MAUEN

Mars Atmosphere and Volatile EvolutioN (MAVEN) Mission

> Solar Wind Electron Analyzer (SWEA) MAVEN Science Community Workshop Dec 02, 2011

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# **IRAP / UCB-SSL Collaboration**





#### IRAP, Toulouse

- Analyzer
- MCP
- Anode
- HVPS
- SSL, Berkeley
  - Pedestal
  - Digital / FPGA
  - LVPS

(same as for STEREO)

### **Instrument Performance**





- Energy range: 5 eV to 5 keV
- FOV: 360° x 120° for energies up to 1.6 keV
- Angular resolution: 22.5° x 20°
- Energy resolution (ΔE/E): 12% below 50 eV, increasing to 18%
- Geometric factor: 0.009 cm<sup>2</sup> ster eV/eV
- Energy fluxes: 10<sup>3</sup> to 10<sup>9</sup> eV/cm<sup>2</sup>-s-ster-eV





 $\Delta E/E$  of the sensor at low energy can be controlled by varying the voltage bias  $U_0$  between the internal and external grids:

$$\Delta E/E_{inc} = (\Delta E/E)_{or} / (1 + U_0/E_{inc})$$

where  $(\Delta E/E)_{or}$  is the energy resolution for zero bias (0.18) and  $E_{inc}$  is the incident energy of the electron

Thus for 1.5x finer energy resolution (and 2x smaller GF), we apply

$$U_0 = E_{inc}/2$$

Works for apoapsis and side segments in all orbit scenarios. Works for Sun-Velocity mode at periapsis (~50% of orbits). Does not work for deep dips.



## **Data Products**



- SWEA provides a single science data product to the PFDPU
  - Counts per accumulation interval for each of the 16 anodes (as a function of analyzer and deflector sweeps)
  - Complete measurement sequence takes 2 seconds.
- PFDPU computes three data products, with cadence depending on altitude

<b>Solar Wind Mode</b> altitude > 500 km	<ul> <li>Energy spectra (64E) every 8 sec</li> <li>Pitch angle distr. (16A x 16E) every 8 sec</li> <li>3D distributions (80A x 16E) every 64 sec</li> </ul>
<b>lonosphere Mode</b> altitude < 500 km	<ul> <li>Energy spectra (64E) every 2 sec</li> <li>Pitch angle distr. (16A x 16E) every 2 sec</li> <li>3D distributions (80A x 16E) every 64 sec</li> </ul>



#### Magnetic Topology & Plasma Regime

Crustal Magnetospheres/Cusps



• Draped Field Lines













#### **Electron Impact Ionization**

- Magnetic Pileup Region
- Ionosphere







#### Mars Express

Photo-ionization of  $CO_2$  by solar  $h_V$  @ 304 Å



Escape associated with heavy ions (M > 16)