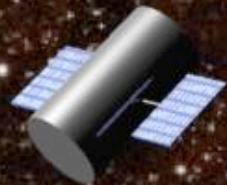


Direct Observation of Exoplanets

**Webster Cash
University of Colorado
January 27, 2007**

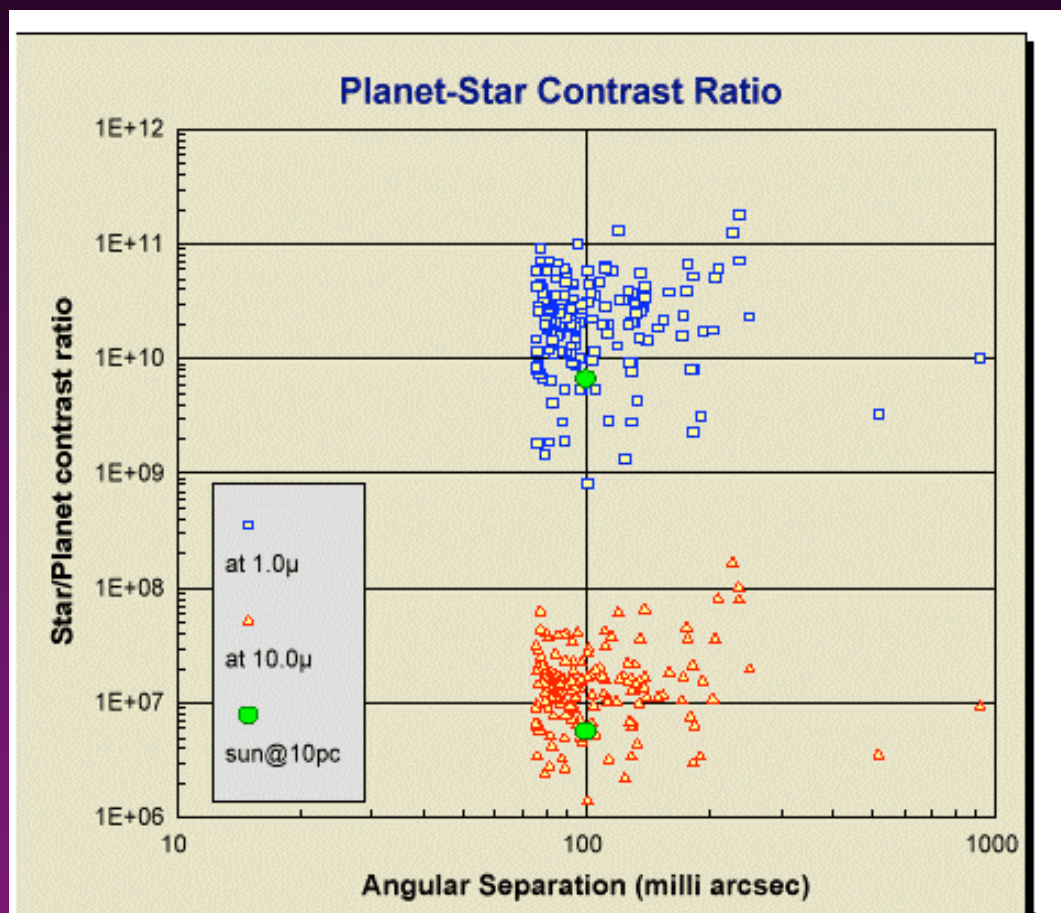


Planet Finding: Extinguish the Star

Contrast ratios better than 10 billion to one needed across a tenth of an arcsecond.

Wow. That's tough!

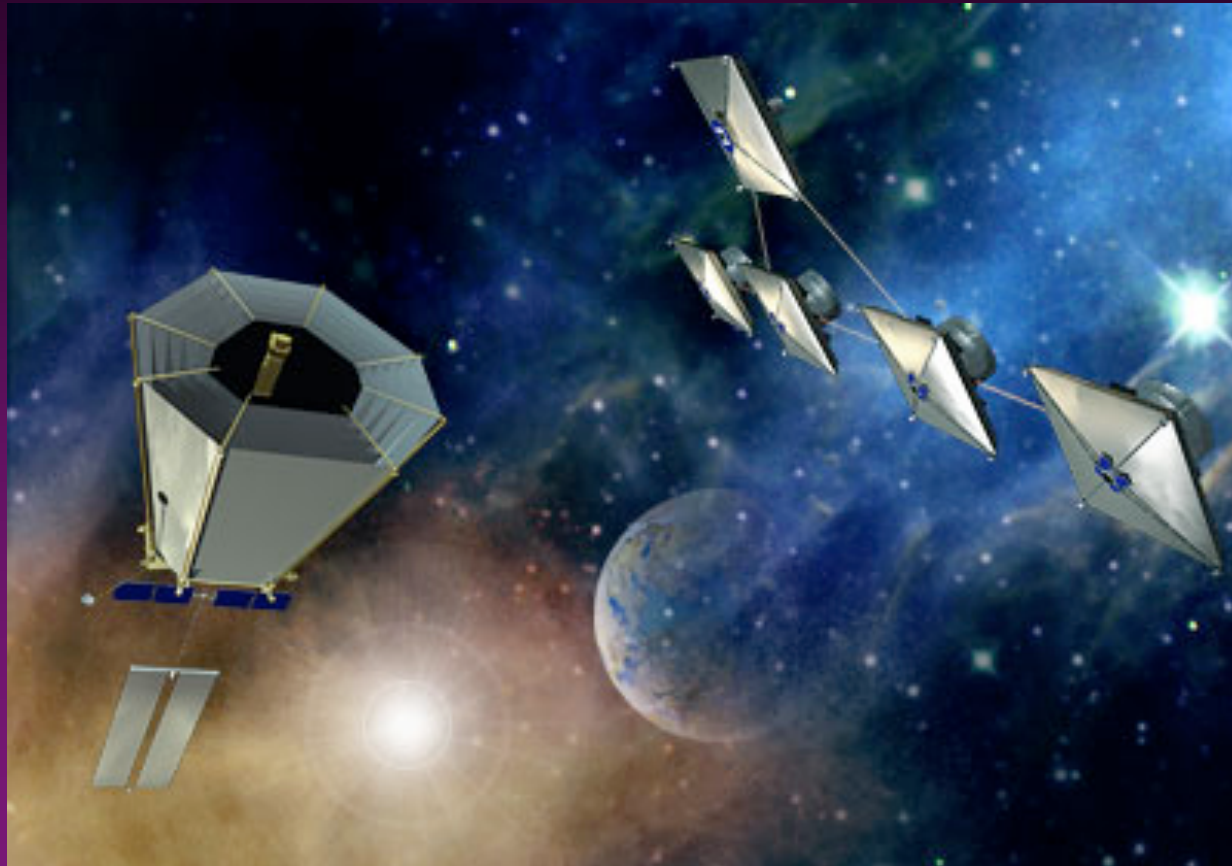
However will we get it done???



Courtesy of N-G

Terrestrial Planet Finder JPL's Answer

TPF-C

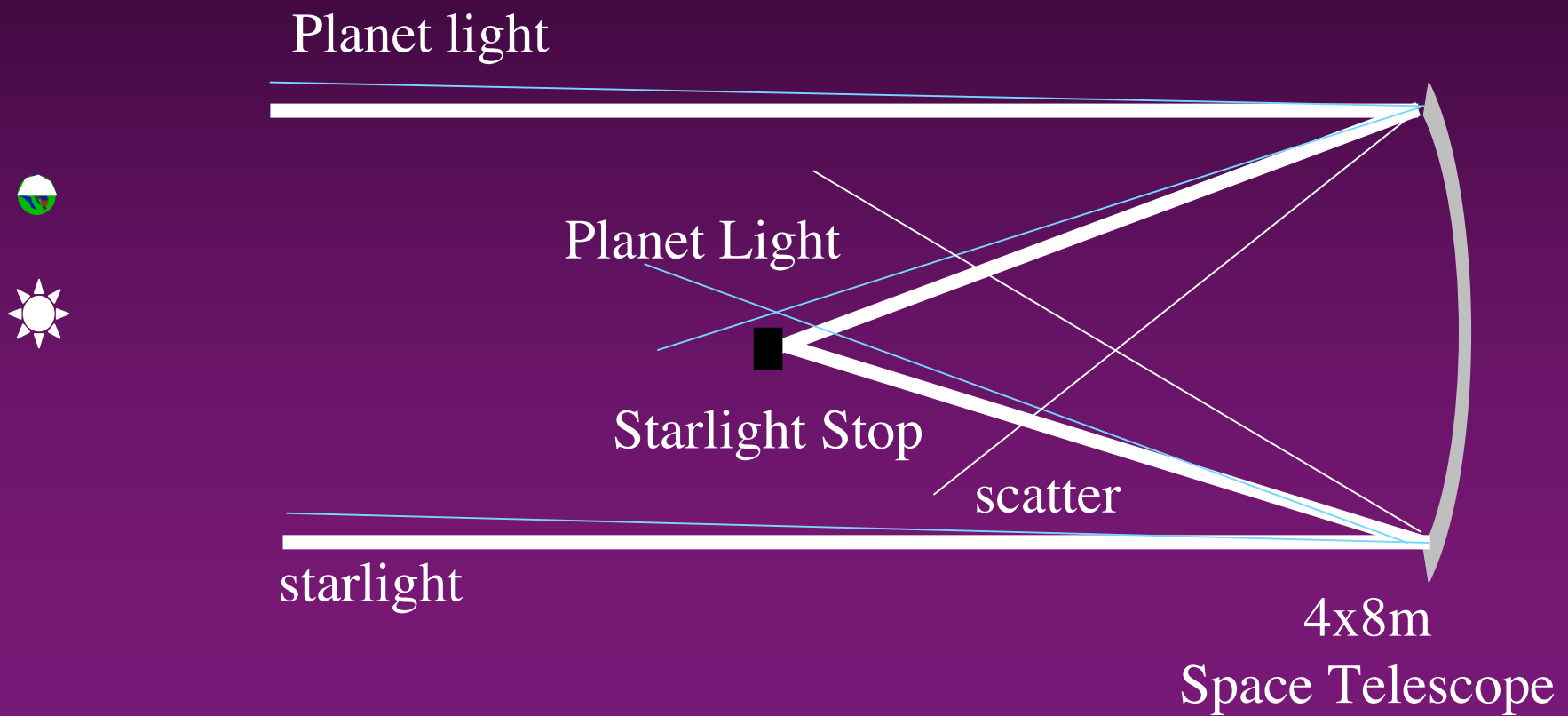


TPF-I

Must Be Done in Space Because of Atmospheric Twinkle

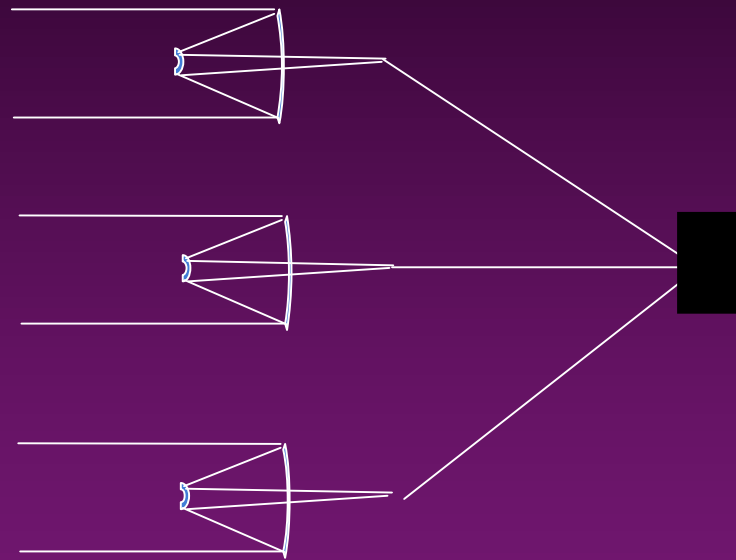
TPF-C

Visible Light Coronagraph



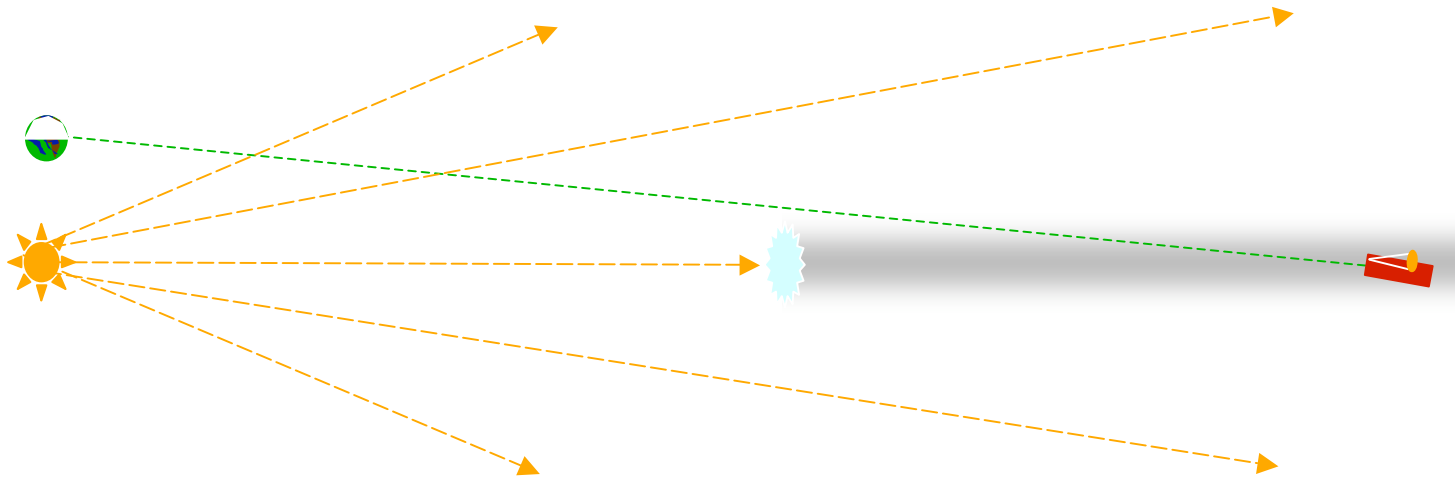
TPF-I & Darwin

Mid-Infrared Interferometer



Multiple Large Cooled-IR Telescopes Combine Beams
Null Out Star at One Angle
Constructive Interference at Nearby (Planet) Angle

