

# Comparison of the 2005 Weimer and HAO Empirical High Latitude Models of Energy Transfer in terms of Poynting Flux

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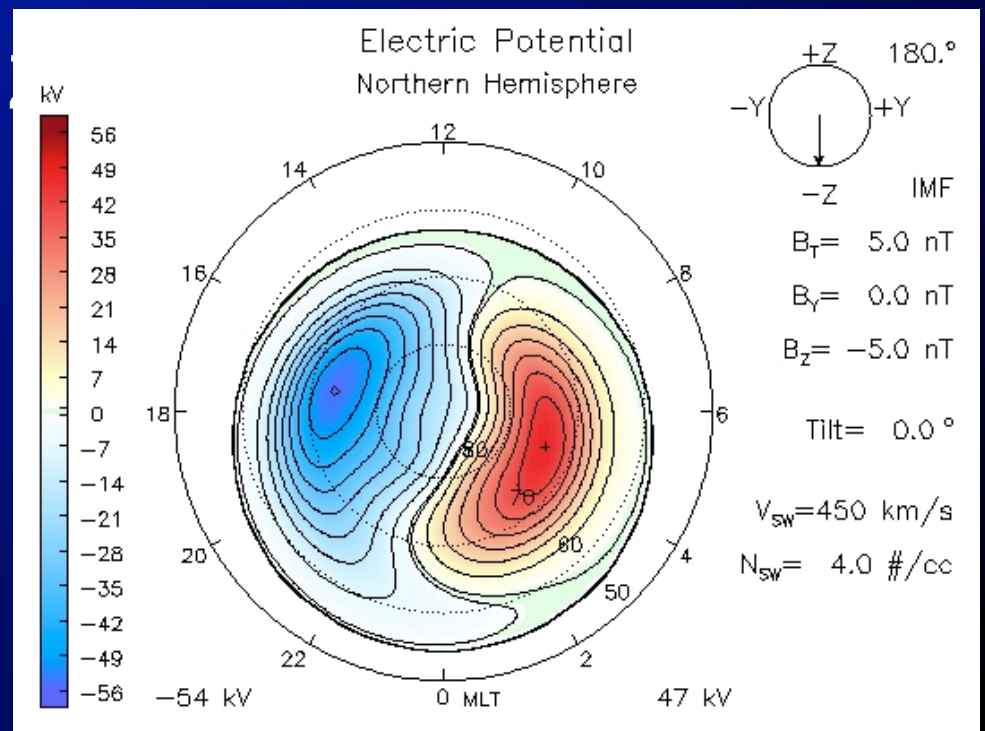
Mentors: Astrid Maute, Yue Deng, Art  
Richmond

# Background

- Two Models
  - Weimer 05
  - High Altitude Observatory (HAO)
- Predictions
  - Electric/Magnetic Potential
  - Electric/Magnetic Field
  - Poynting Flux
  - Joule Heating

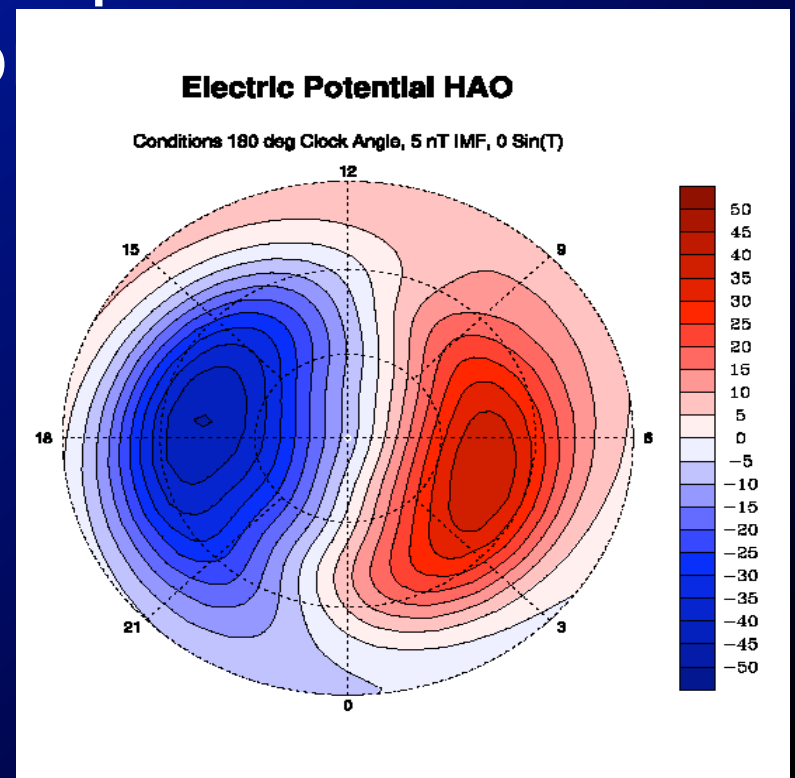
# Weimer 05

- Developed by Dr. Daniel Weimer in 2005
- At the time, Mission Research Corporation
- Models made in 1996 and 2001
- Dynamics Explorer
- IDL



# HAO

- Developed by Astrid Maute and Arthur Richmond
- National Center for Atmospheric Research: High Altitude Observatory
- Dynamics Explorer 2
- FORTRAN



