

# Calculating Hemispheric Power and Joule Heating using Defense Meteorological Satellite Program (DMSP) F-13 data

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Research mentors:

Dr. Barbara Emery & Dr. Astrid Maute



**High Altitude Observatory**



# Outline

## I. Introduction

- Preliminary knowledge
- Definition of Joule heating
- Importance of Joule heating

## II. Research Strategy

- Motivation
- Framework for calculation & analysis
- Goals

## III. Programming methodology & results

- Single day analysis
- Multiple day results

## IV. Key Findings

- Dawn Vs. Dusk comparison
- Equatorward Vs. Poleward comparison
- Hemispheric Power Vs. Joule Heating

## V. Conclusions

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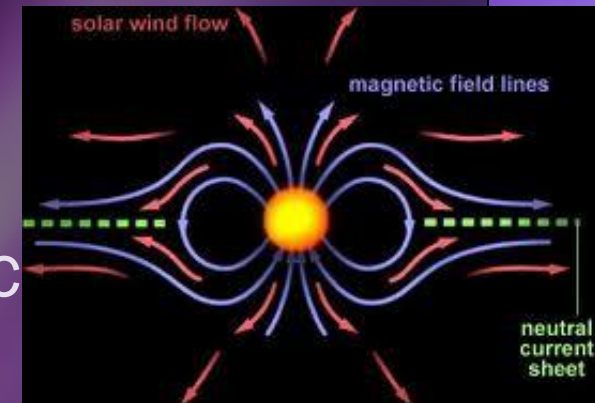
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# Preliminary knowledge

- Interplanetary Magnetic Field (IMF) is the Sun's magnetic field carried by solar wind in interplanetary space.
- IMF is a 3D vector :  $[B_x, B_y, B_z]$   
 $B_x$  &  $B_y$  are parallel to the ecliptic whereas  $B_z$  is perpendicular.
- When  $B_z$  is negative, IMF points south and is anti-parallel to the geomagnetic field. This creates a door for energetic particles to enter Earth's inner magnetosphere.



# Preliminary knowledge

- The DMSP F13 satellite was launched in March 1995 into a Sun synchronous, polar orbit in the 6-18 local time frame.
- We use DMSP data from two of its instruments:
  - (1) Special Sensor Precipitating Electron and Ion Spectrometer (SSJ/4)
  - (2) Ion Drift Meter (IDM).



# Preliminary knowledge

- Ion drift velocity ( $V_i$ ) =  $(E \times B) / B^2$  where E is Electric Field and B is Earth's magnetic field.
- $V_y$  is the horizontal cross-track ion velocity.
- *Convection Reversal Boundary (CRB)* is where  $V_y$  reverses direction.



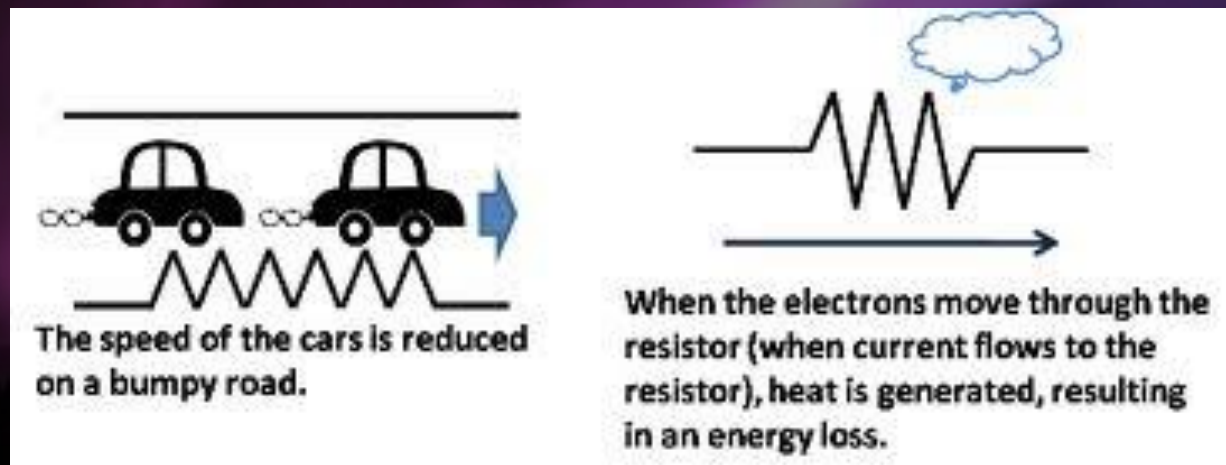


# Preliminary knowledge

- **Weimer 2005** is an empirical model of the high-latitude ion drift velocity. We compare Weimer 2005  $V_y$  with IDM  $V_y$  observations.
- **TIEGCM (Thermosphere Ionosphere Electrodynamics General Circulation Model)** is a numeric simulation model for Earth's upper atmosphere. TIEGCM uses Weimer 2005 model.
- **Hemispheric power (HP)** is the spatially integrated energy flux of precipitating electrons.

# Definition of Joule heating

- *Joule heating* ( $Q_J$ ) is the heat loss due to passage of electric current through a conductor.
- In the ionosphere, it occurs due to the friction of ions moving through neutral atoms.





# Importance of Joule Heating

1. Joule heating is usually the largest heat source in high-latitude regions. During geomagnetic storms, Joule heating can also exceed the global solar heating from UV and EUV radiation [Knipp et al., Solar Physics, 2004].
2. Joule heating is the largest source of uncertainty in the energetics of the thermosphere.

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# Motivation

- Figure 8 from Heelis et al. [JGR,1980] is an estimate of the relative locations of the aurora and the ion drift .
- We aim is to improve the parameterization of the aurora in the TIEGCM so that the resulting Joule heating is approximately correct.

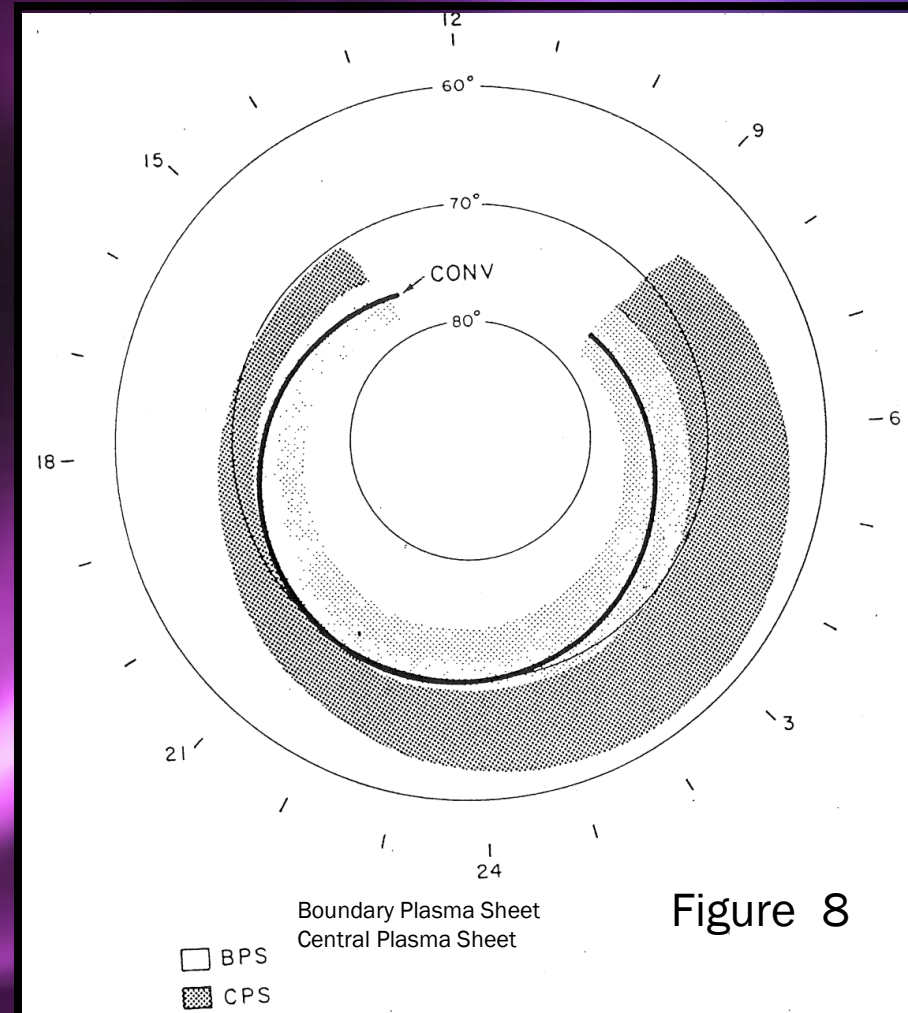


Figure 8

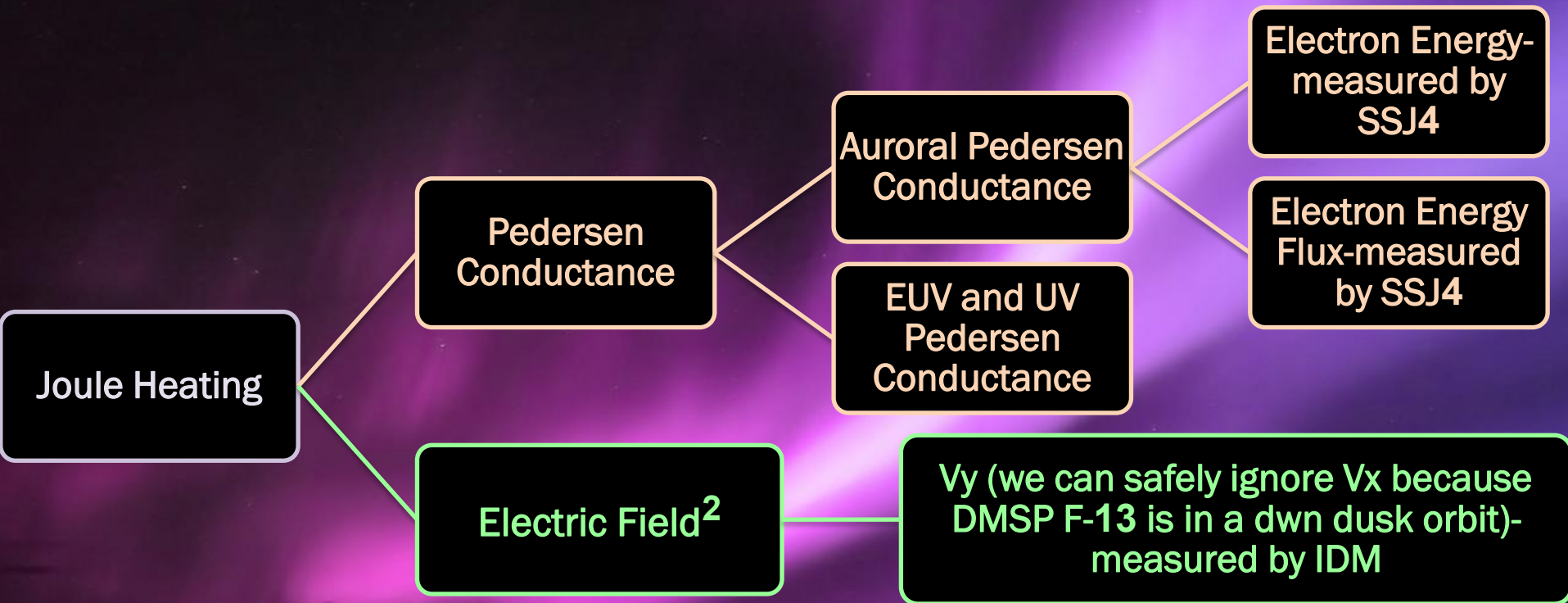


## Framework for calculation & analysis

- *Total Joule heating*  
 $\approx \text{Pedersen Conductance} \times \text{Electric Field}^2$
- *Particle Joule heating*  
 $\approx \text{Auroral Pedersen Conductance} \times \text{Electric Field}^2$
- *Total Joule heating*  
 $= \sqrt{\text{Particle Joule heating}^2 + \text{EUV and UV Joule heating}^2}$

# Framework for calculation & analysis

- Need to analyze the components of Joule heating.