Occulters



Telescope big enough to collect enough light from planet

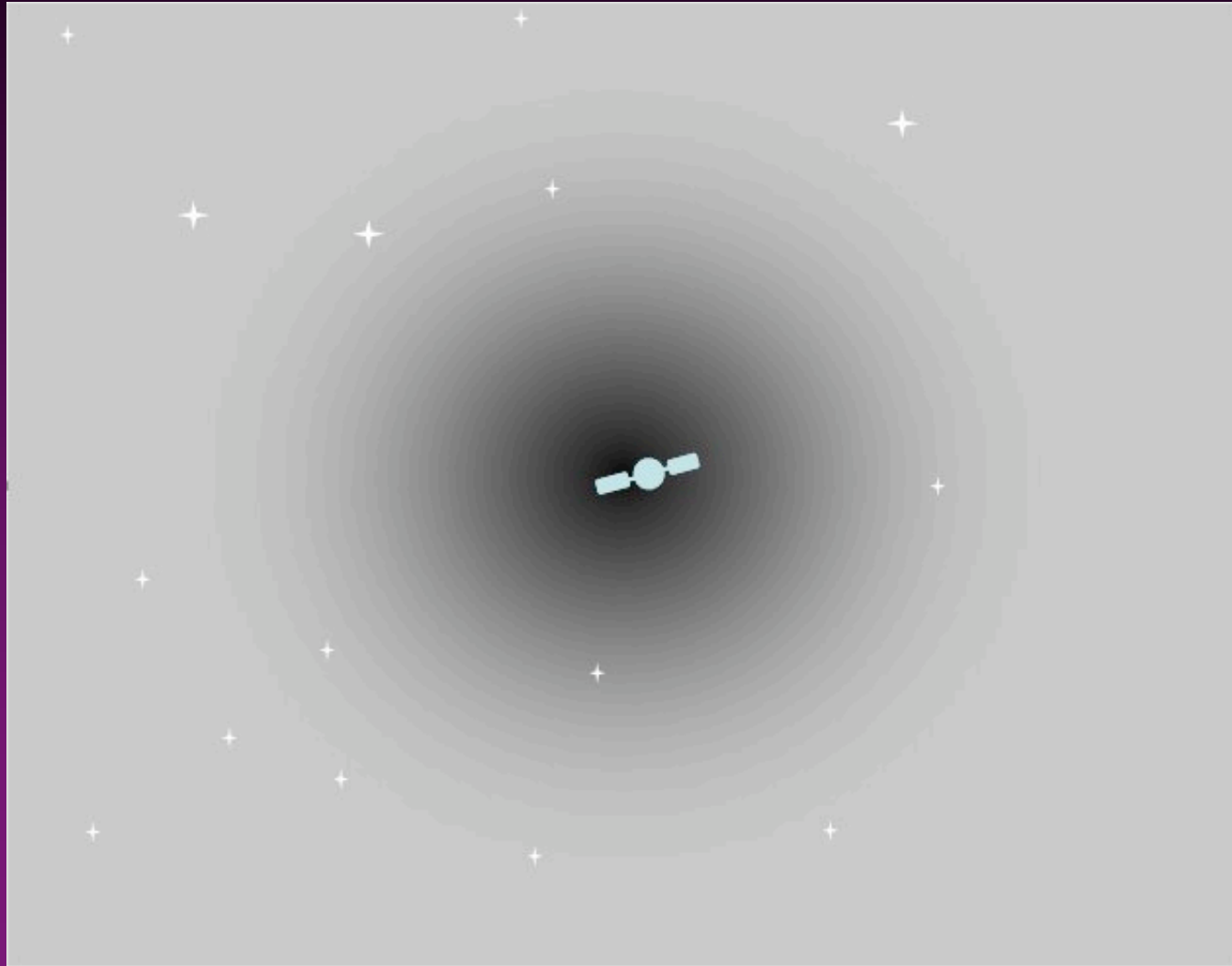
Occulter big enough to block star

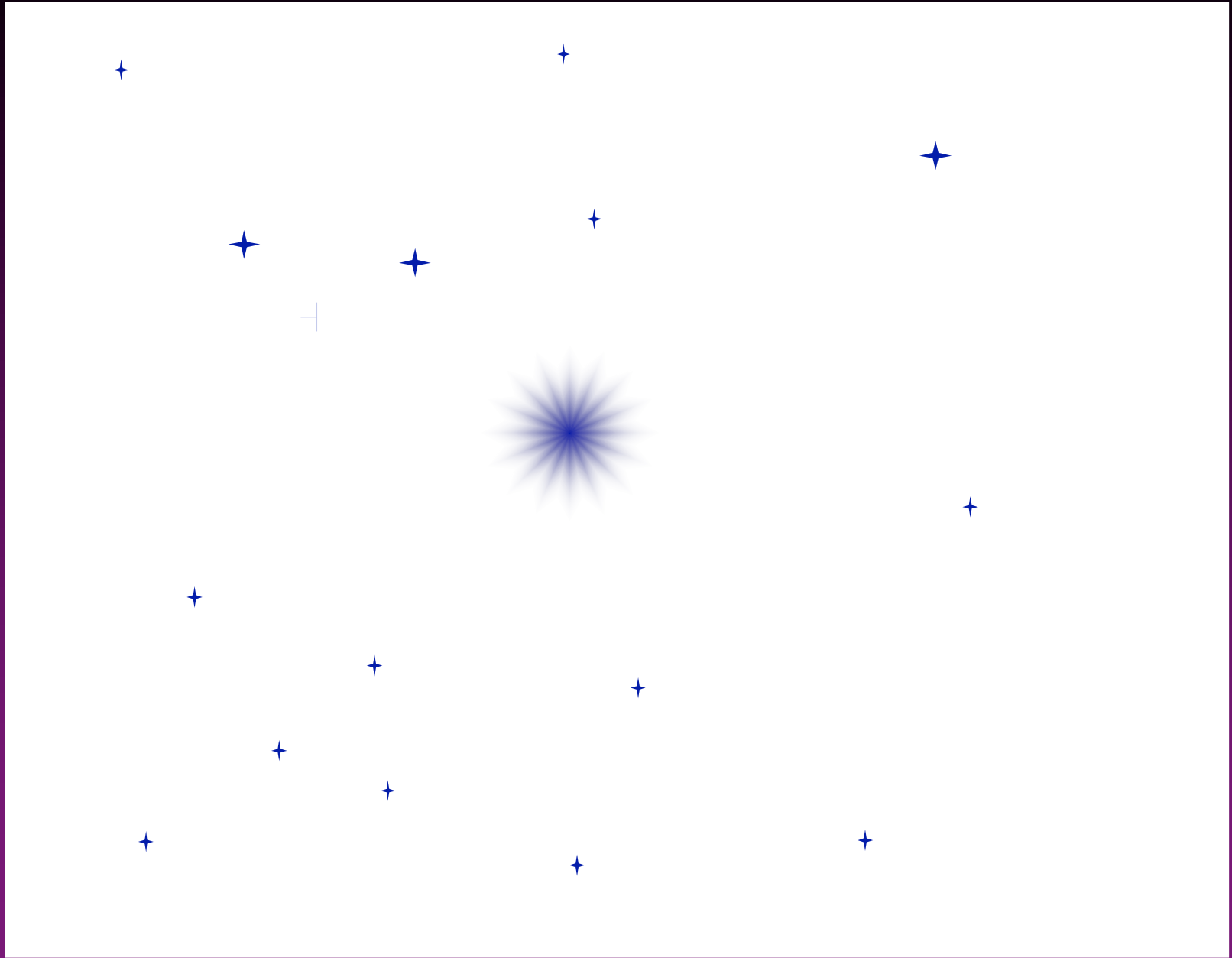
Want low transmission on axis and high transmission off axis

Telescope far enough back to have a properly small IWA

No outer working angle: View entire system at once

Fly the Telescope into the Shadow



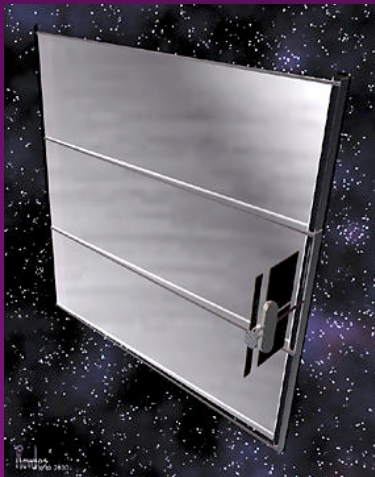


Why Haven't Occulters Been Baselined All Along?


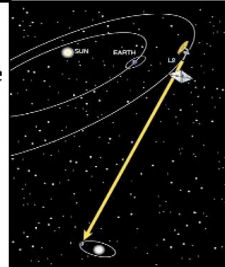
- ☞ Because →
- ☞ Everybody knows that diffraction around an occulter is too severe

Occulters

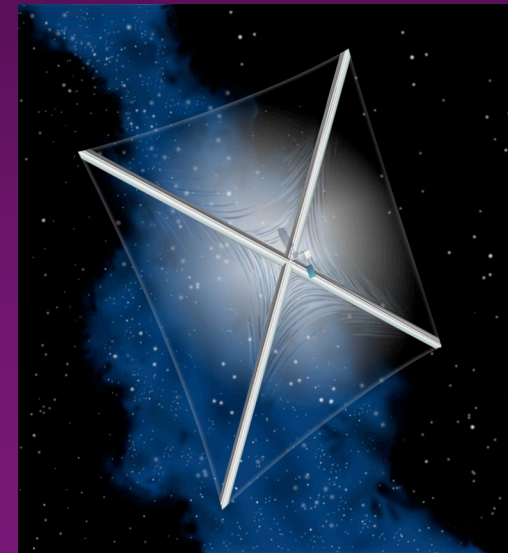
- ☞ **Several previous programs have looked at occulters**
 - First look by Spitzer (1962)
- ☞ **Used simple geometric shapes**
 - Achieved only 10^{-2} suppression across a broad spectral band
- ☞ **With transmissive shades**
 - Achieved only 10^{-4} suppression despite scatter problem



<http://umbras.org/>

	<p>Spokesperson: Glenn Starkman Organization: CWRU</p> <hr/> <p>Phone: (216)368-3660 Email: gds6@po.cwrn.edu URL: http://boss.phys.cwrn.edu Collaborators: Caltech, JPL, L'Garde, Lockheed-Martin Funding: JPL, IPAC, NSF</p>
<p>MISSION CONCEPT:</p> <p>Deploy a large occulting satellite with a space telescope at L2</p> <p>Occult nearby stars to discover and image planets</p> <p>Do ultra-high resolution imaging of target sources</p>	

BOSS



Starkman (TRW ca 2000)

Extinguishing Poisson's Spot



☞ Occulters Have Very Poor Diffraction Performance

- The 1818 Prediction of Fresnel led to the famous episode of:
- Poisson's Spot (variously Arago's Spot)
- Occulters Often Concentrate Light!

☞ Must satisfy Fresnel Equation, Not Just the Fraunhofer Equation

☞ Must Create a Zone That Is:

- Deep Below 10^{-10} diffraction
- Wide A couple meters minimum
- Broad Suppress across at least one octave of spectrum

☞ Must Be Practical

- Binary Non-transmitting to avoid scatter
- Size Below 150m Diameter
- Tolerance Insensitive to microscopic errors

The Apodization Function

Found this in April. Extended in June.