# Why Model Comparison?

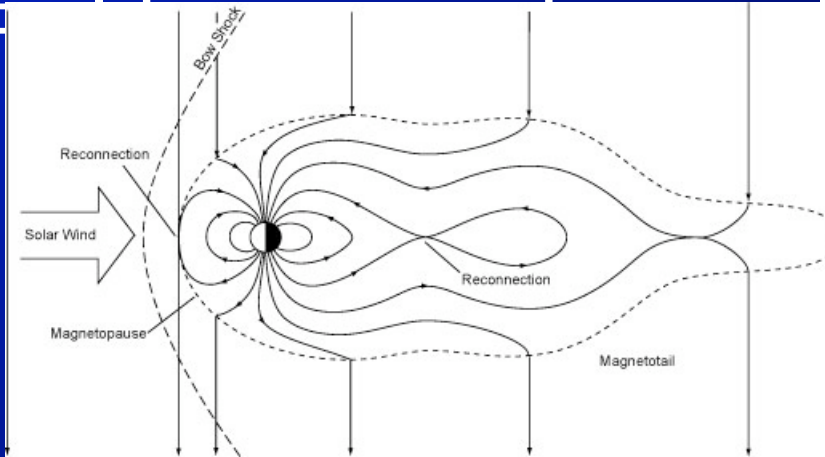
- Check for new model
  - Debugging
  - Biases
  - General behavior
- Check for old model
  - Still viable
  - Debugging
  - Biases
- Better Option
- Understanding



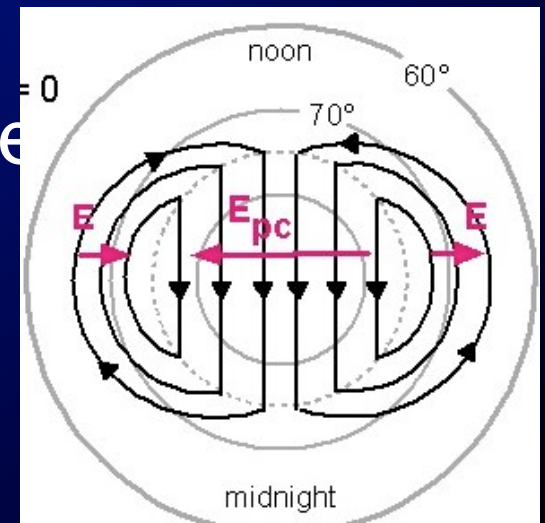
# Why do we have E Fields?

- Magnetic Reconnection causes Geomagnetic Field lines to interact with IMF
- IMF feels a electric field
- Geomagnetic Lines are equipotential; feel the field

• F



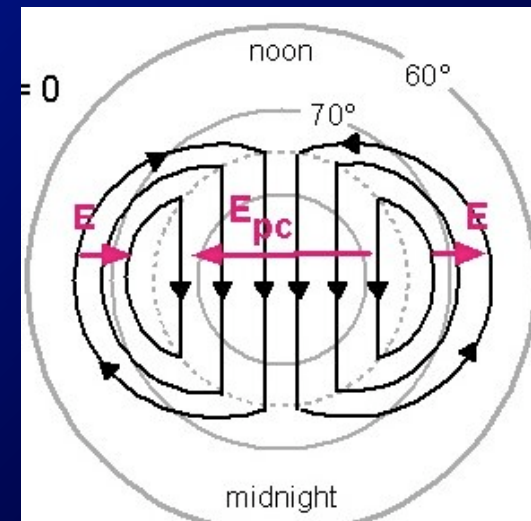
osier line



# Why do we have B Fields?

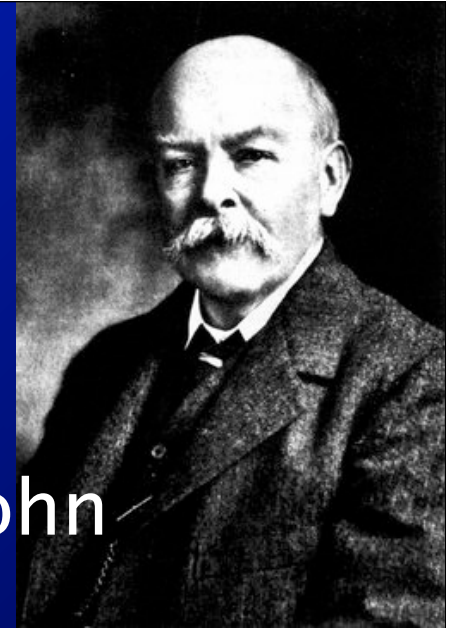
- E fields cause converges and diverges; currents form
- Using Ampere's Law, you got induced magnetic fields

$$\oint \mathbf{B} \cdot d\mathbf{s} = \mu_0 \epsilon_0 \frac{d\Phi_{\mathbf{E}}}{dt} + \mu_0 i_{enc}$$
$$\nabla \times \mathbf{B} = \mu_0 \epsilon_0 \frac{\partial \mathbf{E}}{\partial t} + \mu_0 \mathbf{j}_c$$



# Poynting Flux

- Representation of energy flux
- Independently co-discovered by John Henry Poynting, Oliver Heaviside
- Joule Heating can be estimated by Poynting's Theorem



$$\frac{\partial u}{\partial t} + \nabla \cdot \mathbf{S} = -\mathbf{J} \cdot \mathbf{E}$$