## Goals

1. Analyze the local time variation in Joule heating, i.e. compare Joule heating during dawn, dusk, midnight and noon.
2. Study the spatial distribution of Joule heating  
In particular, compare Joule heating in the polar cap (anti-sunward ion flow) with equatorward Joule heating (sunward ion flow).
3. Analyze the relative location of electron energy flux with respect to  $V_y$ .
4. Quantitatively compare hemispheric power, particle Joule heating, and total Joule heating for different IMF values,.



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## III. Programming methodology & results

- Single day analysis ( $A_p=84$ , very stormy)
- Multiple day results

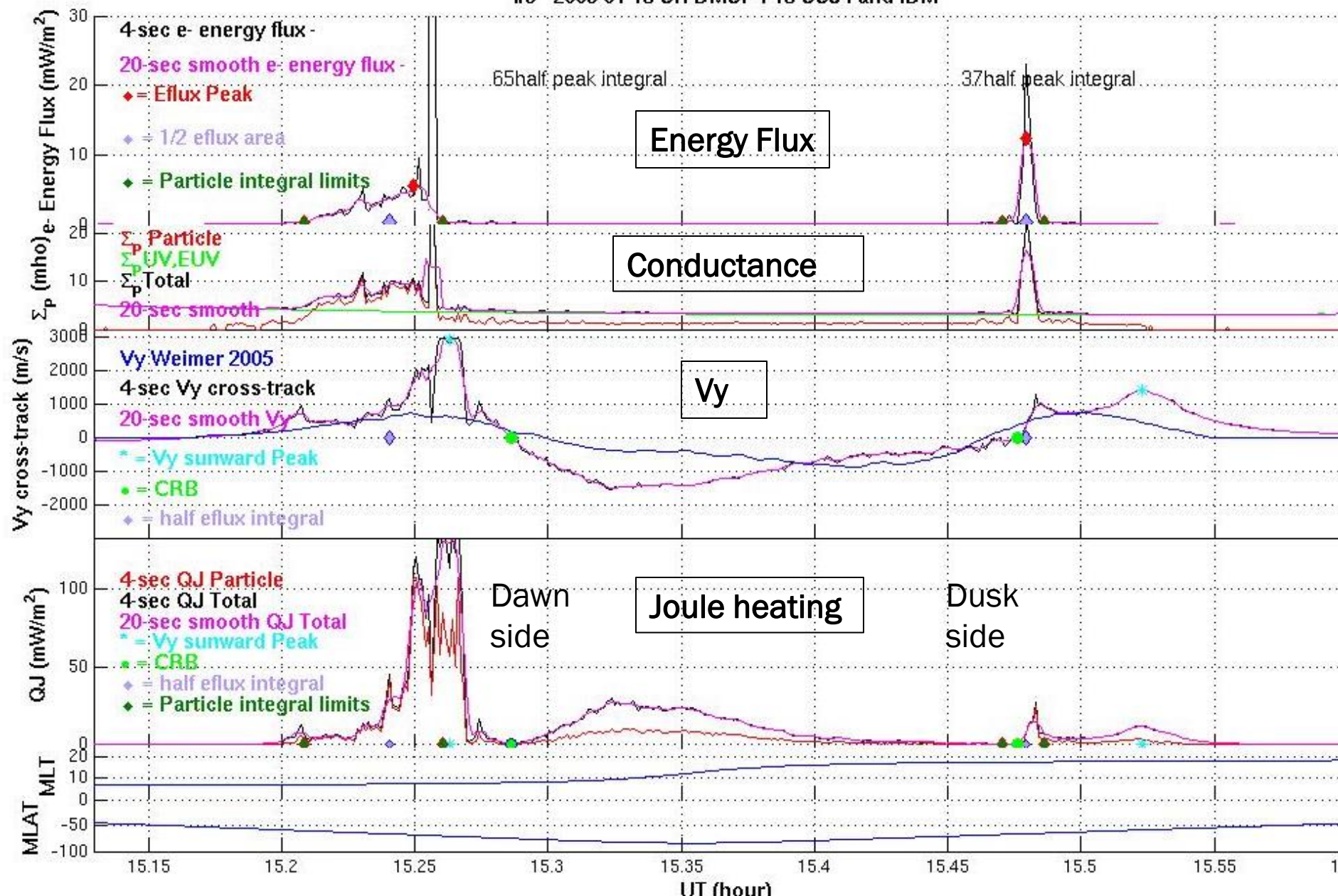
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## Single day analysis: One Orbit

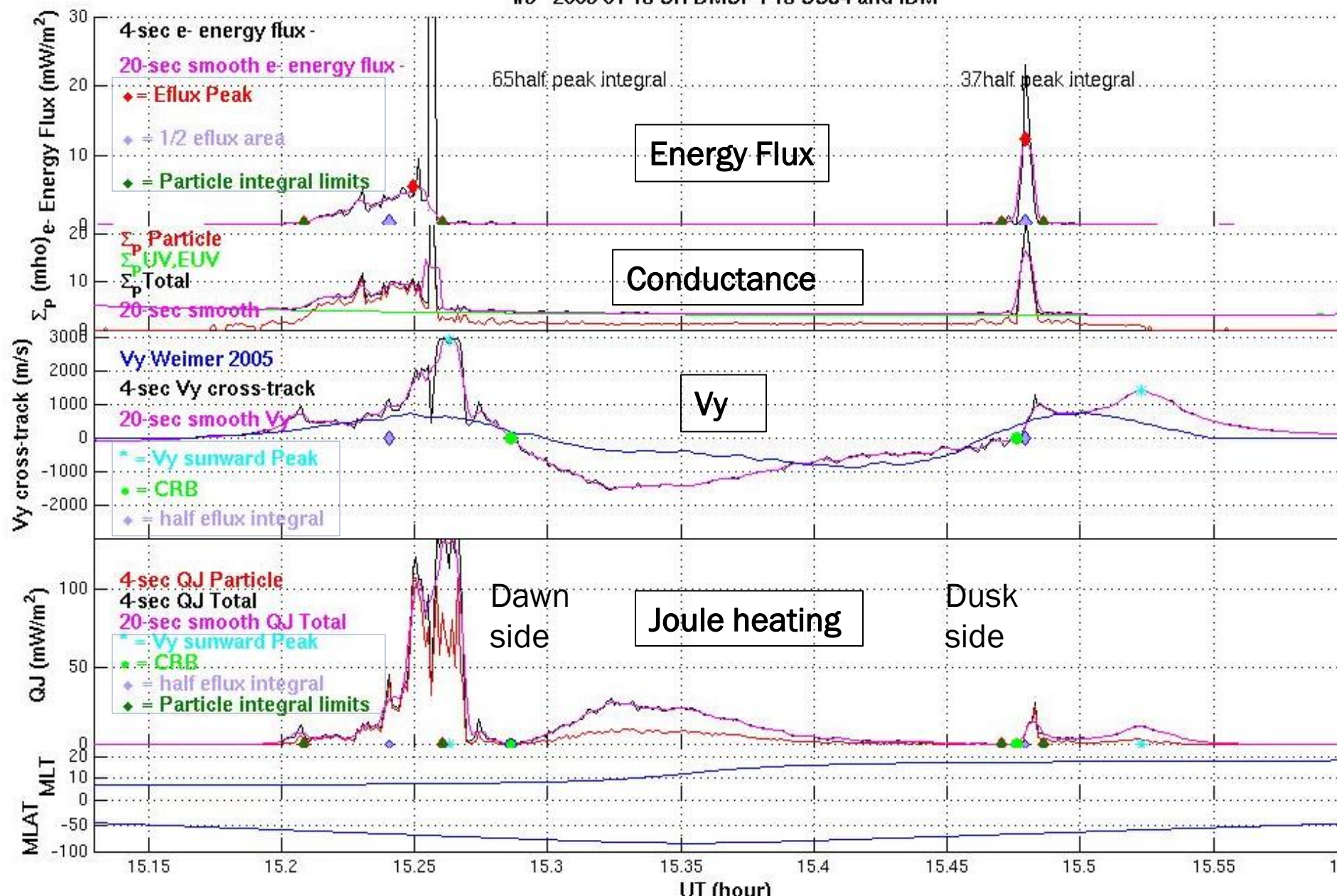
#5 2005 01 18 SH DMSP-F13 SSJ4 and IDM





## Single day analysis: One Orbit

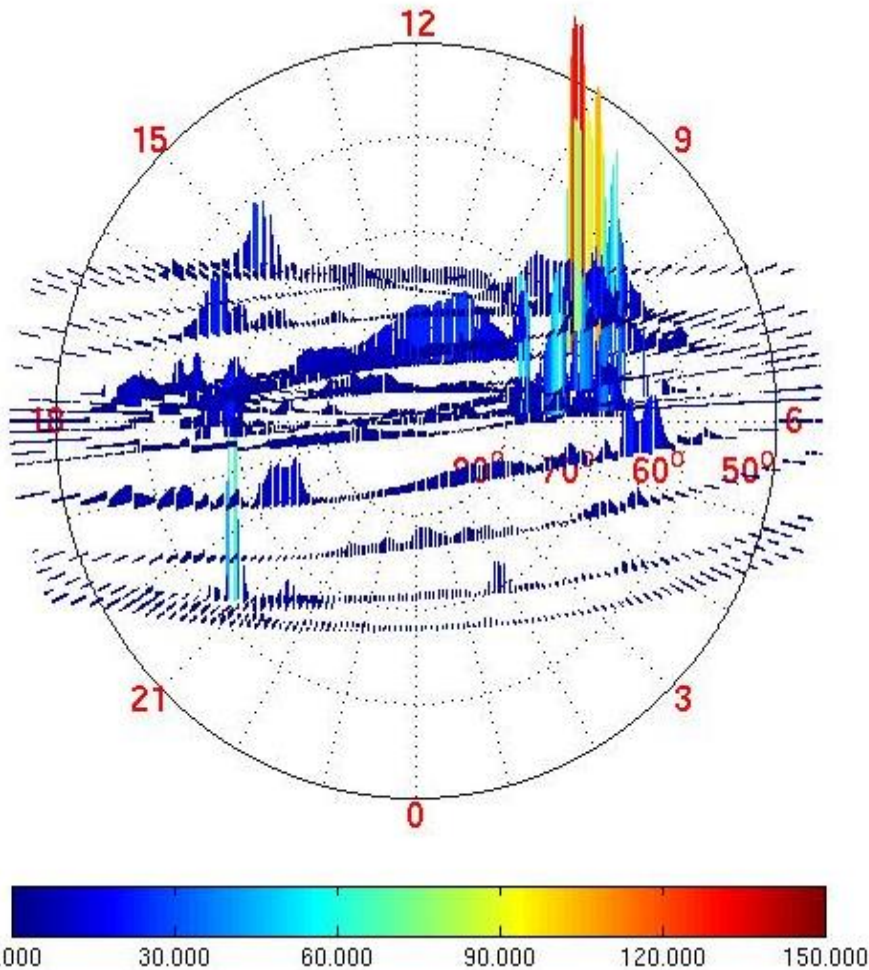
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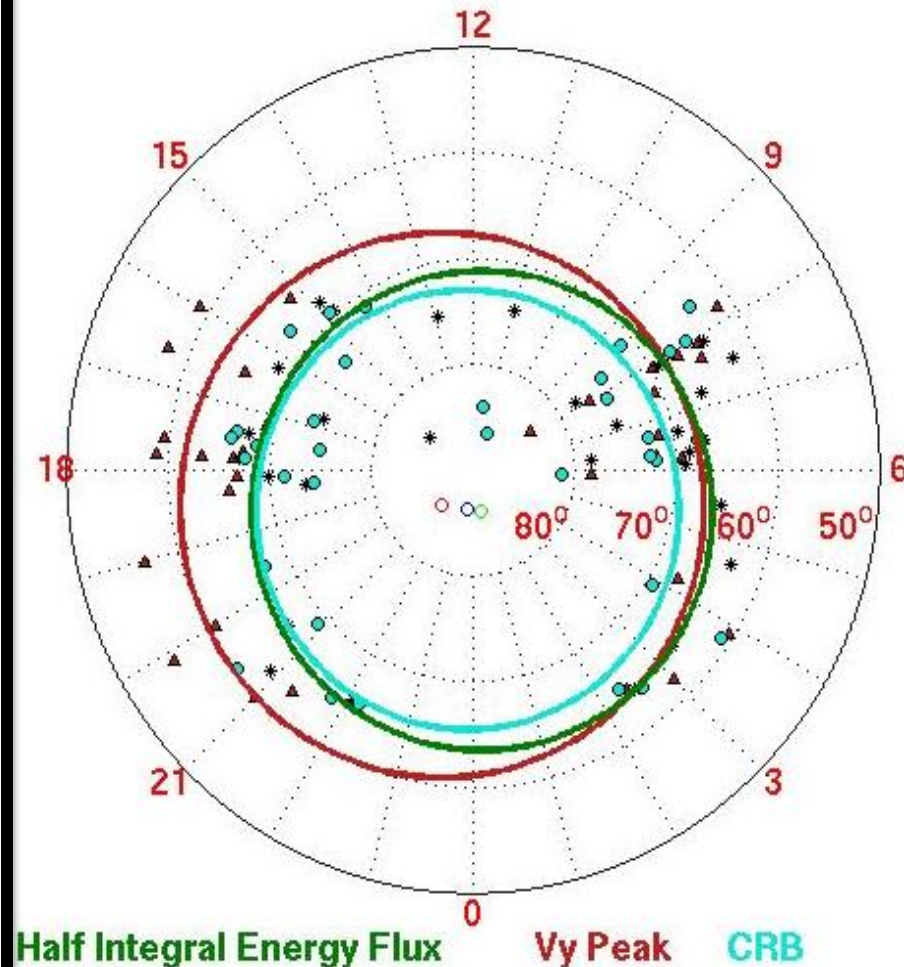


