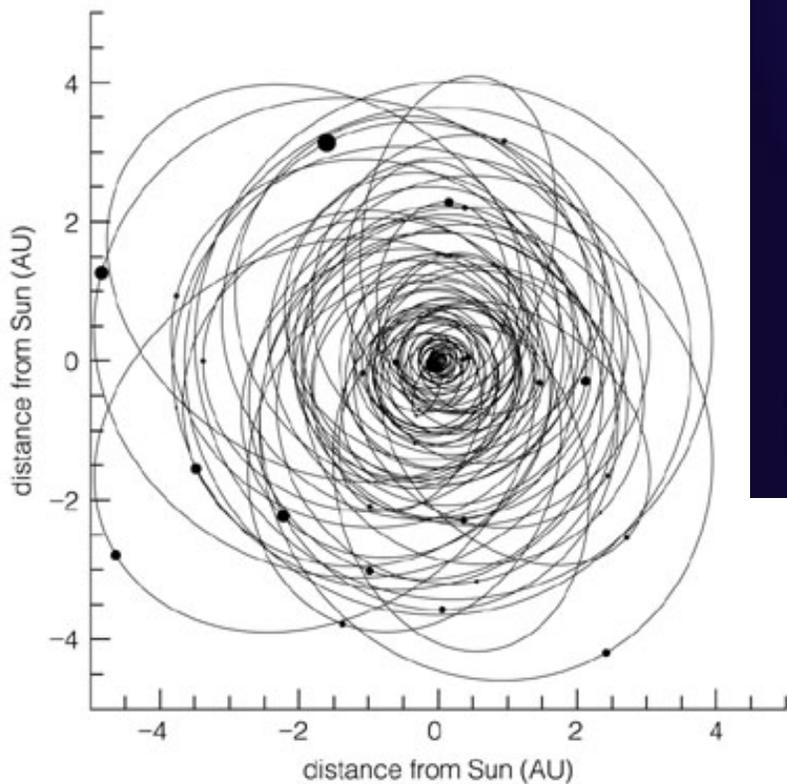
What happens to the gas?

Young stars go through phase of very
strong solar wind - blows away gas

But there's still (a) gas and (b) junk between planets

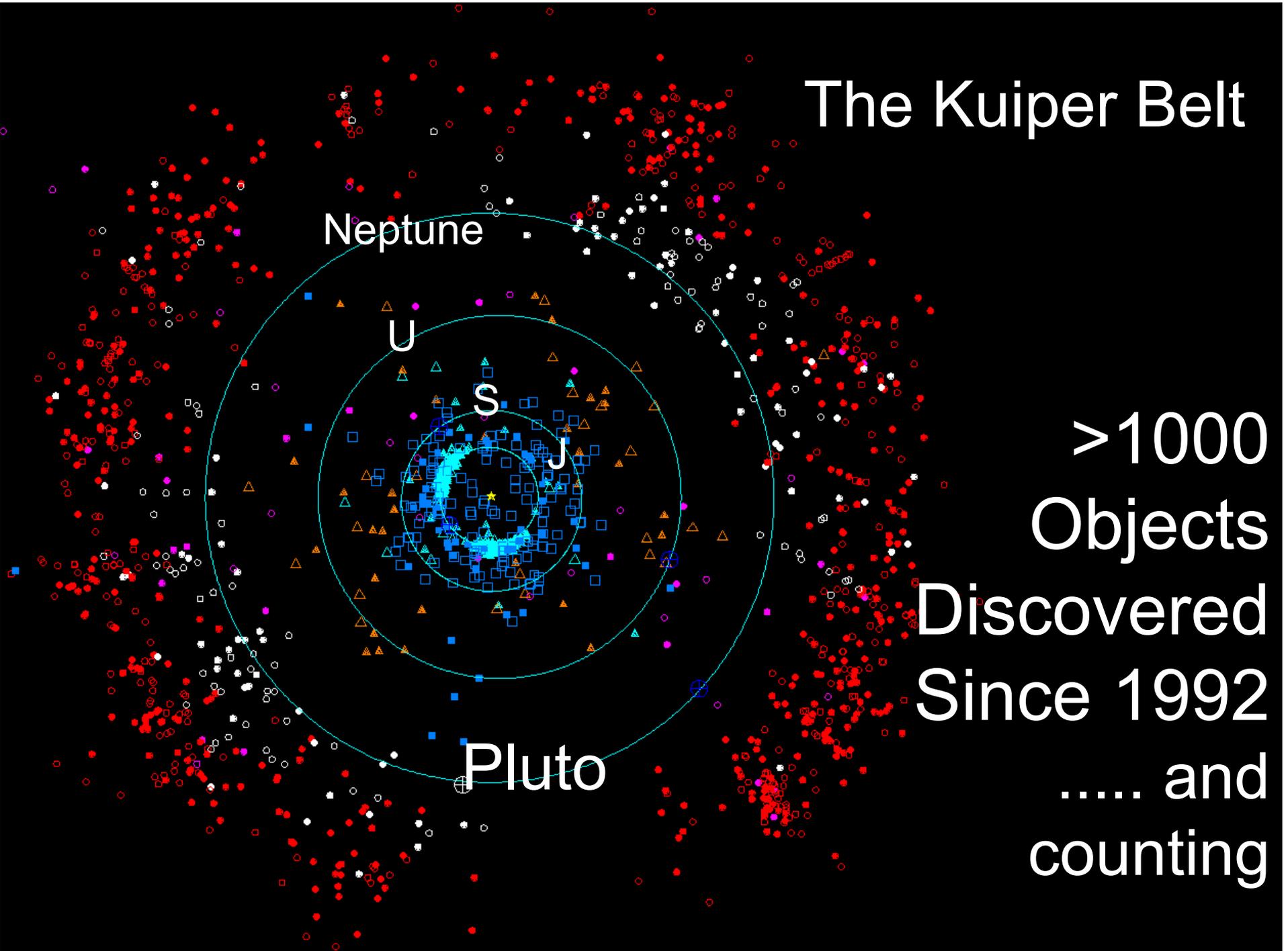
What happens to the junk?

- Gets kicked about - mostly by JSUN
- Kicked out of solar system
- Herded - asteroid belt, Kuiper Belt
- Captured as moons
- Bashed into young planets -> Moon, Charon, others?
- Delivers water (+other volatiles) to Earth

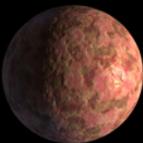


100s of thousands of asteroids
 100s Near Earth Objects
 100s Trojans $\pm 60^\circ$ of Jupiter

The Kuiper Belt



Xena & Gabriella - 10th Planet?



Sedna
800-1100 miles
in diameter



Quaoar
(800 miles)



Pluto
(1400 miles)



Moon
(2100 miles)



Earth
(8000 miles)

*All the more
reason to get
out there and
explore!*

Summary of Solar System Formation - Part 4

Accretion of Planetesimals: Solid "seeds" collide and stick together. Larger ones attract others with their gravity, growing bigger still.

Terrestrial planets are built from metal and rock.

The seeds of jovian planets grow large enough to attract hydrogen and helium gas, making them into giant, mostly gaseous planets; moons form in disks of dust and gas that

Clearing the Nebula: The solar wind blows remaining gas into interstellar space.

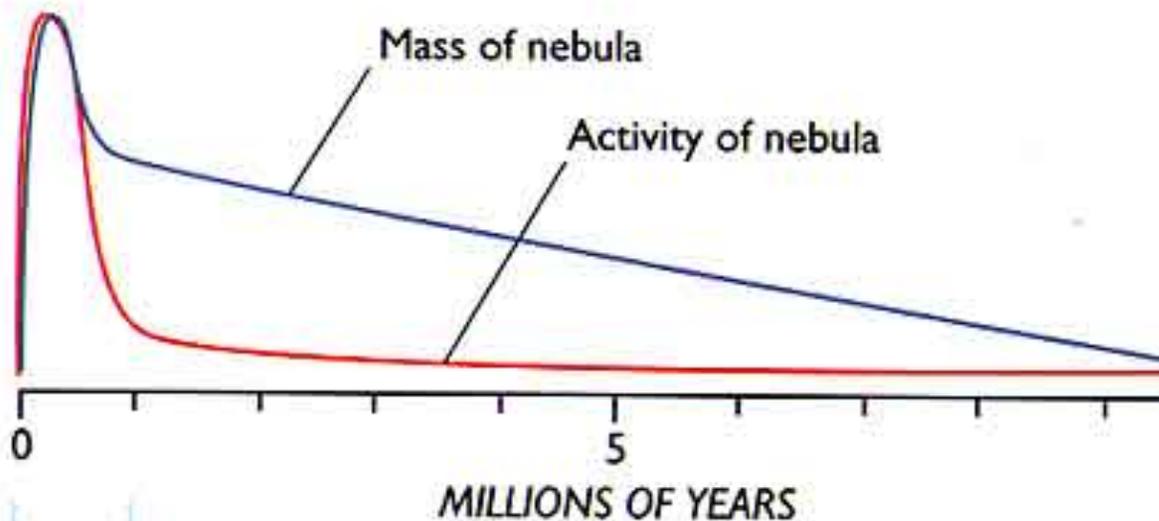
Terrestrial planets remain in the inner solar

Jovian planets remain in the outer solar

"Leftovers" from the formation process become asteroids (metal/rock) and comets (mostly

Not to scale

How long does this whole process take?



- Formation of Jovian gas protoplanets
- Formation of chondrites
- Accretion of rocky and icy planetesimals
Growth of terrestrial planets

- Capture of icy planetesimals by Jovian planets
- Gravitational scattering of asteroids and comets

- Accretion of icy planetesimals

- Jupiter and Saturn capture nebular gas

- Formation of chondrites
- Accretion of rocky planetesimals
- Growth of terrestrial planets

- Uranus and Neptune capture nebular gas

How long did all this take?

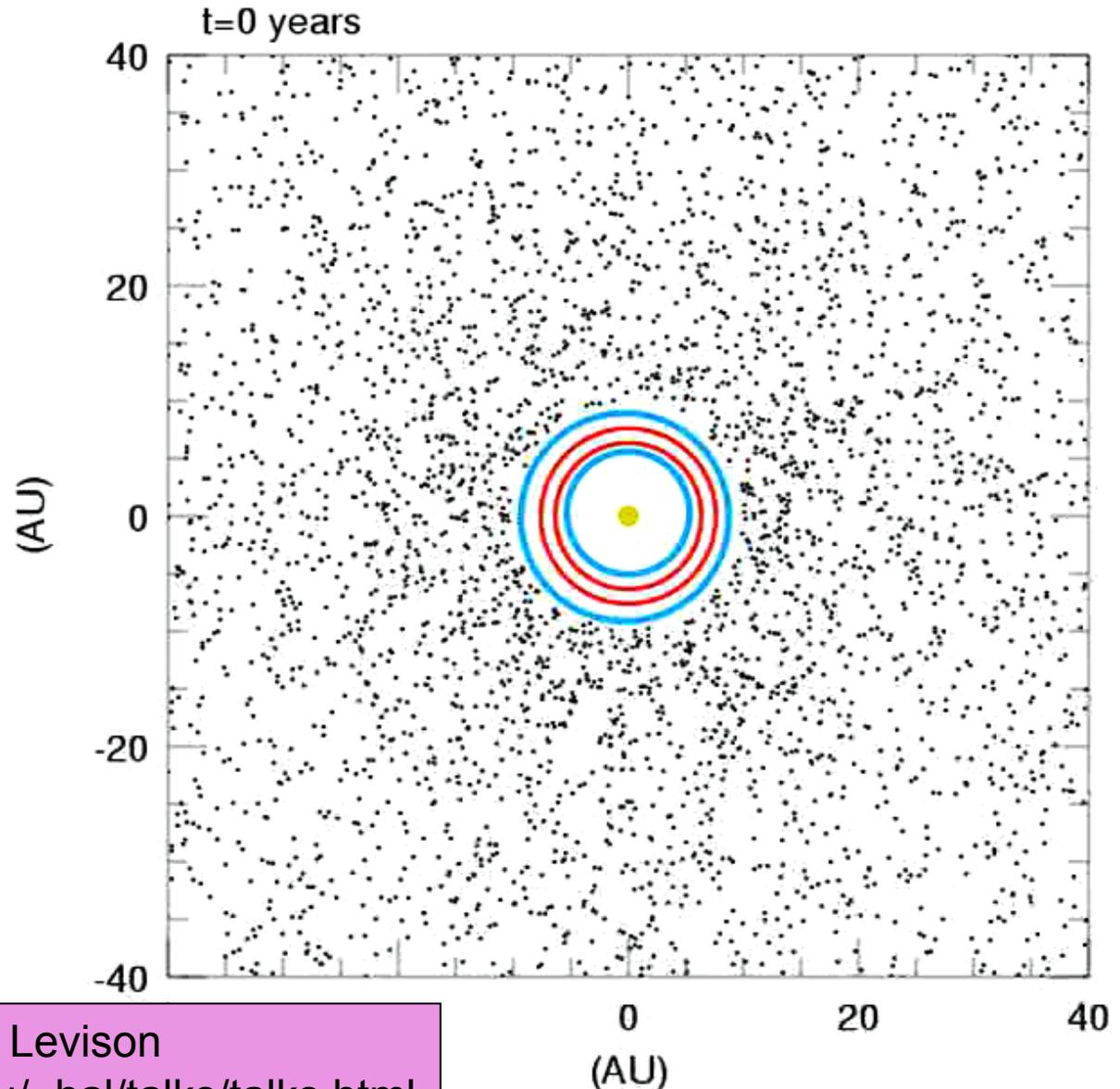
- Nebula collapse <1 MY
- planetesimal formation in 1MY
- J, S < 2 MY
- Terrestrials <4 MY
- Uranus & Neptune?

But it's not so simple...