This Function Extinguishes Poisson's Spot to High Precision

$$A(\rho) = 0$$

for

$$\rho < a$$

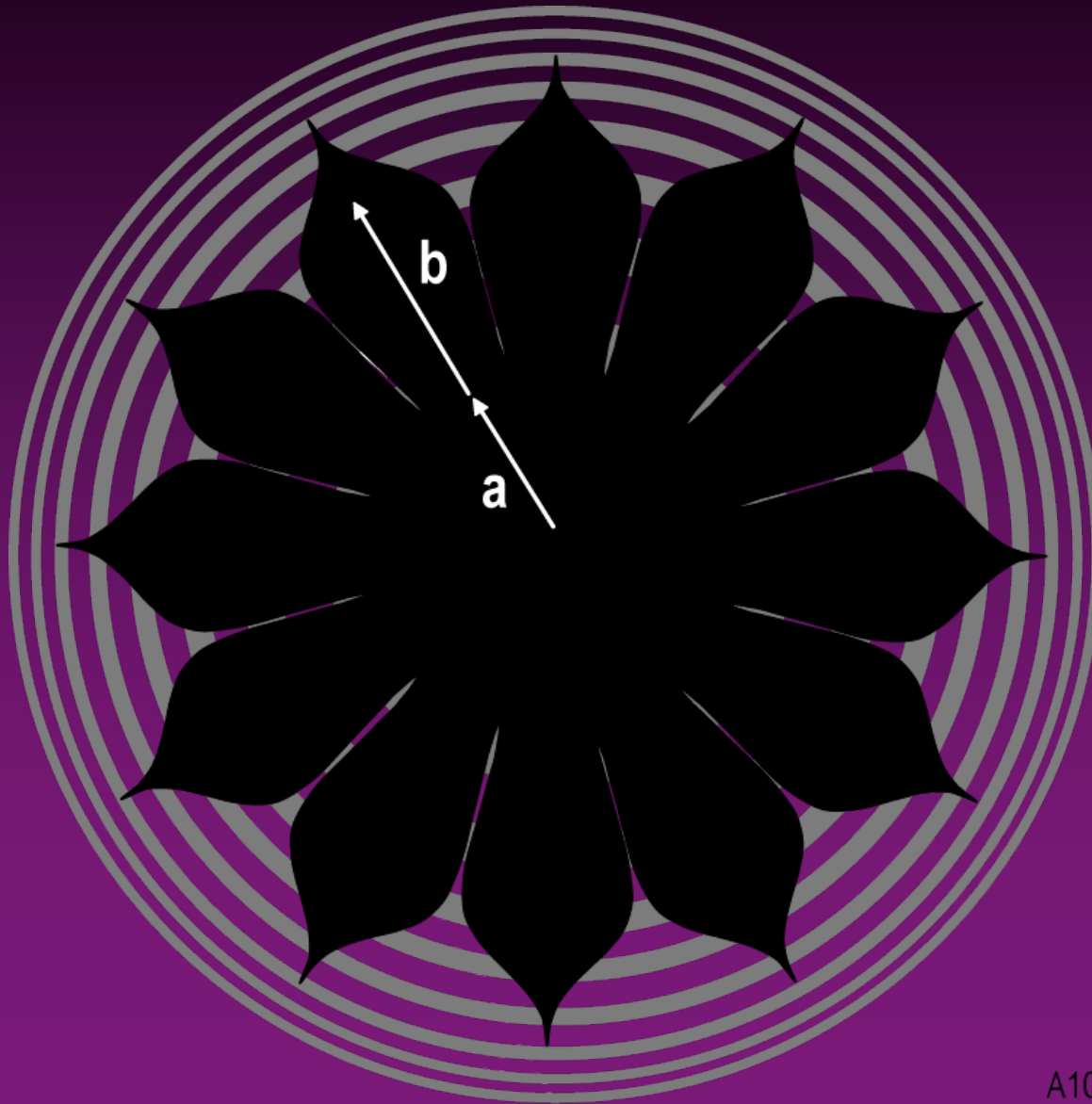
and

$$A(\rho) = 1 - e^{-\left(\frac{\rho-a}{b}\right)^n}$$

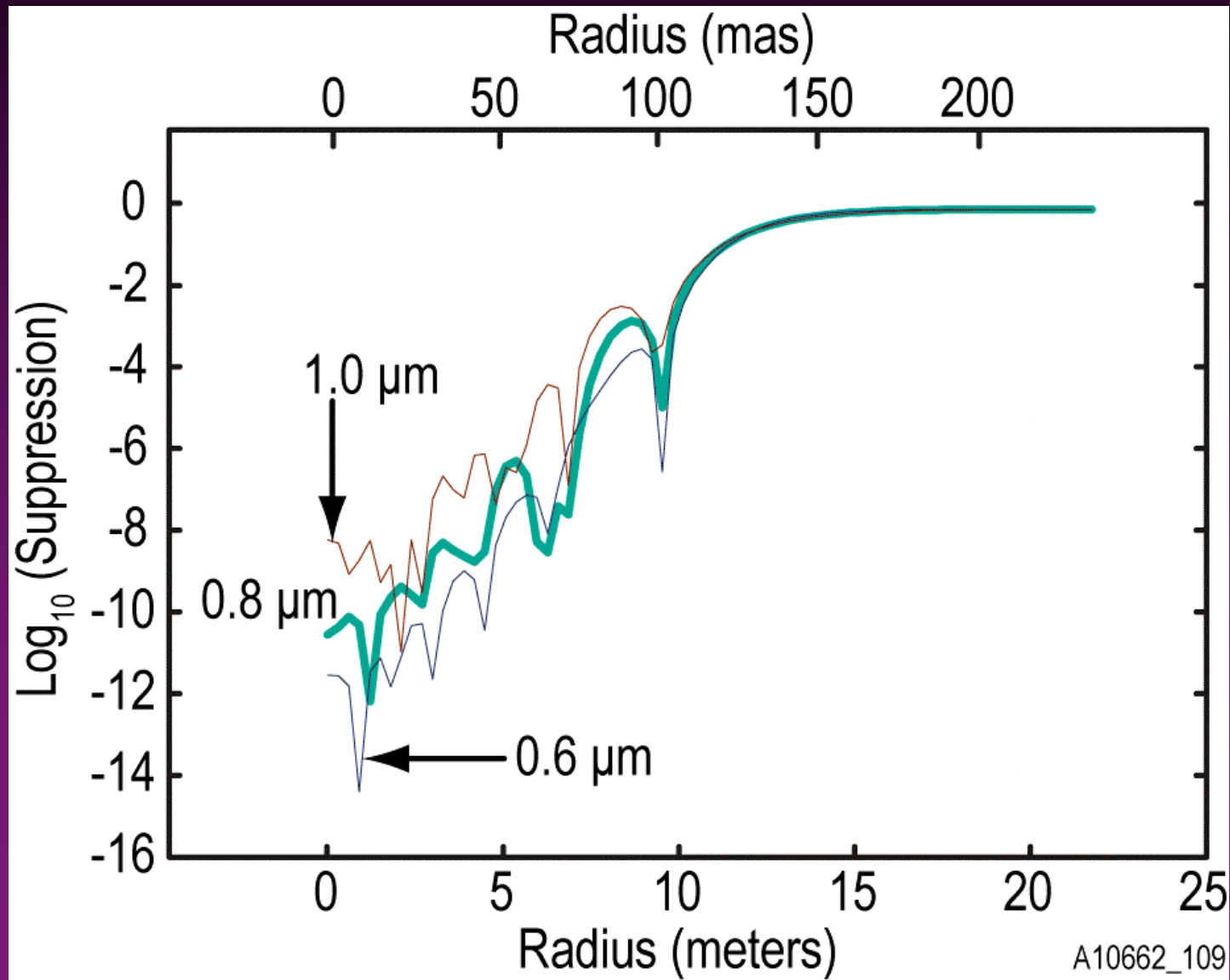
for

$$\rho > a$$

Binary Shape



Shadow Depth



A10662_109

Doing the Math (Cash, Nature 2006)

☞ The Residual Intensity in the Shadow is

$$I_s = E_s^2$$

☞ By Babinet's Principle

$$E_s = 1 - E_A$$

where E_A is field over Aperture

☞ So We Must Show

$$\frac{k}{2\pi F} \left[\int_0^{2\pi} \int_0^a e^{\frac{ik\rho^2}{2F}} e^{-\frac{ik\rho s \cos\theta}{F}} \rho d\rho d\theta + \int_0^{2\pi} \int_a^\infty e^{\frac{ik\rho^2}{2F}} e^{-\frac{ik\rho s \cos\theta}{F}} e^{-\left(\frac{\rho-a}{b}\right)^n} \rho d\rho d\theta \right] = i$$

☞ F is distance to starshade, s is radius of hole, k is $2\pi/\lambda$

☞ To one part in

$$\sqrt{C} \approx 10^{-5}$$

Contrast Ratio

☞ Preceding integral shows the contrast ratio is

–

$$R = \left[\frac{n!}{a^n b^n} \left(\frac{F\lambda}{2\pi} \right)^n \right]^2$$

– n is an integer parameter, typically n=6

☞ **To keep R small a~b**

– this is the reason the occulter has that symmetric look

Starshade Tolerances

☞ Position

- Lateral Several Meters
- Distance Many Kilometers

☞ Angle

- Rotational None
- Pitch/Yaw Many Degrees

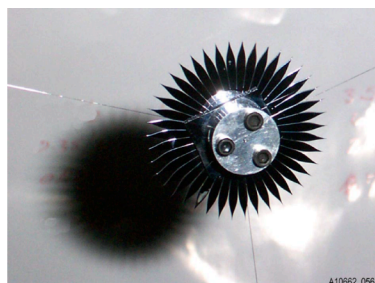
☞ Shape

- Truncation 1mm
- Scale 10%
- Blob 3cm^2 or greater

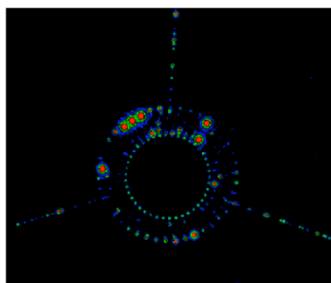
☞ Holes

- Single Hole 3cm^2
- Pinholes 3cm^2 total

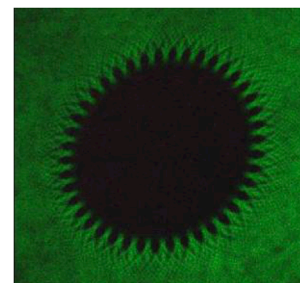
Lab Studies



a)

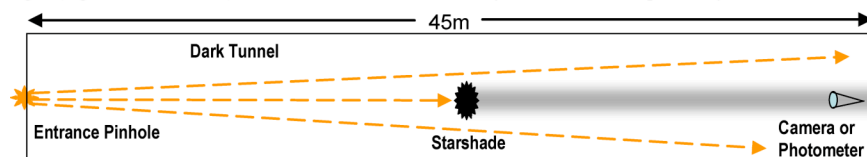


b)



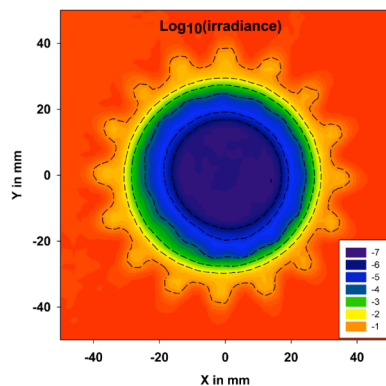
c)

Figure D-12 a): A test starshade, 35mm tip to tip made from silicon by lithographic techniques and suspended by three thin wires with $a = b = 8\text{mm}$, $n = 6$, and 42 petals; b) image of starshade back-illuminated by diverging beam of sunlight with approximately the same Fresnel number as expected in flight; clearly visible are scattering off support wires, two rings of bright points that correlate with tips of petals and gaps at bases of petals; c) when illuminated with coherent light (a green HeNe laser), one can see that diffraction by the starshade is primarily in the azimuthal direction.

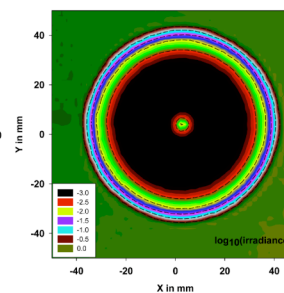
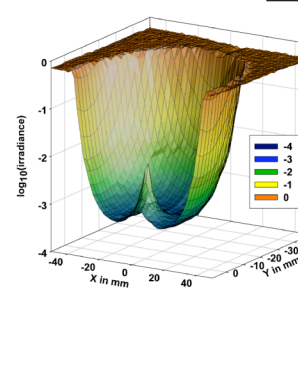
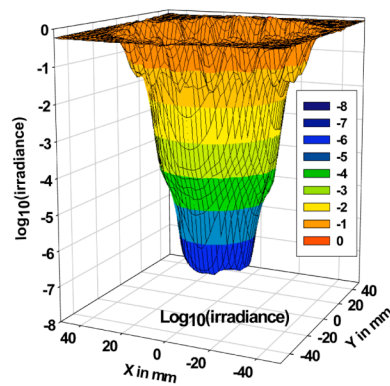


d) Schematic of test beamline; a heliostat feeds sunlight into the dark tunnel; a photon counting photometer scans the shadow to measure its depth; a camera can image the diffraction and scatter by defects around the starshade.

e) Students near one end of dark tunnel at University of Colorado



f) Photometer mapping of shadow from 16 petal silicon starshade in white light shows suppression of several orders of magnitude of irradiance within a few millimeters, with an ultimate suppression level of $< 1 \times 10^{-7}$.



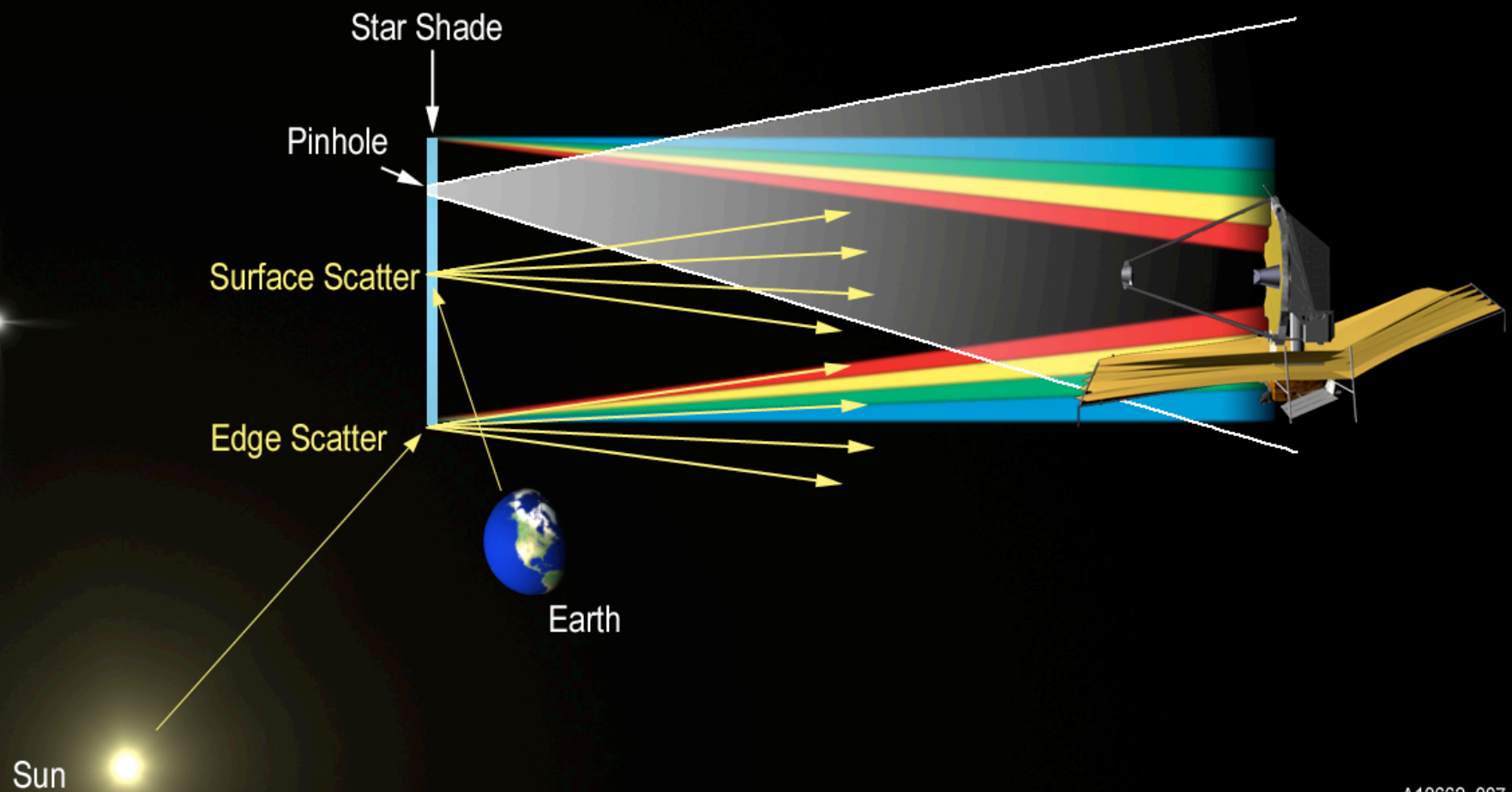
g) Map of shadow from a precision circular disc in white light showing amplitude of Poisson's spot at a predicted irradiance level 1% of incident level, verifying the photometric accuracy of measurement facility.