# Single day analysis: Format for All Orbits

05 01 18 DMSP-F13 smoothed Total Joule Heating ( $\text{mW/m}^2$ )



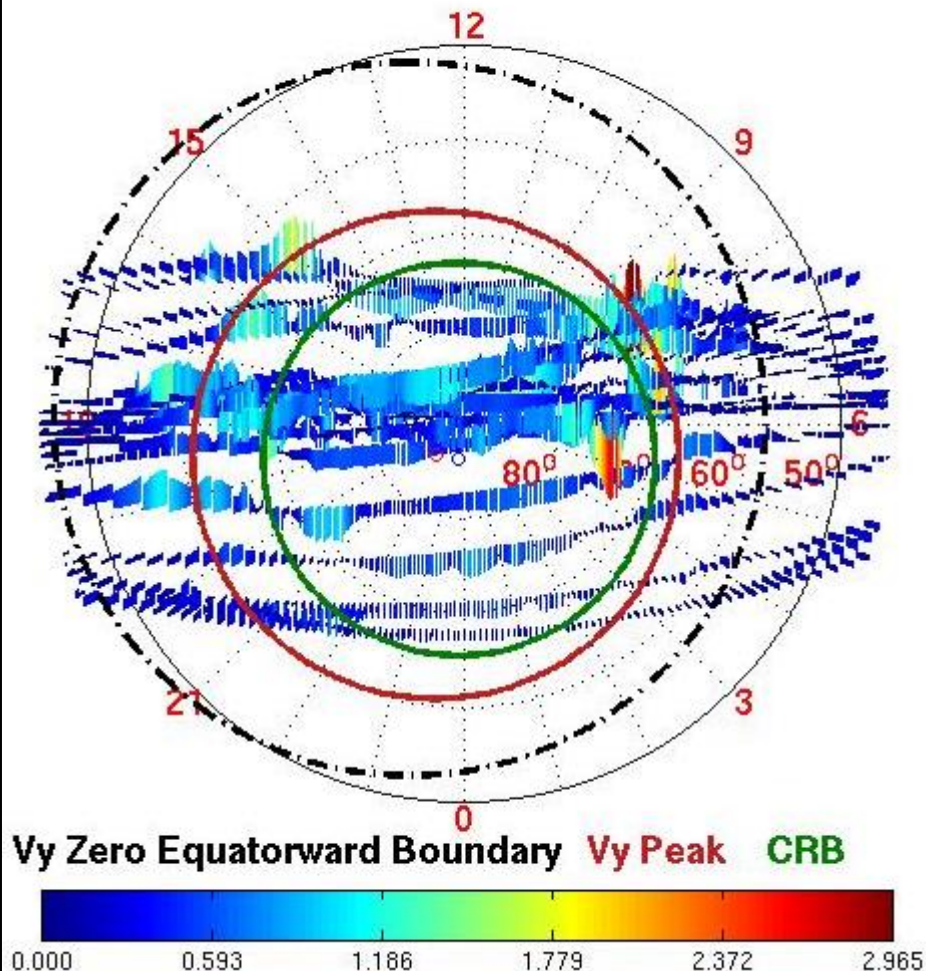
2005 01 18 DMSP-F13



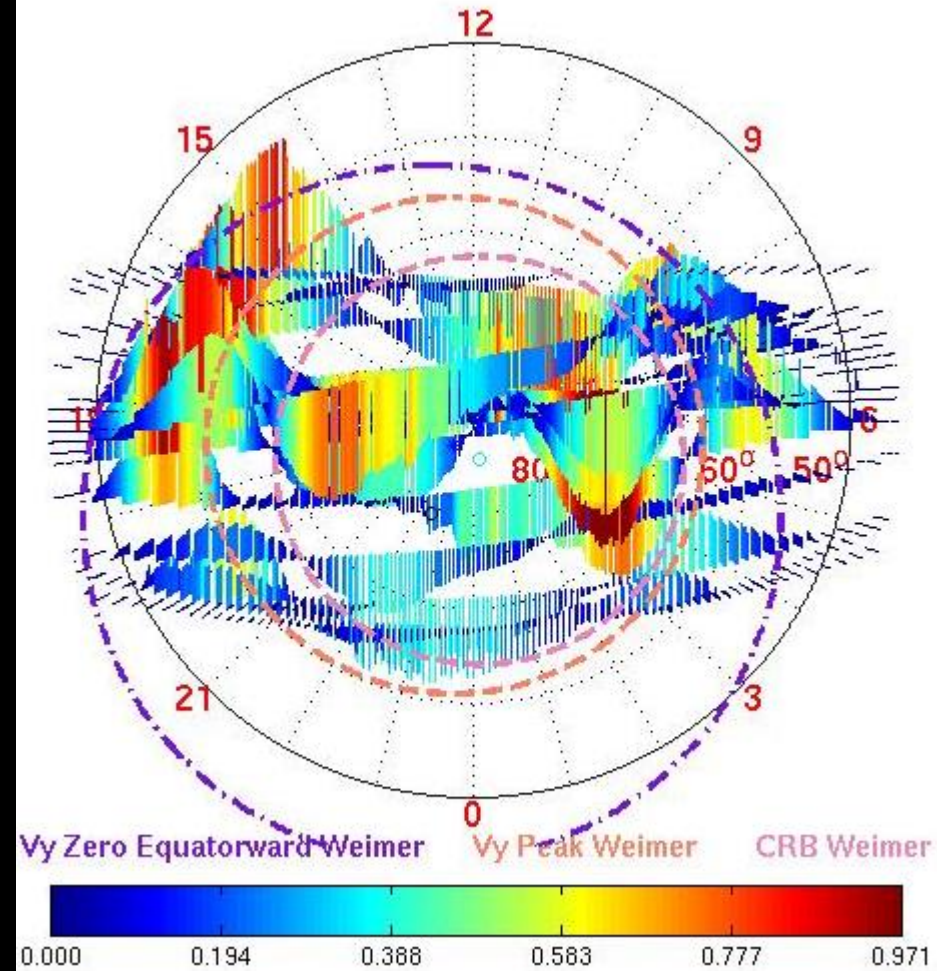
### III. Programming methodology & results

# Single day analysis: All Orbits

2005 01 18 DMSP-F13 smoothed IDM Vy (km/s)



2005 01 18 Weimer 2005 smoothed Vy (km/s)

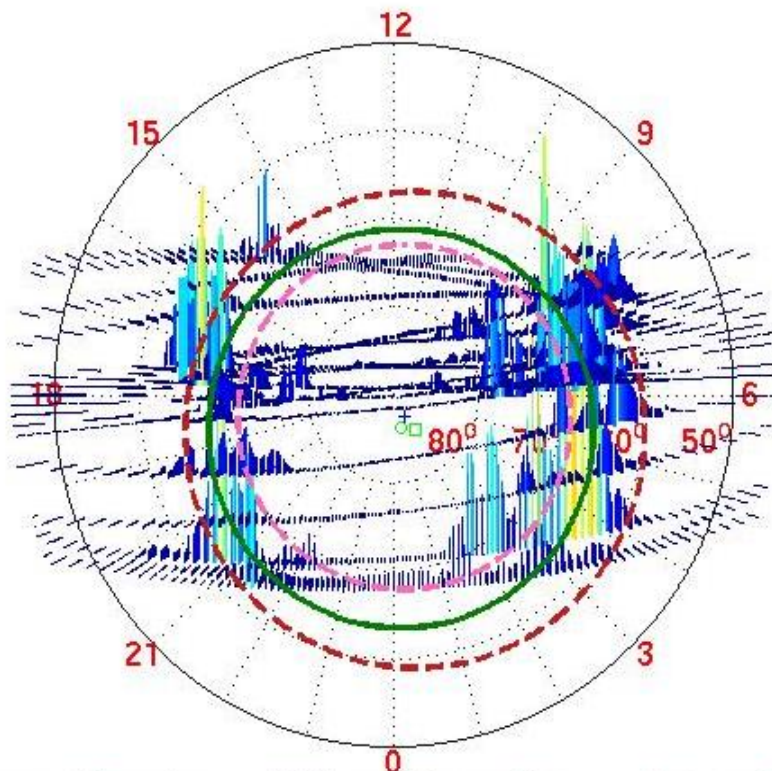




### III. Programming methodology & results

# Single day analysis: All Orbits

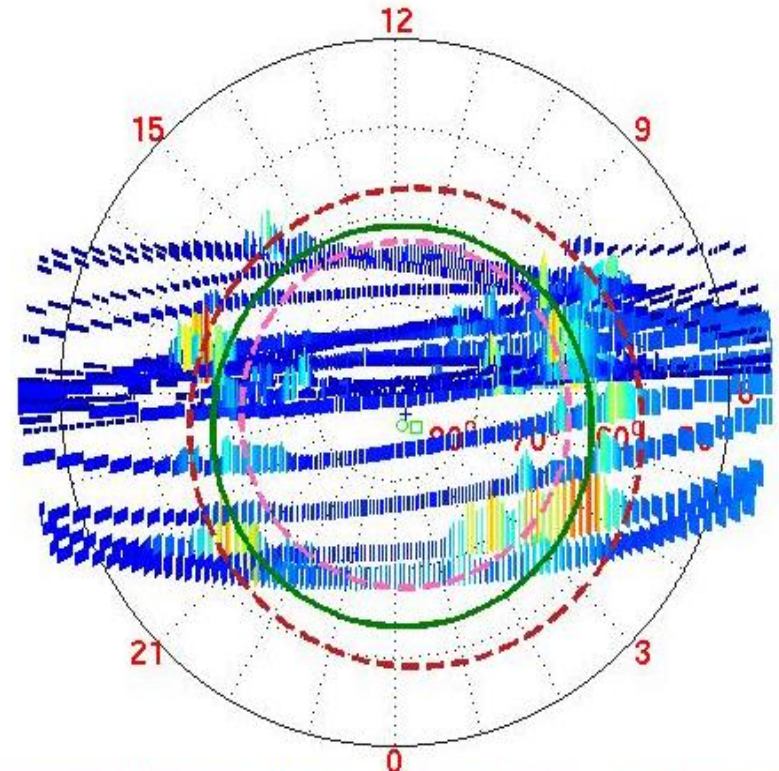
2005 01 18 DMSP-F13 smoothed SSJ4 eflux ( $\text{mW/m}^2$ )



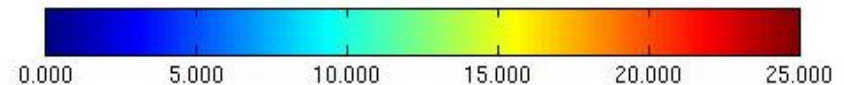
Equatorward Boundary Half Integral Energy Flux Poleward Boundary



01 18 DMSP-F13 smoothed SSJ4 Pedersen conductance (S)