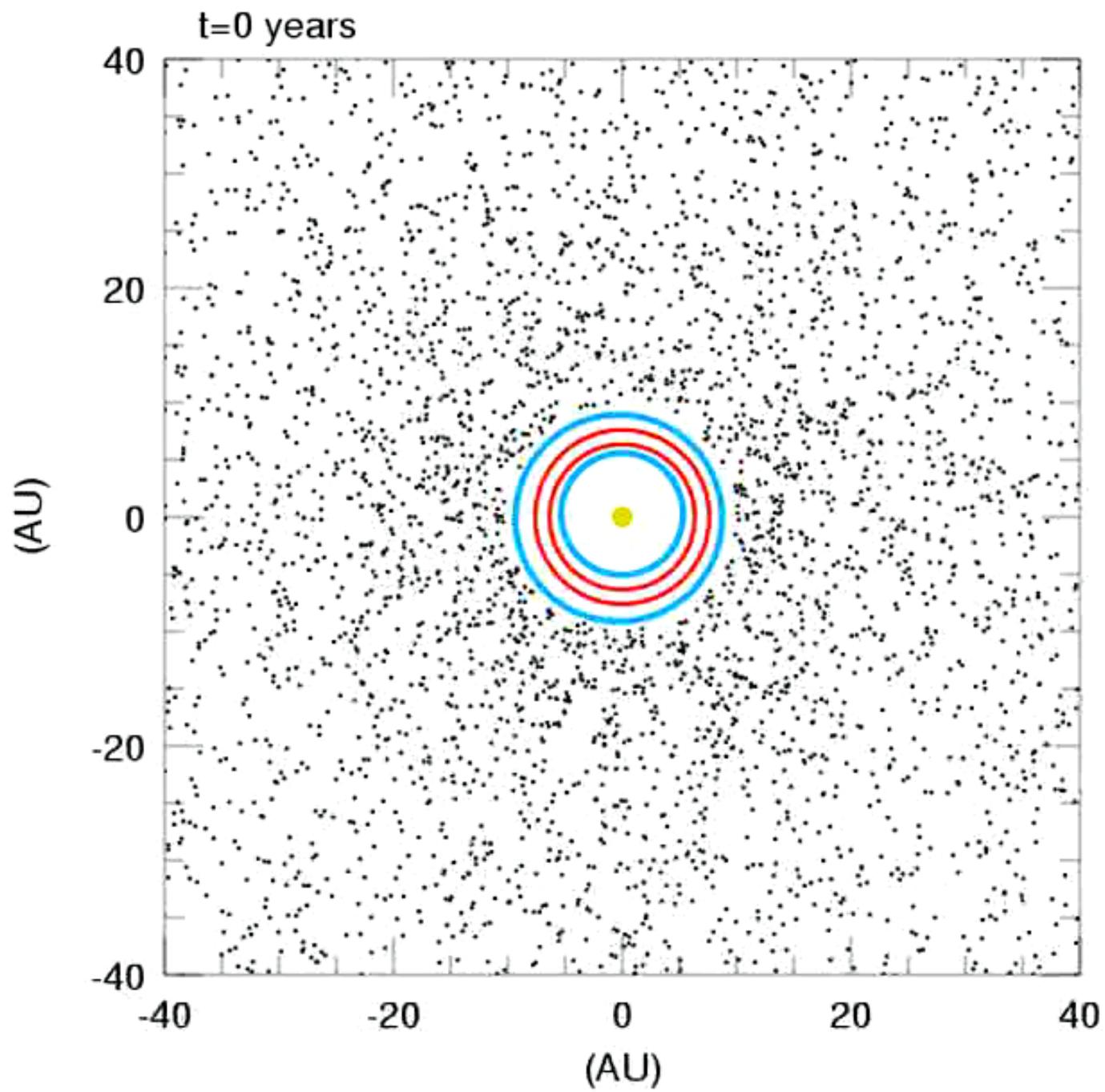
Jupiter & Saturn
& Neptune Uranus

- Hard to make U & N at present orbits
- Where did earth's oceans come from?
- How to make Kuiper Belt?

=> Migration?



See movies in talks by Hal Levison
<http://www.boulder.swri.edu/~hal/talks/talks.html>



Comets contain the most primitive material - they tell us about the solar nebula composition and the earliest steps in the process





Nucleus of Comet 19P/Borrelly
8 x 3.2 km in diameter
Imaged by Deep Space 1 2001
~43 m/pixel



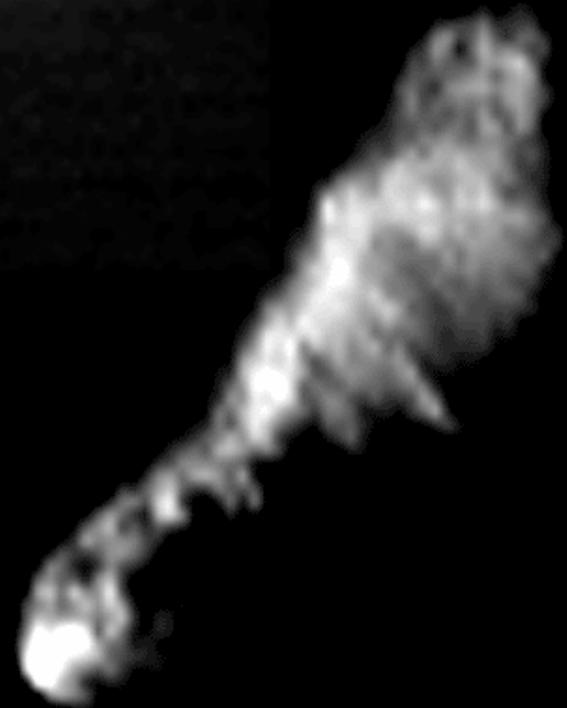
Nucleus of Comet 1P/Halley 16
x 8 km in diameter
Imaged by Giotto, 1986
~100 m/pixel (but only at top left)

Comet Borrelly

Deep Space 1

Comet Wild 2

Stardust

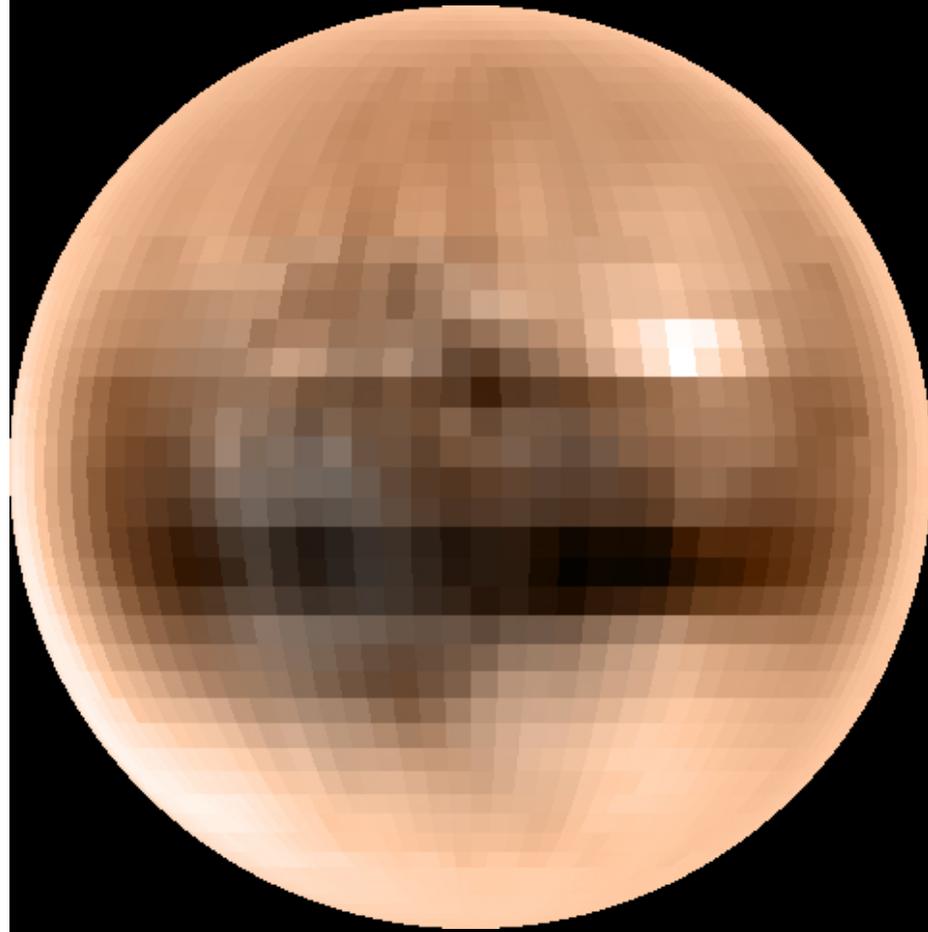




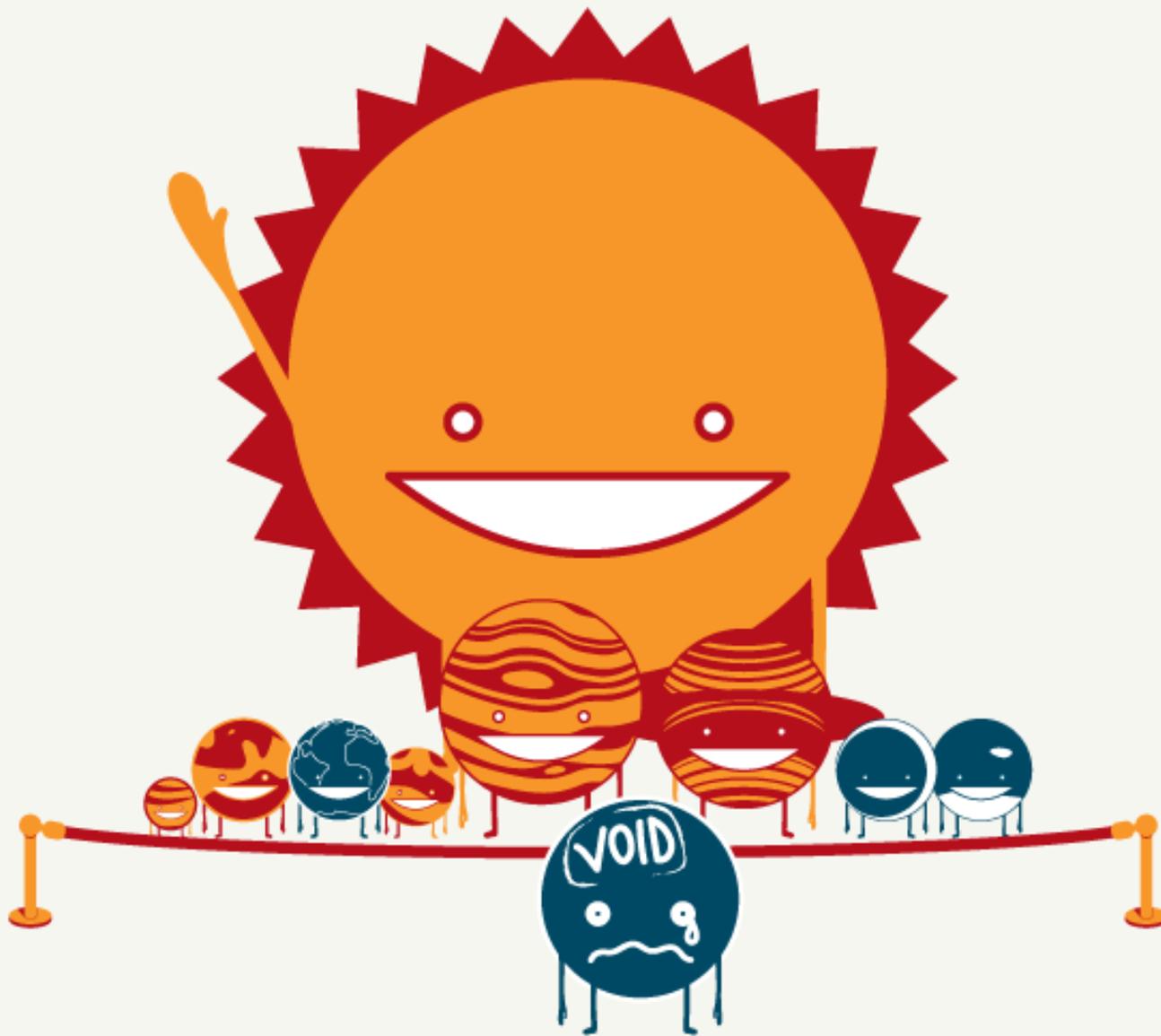
Pandora - moon of Saturn
114 x 84 x 62 km



Asteroid Itokawa
Hayabusa

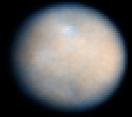


Objects like Pluto -
planetesimals or
planetary embryos -
tell us about
planetary growth
stages