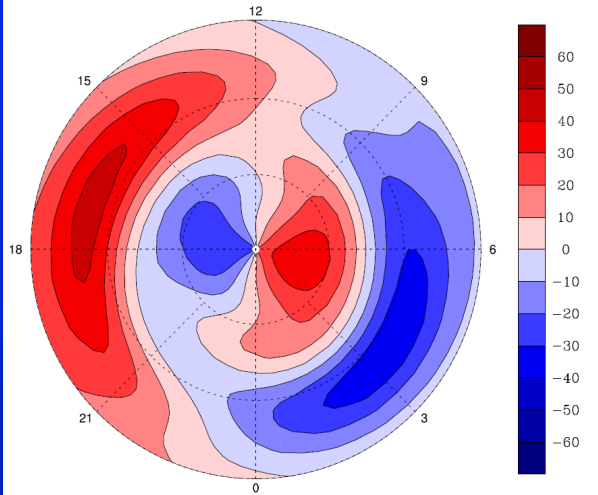
$$\mathbf{S} = \frac{\mathbf{E} \times \mathbf{B}}{\mu_0}$$



# Electric Field (EF)

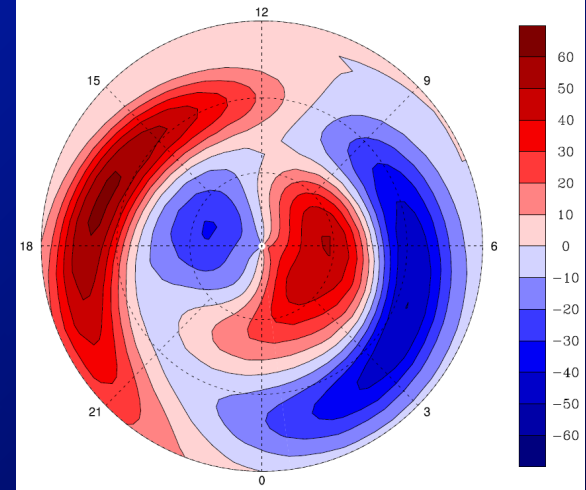
### Electric North Field HAO

Conditions 180 deg Clock Angle, 5 nT IMF, 0 Sin(T) mV/m



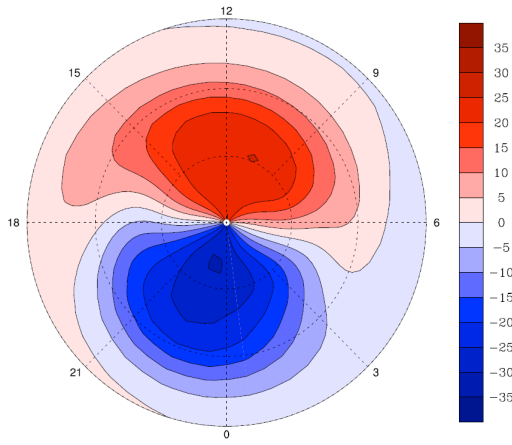
### Electric North Field Weimer

Conditions 180 deg Clock Angle, 5 nT IMF, 0 Sin(T) mV/m



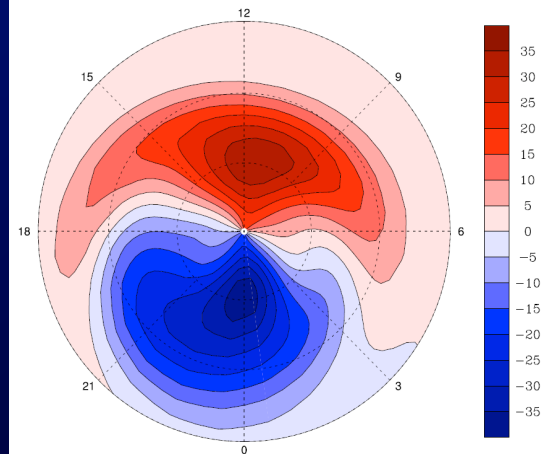
### Electric East Field HAO

Conditions 180 deg Clock Angle, 5 nT IMF, 0 Sin(T) mV/m



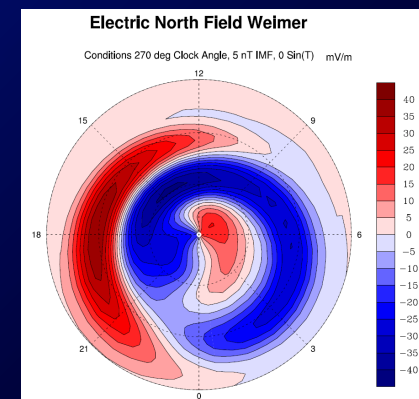
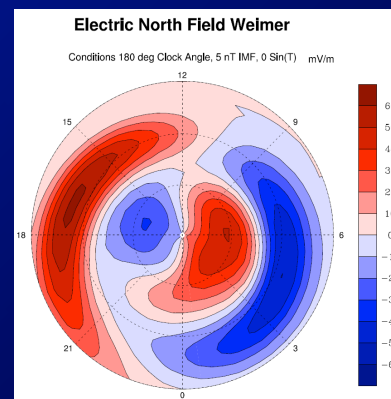