Equatorward Boundary Half Integral Energy Flux Poleward Boundary

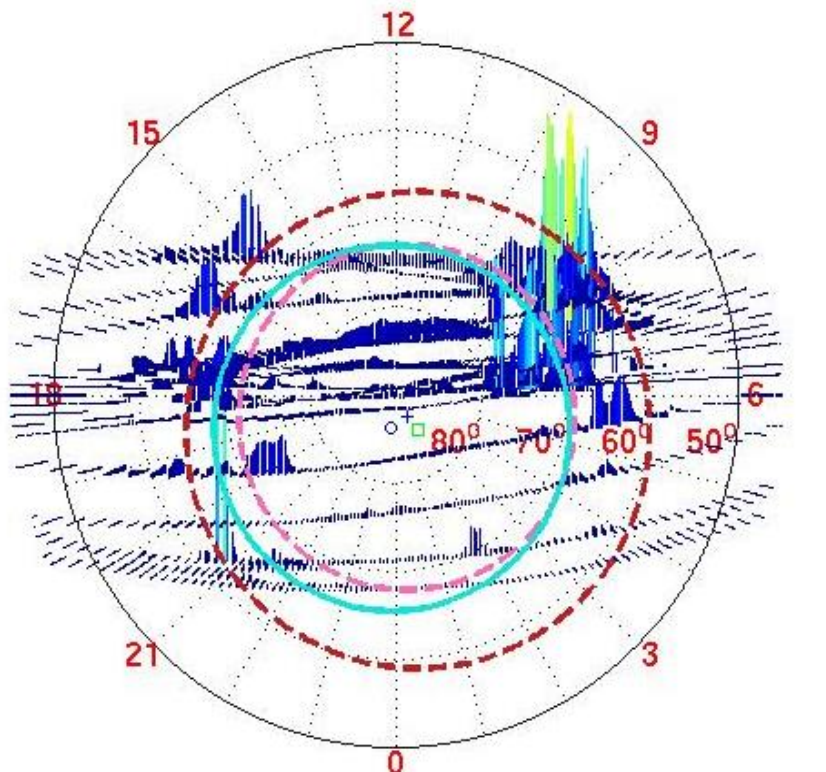


Particle Joule heating and Hemispheric Power are calculated for the region between the Poleward and Equatorward boundary.

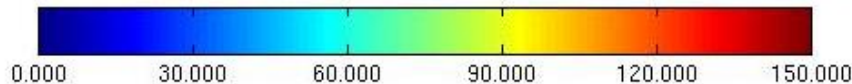


# Single day analysis: All Orbits

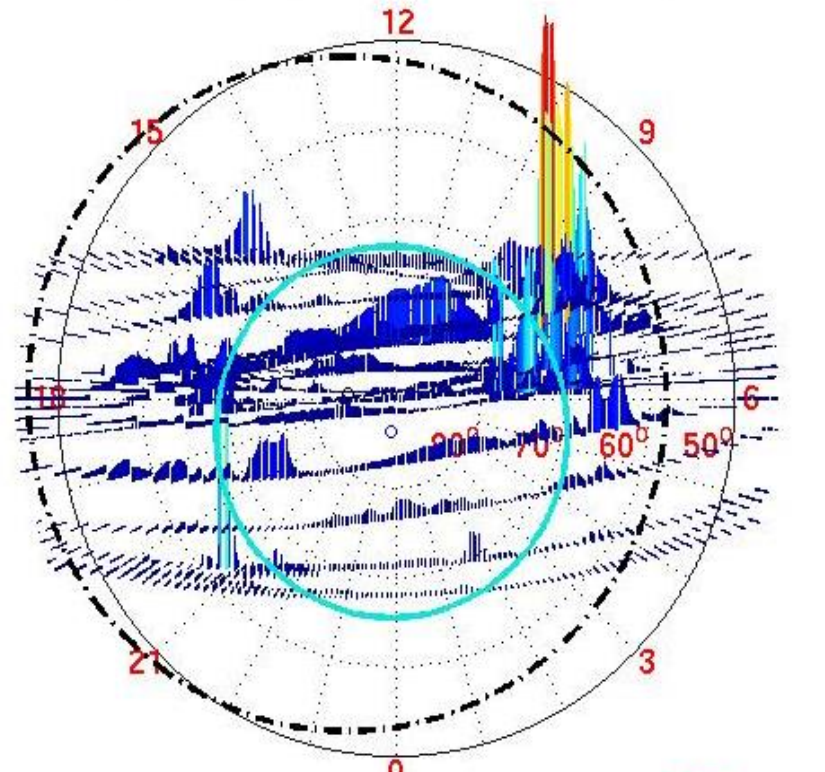
05 01 18 DMSP-F13 smoothed Particle Joule Heating ( $\text{mW}/\text{m}^2$ )



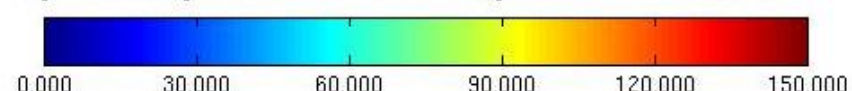
Equatorward Boundary CRB Poleward Boundary



05 01 18 DMSP-F13 smoothed Total Joule Heating ( $\text{mW}/\text{m}^2$ )



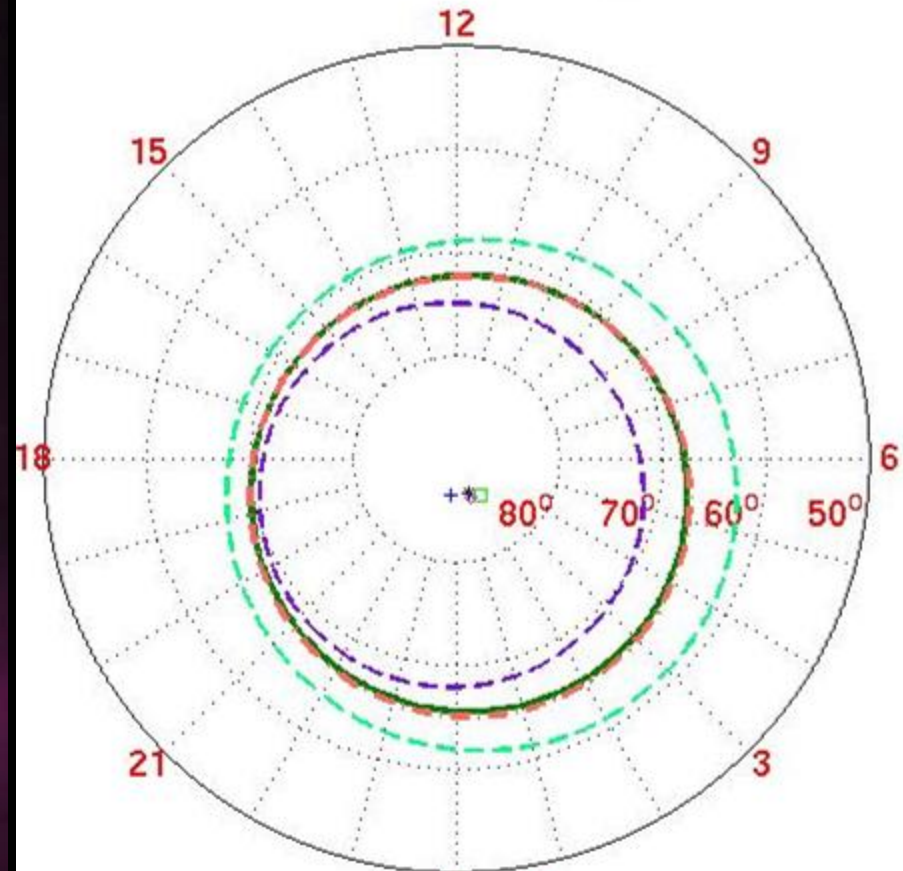
Vy Zero Equatorward Boundary CRB



Region inside CRB circle has *poleward Joule heating* due to anti-sunward ion flow, whereas the region between Vy zero Equatorward Boundary and CRB has *equatorward Joule heating* due to sunward ion flow.