Dwarf Planets

Ceres



Dysnomia



Eris

Charon



Pluto



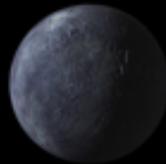
Makemake



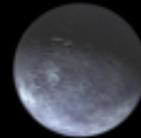
Haumea



Sedna



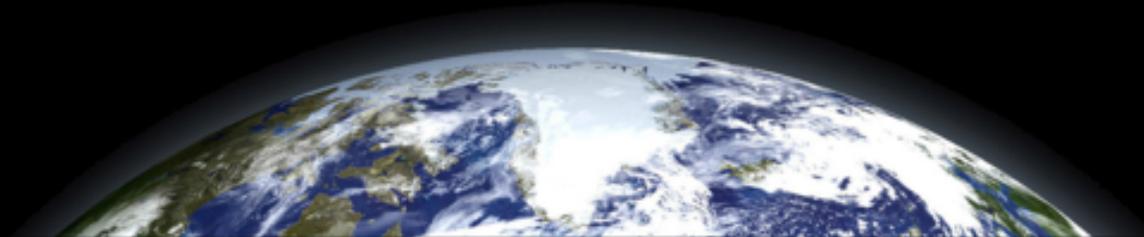
Orcus



Quaoar



Varuna

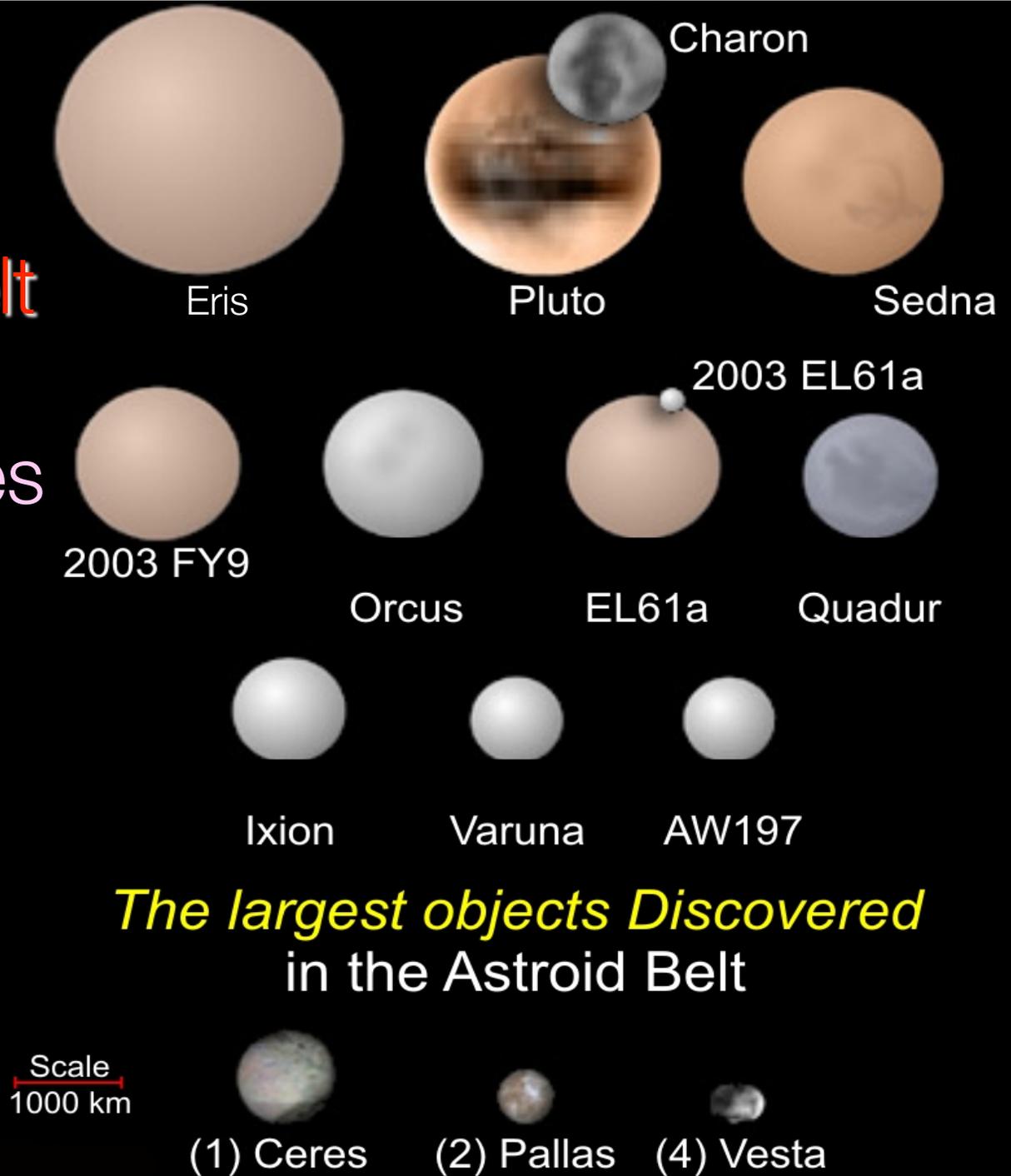


Dwarf Planets: The Most Populous Class of Planet

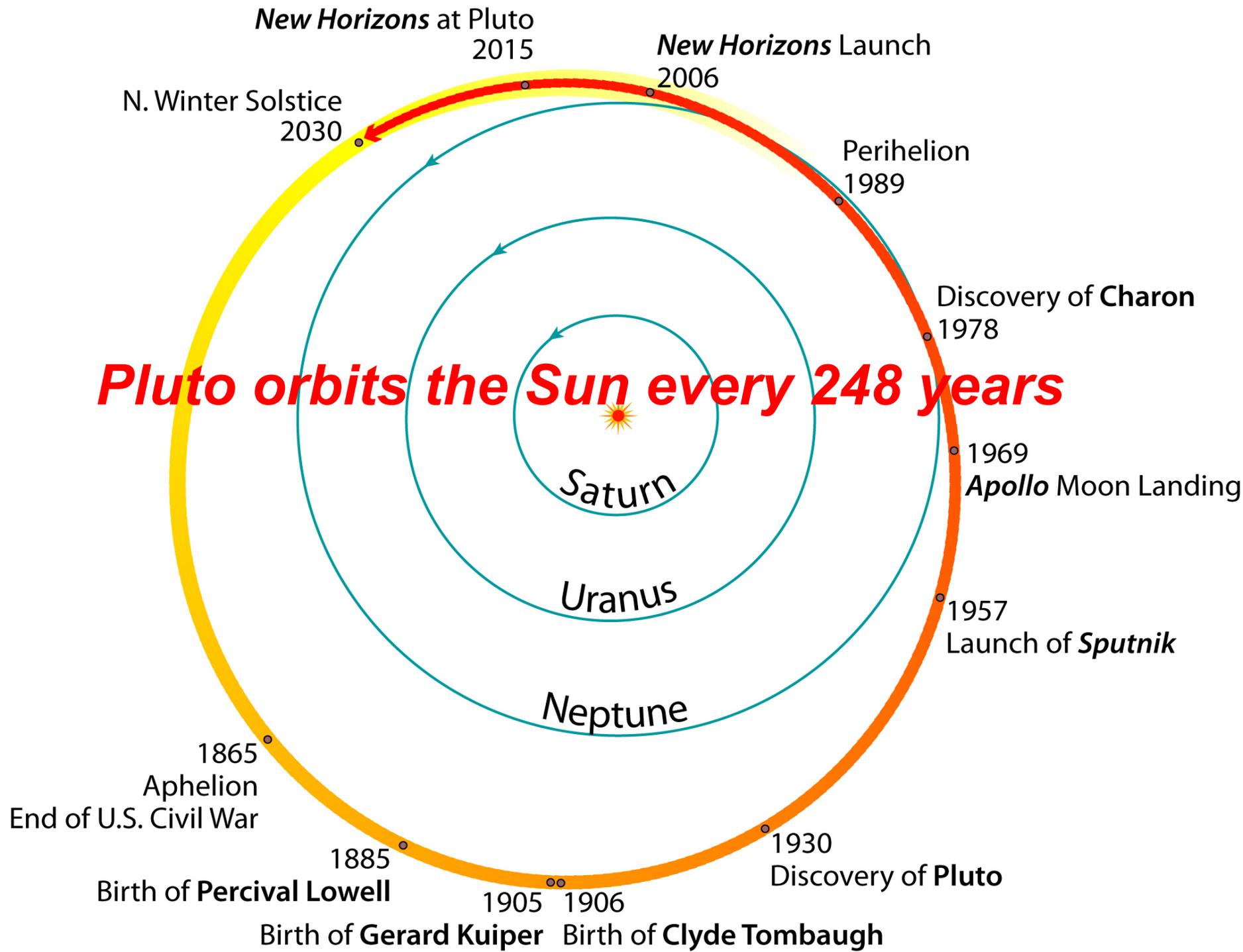


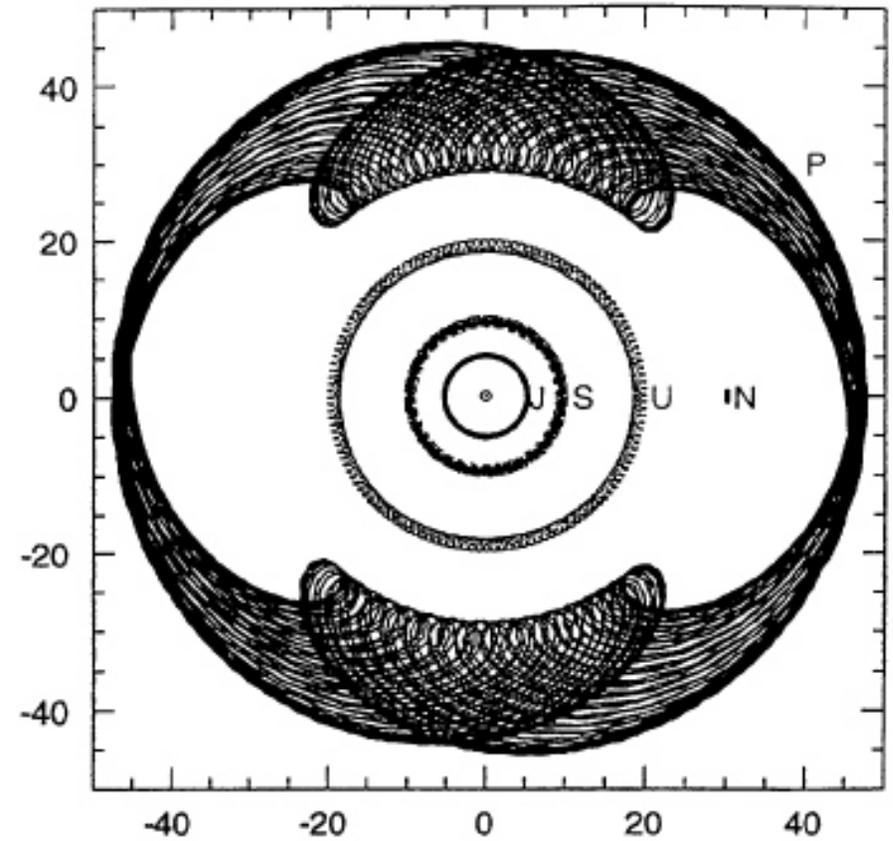
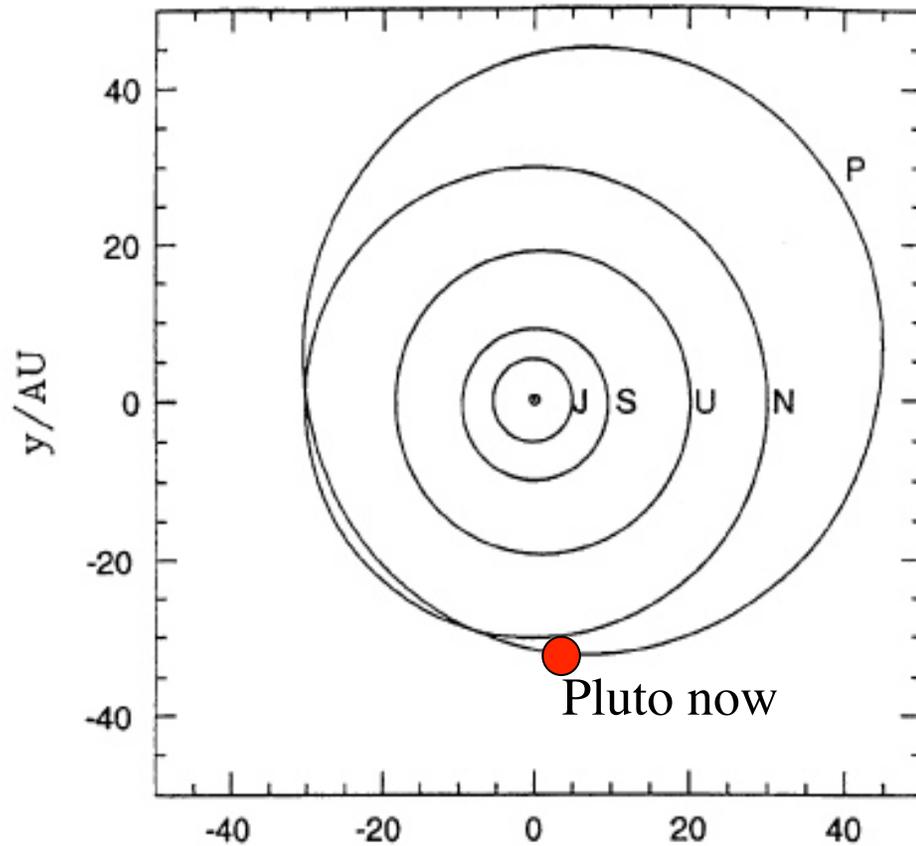
The Kuiper Belt