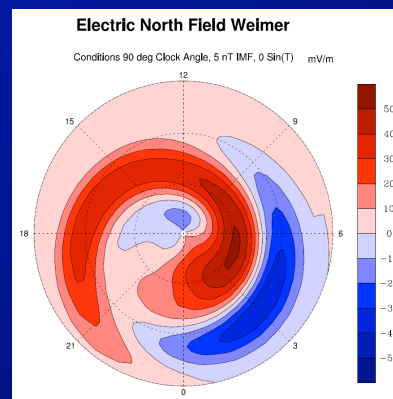
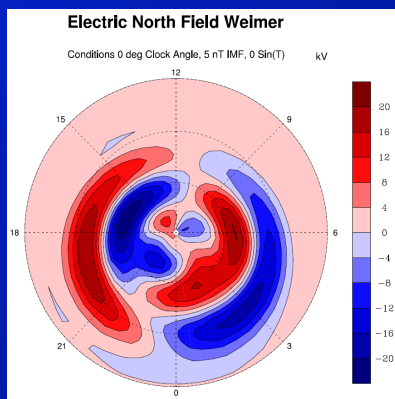
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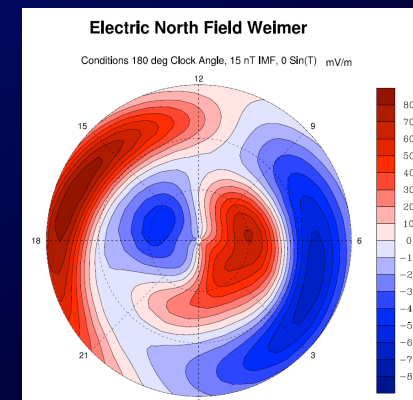
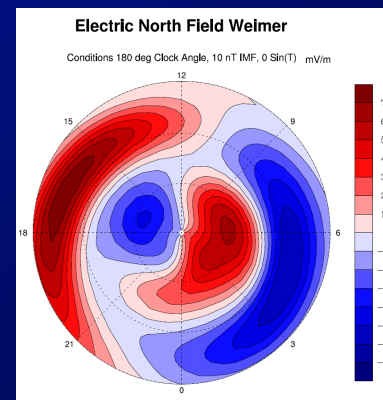
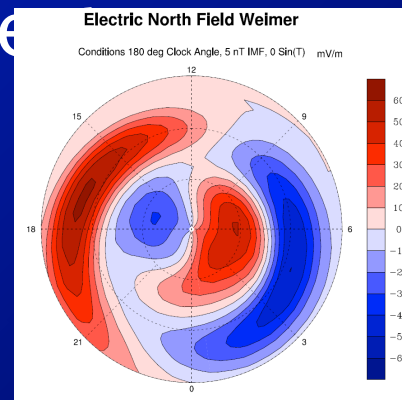
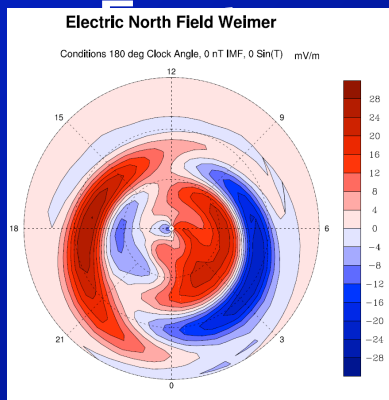
# EF IMF Clock Angle Summery

- Weimer is consistently stronger than HAO
- Both show similar patterns
- Patterns are those that are expected
- Difference plot values not too large



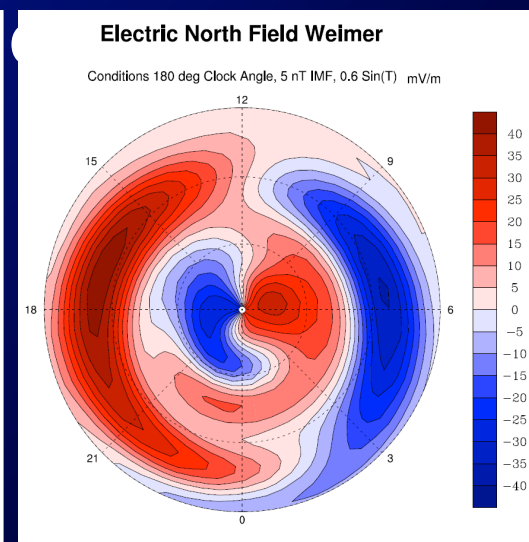
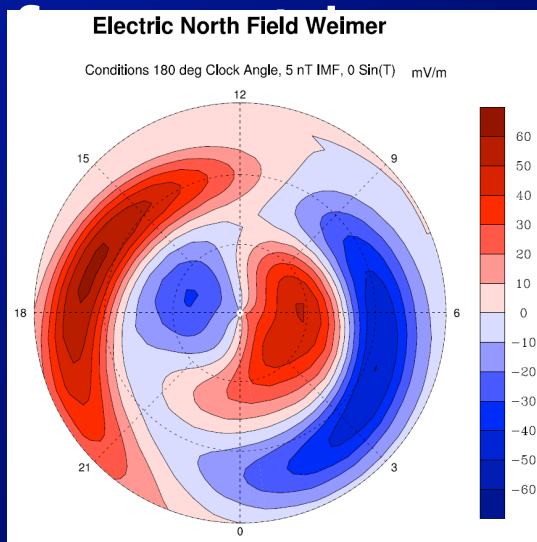
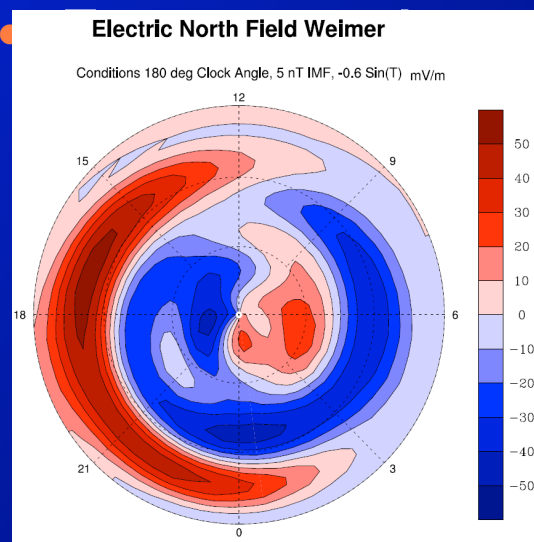
# EF IMF Strength Summery