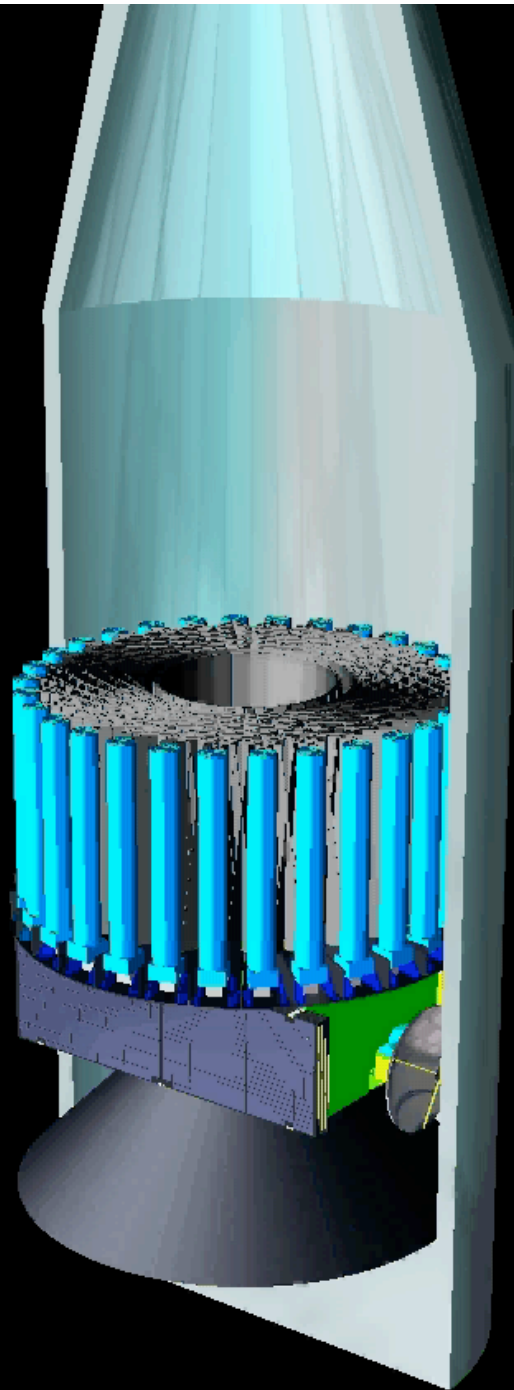
Tall Poles

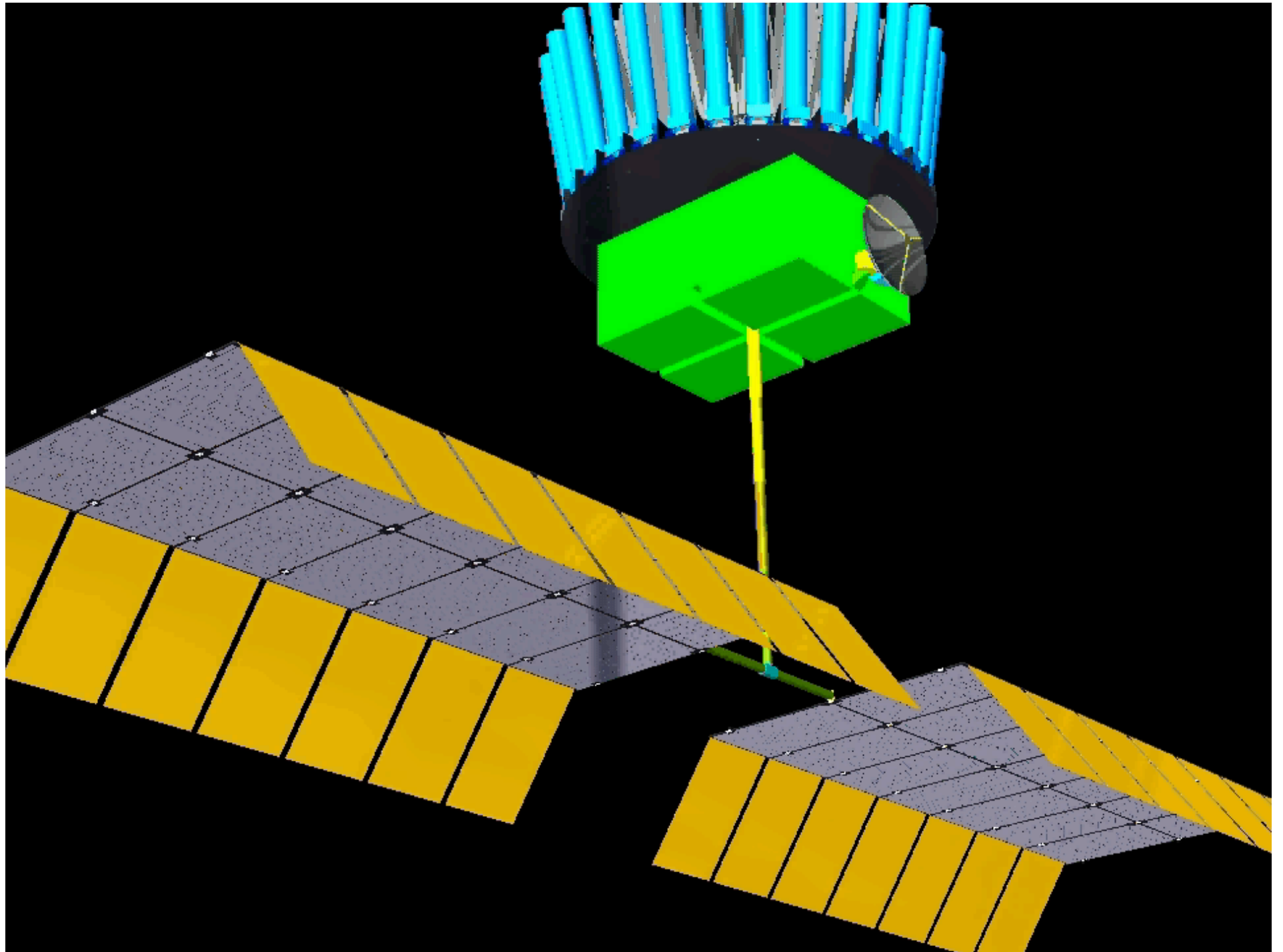
- ☞ **Deployment of 35m shade to mm class tolerance**
- ☞ **Acquiring and holding line of sight**
- ☞ **Fuel usage, orbits and number of targets**
- ☞ **Stray Light – particularly solar**

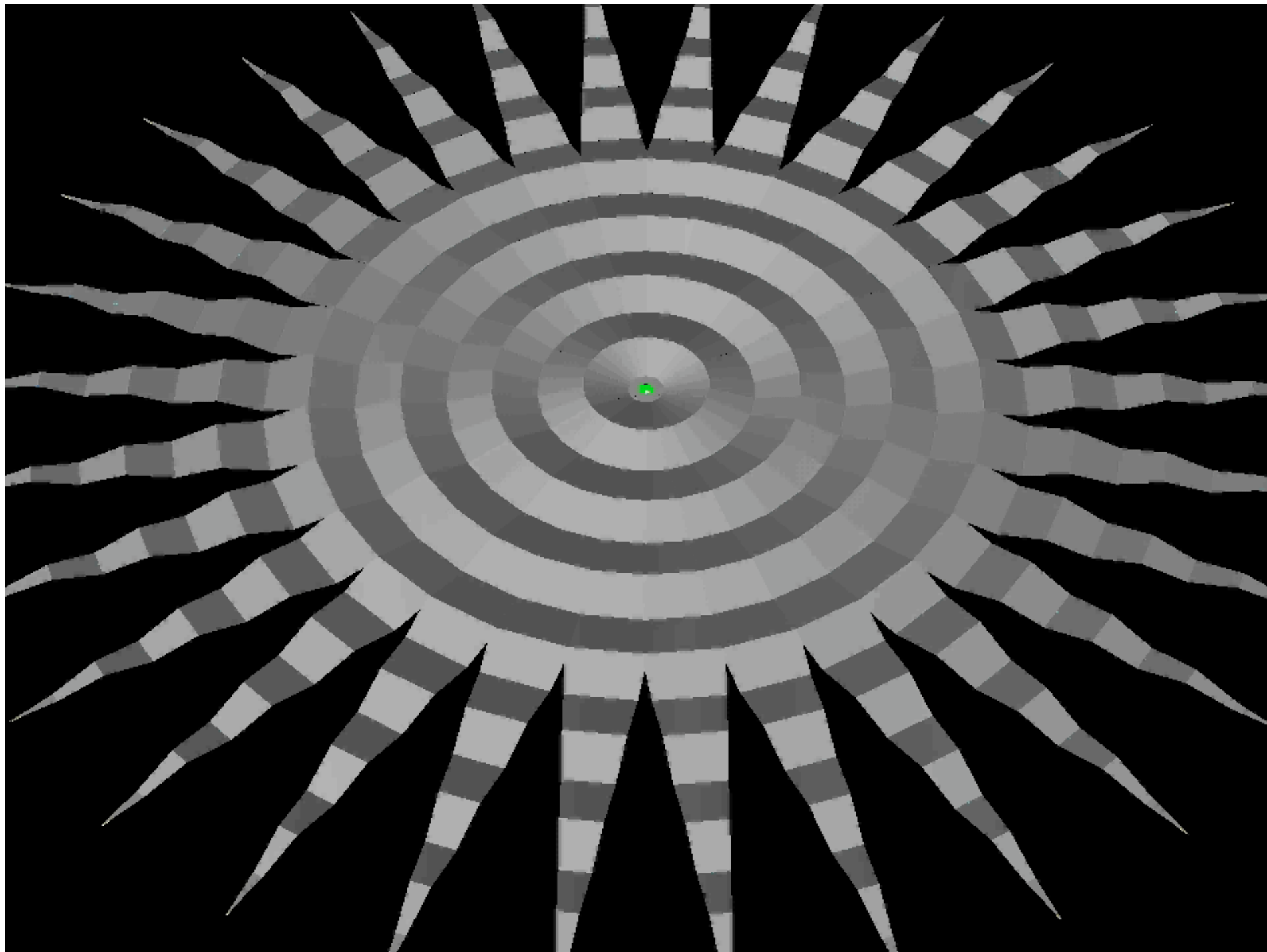
Scatter Control

View Nightside



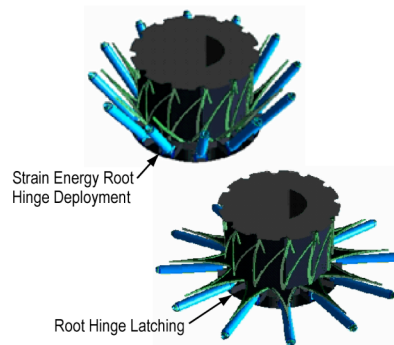
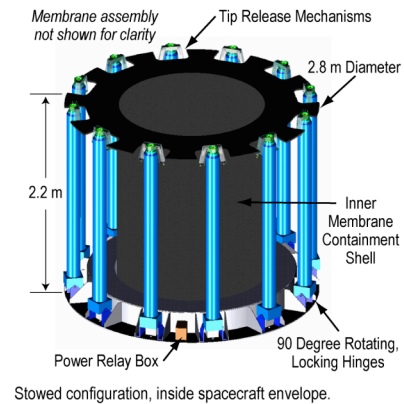






Deployment

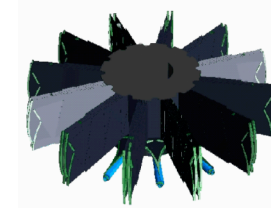
Starshade Deployment Sequence



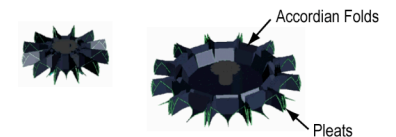
The booms fold down, each boom supports one petal, and there are twelve booms.



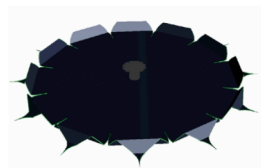
The folds of the occulter are swirled around a central holding cap, similar to the folding method of a collapsible umbrella.



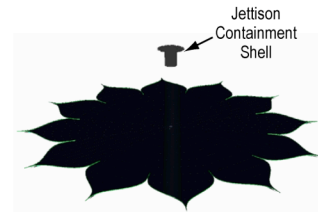
The swirled folds are straightened and ready for the booms to extend



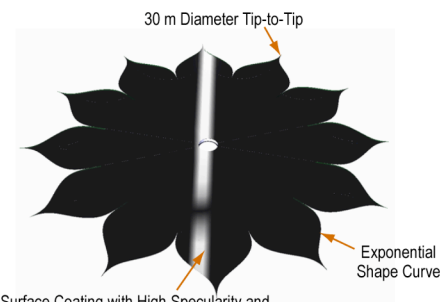
The stem drives extend the booms segment by segment. The occulter edge is fixed to the boom segment. Each segment locks into place when it becomes fully deployed.



There are seven segments to the boom for this 25 meter occulter. The central object is a holder for the occulter while it is inside the spacecraft.

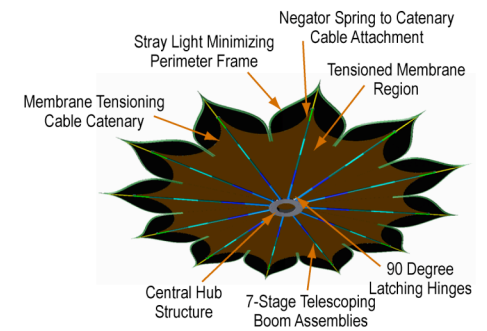


After full deployment, the holder is jettisoned.



Surface Coating with High Specularity and Low Solar Absorption/Emissivity Ratio

The tips of the petals are composed of a solid rod of the appropriate shape, which folds out and locks onto the last boom segment.



A view of the underside of the occulter, each segment of the boom is represented by a different color. Cables run along the perimeter to hold the petal edges in proper tension.

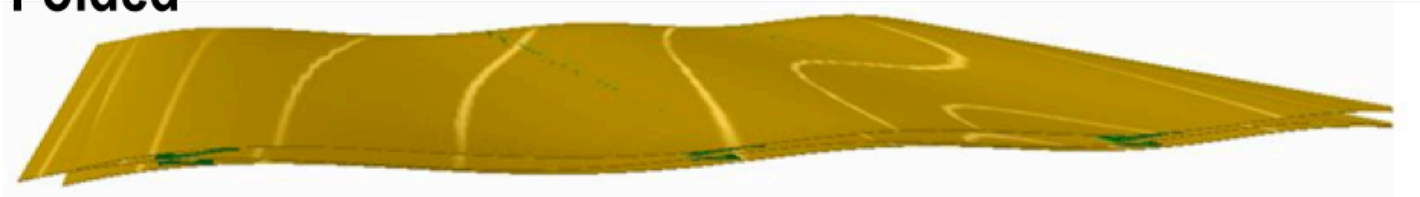
Table Top Demonstrator

This paper model shows that all the planned folds can be accomplished with no interference or binding. The paper model is a standard step in the development of a viable deployable. All of the salient folds are included in this subscale demonstrator.