- Different sizes and colors
- Bigger than the largest asteroids

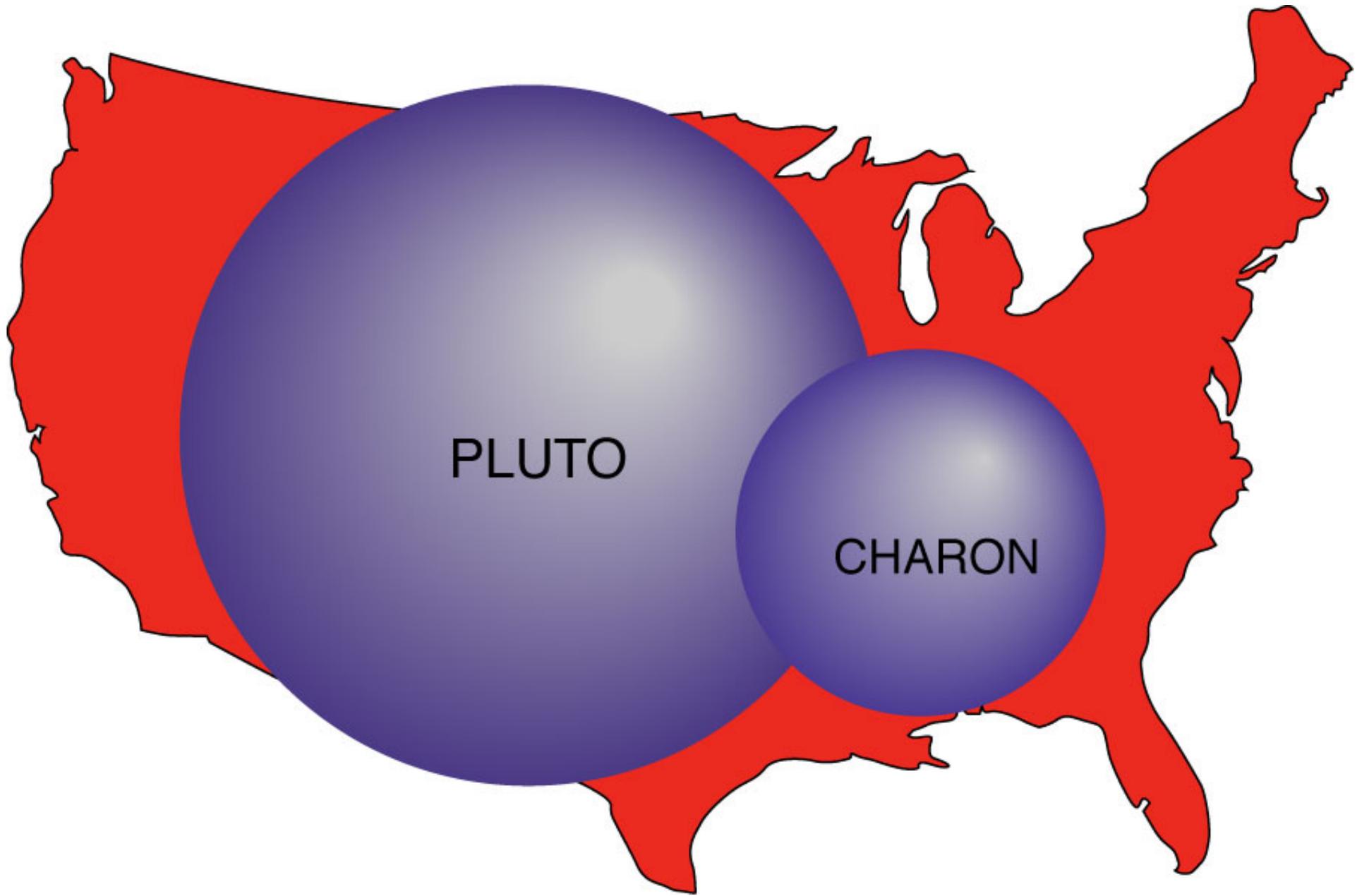


Pluto orbits the Sun every 248 years



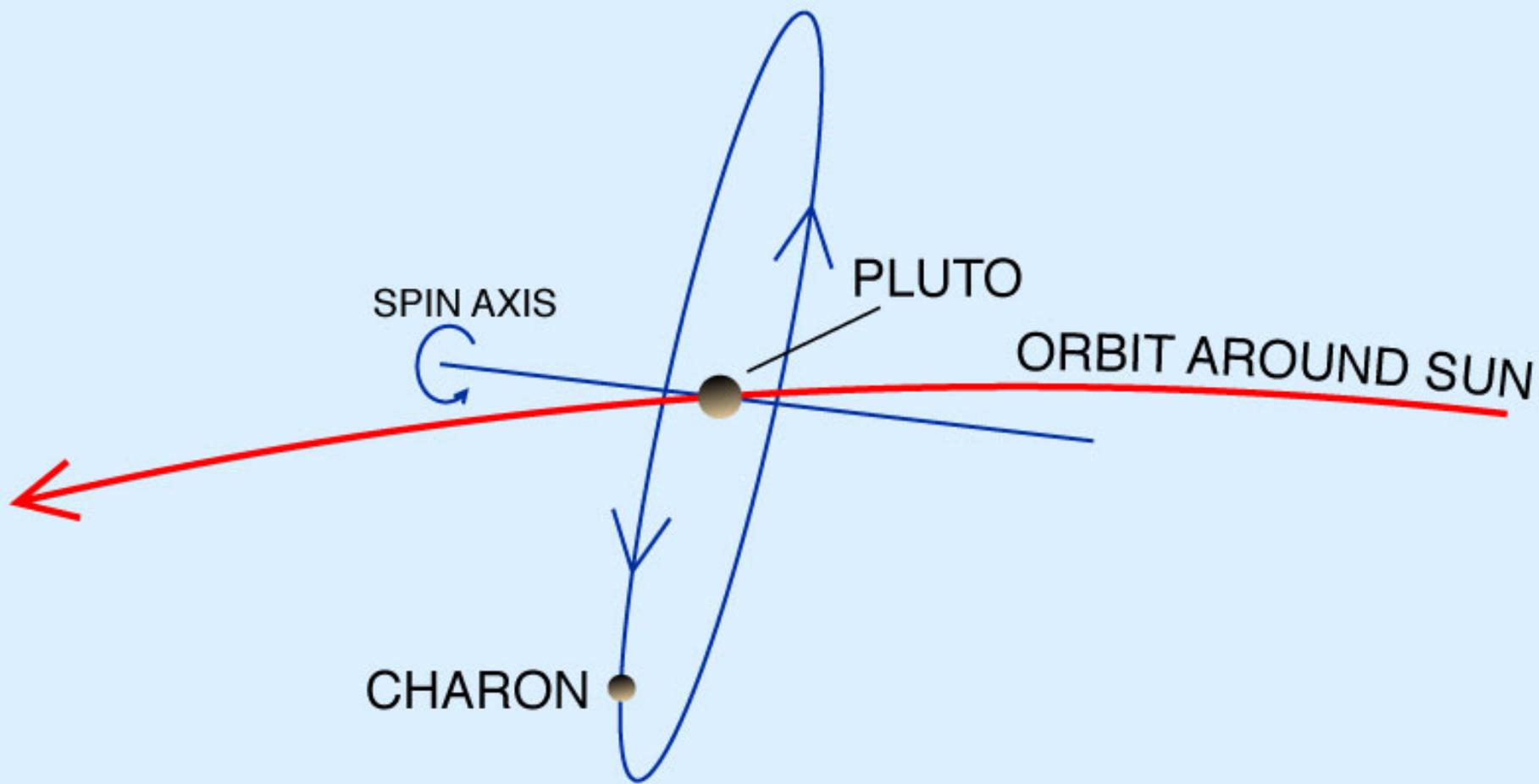


*Neptune and Pluto dance together
- but never get close*



PLUTO

CHARON



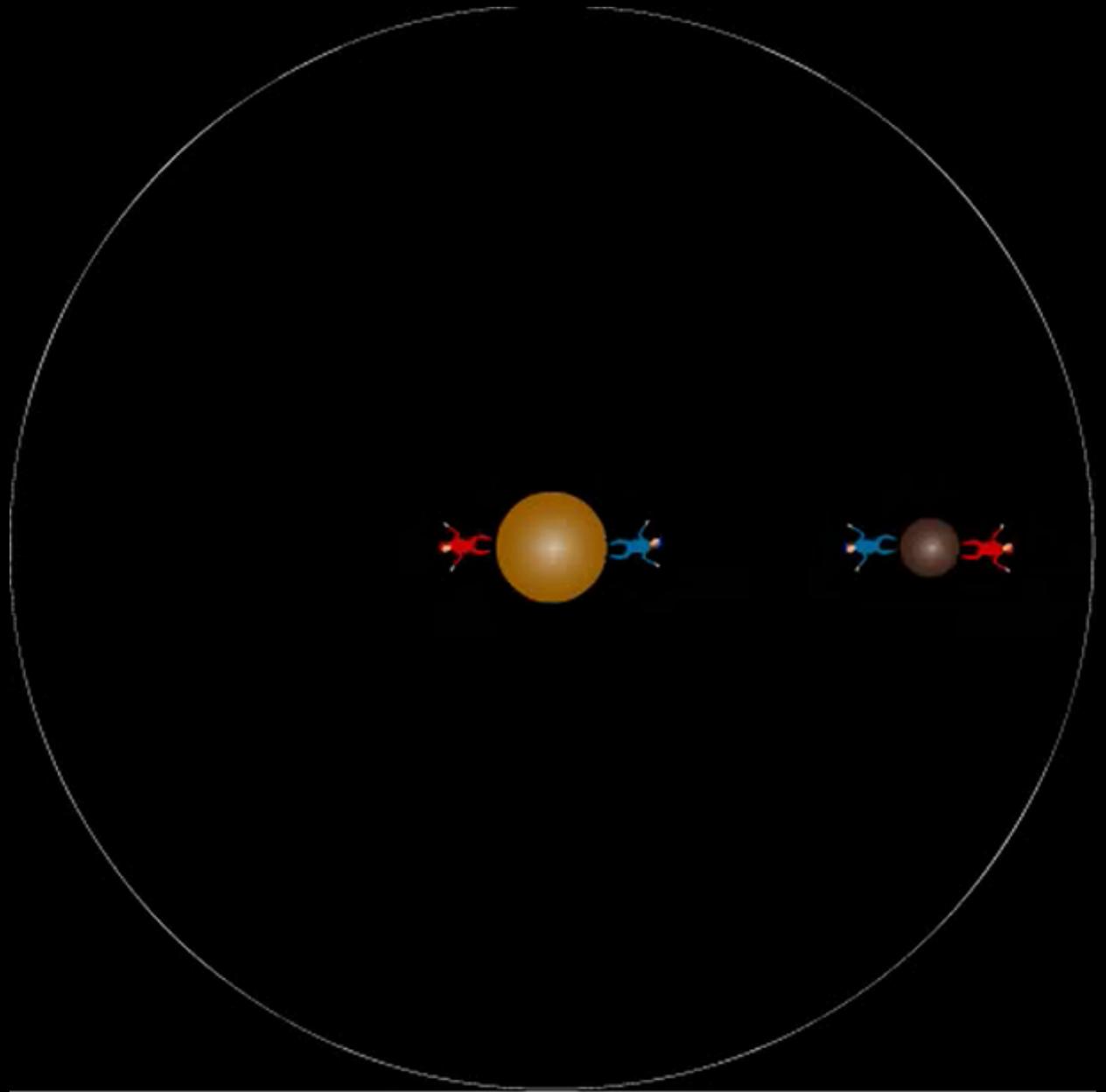
Binary Planet System

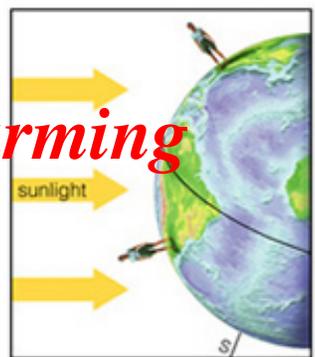
Mass of Pluto ~ 6 x Mass of Charon



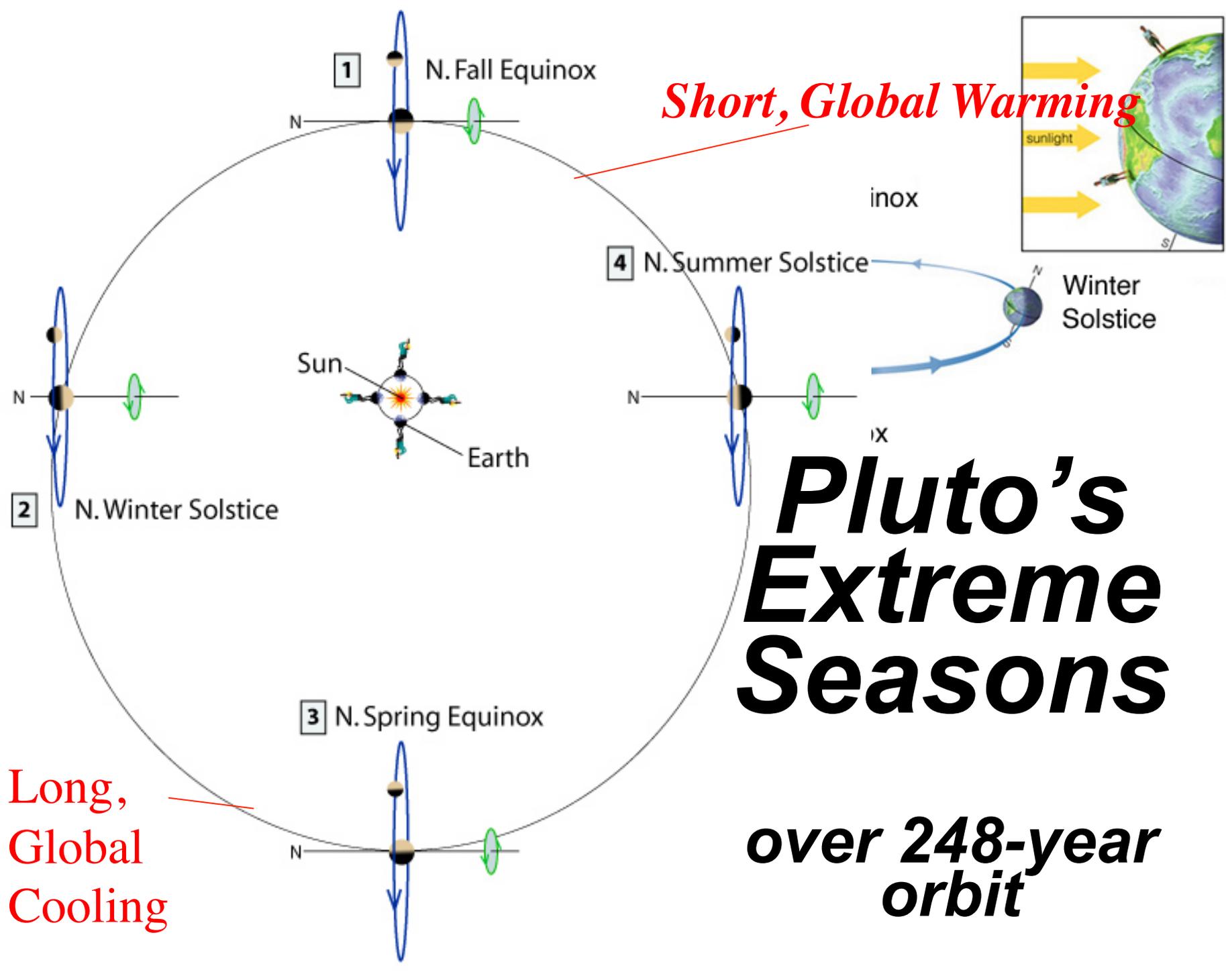
Spin-Orbit Phase Lock

Pluto &
Charon
both spin
and orbit
every 6.4
days





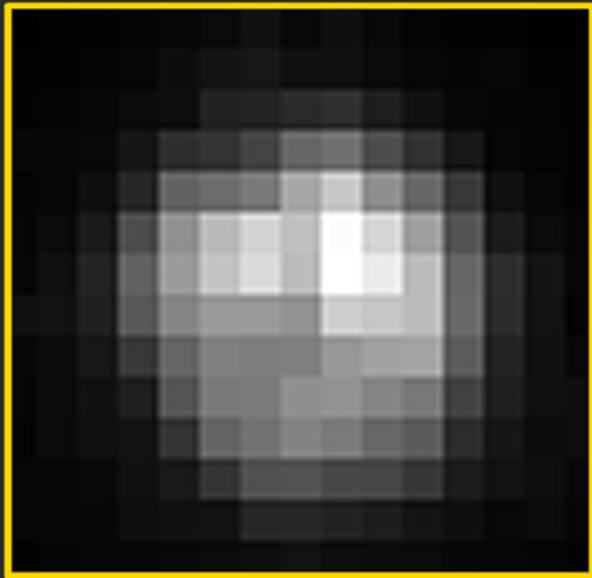