- Weimer Stronger peak values than HAO
- Models closer at 5 nT and 10 nT than 15 nT
- 0 nT patterns/strengths quite different
- More variation in Northern vector than



# EF Season Summery

- Season causes great changes in E field
- Rotation around Midnight/noon (MN) line
- $\text{Sin}(T) = -0.6$  has larger differences than  $\text{Sin}(T) = 0.6$



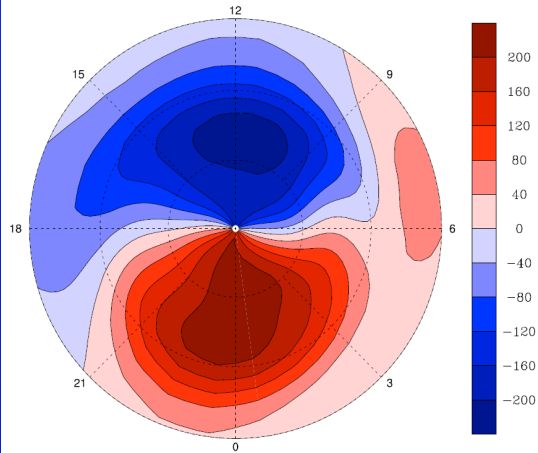
# EF Summery

- Weimer consistently stronger than HAO
- Though there are areas of great differences, overall they are quite similar
- Pattern variations between the two models show up in a lot of the plots
- Some strength differences