# Multiple day results: Jan-June 2005

**-4<Bz<-3 Electron Energy Flux**



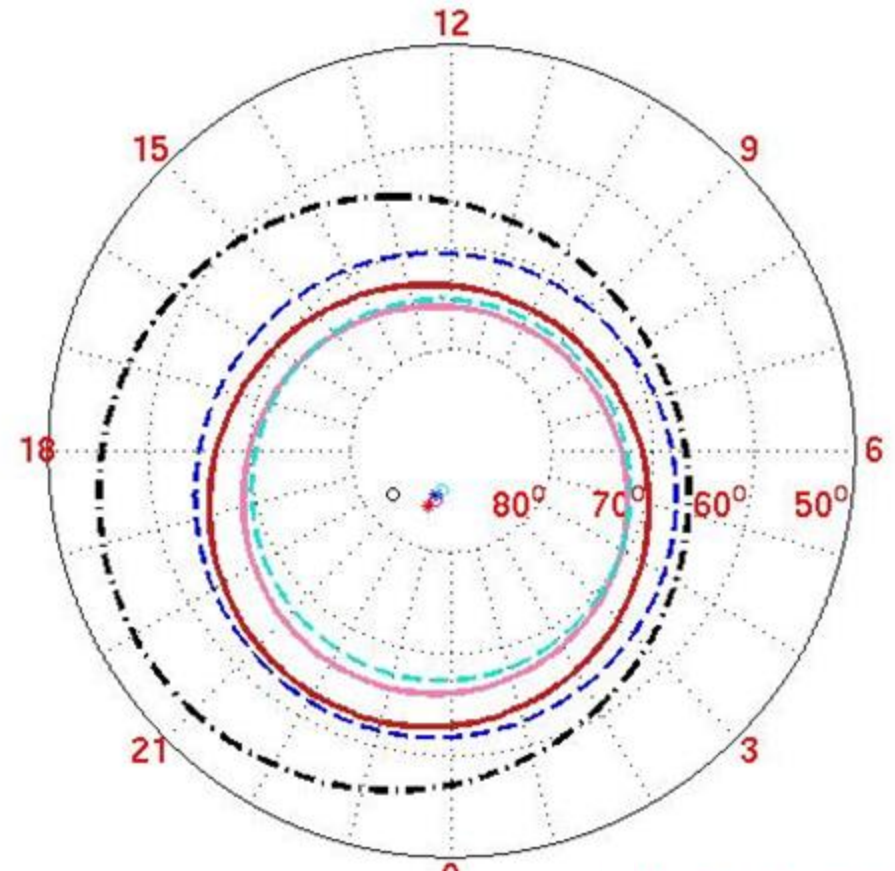
Half Integral Energy Flux

Peak e- Energy flux

Equatorward Boundary

Poleward Boundary

**-4<Bz<-3 DMSP F13 & Weimer 2005 Cross Track Ion drift**



Vy Zero Equatorward Boundary

Vy Peak CRB

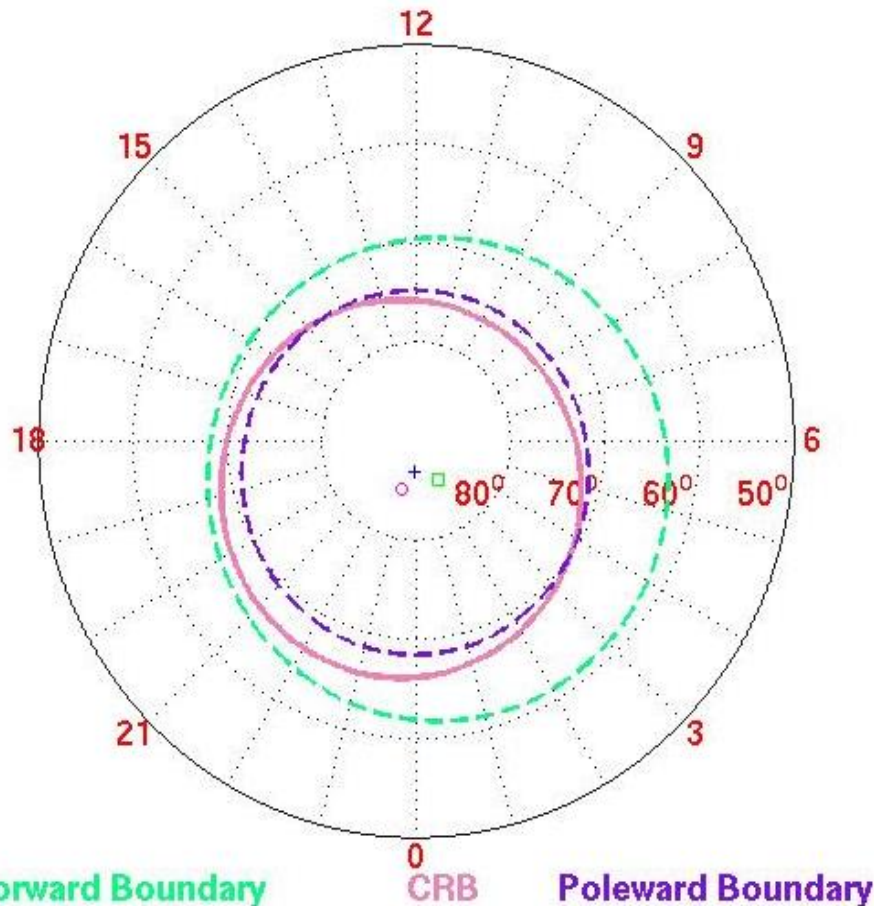
CRB Weimer

Vy Peak Weimer



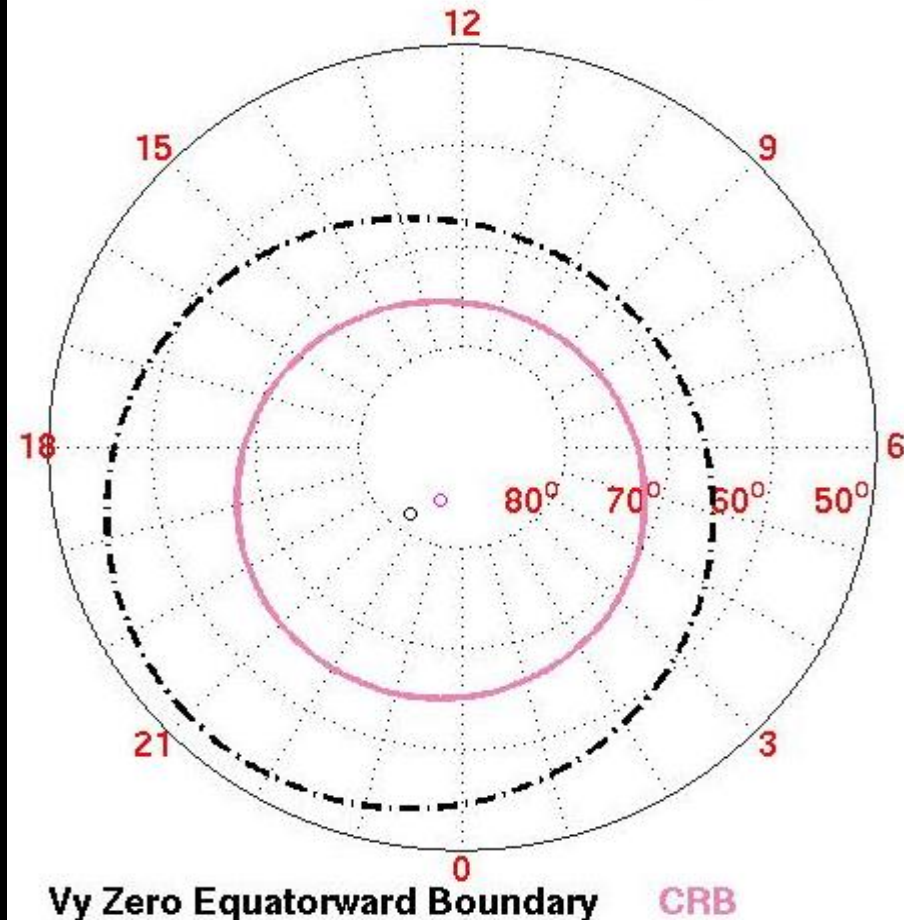
# Multiple day results: Jan-June 2005

-4<Bz<-3 DMSP F13 Particle Joule Heating



Area for particle Joule heating is bigger on the **dawn** side compared to the **dusk** side.

-4<Bz<-3 Total Joule heating

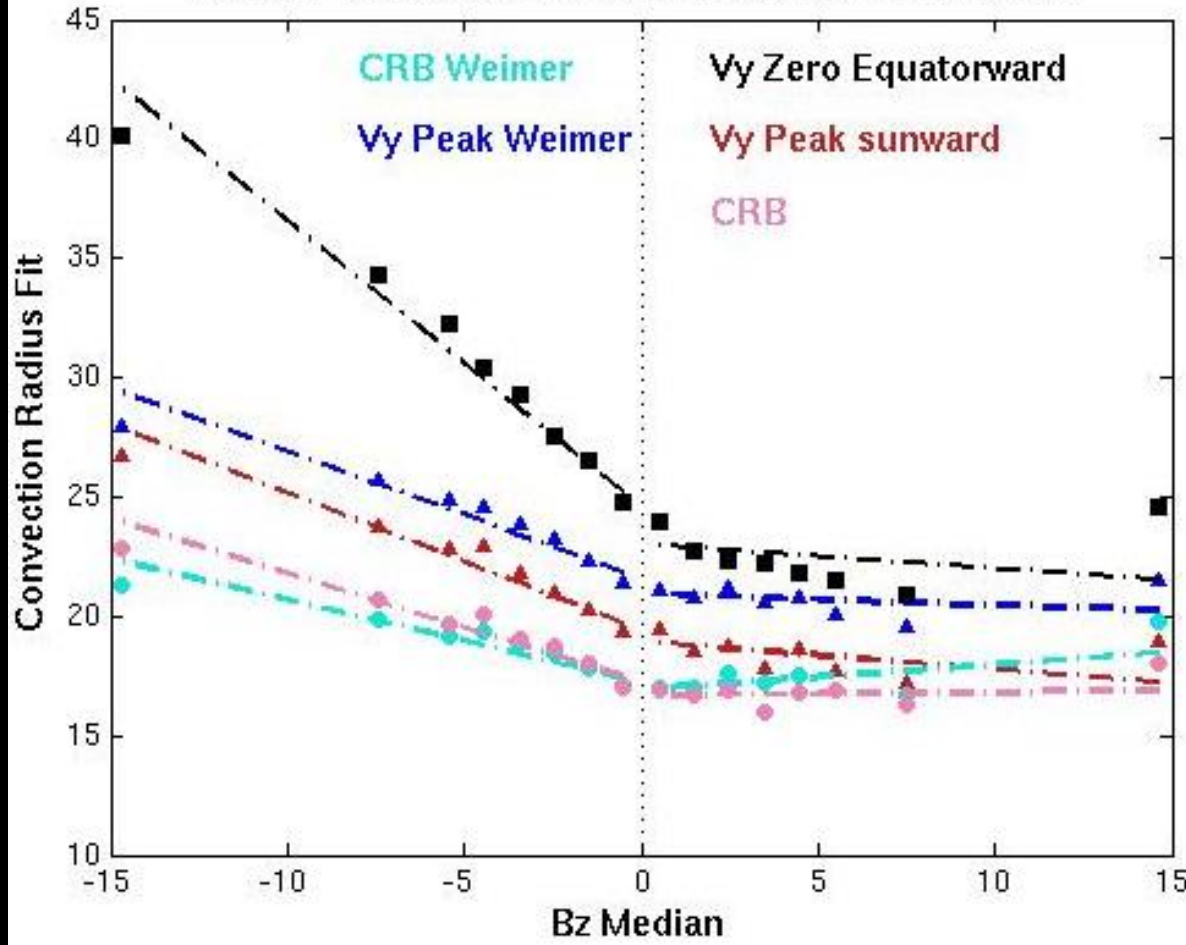


Area for Total Joule heating is bigger on the **dusk** side compared to the **dawn** side.



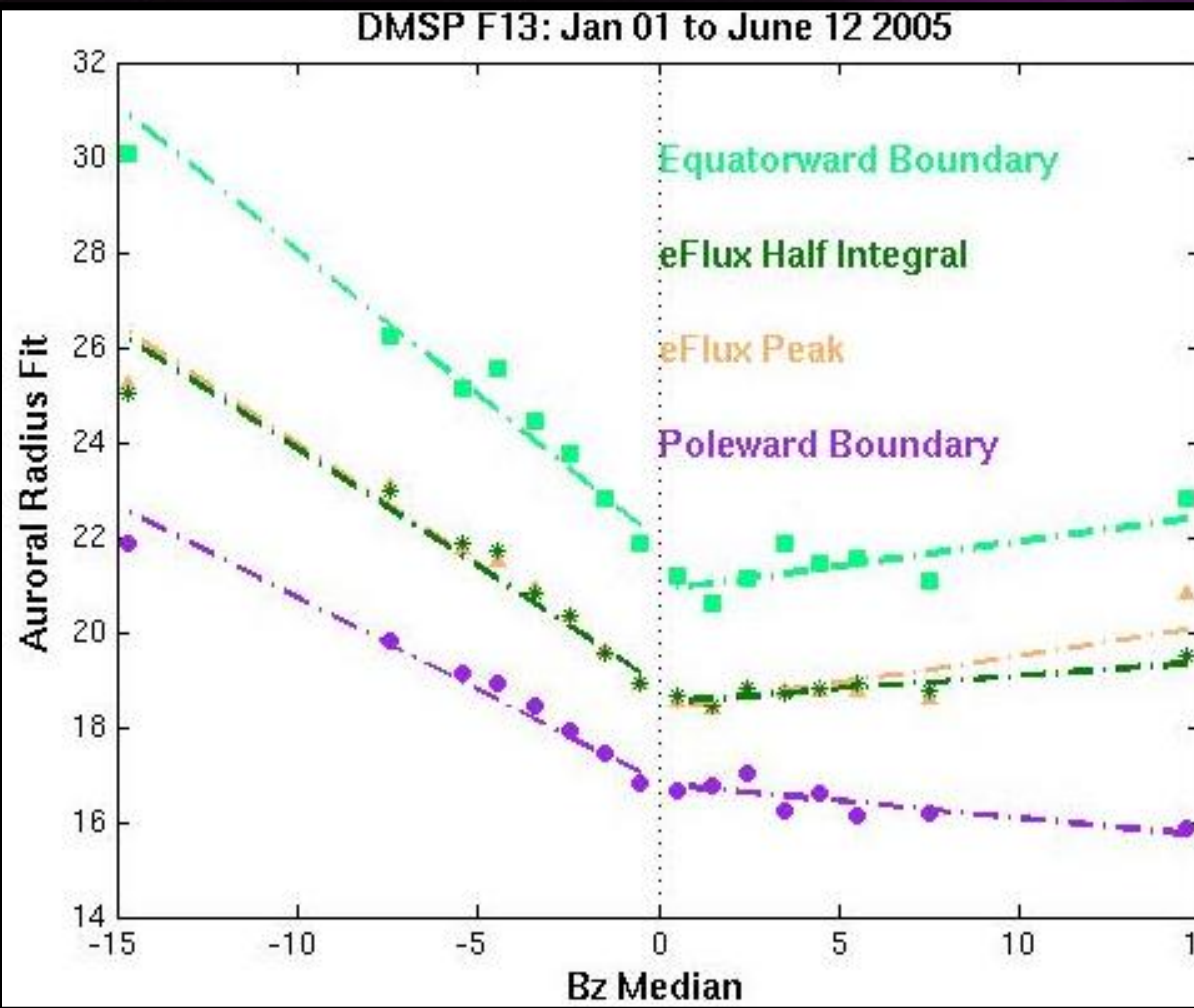
# Multiple day results: Jan-June 2005

DMSP F13 and Weimer 2005: Jan 01 to June 12 2005



Difference between radii of **CRB** and **Vy Zero Equatorward Boundary** increases as  $B_z$  decreases. *This means area for equatorward Joule heating increases as  $B_z$  becomes more negative. Also, as  $B_z$  becomes more negative, CRB radius increases and so does the area for poleward Joule heating.*

# Multiple day results: Jan-June 2005



Difference between the radii of **Equatorward** and **Poleward** Boundaries increases with the absolute value of  $B_z$ .