Short, Global Warming



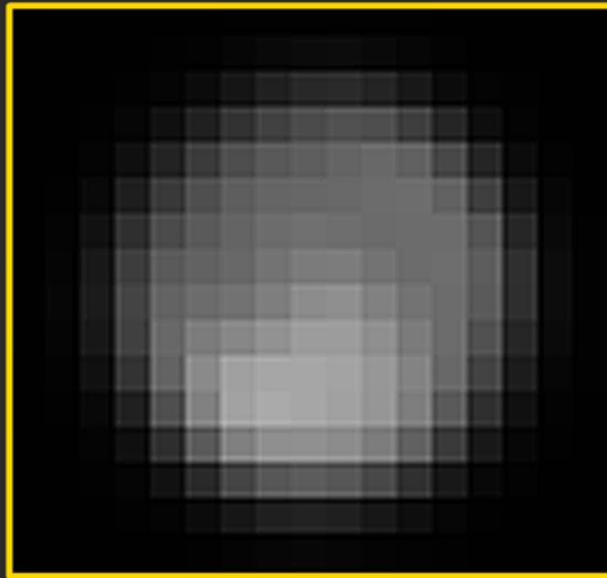
Long, Global Cooling

Pluto's Extreme Seasons

over 248-year orbit



**Pluto at best
Hubble
resolution**

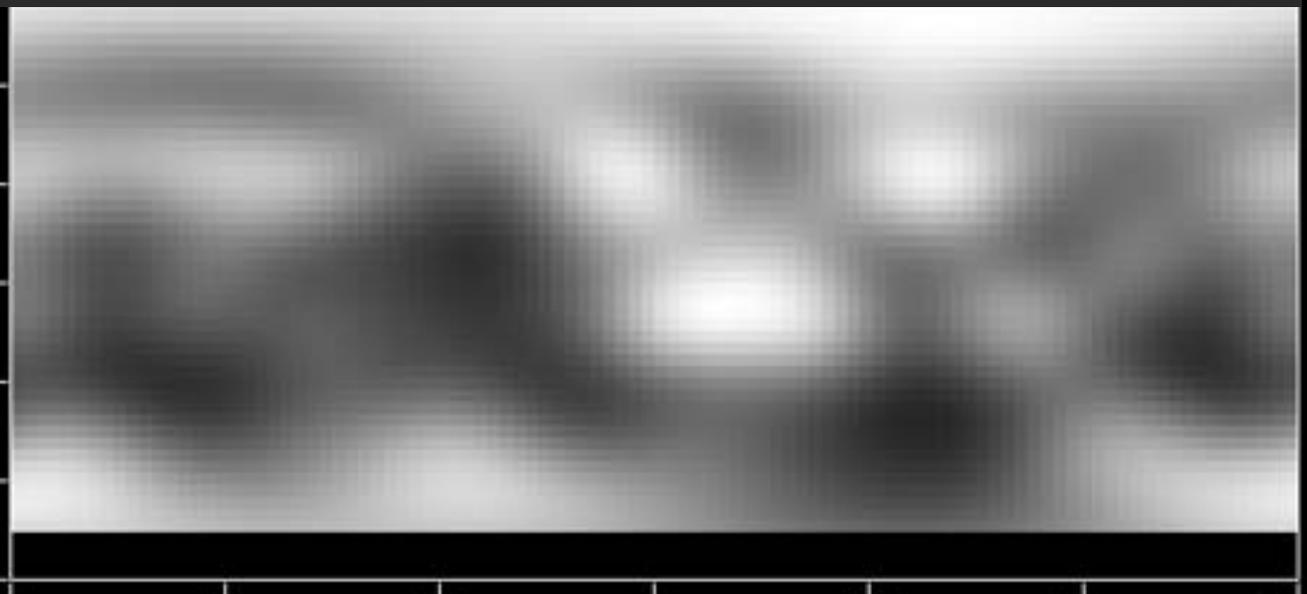


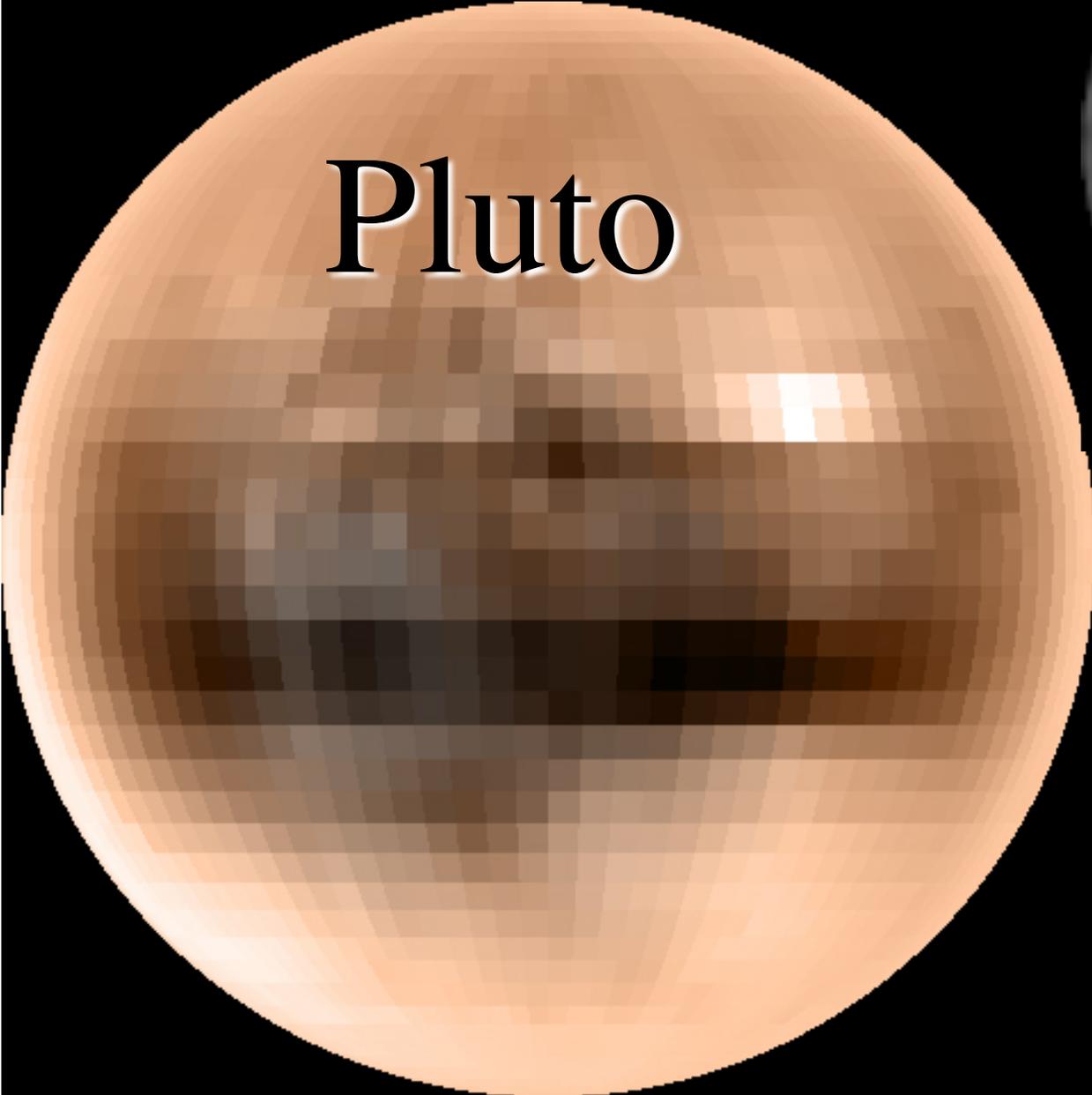
**Earth's Moon
at the same
resolution**



**Earth's Moon
at 5 km
per pixel**

**Best Map
of Pluto ⇒**





Pluto



Charon

Temperature:

40 Kelvin

-233 Celsius

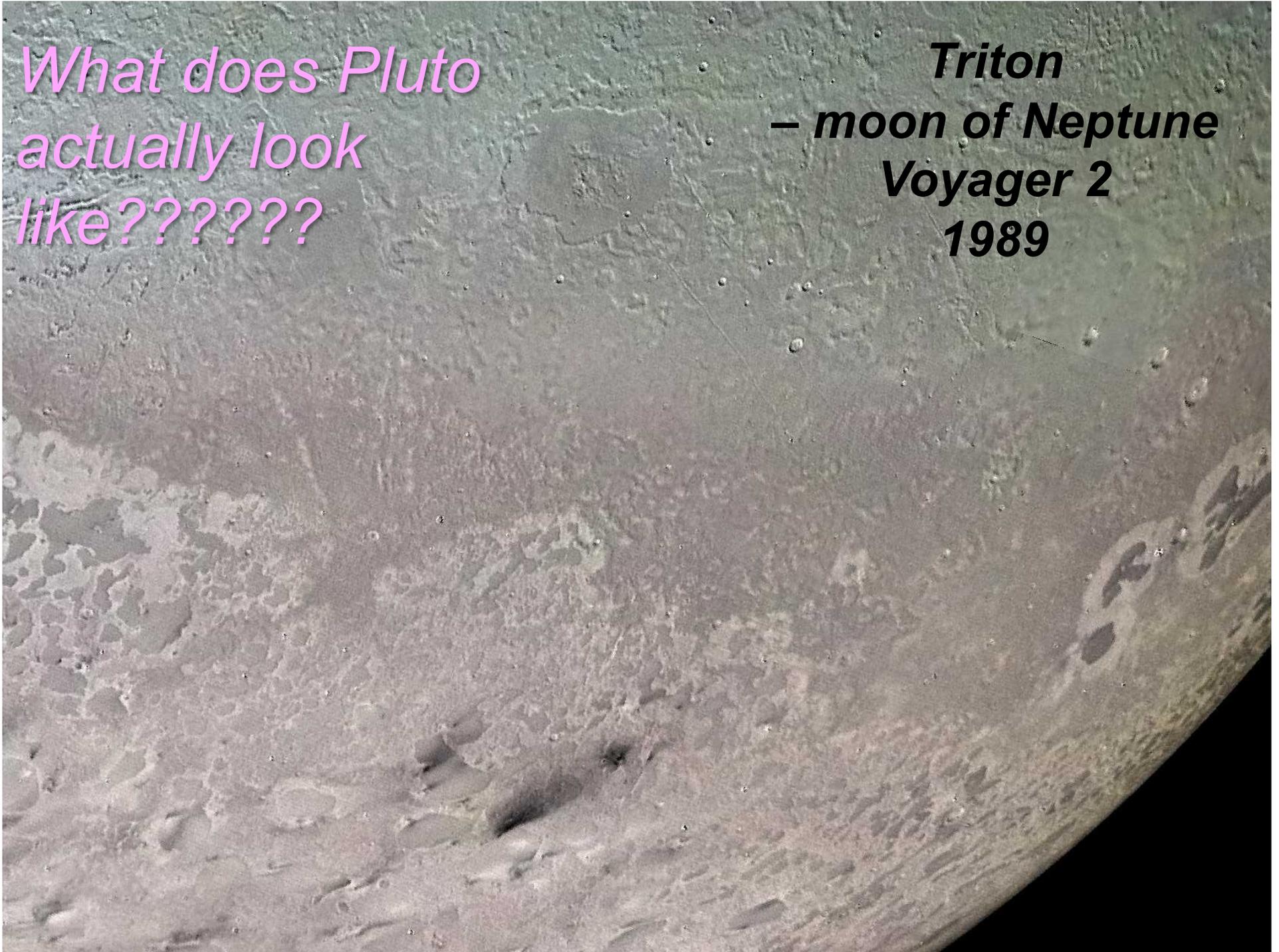
-340 F

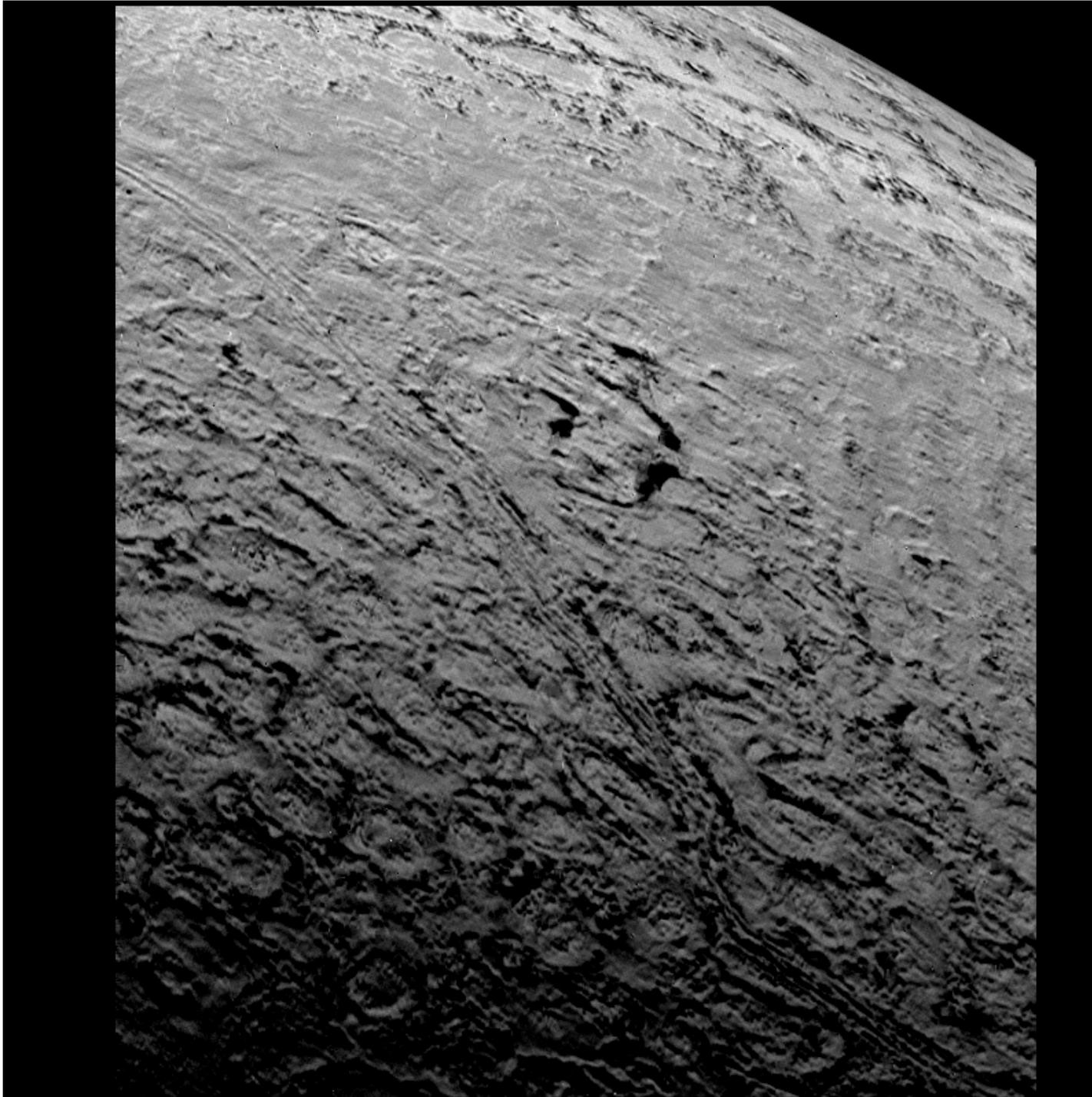
Surface = Ice

+ ... ???