# Magnetic Field Perturbations (BF)

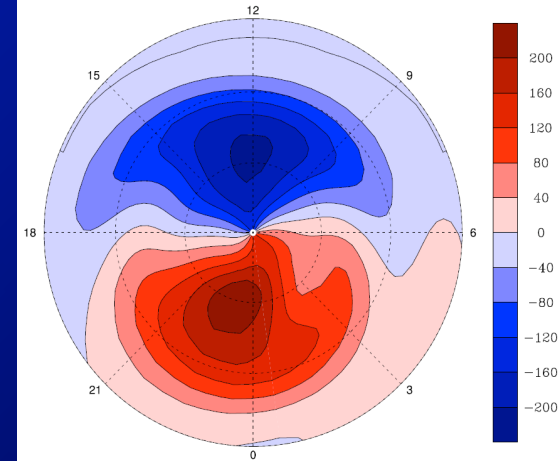
## Magnetic North Field HAO

Conditions 180 deg Clock Angle, 5 nT IMF, 0 Sin(T) cTm



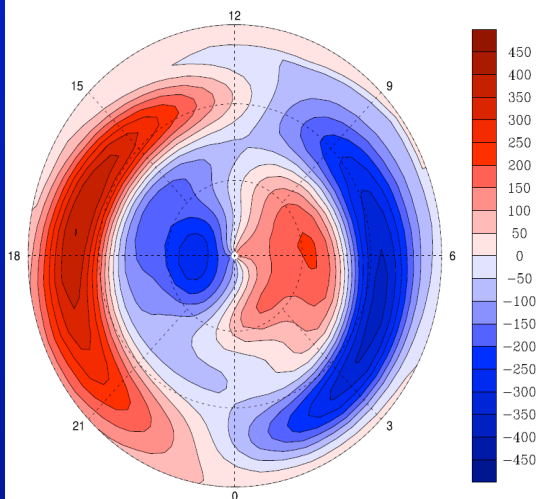
## Magnetic North Field Weimer

Conditions 180 deg Clock Angle, 5 nT IMF, 0 Sin(T) cTm



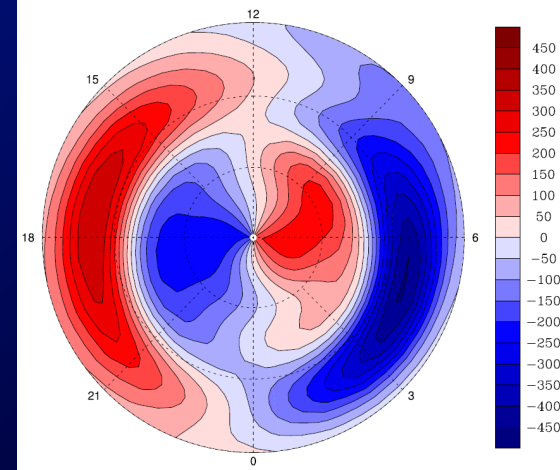
## Magnetic East Field Weimer

Conditions 180 deg Clock Angle, 5 nT IMF, 0 Sin(T) cTm



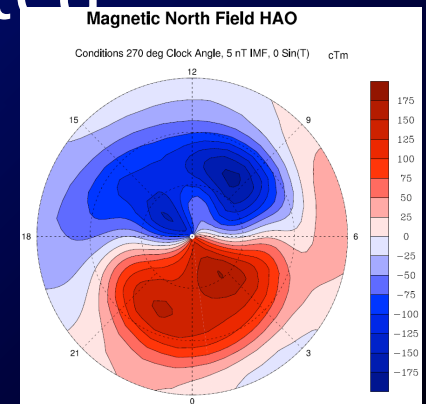
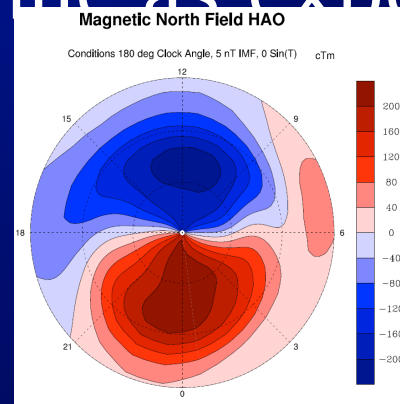
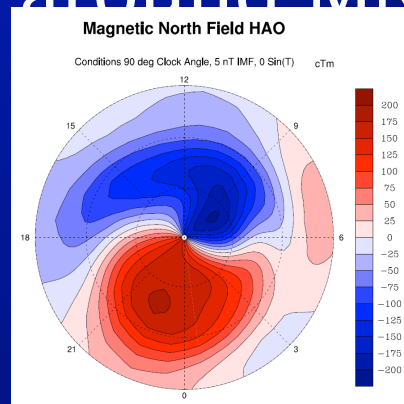
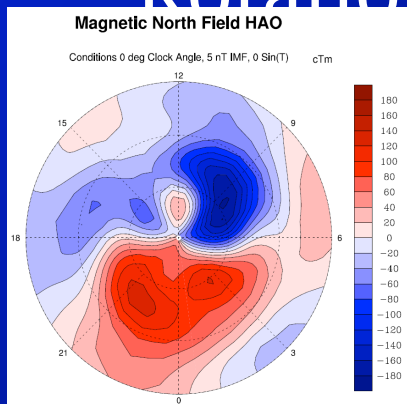
## Magnetic East Field HAO

Conditions 180 deg Clock Angle, 5 nT IMF, 0 Sin(T) cTm



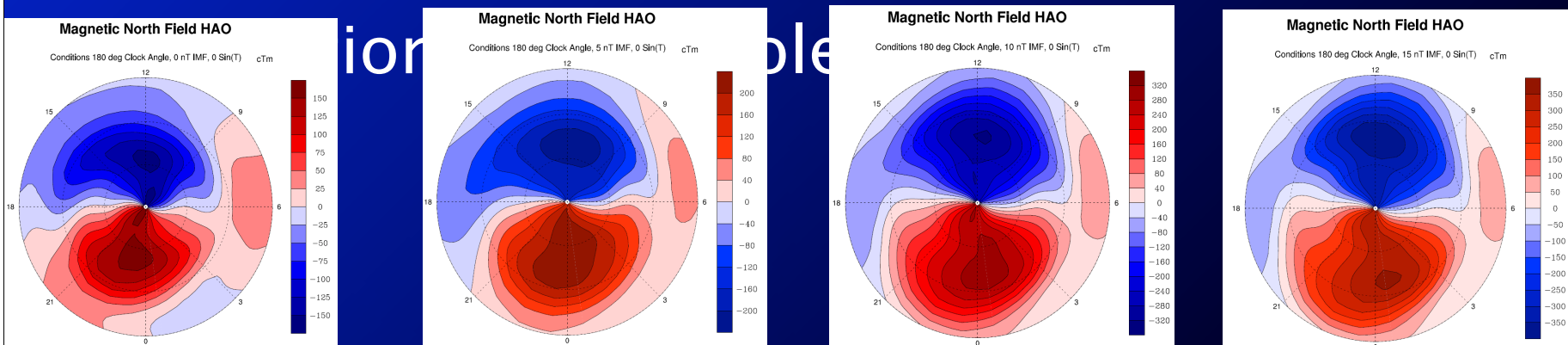
# BF IMF Clock Angle Summery

- HAO peaks always stronger than Weimer
- $180^\circ$  is strongest of all the clock angles
- $0^\circ$  is weakest and has the greatest difference
- Rotation around MN line as expected



# BF IMF Strength Summary

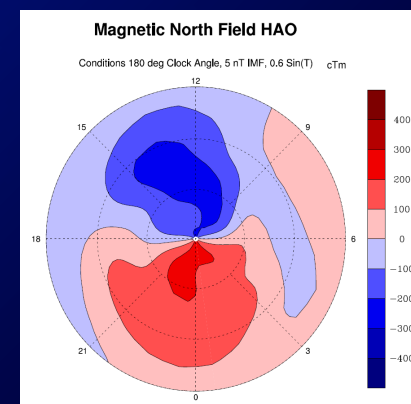
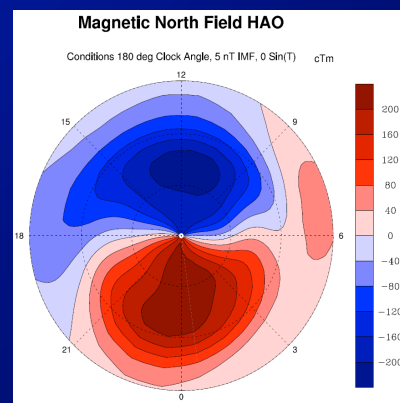
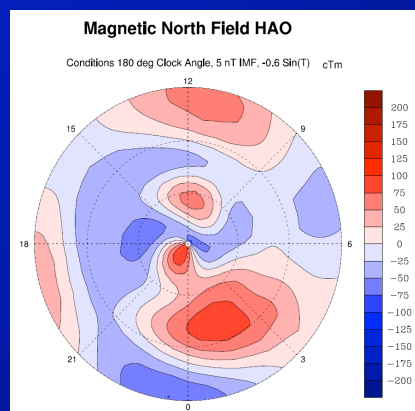
- HAO peaks always stronger than Weimer
- Some variation in pattern, but mostly strength
- Same patterns, with some expansion
- As IMF strength goes up, the differences in strength/pattern go up