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Aha, finally ☺



## IV. Key Findings

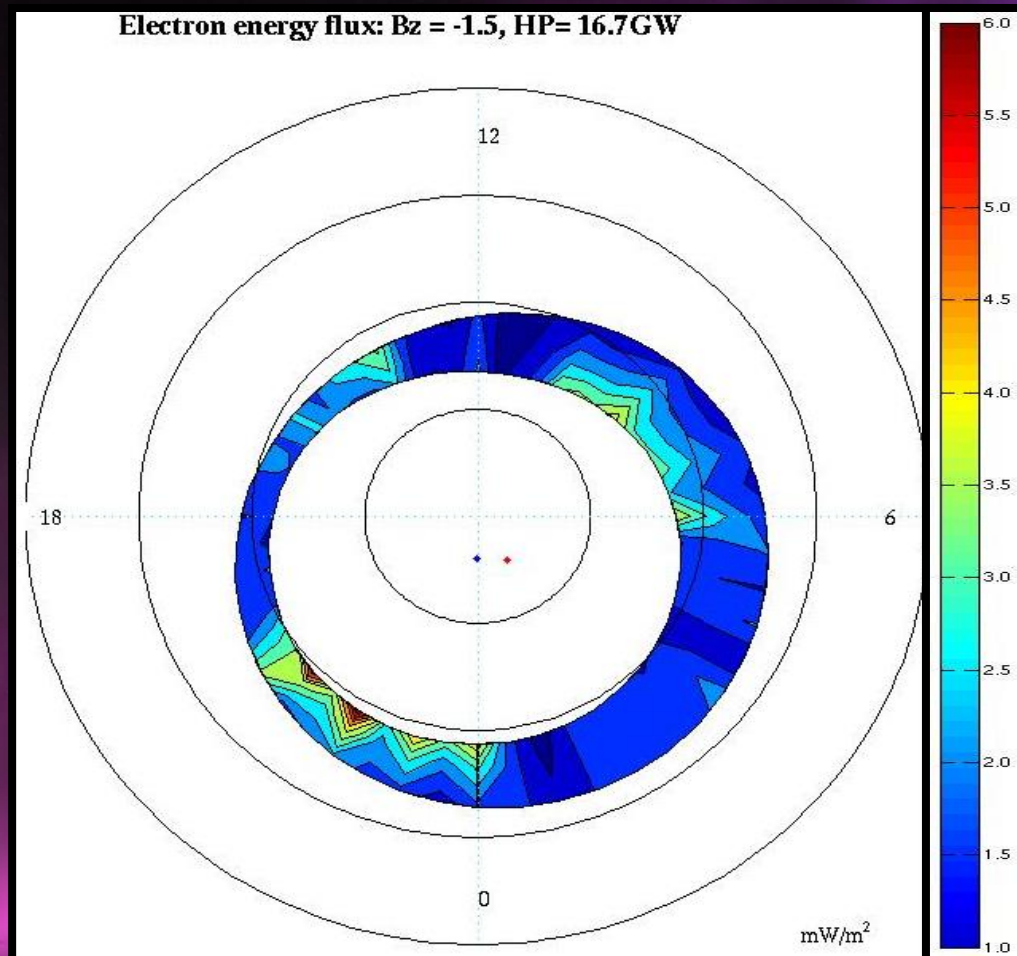
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## V. Conclusions



#### IV. Key Findings: Dawn Vs. Dusk

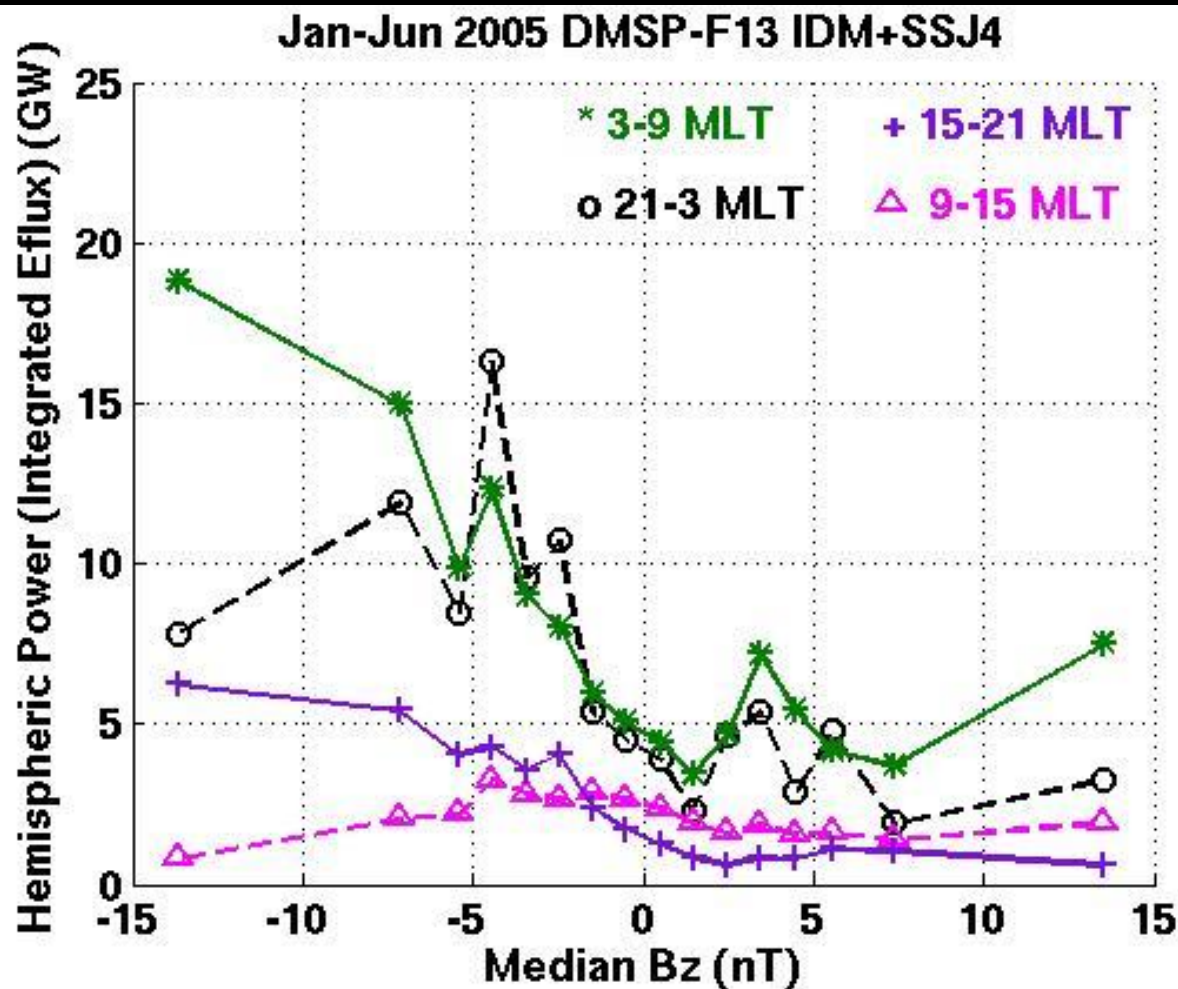
# Electron Energy Flux and Hemispheric Power



Area for Hemispheric Power is mostly bigger on the **dawn** side compared to the **dusk** side.

## IV. Key Findings: Dawn Vs. Dusk

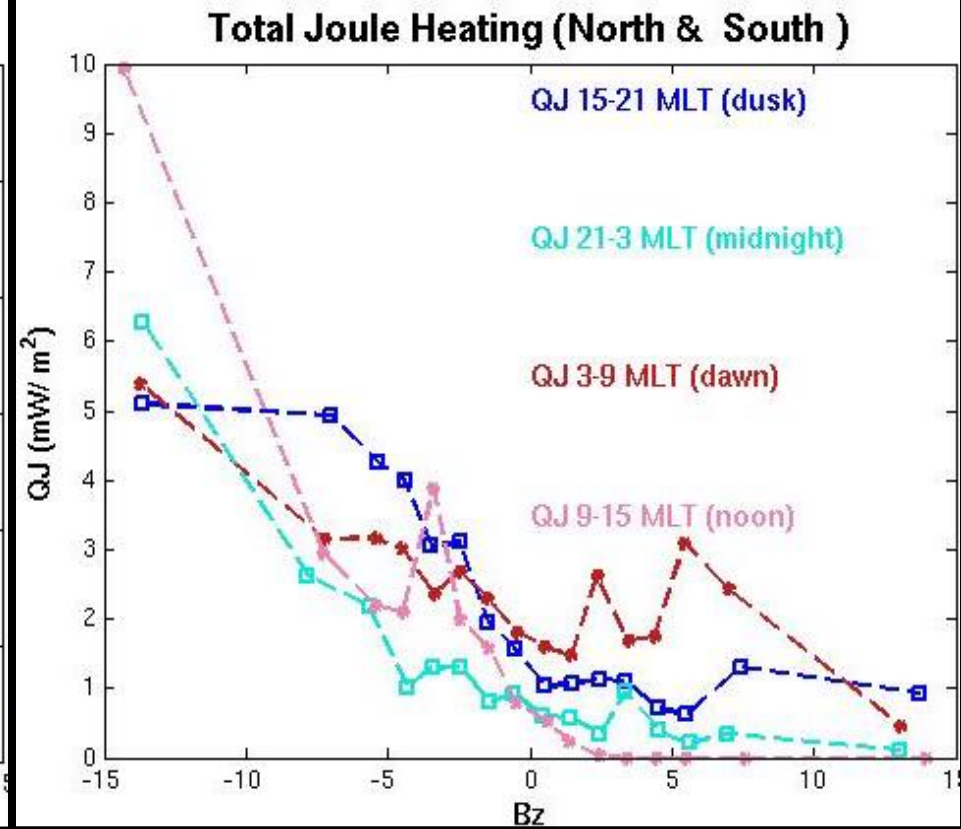
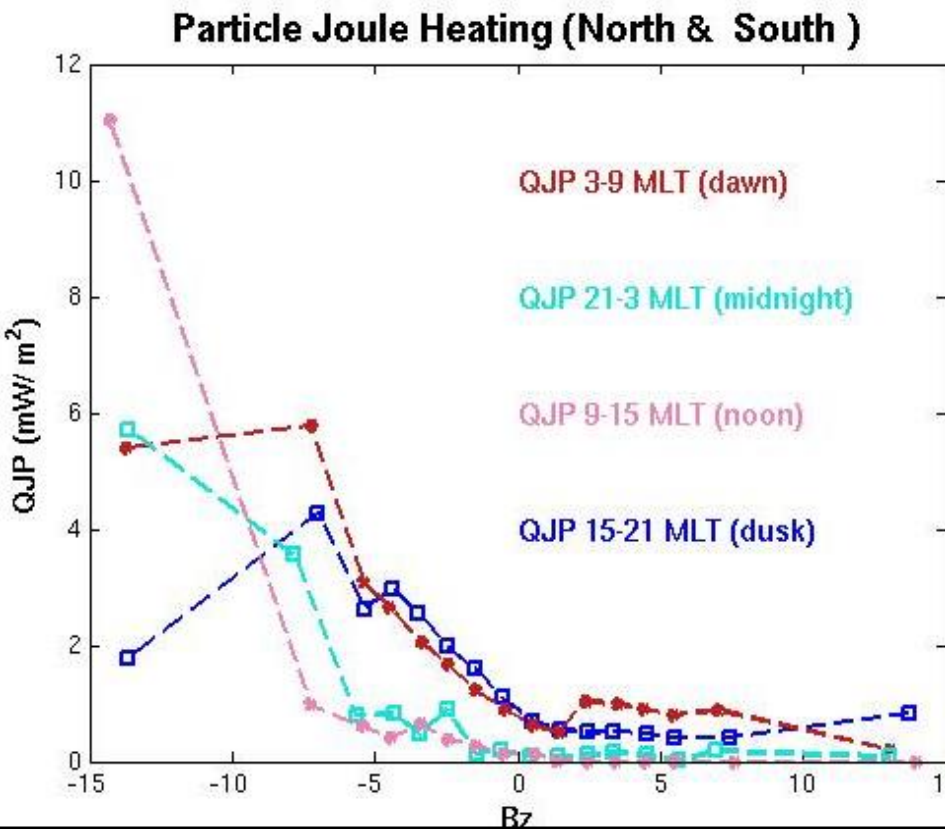
# Hemispheric power



HP is **highest** on the **dawn side**, and HP for the **dusk side** is relatively small.

#### IV. Key Findings: Dawn Vs. Dusk

## Average Particle & Total Joule heating



Average particle Joule heating on dawn side is almost equal to average particle Joule heating on dusk side.