*What does Pluto
actually look
like??????*

Triton
– moon of Neptune
Voyager 2
1989





Cantaloupe
terrain:

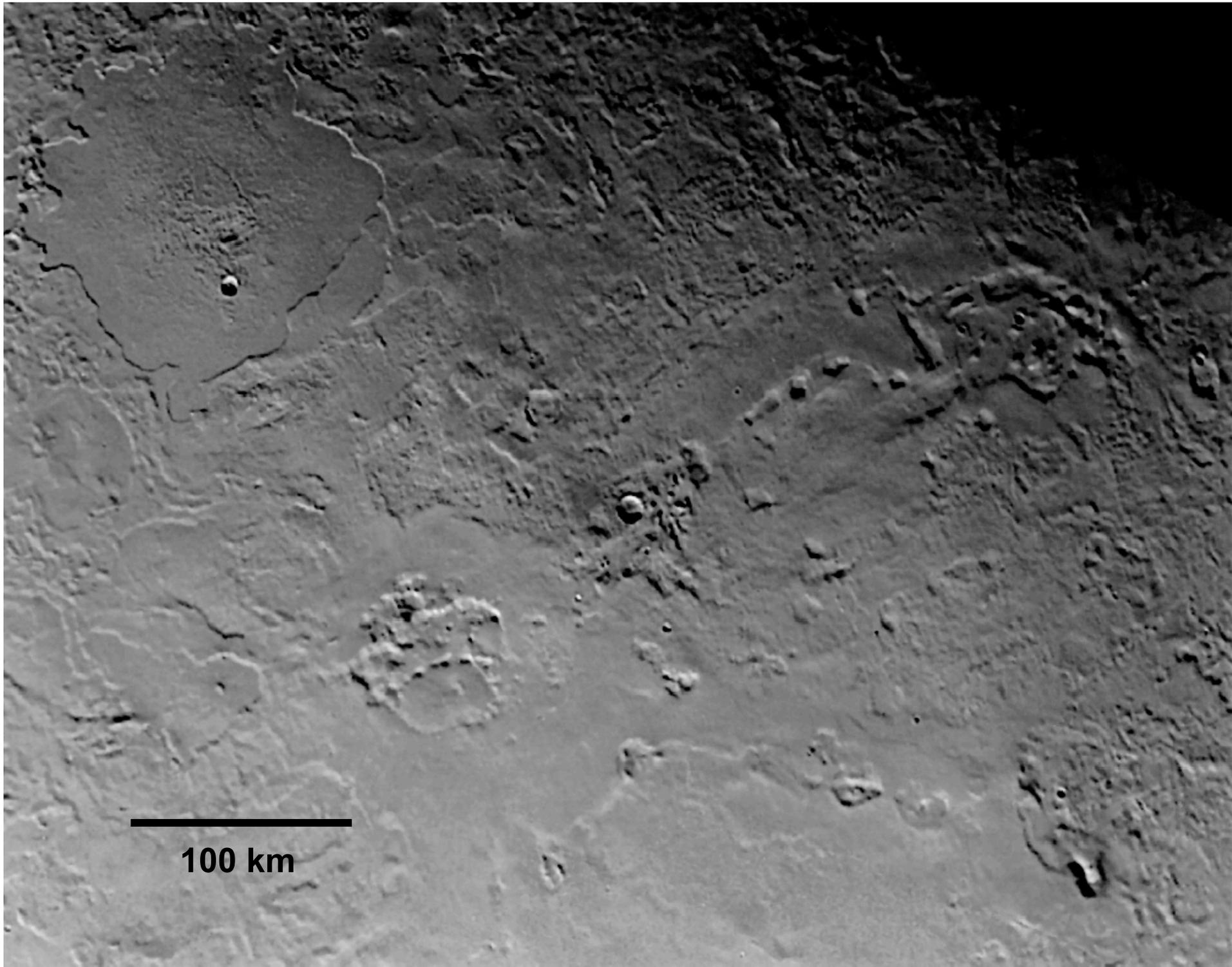
Cavi (dimples),
~25-35 km across

cross-cutting
ridges

& polar terrain
at top

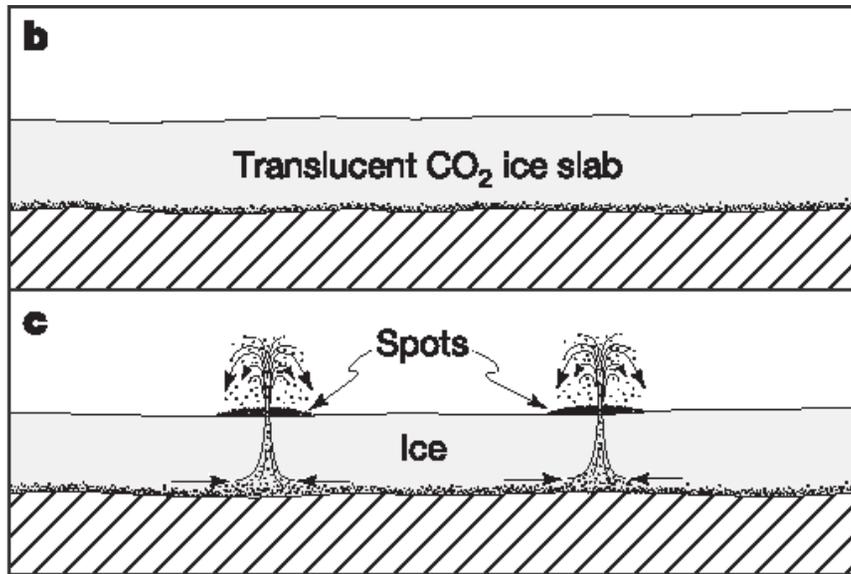
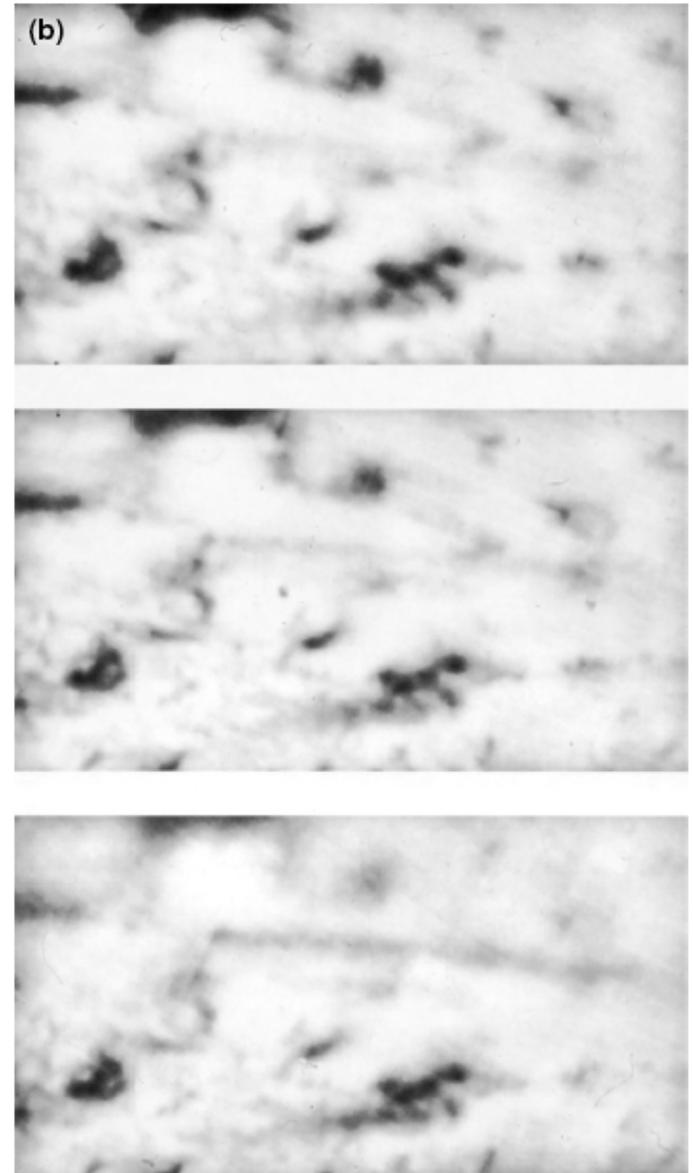
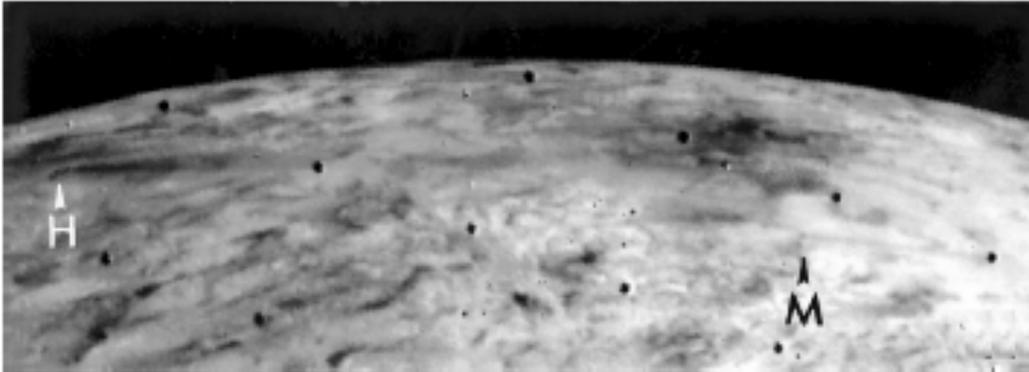
Oldest terrain?

Where are the
craters?



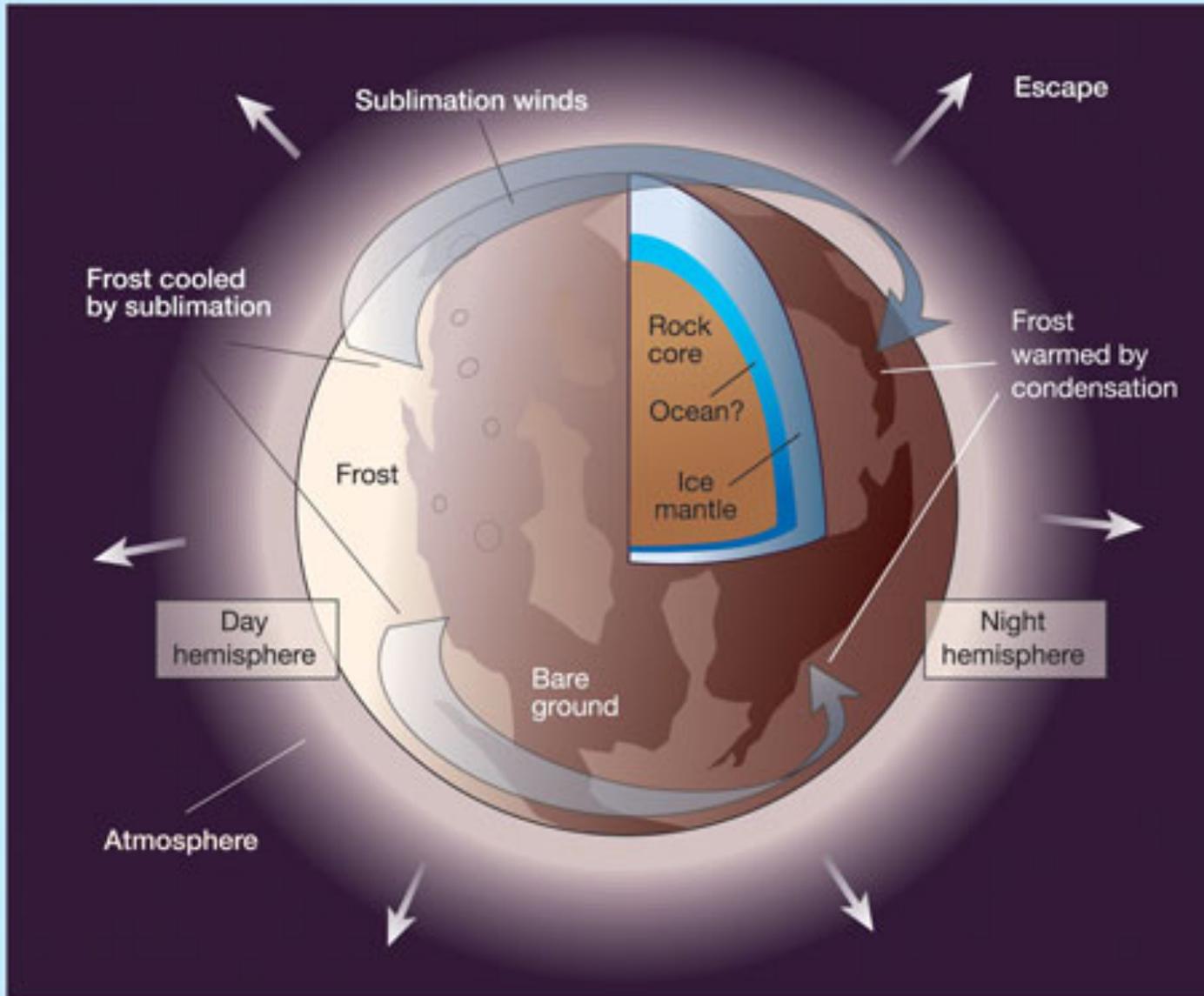
100 km

Nitrogen Plumes (or Geysers)