# BF Season Summery

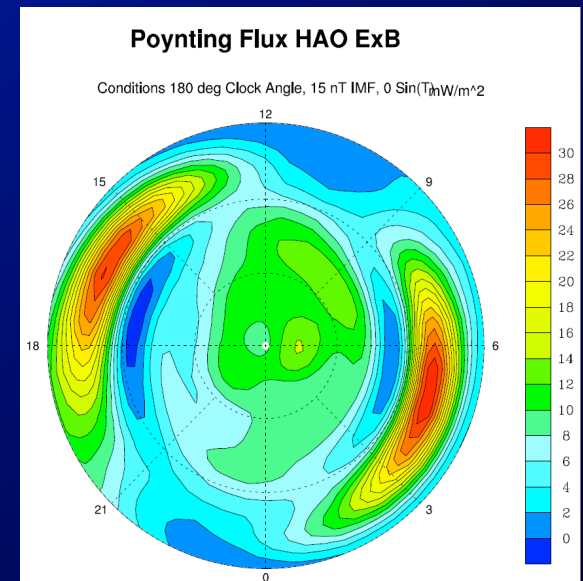
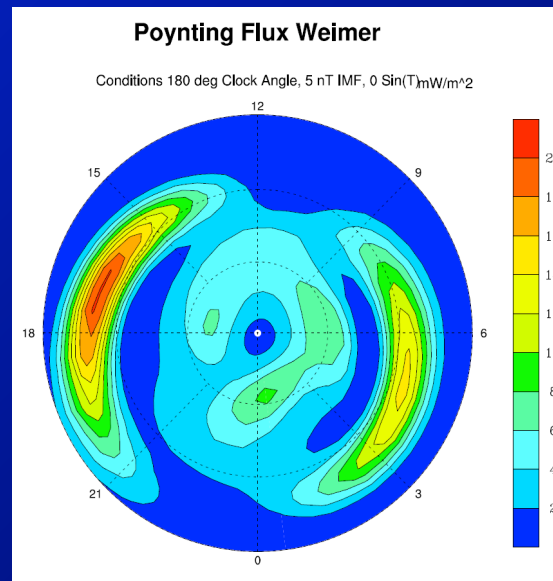
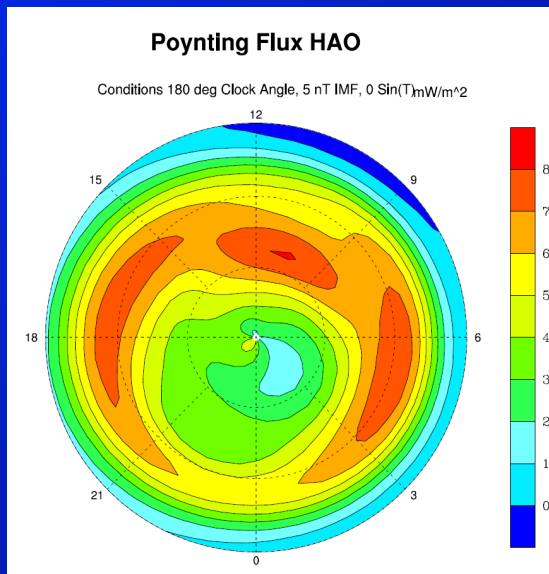
- Weimer is much larger than HAO when not at equinox
- Regions and patterns between the models vary
- Models are most alike at equinoxes



# BF Summery

- HAO is stronger than Weimer, except away from equinox
- Pattern variation is small
- Strength variation is normal
- Behaves almost like E Field

# Poynting Flux



# Poynting Flux IMF Clock Angle Summery

- HAO's ExB and Weimer have similar structure and values for the Poynting flux
- HAO's Data Fitted values are larger than both of the other models
- Rotation around MN line can be seen between the different clock angles; except HAO's Data Fitted

# Poynting Flux IMF Strength Summery

- Flux increase with IMF strength
- HAO's ExB and Weimer show similar structure, location varies
- HAO's Data Fitted becomes rings as saturation is reached
- Weimer has the highest peak values

# Poynting Flux Season Summery

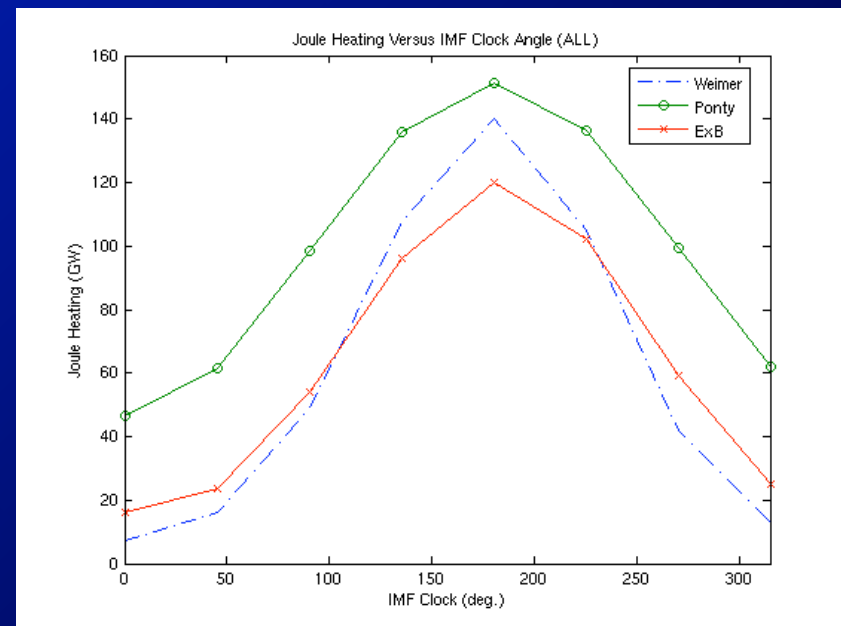
- Three model become more similar away from equinox
- HAO's ExB and Weimer peak at equinox while HAO's Data Fitted peak at extreme summer
- Large rotations around MN line with season change