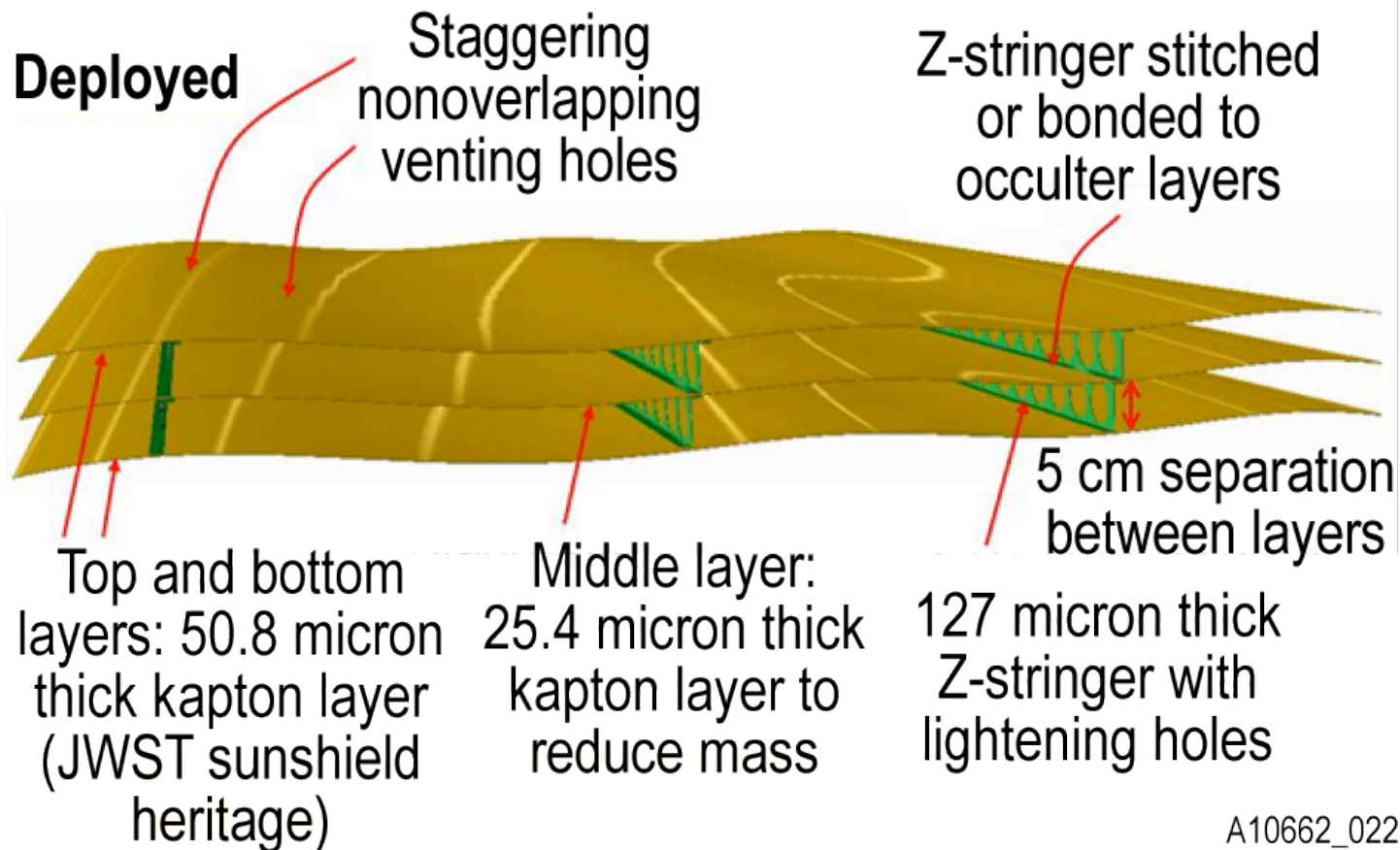


Triple Layered Micro-meteor Protection

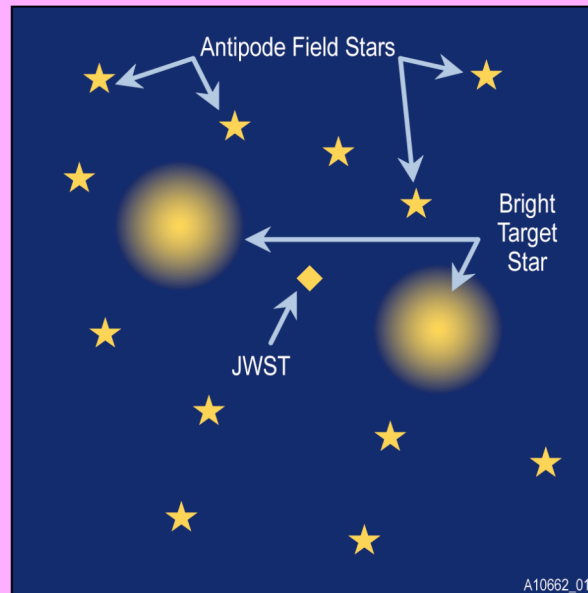
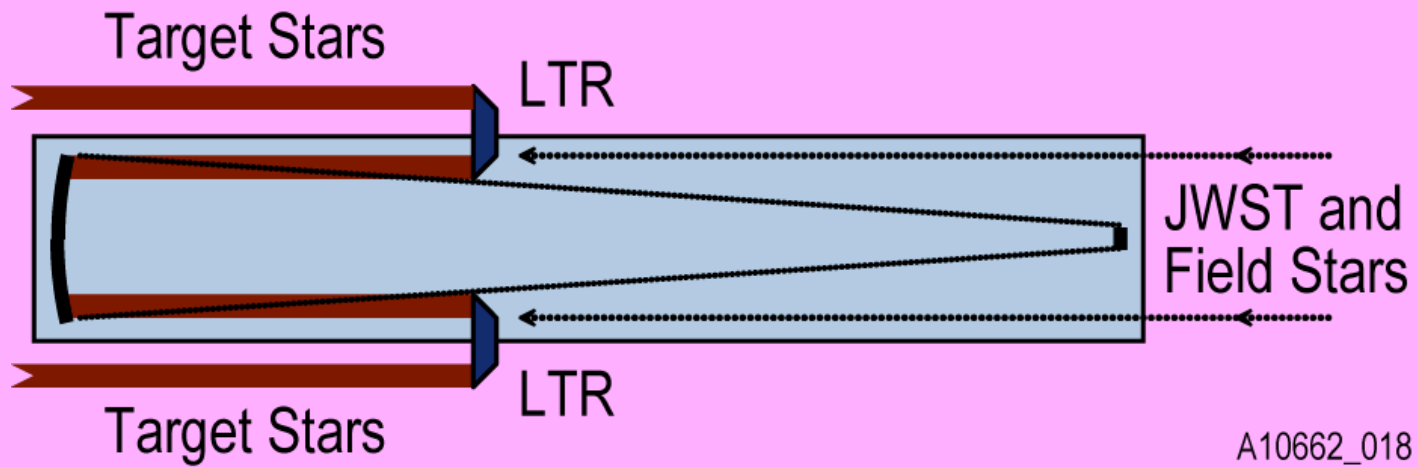
Folded