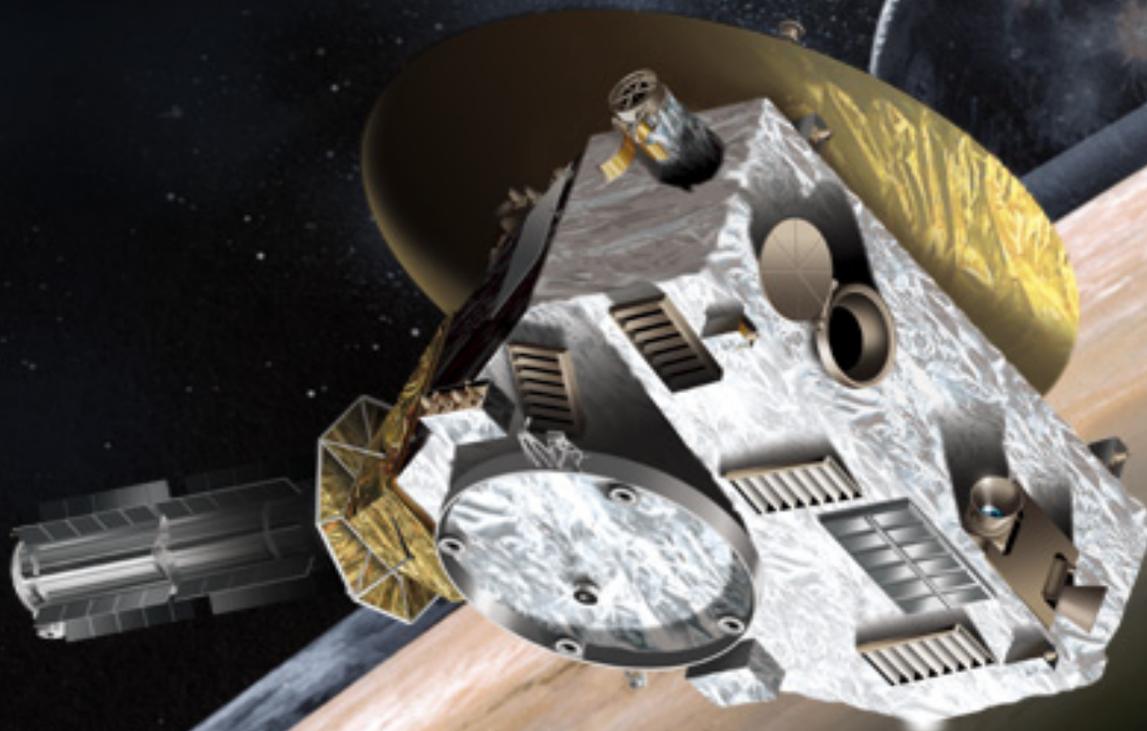


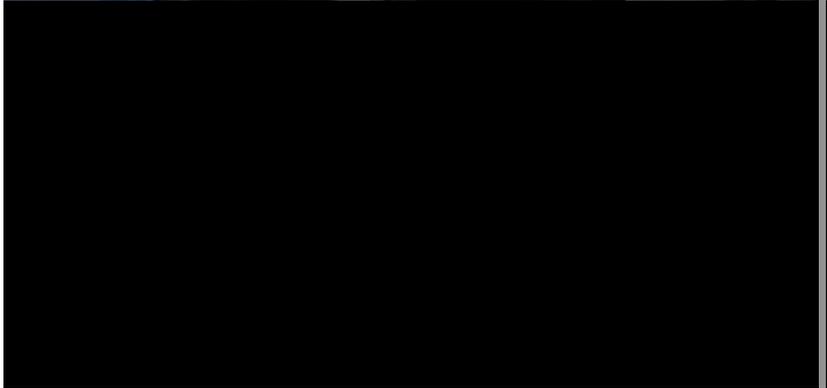
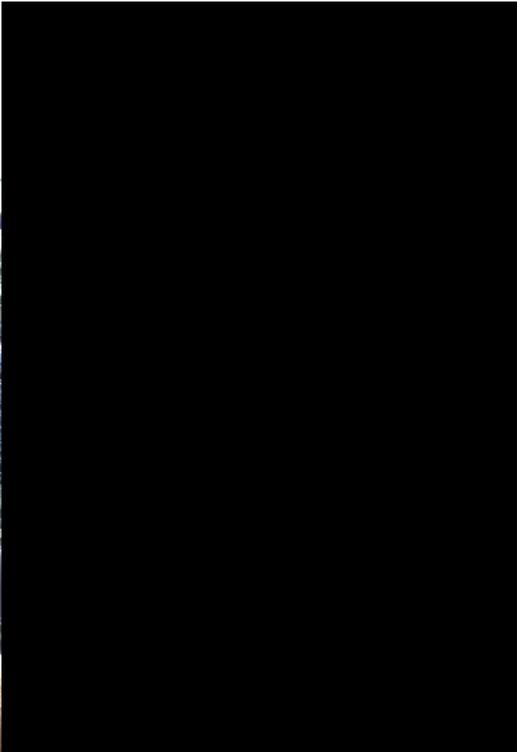
Kieffer et al. 06; Mars "spiders"

Pluto?



New Horizons
Mission To Pluto







Where is New Horizons NOW?!

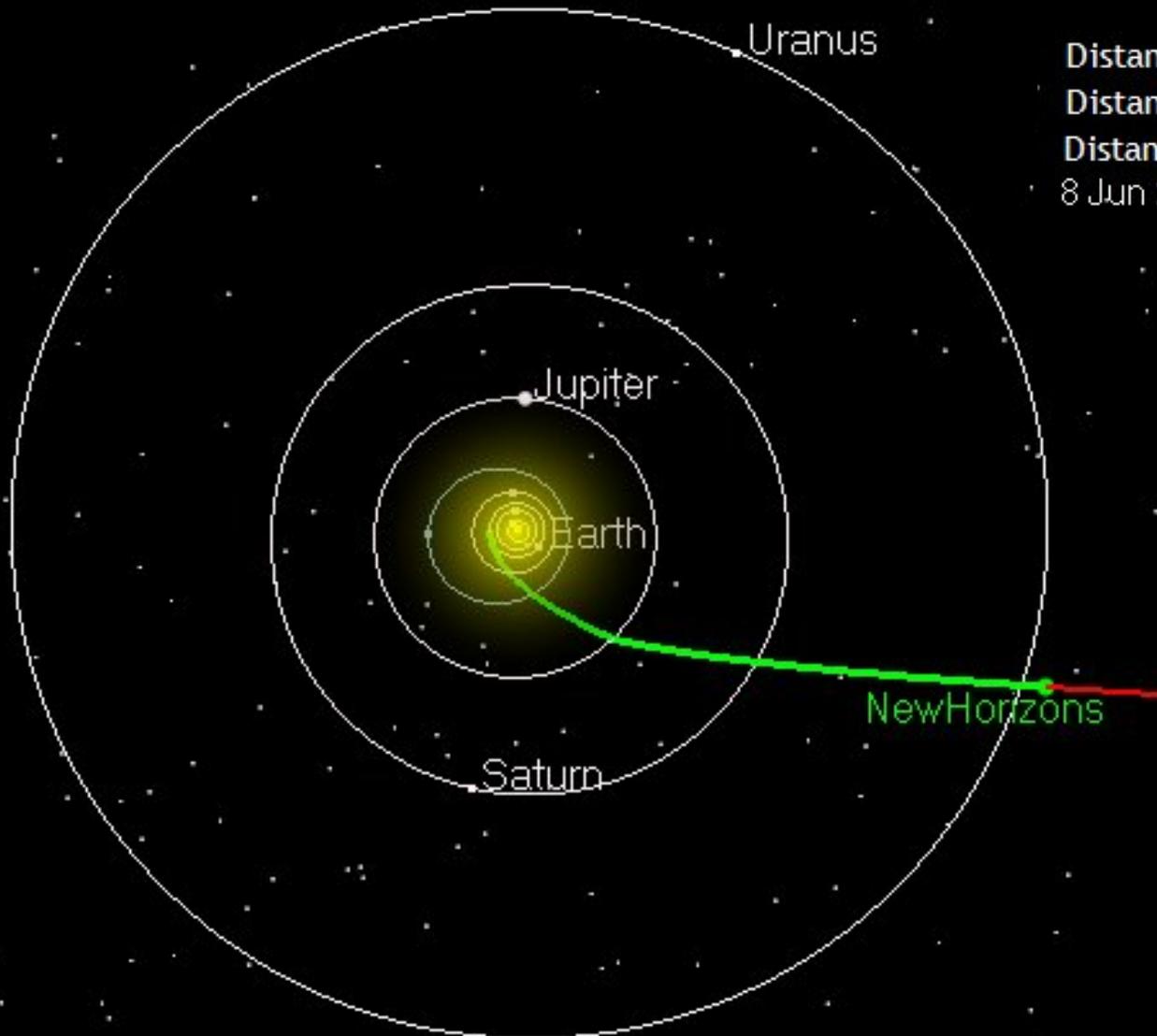
Distance from Sun (AU): 20.15 Heliocentric Velocity (km/s): 15.64

Distance from Earth (AU): 19.21

Distance from Jupiter (AU): 21.82

Distance from Pluto (AU): 12.02

8 Jun 2011 16:00:00 UTC

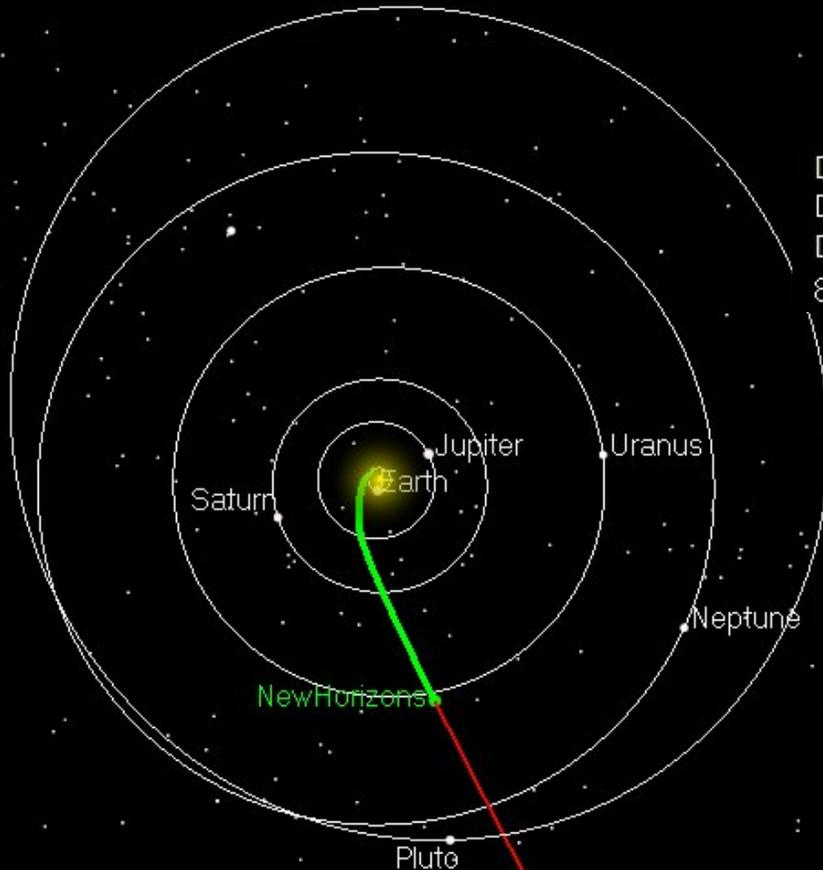


New Horizons Full Trajectory - Side View



Distance from Earth (AU): 19.21
Distance from Sun (AU): 20.15
Distance from Pluto (AU): 12.02
8 Jun 2011-16:00:00 UTC

New Horizons Full Trajectory - Overhead View



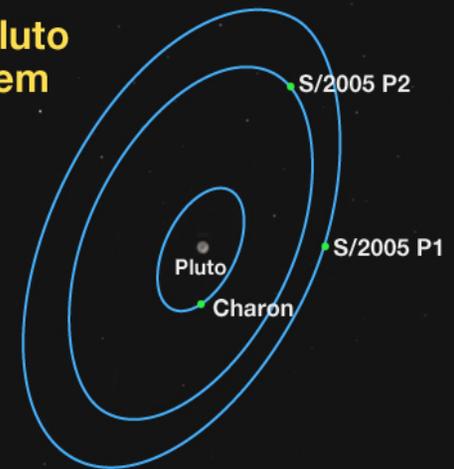
Distance from Earth (AU): 19.21
Distance from Sun (AU): 20.15
Distance from Pluto (AU): 12.02
8 Jun 2011 16:00:00 UTC

There's still
a long way
to go...

And then in 2015...