# Poynting Flux Summary

- Weimer values are almost always larger than HAO's  $E \times B$
- Weimer and HAO's  $E \times B$  show similar structure
- HAO's Data Fitted forms rings
- As expected, the models behave like E field and B field

# Joule Heat v. IMF Clock Angle

- HAO's Data Fitted is largest over all clock angles
- HAO's ExB and Weimer are close together
- All peak at  $180^\circ$
- Behaviour expected

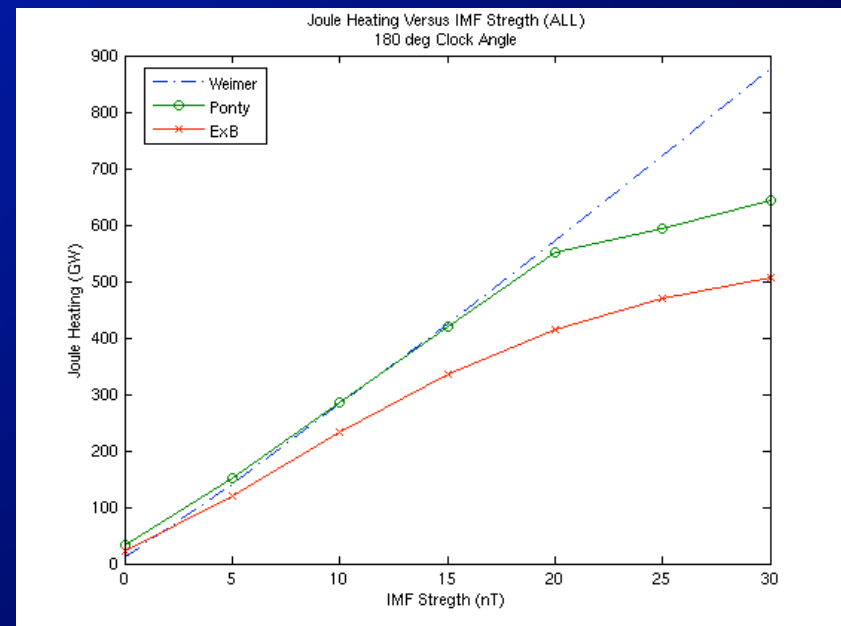


IMF Strength: 5 nT  
Dipole Tilt Angle:  $0^\circ$



# Joule Heat IMF

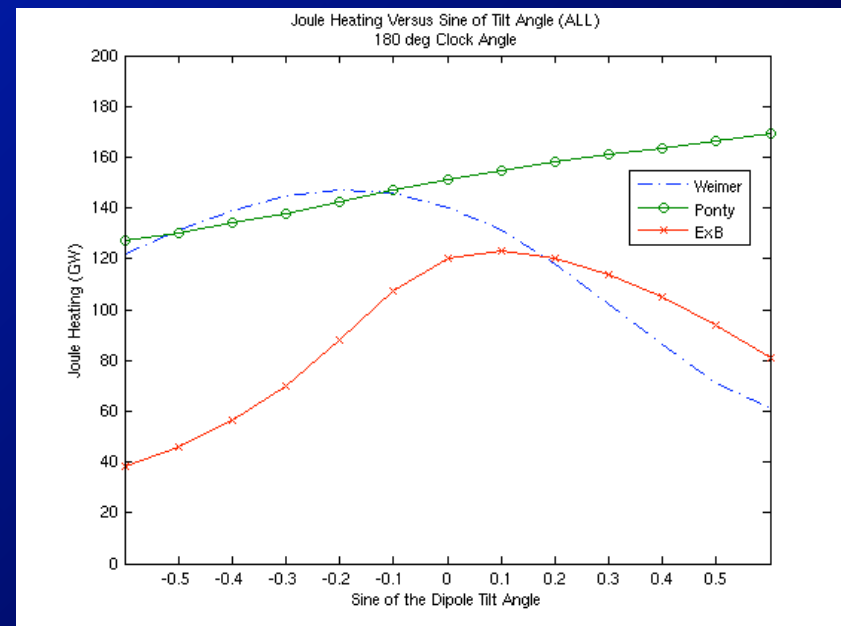
- Weimer and HAO's Data Fitted are close until around 20 nT
- Weimer appears linear
- Both HAO's Data Fitted and HAO's ExB level off



Clock Angle:  $180^\circ$   
Dipole Tilt Angle:  $0^\circ$

# Joule Heat Season

- HAO's Data Fitted appears linear; small bump around equinox
- Both Weimer and HAO's ExB have peaks
- Weimer peaks around  $\text{Sin}(T) = -0.2$



Clock Angle:  $180^\circ$   
IMF Strength: 5 nT

# Conclusions

- Two models show differences as conditions are varied (clock angle, IMF strength, dipole tilt)
- Strength and pattern variations
- Though there are local areas of great difference, globally the values are small
- With residuals, no major problems were seen except in Poynting flux

# Future Plans

- HAO's Data Fitted Poynting flux being reworked
- Incorporate model into a General-Circulation Model to study effects on Thermosphere
- Use model to find spatial and temporal properties of the energy input

# References

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