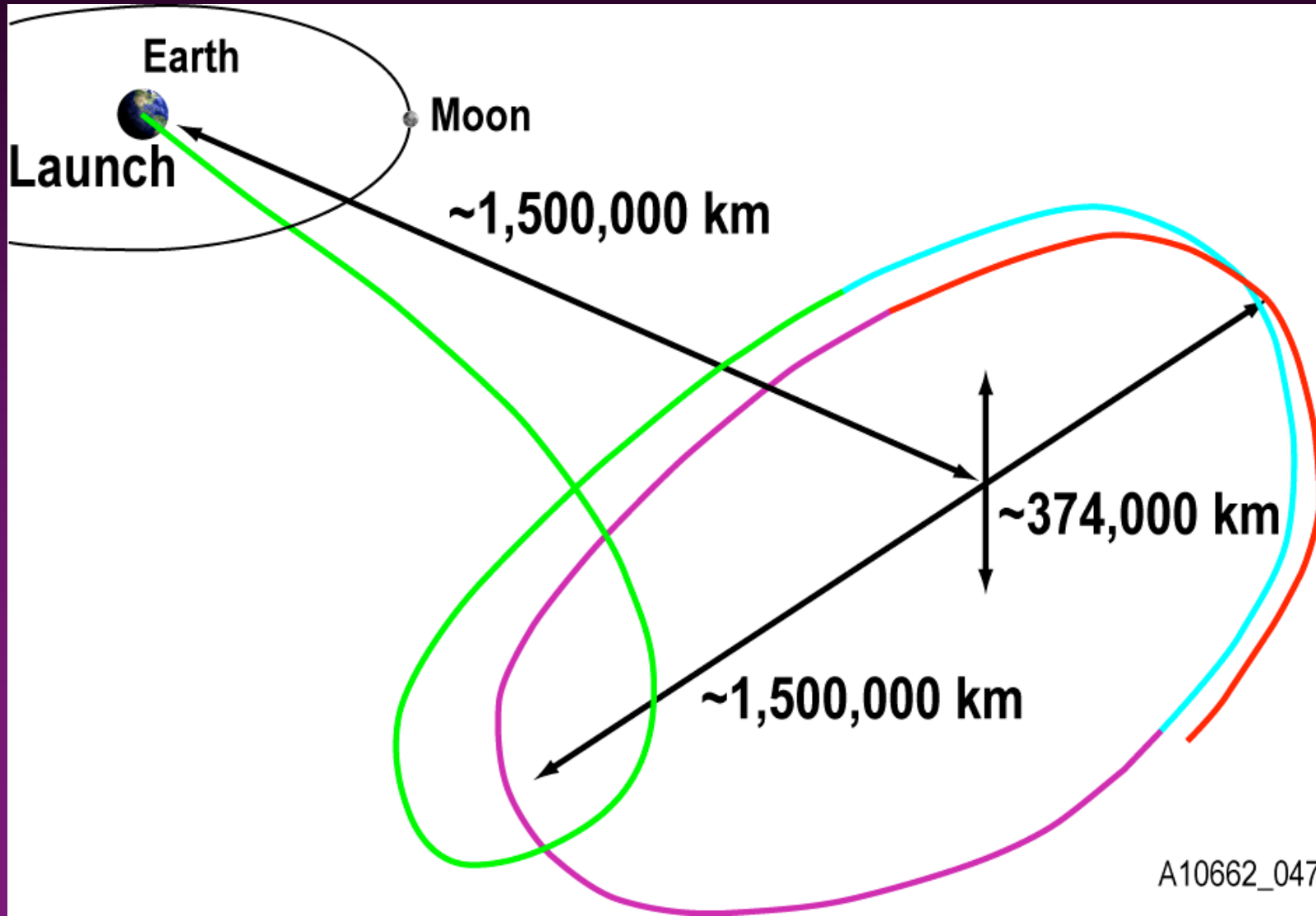
Deployed



Alignment acquisition & hold



L2 Orbit Favorable

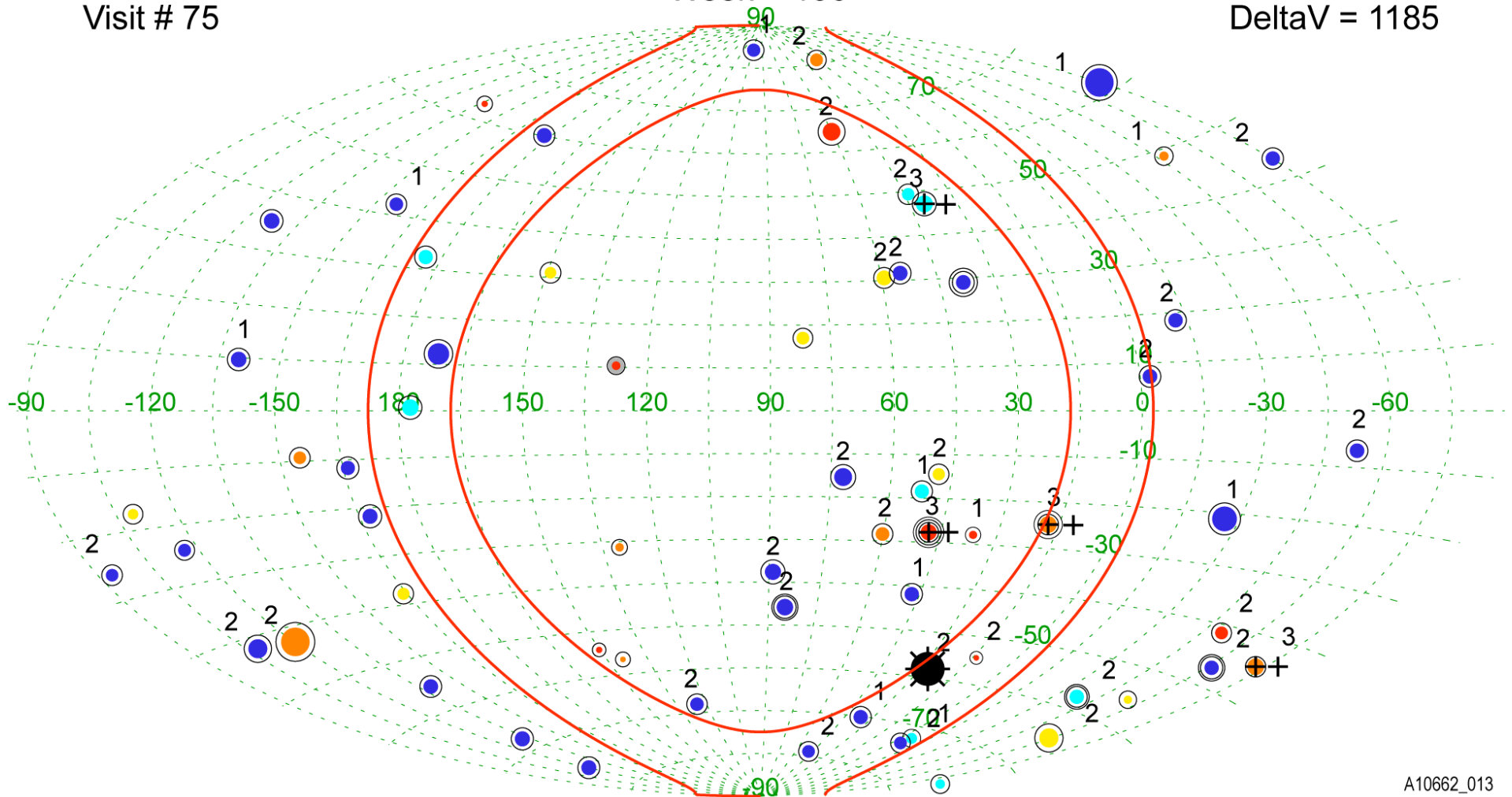


Planning the Mission

Week = 156

DeltaV = 1185

Visit # 75

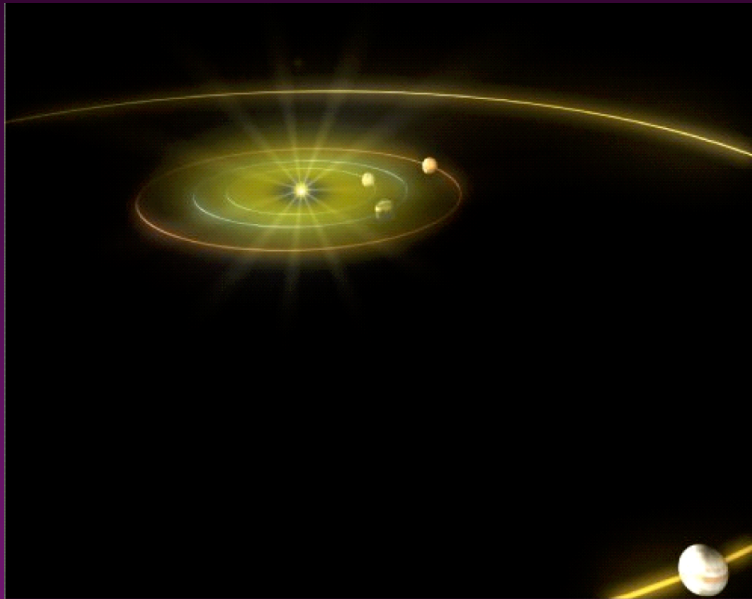


New Worlds Discoverer

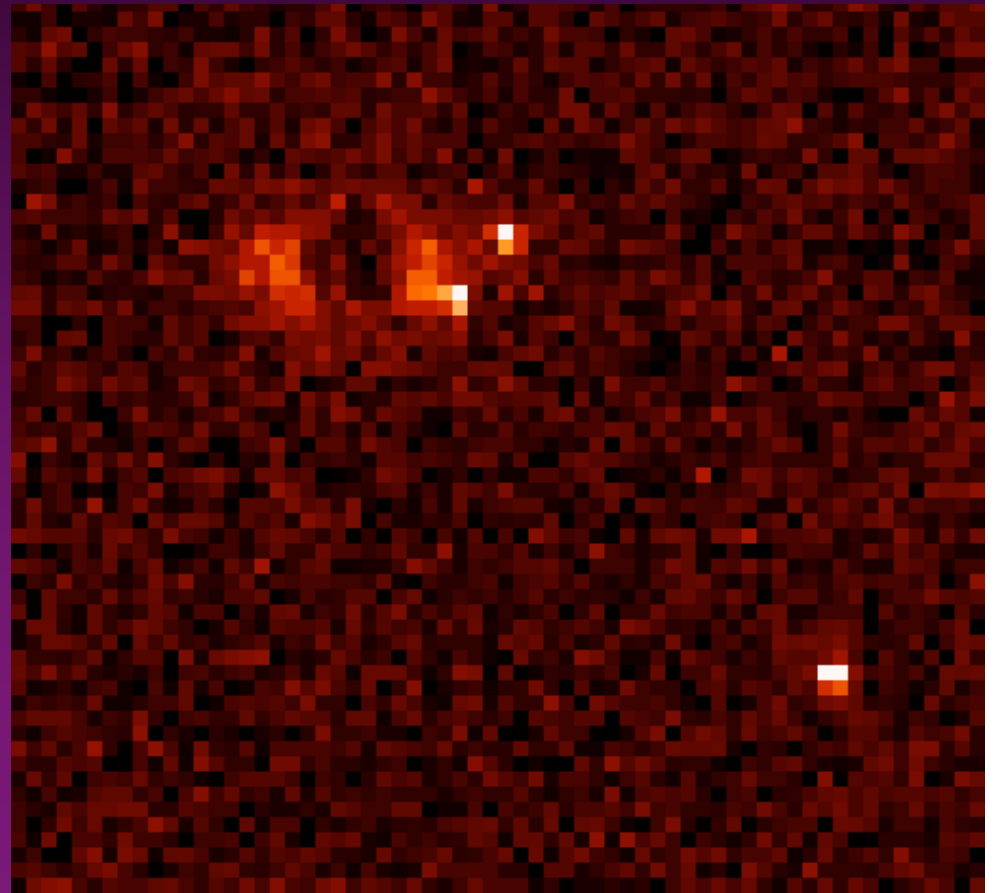
- ➡ **Could launch with JWST in 2013**
- ➡ **All formation flying requirements on starshade**
 - JWST passive, just points
- ➡ **Meets cost cap and technology readiness requirements**
- ➡ **Three year mission – over 100 lines of sight**
- ➡ **Capable of detecting Earth to 10pc**
- ➡ **Spectroscopy of Jovian planets**
- ➡ **Earth spectroscopy marginal**

- ➡ **Constitutes a low cost TPF**
- ➡ **Not Selected on this round**

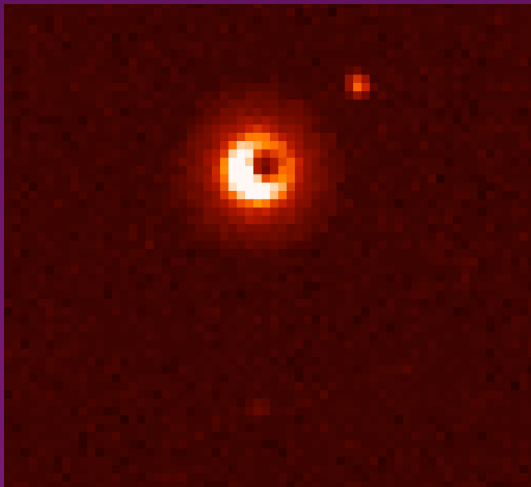
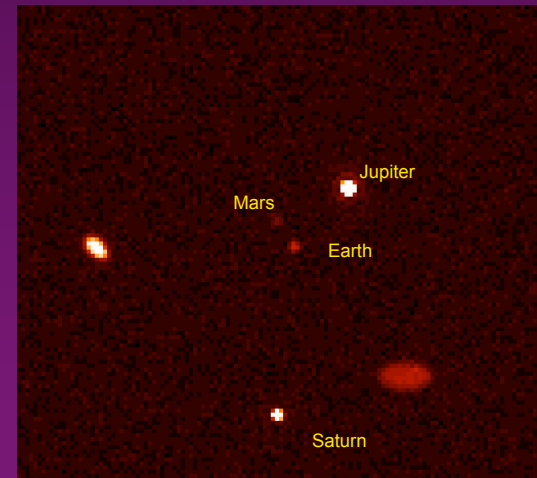
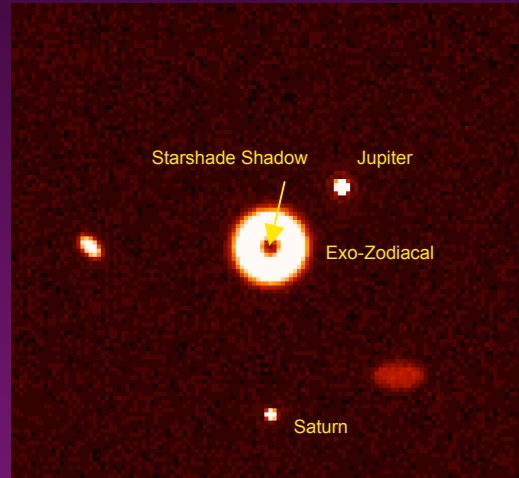
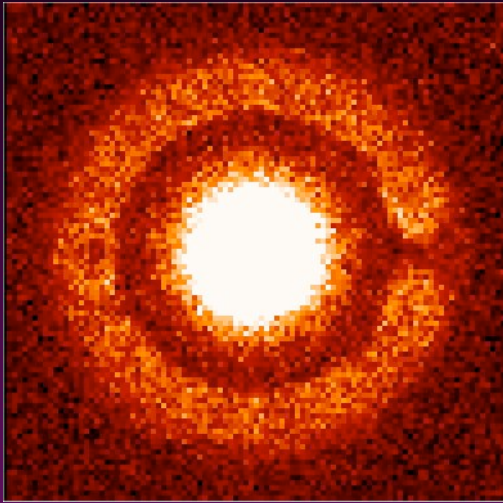
Simulated Solar System



A10662_038

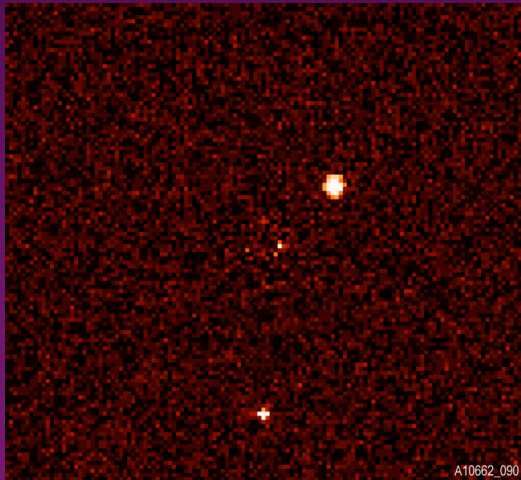


Discoverer Science Simulations

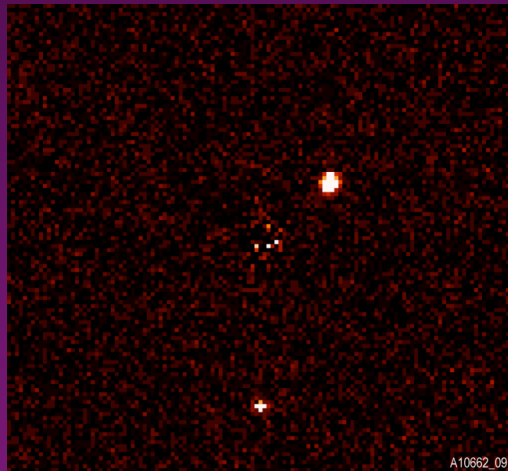


Additional Contrast from Telescope

10^{-8}



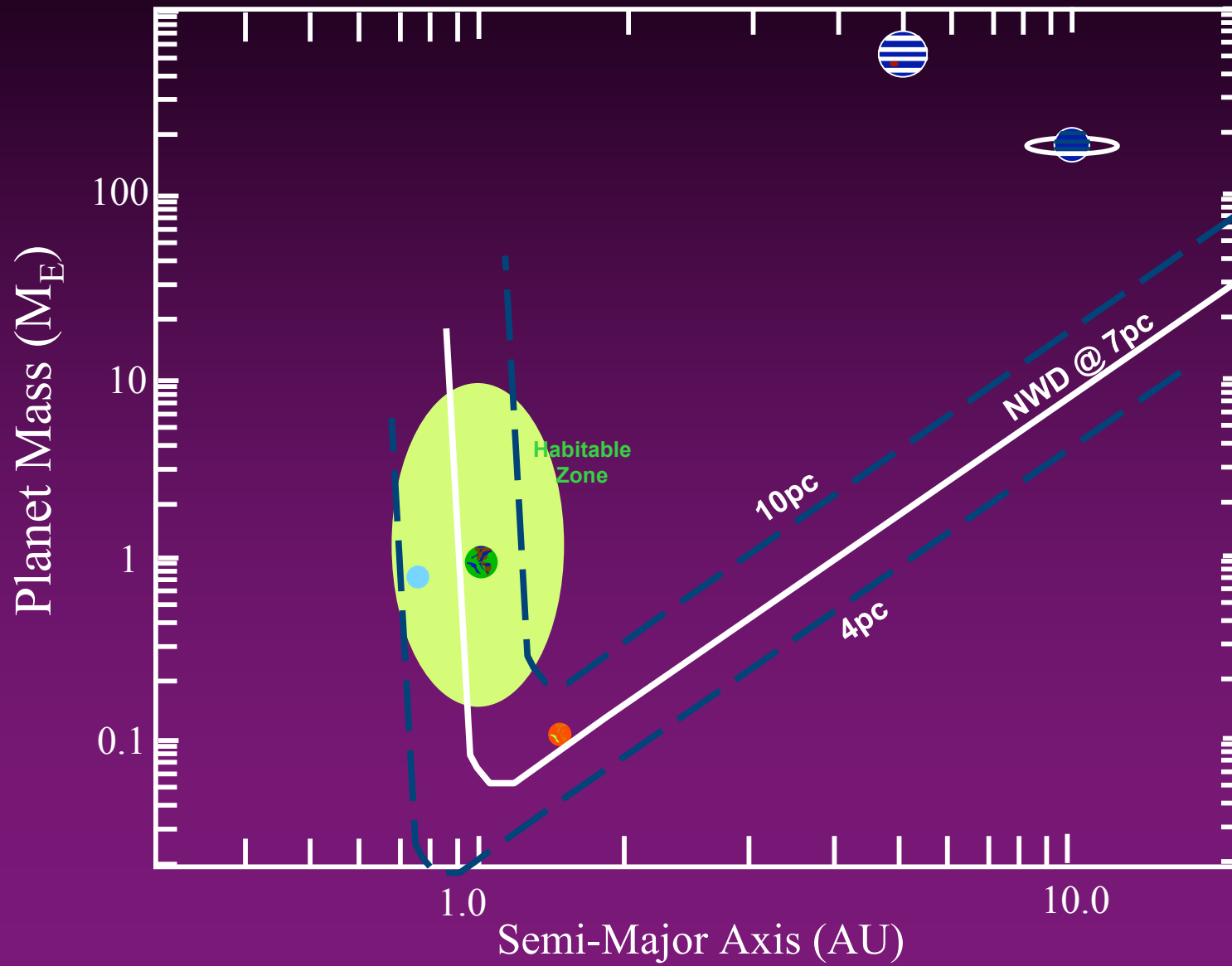
10^{-7}



10^{-6}



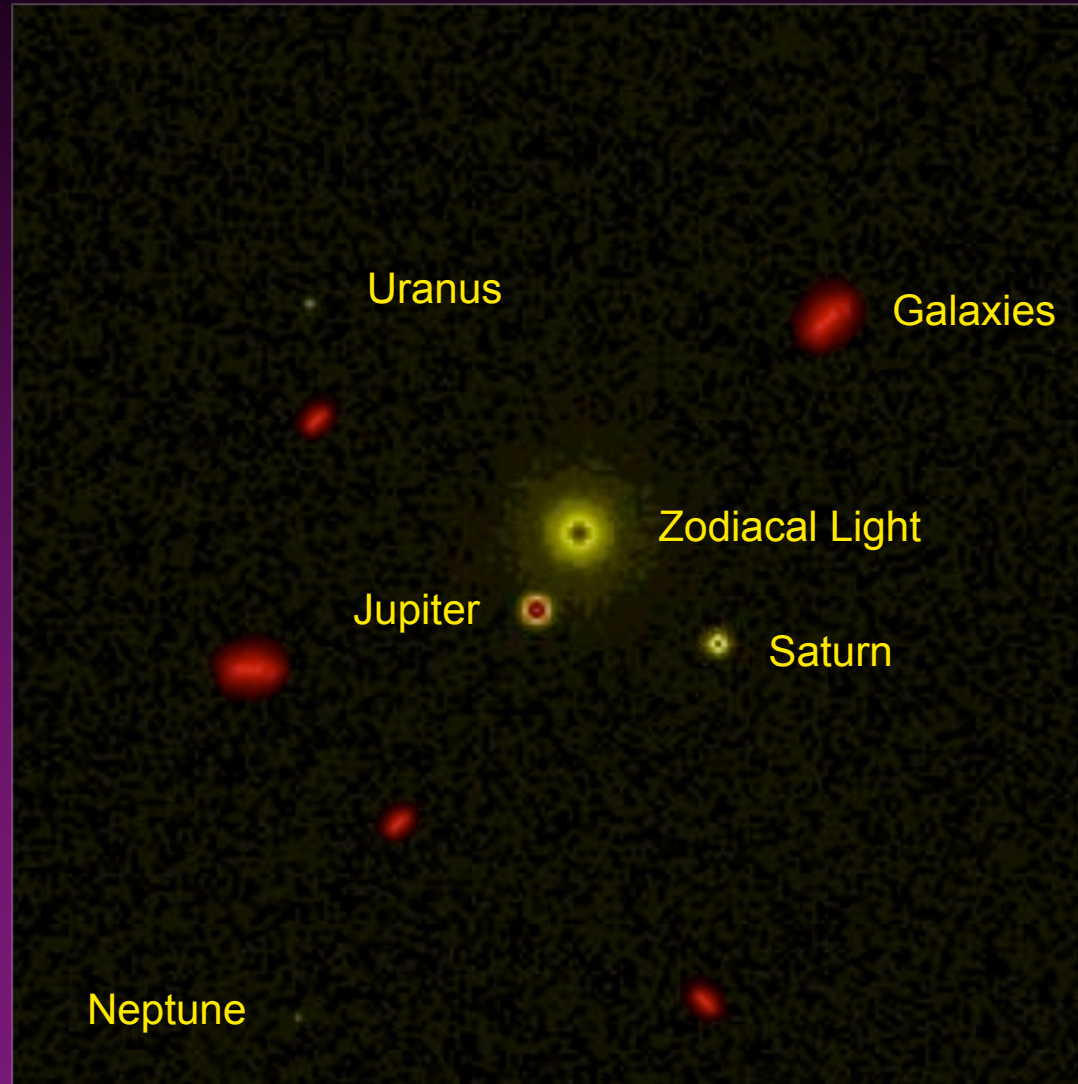
NWD Sensitivity



New World Observer Architecture

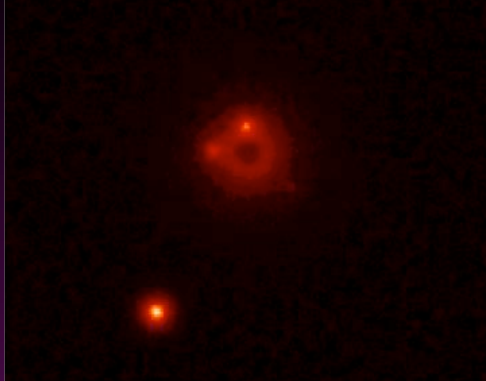
- ☞ **After NWD Proposal Submitted NGST looked at full-up system**
- ☞ **4m Telescope Diameter Breakpoint**
- ☞ **Two Starshades – one small and fast**
- ☞ **Very Powerful Scientifically**
- ☞ **Cost comparable to other missions on table**