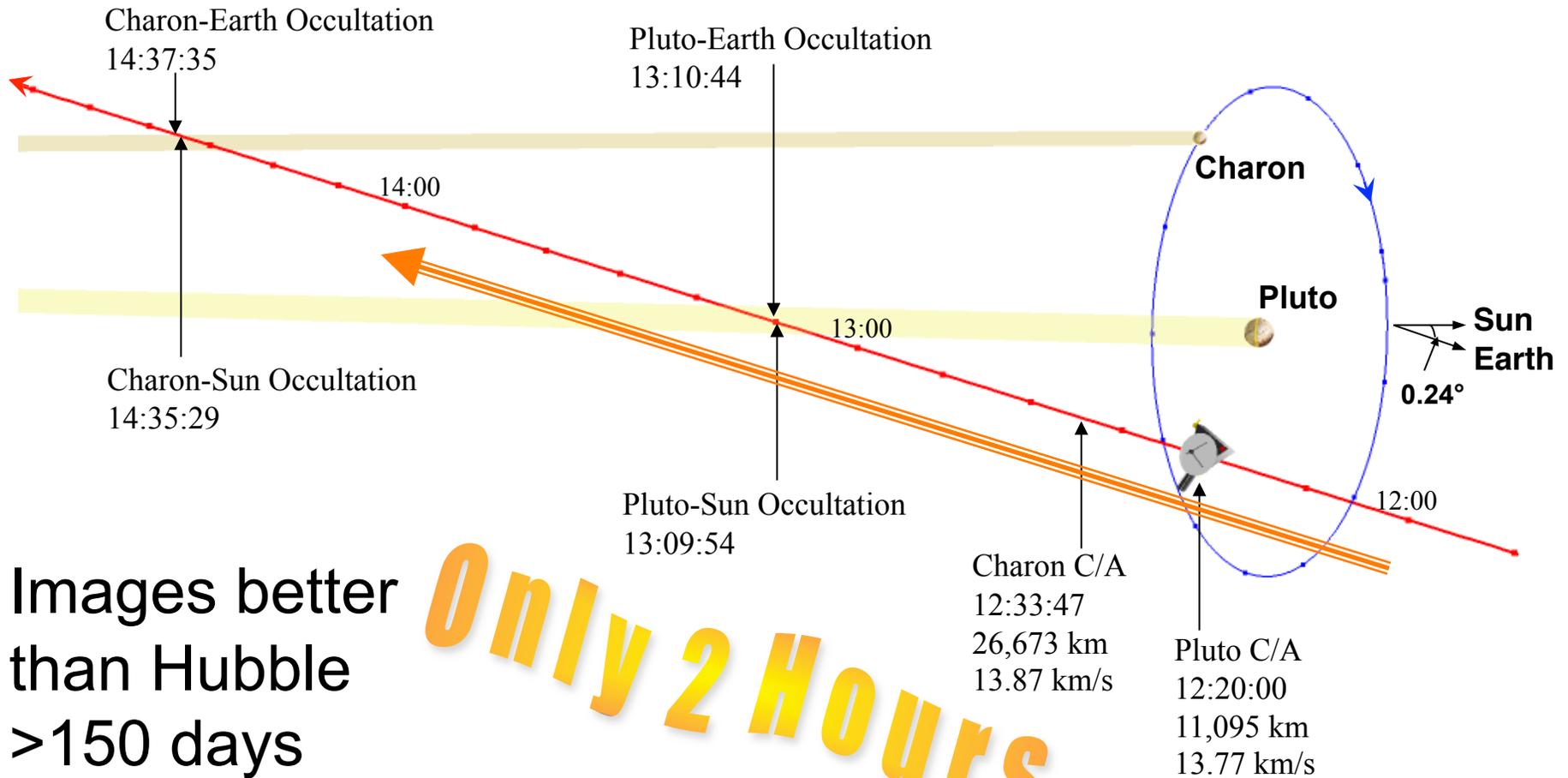
Pluto: The Main Attraction

The Pluto System



Pluto-Charon Encounter Geometry

Arrival July 14, 2015



Images better
than Hubble
>150 days
before arrival

Only 2 Hours

NEW HORIZONS

NASA's Pluto-Kuiper Belt Mission



Let's talk money.....

For example, how much does NASA's New Horizons mission to Pluto cost?

But first... What do you think is the fraction of the federal budget that goes to NASA?

- A. 0.05%
- B. 0.5%
- C. 1%
- D. 5%
- E. 10%

NEW HORIZONS

NASA's Pluto-Kuiper Belt Mission



How much does NASA's New Horizons mission to Pluto cost?

How many people in the US today? <http://www.census.gov/>

About 310 million $\sim 3 \times 10^8$

What is total US federal budget? <http://www.gpoaccess.gov/usbudget/fy11/index.html>

About \$3.8 trillion $\sim \$4 \times 10^{12}$

How much is 1 trillion bucks?

Let's start with a \$100 dollar bill. It's currently the largest US denomination in general circulation. They're guaranteed to make friends wherever they go.



A packet of one hundred \$100 bills is less than 1/2" thick and contains \$10,000. It fits in your pocket and is more than enough for week or two fun.



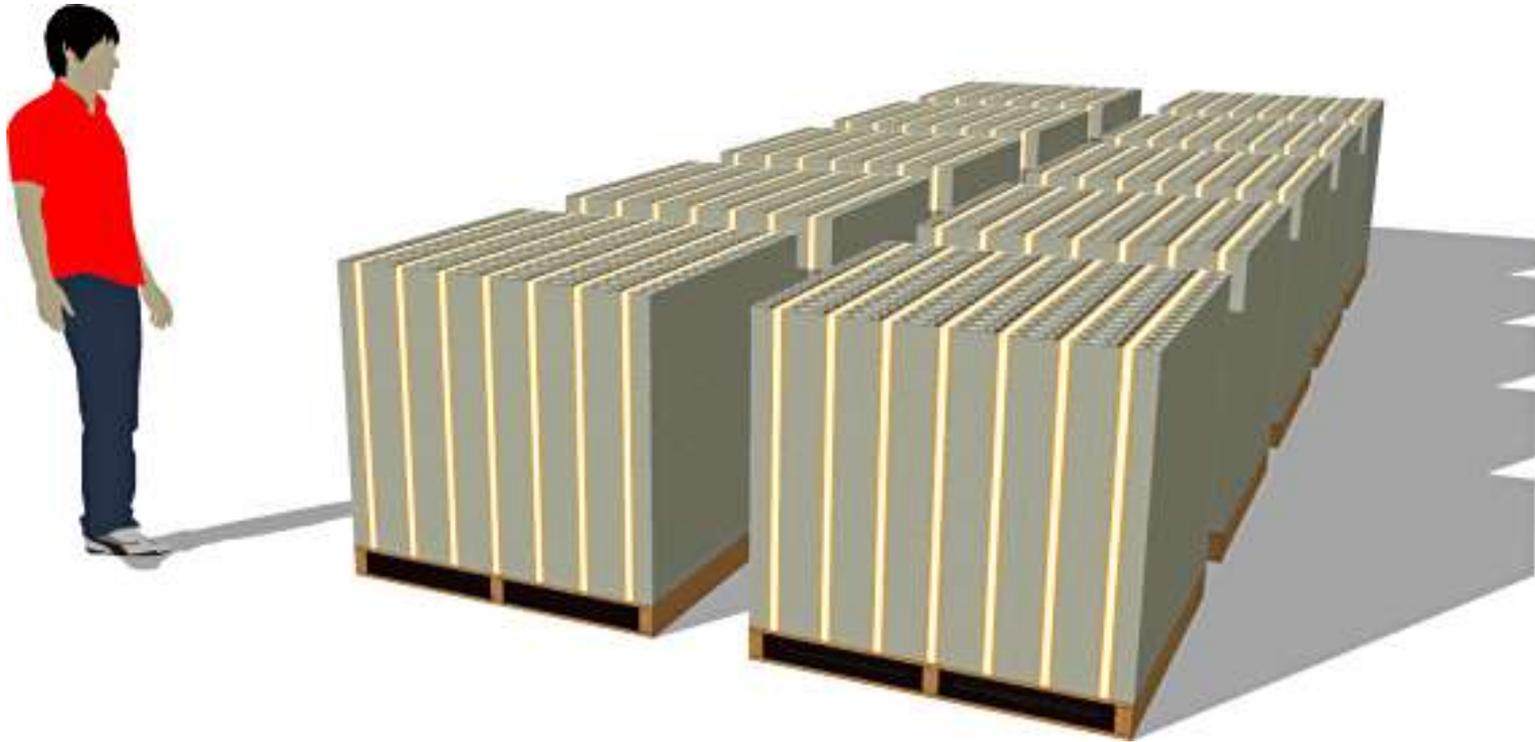


Believe it or not, this next little pile is \$1 million (100 packets of \$10,000). You could stuff that into a grocery bag and walk around with it.

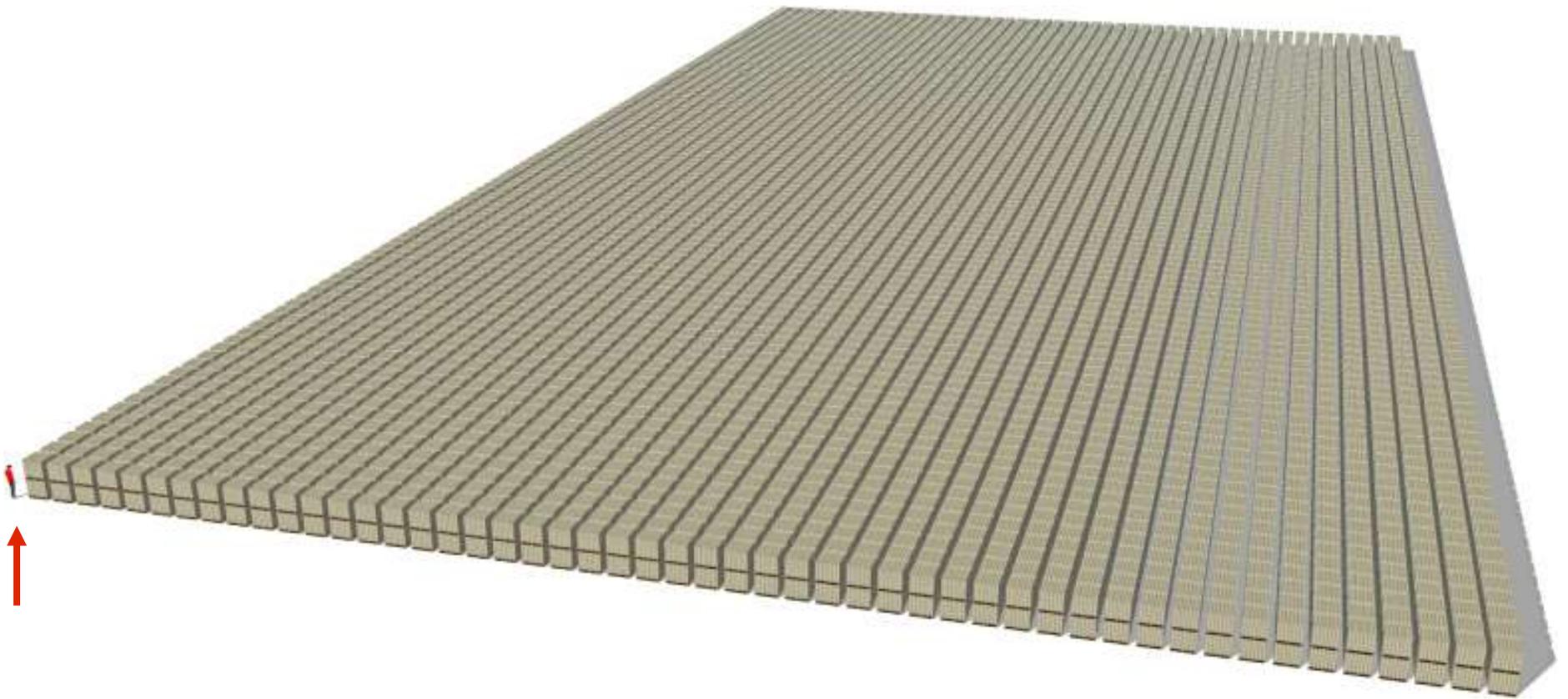
While a measly \$1 million looked a little unimpressive, \$100 million is more respectable. It fits neatly on a standard pallet.



And here's \$1 *billion* which is really impressive.



Okay, so let's look at *one trillion dollars*. It's the number we've been hearing so much about. What is a trillion dollars? Well, it's a million million. It's a thousand billion. It's a one followed by 12 zeros. Are you ready for this?



notice that those pallets are *double stacked*

US federal budget \$3.5 trillion

~ $\$4 \times 10^{12}$

How much is that per person?

$\$4 \times 10^{12}$ ~ $\$12,000$

3×10^8

How much is NASA's budget?

About \$19 billion ~ $\$2 \times 10^{10}$

What fraction of federal budget goes to NASA?

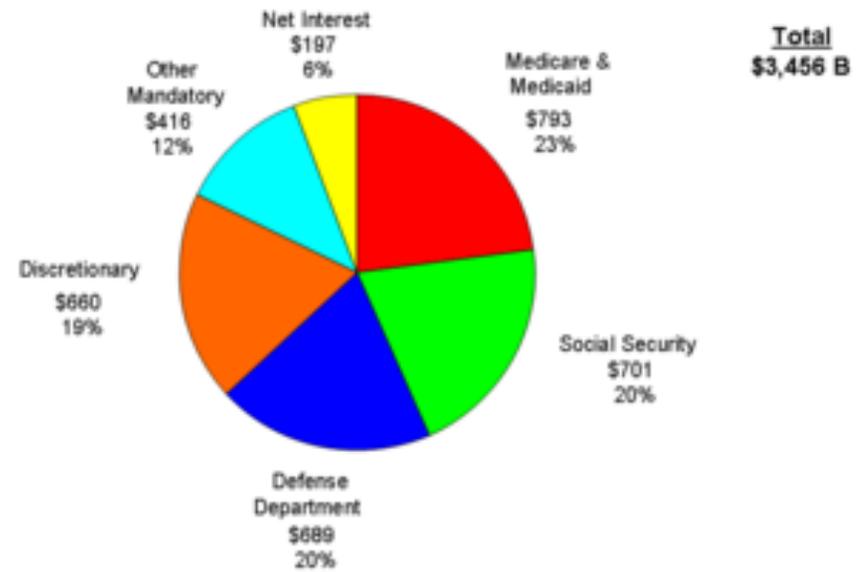
$\$2 \times 10^{10}$ ~ $1/200$ ~ 0.5%

$\$4 \times 10^{12}$

How much is this per person?

$\$2 \times 10^{10}$ ~ $\$61$

3×10^8



Source Data: CBO Historical Tables

<http://www.nasa.gov/news/budget/index.html>

NEW HORIZONS

NASA's Pluto-Kuiper Belt Mission



How much does NASA's New Horizons mission to Pluto cost?

About \$1 billion $\sim \$1 \times 10^9$

Over 10-year mission, how much is this per person?

About $\frac{\$1 \times 10^9}{3 \times 10^8} \sim \3