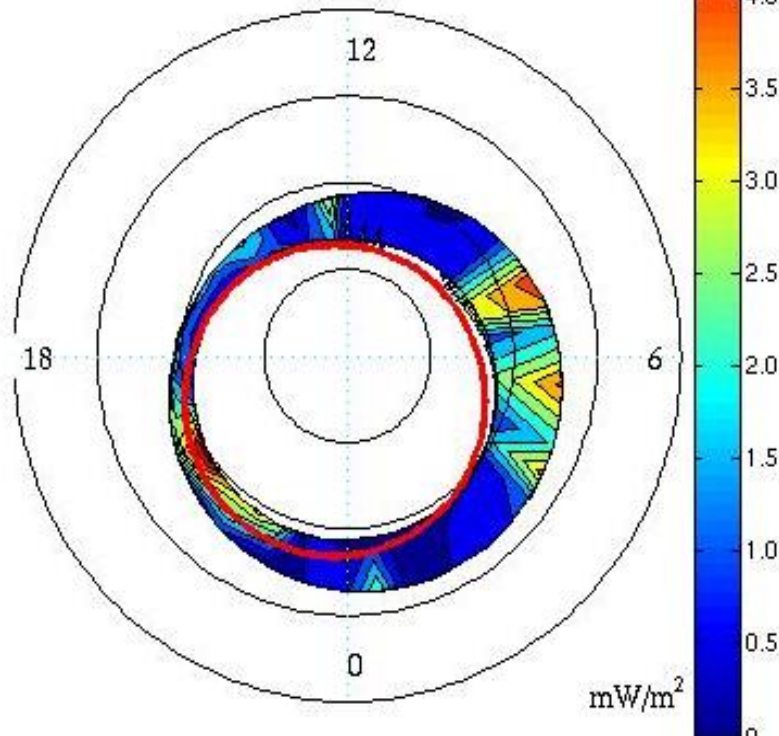
Average Joule heating for dawn side is greater than that for dusk side when  $B_z > 0$ , and vice versa for  $B_z < 0$ .



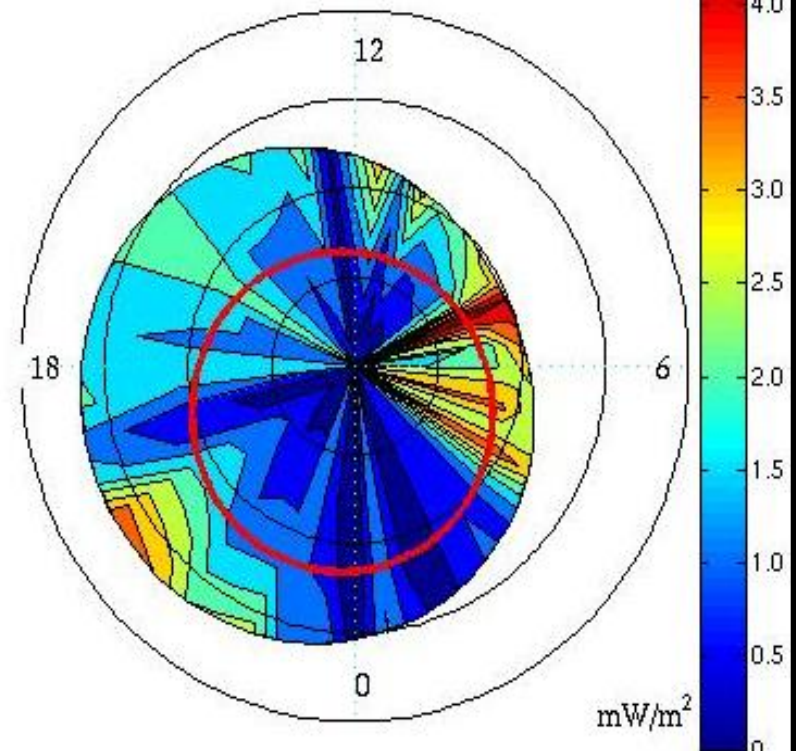
#### IV. Key Findings: Dawn Vs. Dusk

## Area for Particle & Total Joule heating

Particle Joule Heat:  $B_z = -1.5$ ,  $Q_{JP} = 15.1 \text{ GW}$

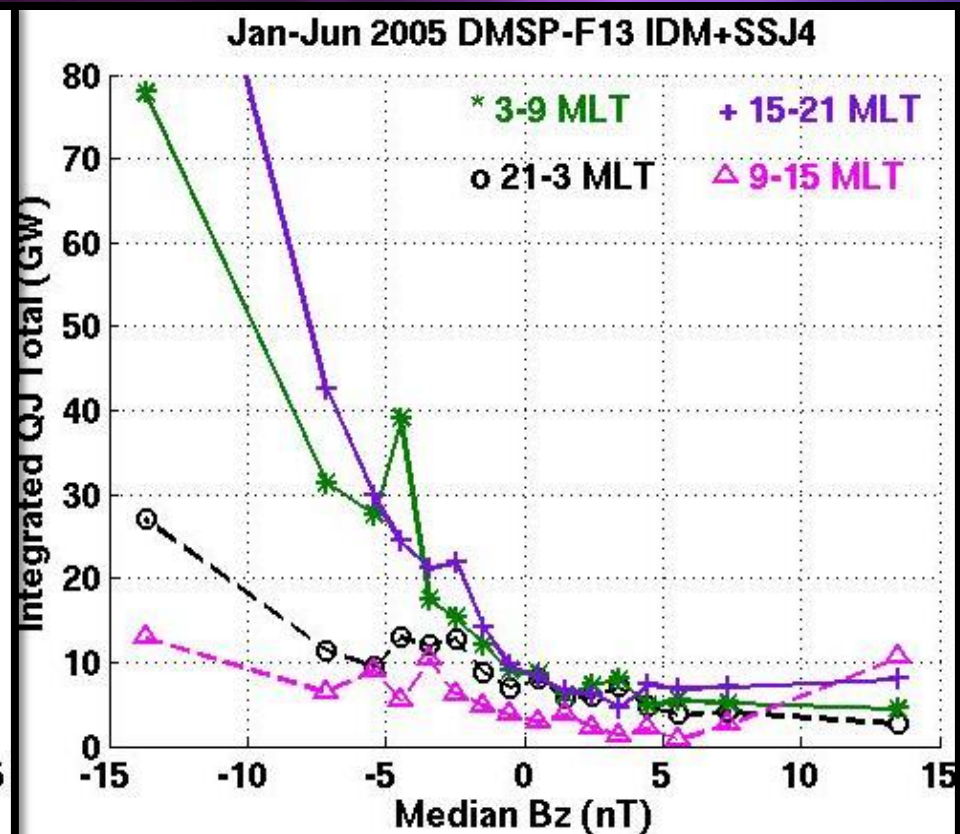
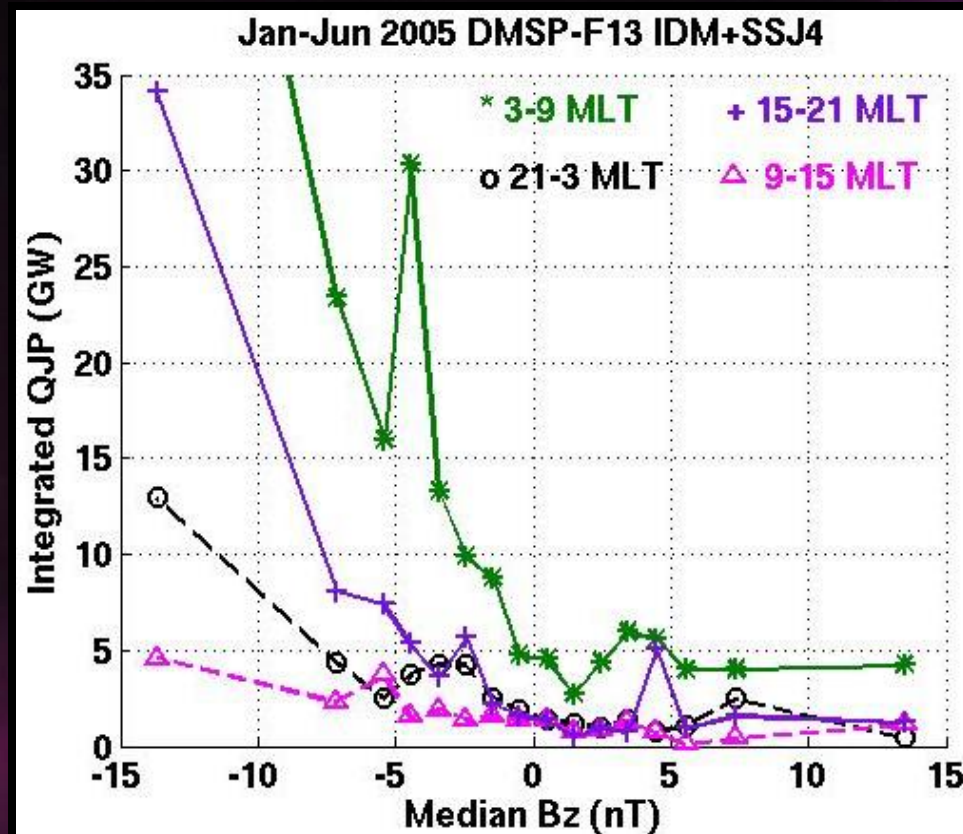


Total Joule Heat:  $B_z = -1.5$ ,  $Q_J = 40.2 \text{ GW}$



#### IV. Key Findings: Dawn Vs. Dusk

## Area Integrated Particle & Total Joule heating



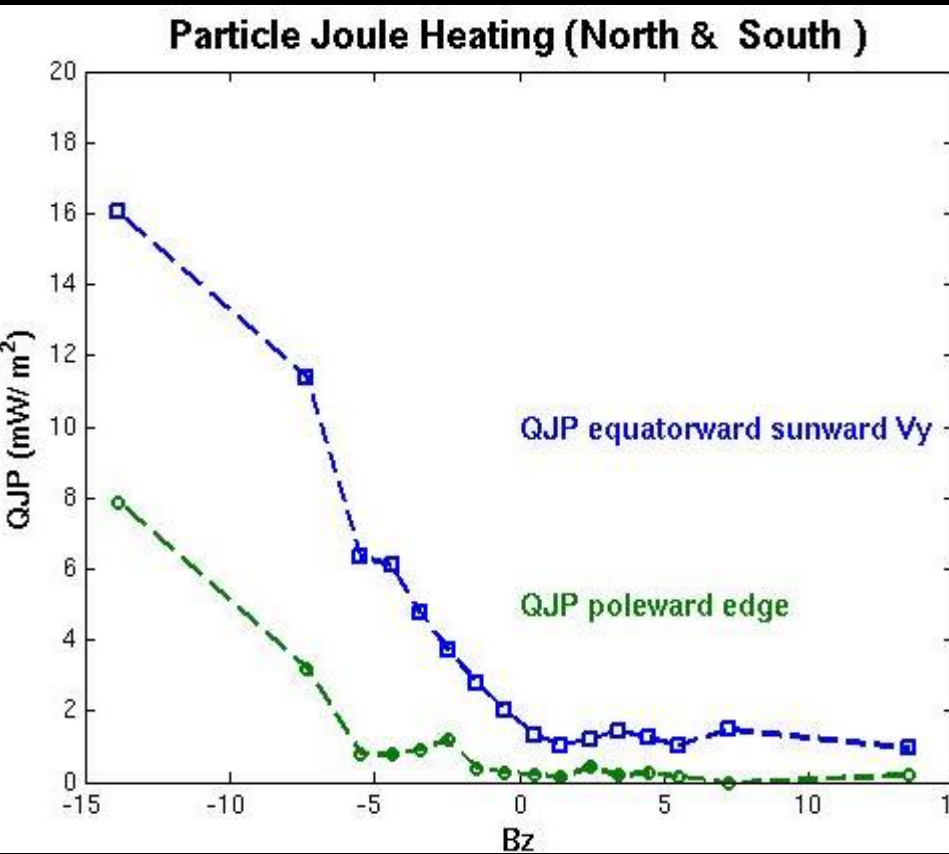
Integrated particle Joule heating is higher on the dawn side than on the dusk.

Integrated QJ on the dawn side is almost equal to QJ on the dusk side.