The First Image of Solar System



← 10 arcseconds →

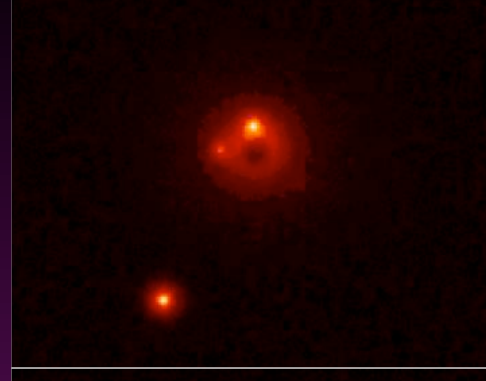
90°



80°



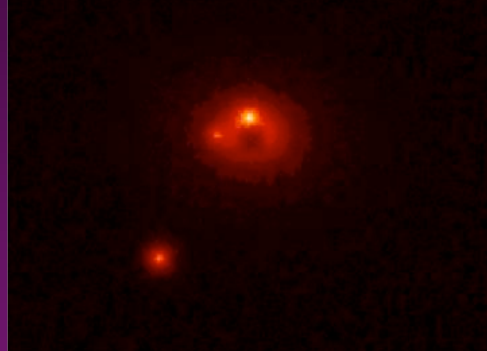
70°



60°



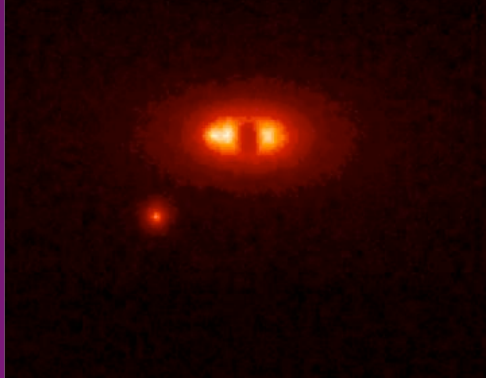
50°



40°



30°



20°

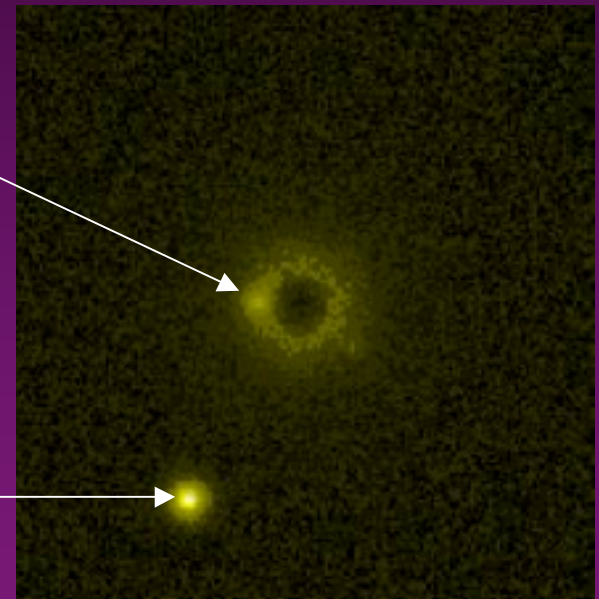
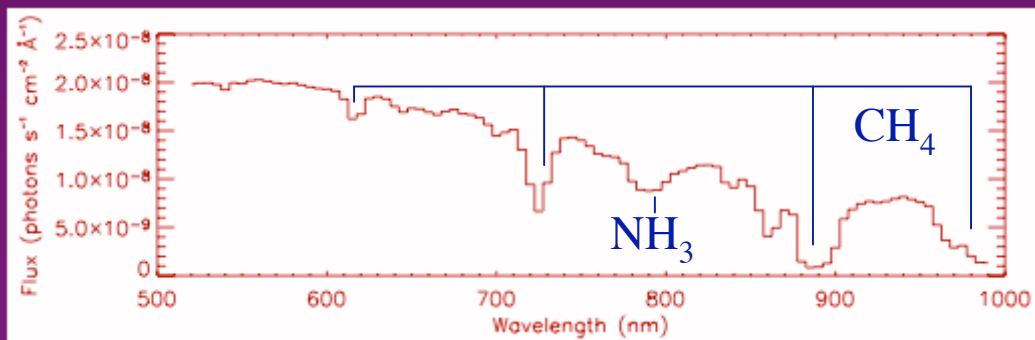
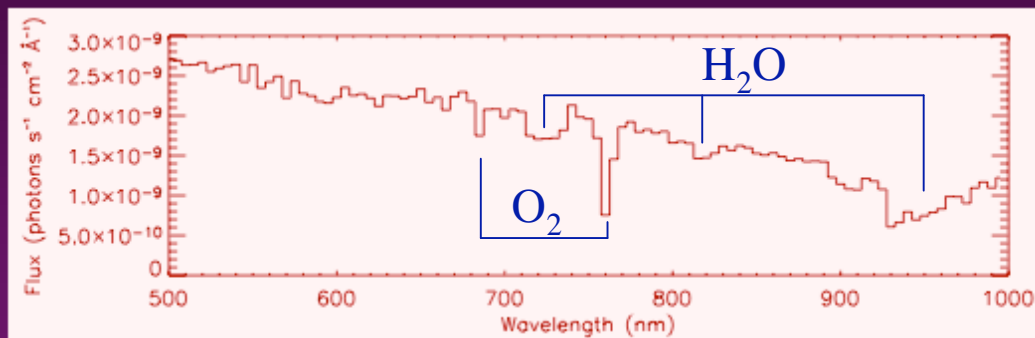


10°



Spectroscopy

☞ $R > 100$ spectroscopy will distinguish terrestrial atmospheres from Jovian with modeling

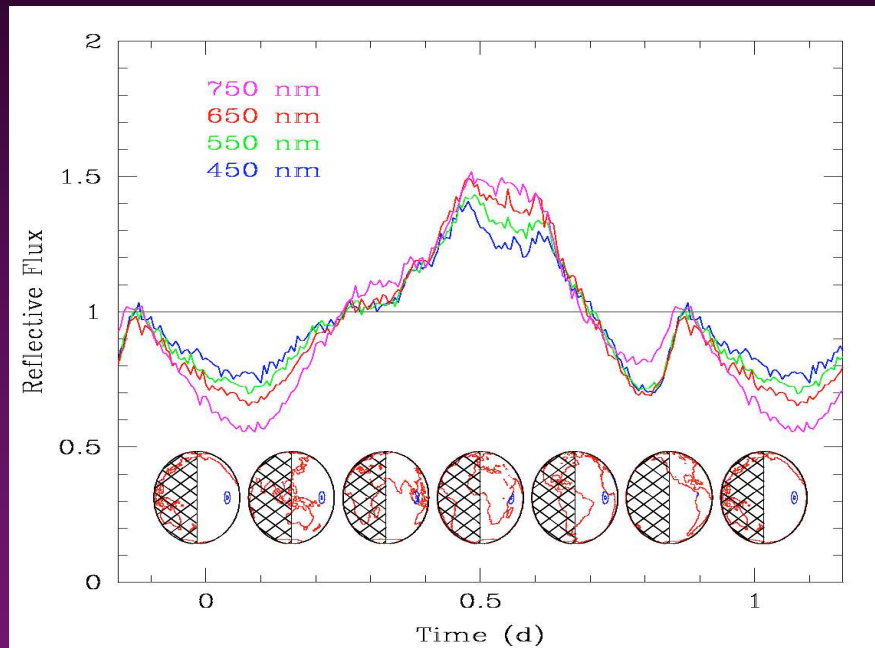


S. Seager

Spectroscopic Biomarkers

Water	Necessary for habitability
Oxygen	Free oxygen results only from active plant life
Ozone	Results from free oxygen
Nitrous Oxide	Another gas produced by living organisms
Methane	Life indicator if oxygen also present
Vegetation	Red edge of vegetation at 750nm

Photometry

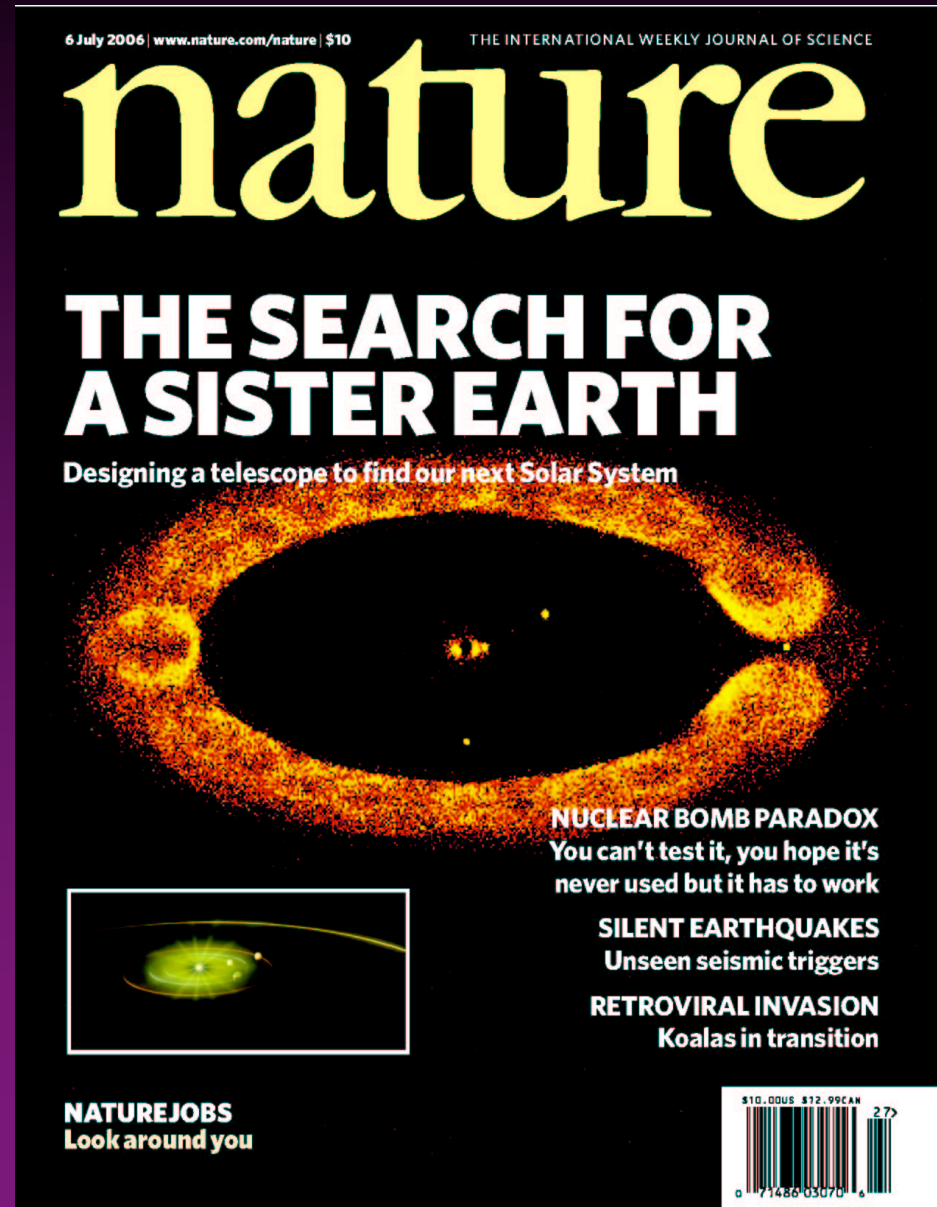


*Calculated Photometry of
Cloudless Earth as it
Rotates*

It Should Be Possible to Detect Oceans and Continents!

NWO Science

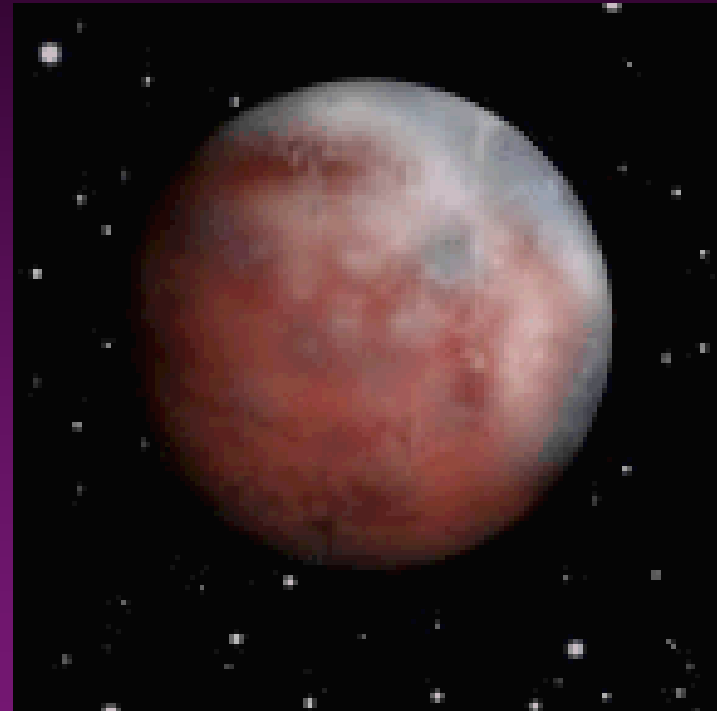
- ☞ **Result of Nature interviews**
 - Many discussions with press and other interested parties
- ☞ **It is Life Seeking that EVERYBODY wants**
 - Just finding water planets enough, but its not what motivates the public
- ☞ **Can there be a bigger or more important question for astronomers?**
- ☞ **New Worlds Observer can do it**
 - \$2-3 Billion and 10 years



The New Worlds Imager



Earth at 200km resolution. Oceans, continents and clouds are visible.

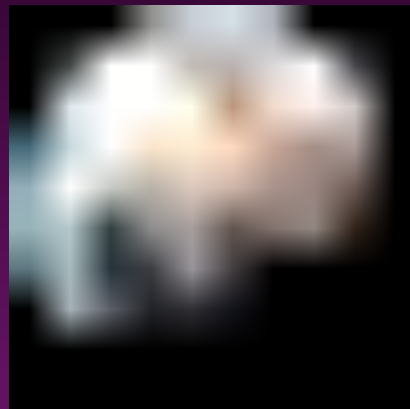


A simulated exo-planet at 500 km resolution.

TRUE PLANET IMAGING



3000 km



1000 km



300 km



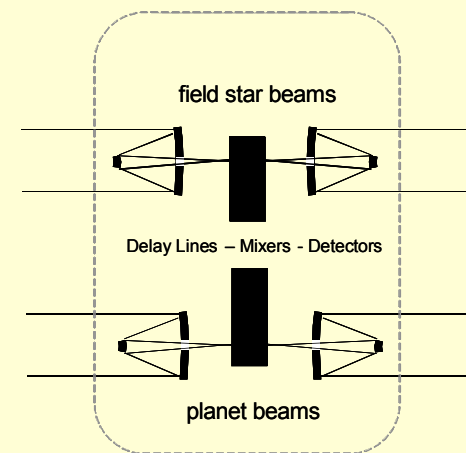
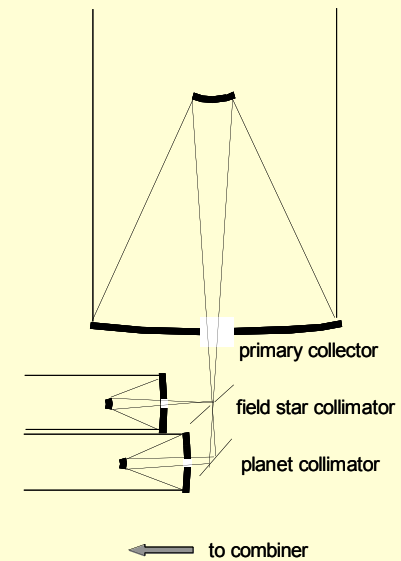
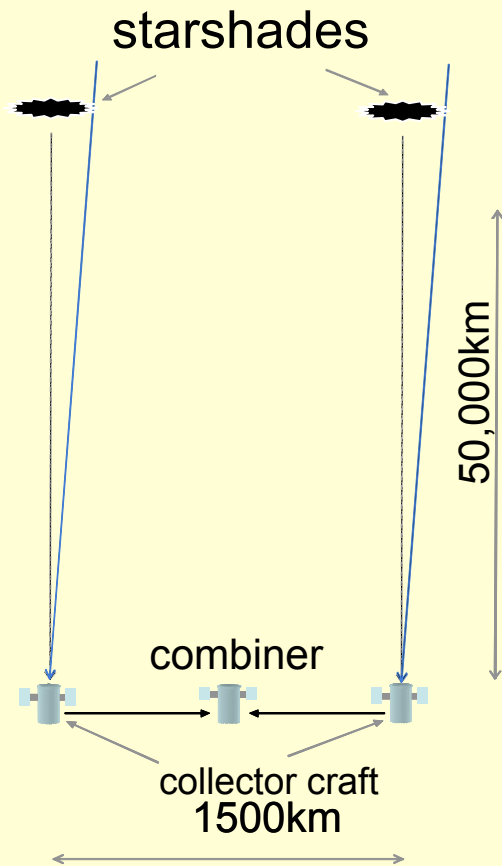
100 km

**Earth Viewed at Improving
Resolution**

Solar System Survey at 300km Resolution



NWI Concept



Hypertelescope Problem

☞ How Many Apertures Needed?

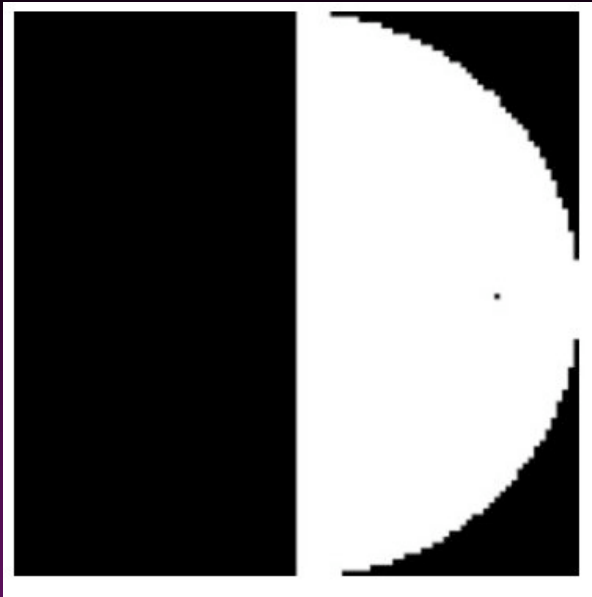
- One per pixel (no!)

☞ Cost control of multiple craft

☞ Formation Flying to Tolerance

☞ Labeyrie has worked on this

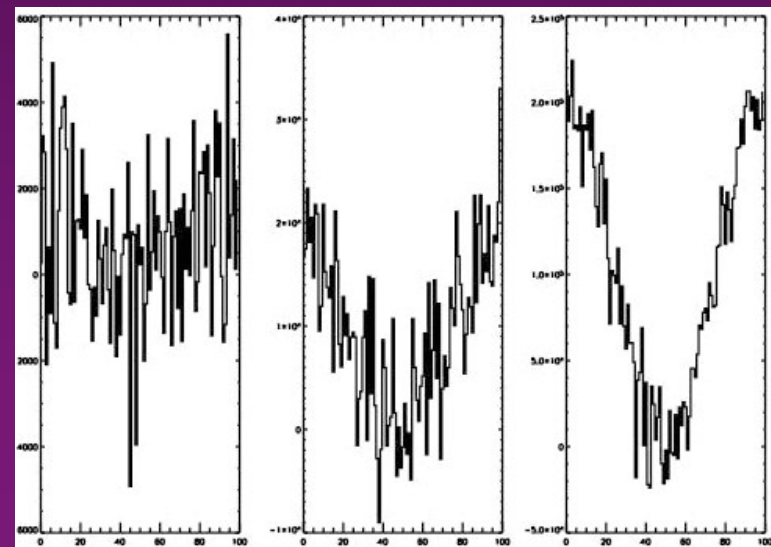
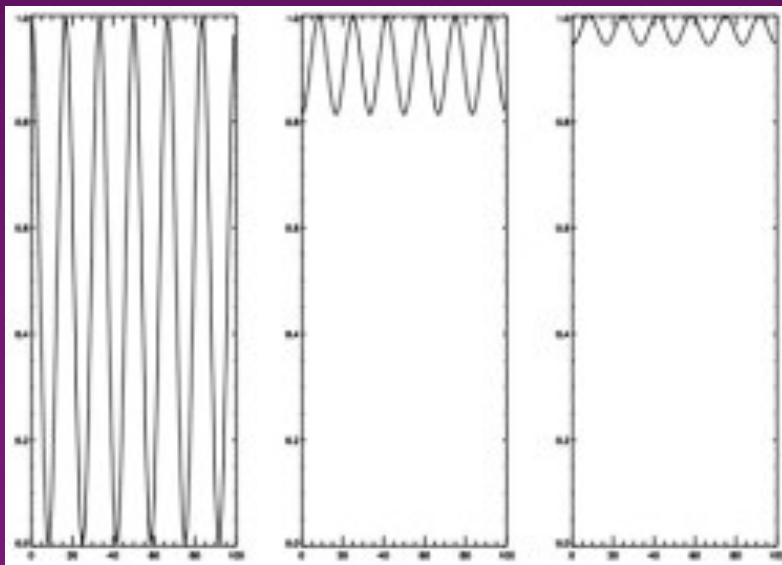
- Amazing telescope even without starshades



Sims

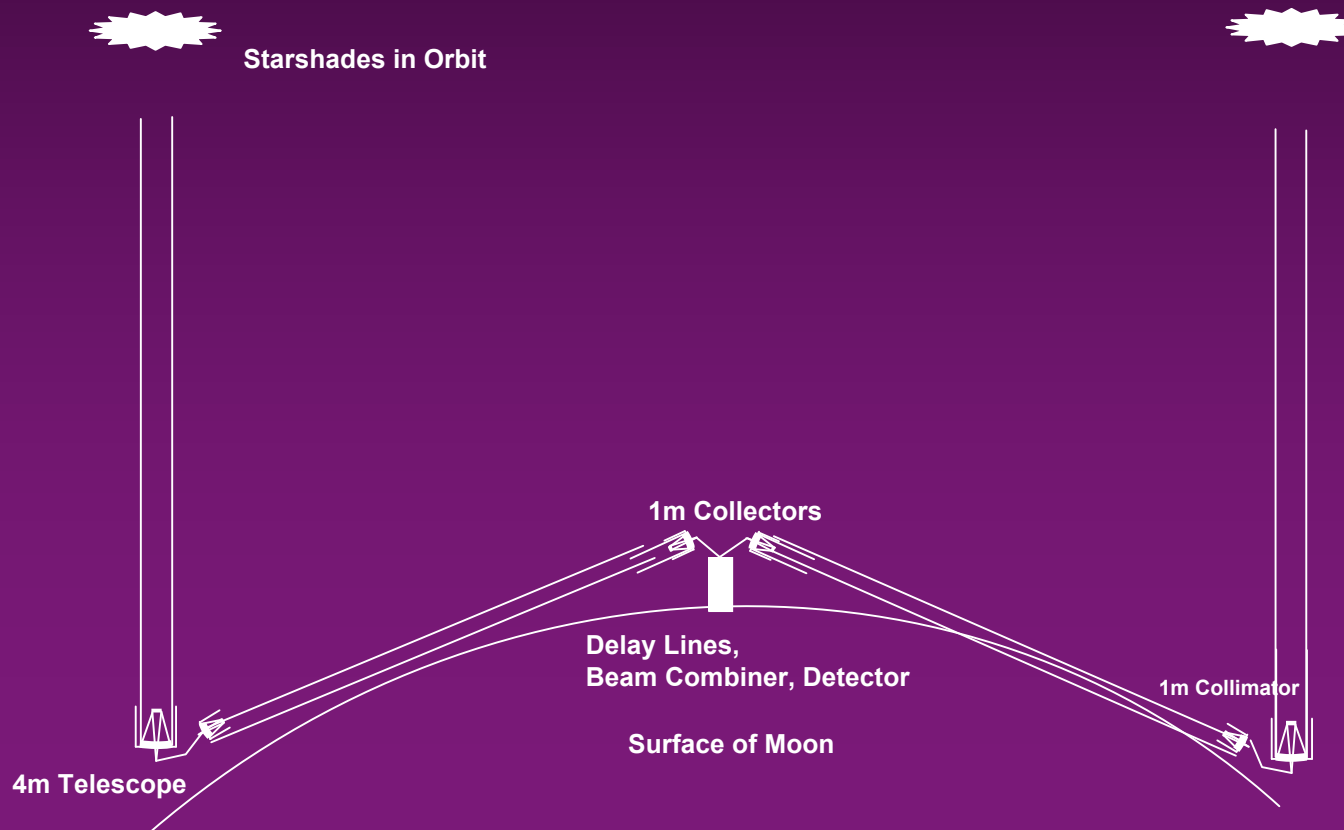
Established that information is present in the fringes and detectable.

How do we invert into images?
Is this enough?



Lunar Option

- ☞ Planet Imaging is exciting enough to justify the expense level
- ☞ Appropriate Level of Challenge for 20+ years from now



Tradeoff

Pro

Infrastructure

Moon Stable Bench

Refuel Starshades

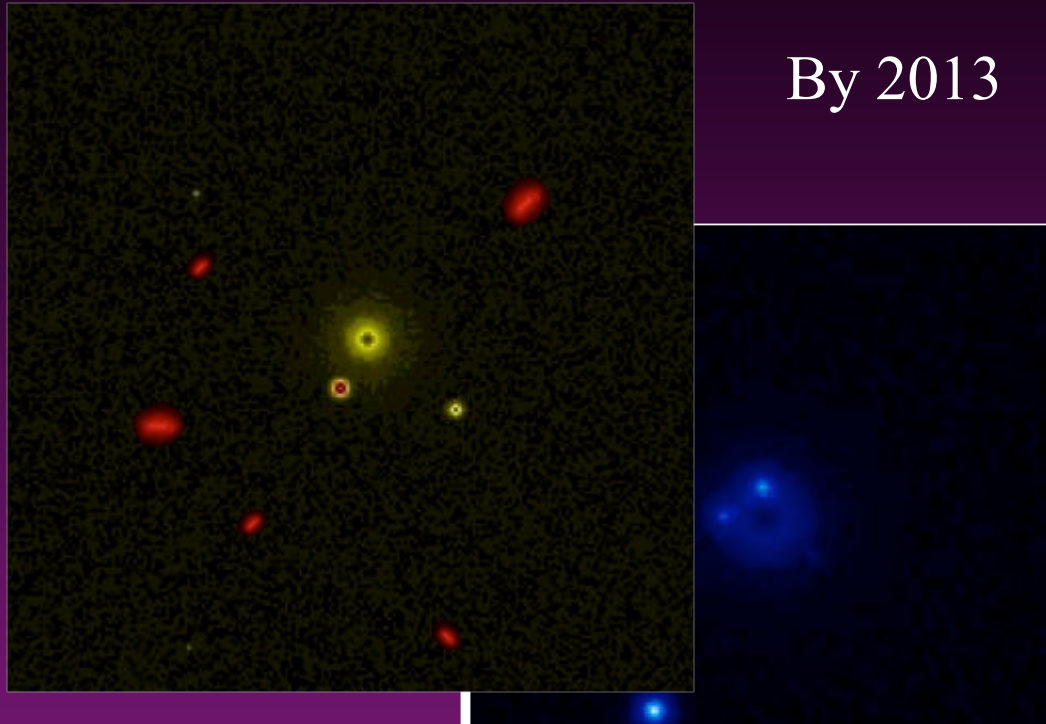
Con

Moon Rotates

100km class delay lines

Conclusion

By 2013



By 2025