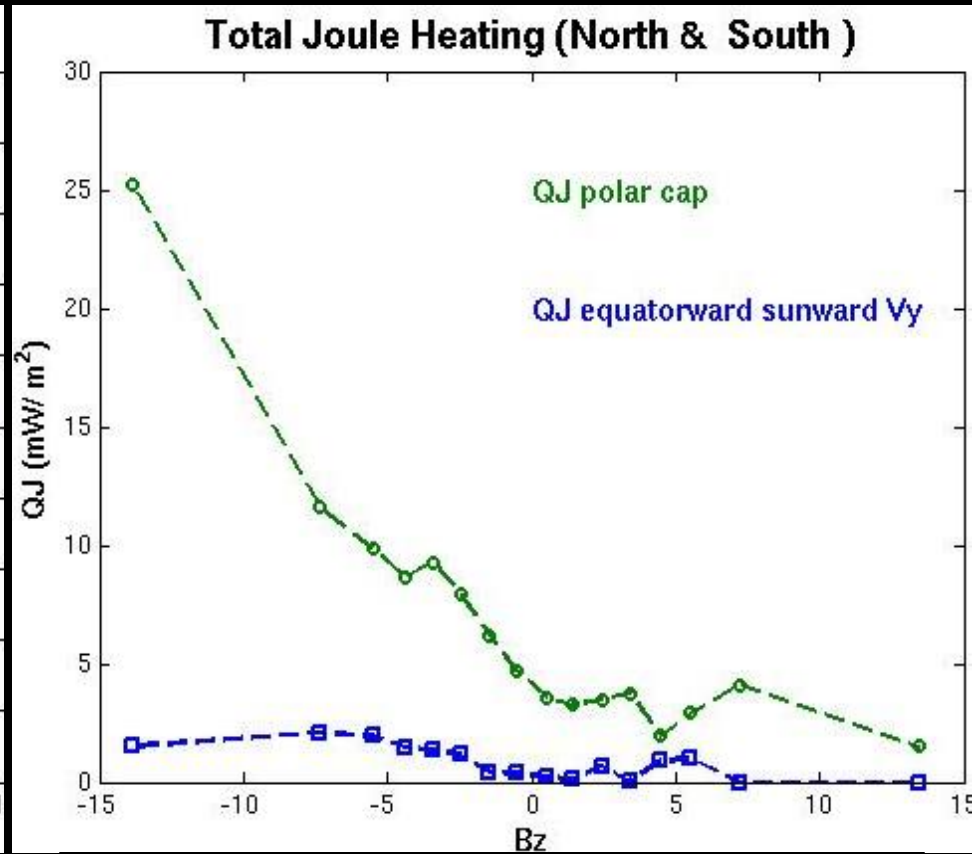


# IV. Key Findings: Equatorward Vs. Poleward

## Average Particle & Total Joule heating



Average equatorward particle Joule heating for sunward  $V_y$  is greater than average poleward particle Joule heating.

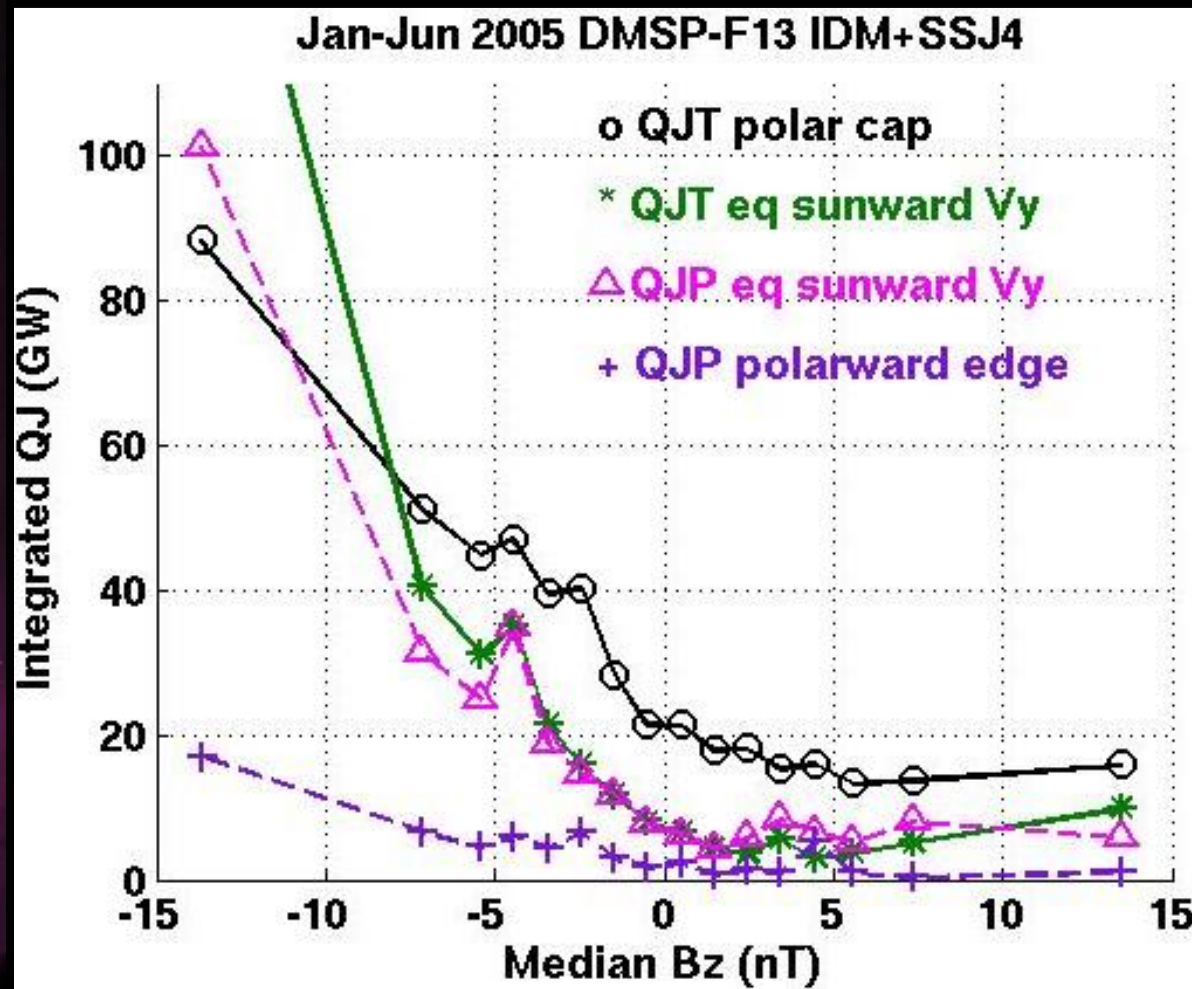


Average Joule heating in the polar cap for anti-sunward  $V_y$  is greater than the average equatorward Joule heating .



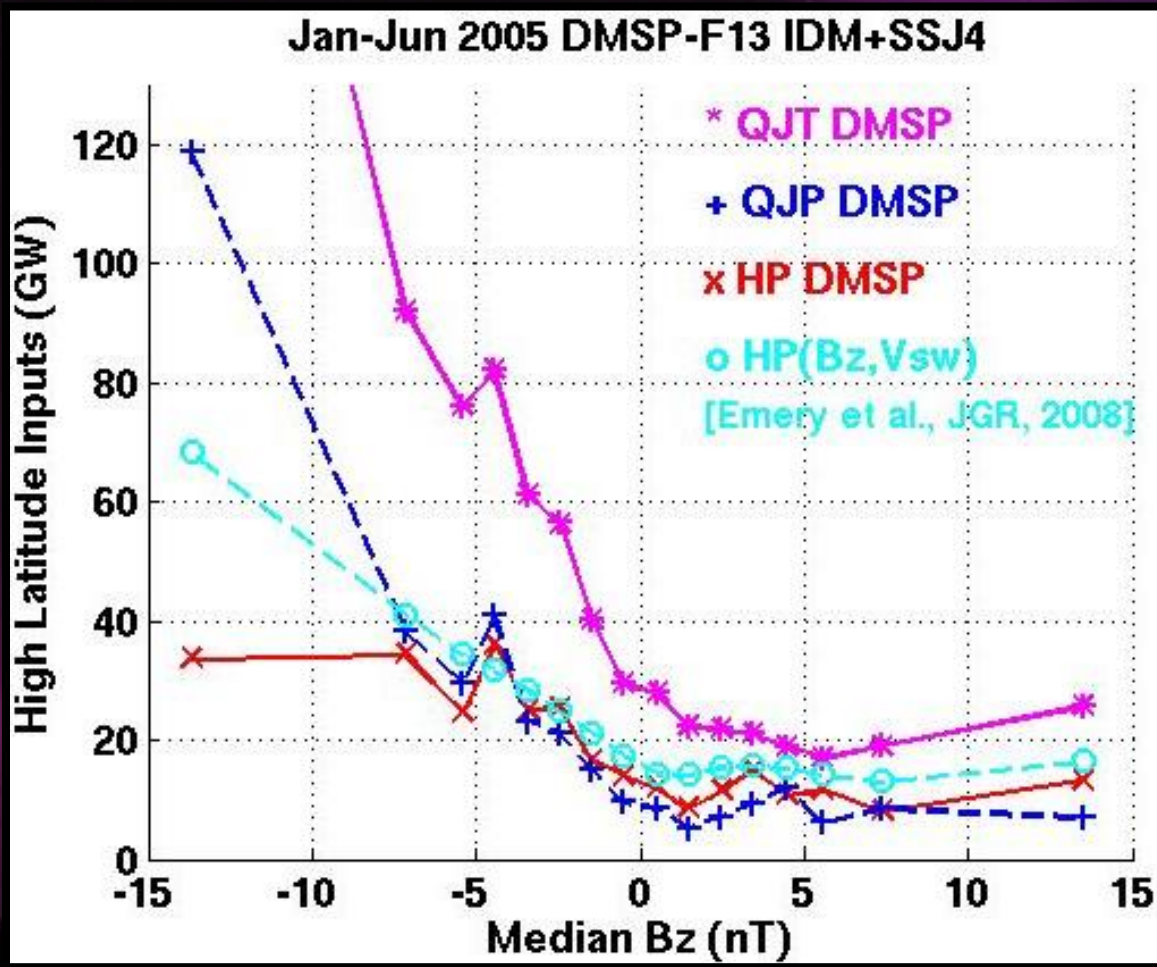
#### IV. Key Findings: Equatorward Vs. Poleward

## Area Integrated Particle & Total Joule heating



- Integrated QJ in the polar cap is more than integrated equatorward QJ for sunward  $V_y$ .
- Integrated equatorward QJP for sunward  $V_y$  is more than integrated poleward QJP.
- On the equatorward side, QJP is mostly equal to QJ, indicating the importance of the auroral Pedersen conductance.

# Integrated Joule heating Vs. Hemispheric Power



- Integrated total Joule heating is much higher than Hemispheric Power.
- Integrated Particle Joule heating has almost the same magnitude as Hemispheric Power .



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PARTICLE JOULE HEATING (QJP)	TOTAL JOULE HEATING (QJ)
<p><b>Average QJP:</b> Mostly, Dawn side <math>\approx</math> Dusk side.</p> <p><b>Integrated QJP:</b> Dawn side <math>&gt;</math> Dusk side since dawn area is larger than dusk area</p>	<p><b>Average QJ:</b> when <math>B_z &gt; 0</math>, Dawn side <math>&gt;</math> Dusk side when <math>B_z &lt; 0</math>, Dawn side <math>&lt;</math> Dusk side</p> <p><b>Integrated QJ:</b> Dawn side <math>\approx</math> Dusk side although dusk area is mostly greater than dawn area</p>
<p><b>Average QJP:</b> Equatorward <math>&gt;</math> Poleward.</p> <p><b>Integrated QJP:</b> Equatorward <math>&gt;</math> Poleward. <math>QJP \approx HP</math></p>	<p><b>Average QJ:</b> Poleward (anti-sunward <math>V_y</math>) <math>&gt;</math> Equatorward (sunward <math>V_y</math>)</p> <p><b>Integrated QJ:</b> <math>QJ \gg HP</math> Poleward <math>&gt;</math> Equatorward. On the equatorward side, <math>QJP \approx QJ</math>.</p>
<p>Area for QJP and HP increases as the absolute value of <math>B_z</math> increases.</p>	<p>Area for QJ, especially equatorward QJ, increases as <math>B_z</math> becomes more negative</p>



# Acknowledgements

- Thank you to Dr. Barbara Emery for her time, support & guidance.
- Thank you to Dr. Astrid Maute for her help and encouragement.
- Thank you to LASP and HAO for giving me this amazing research opportunity.
- Thank you to Marty, Erin and all REU students. It was wonderful meeting you all!

Thank you for your attention!

Questions?

