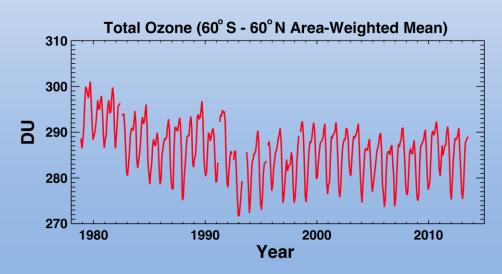
The impact of solar spectral irradiance (SSI) variations on stratospheric composition: Theory and observations

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William H. Swartz
JHU/Applied Physics Lab



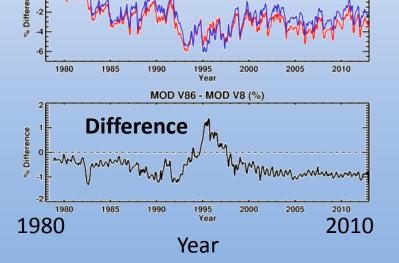
SBUV (Solar Backscatter UltraViolet) Instruments

- Nadir-viewing; use solar UV radiation backscattered from the atmosphere to measure ozone
- New Version 8.6 (replacing version 8)
 - Total ozone is the sum of layer amounts
 - Early instrument calibration to SSBUV; late instrument calibration to NOAA 17
- Merged ozone data set (MOD) SBUV only: no TOMS data

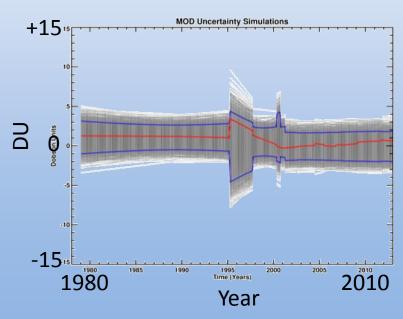
Version 8.6 to 8

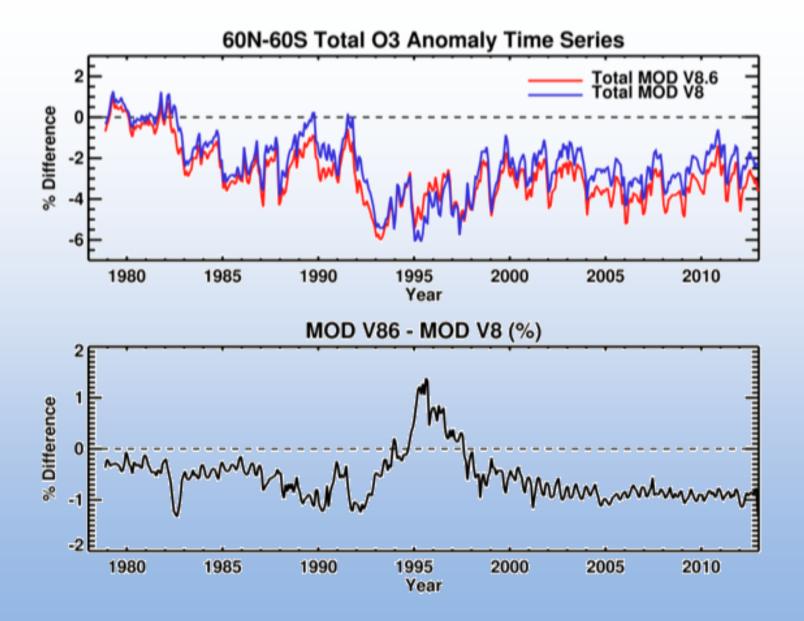
60N-60S Total O3 Anomaly Time Series

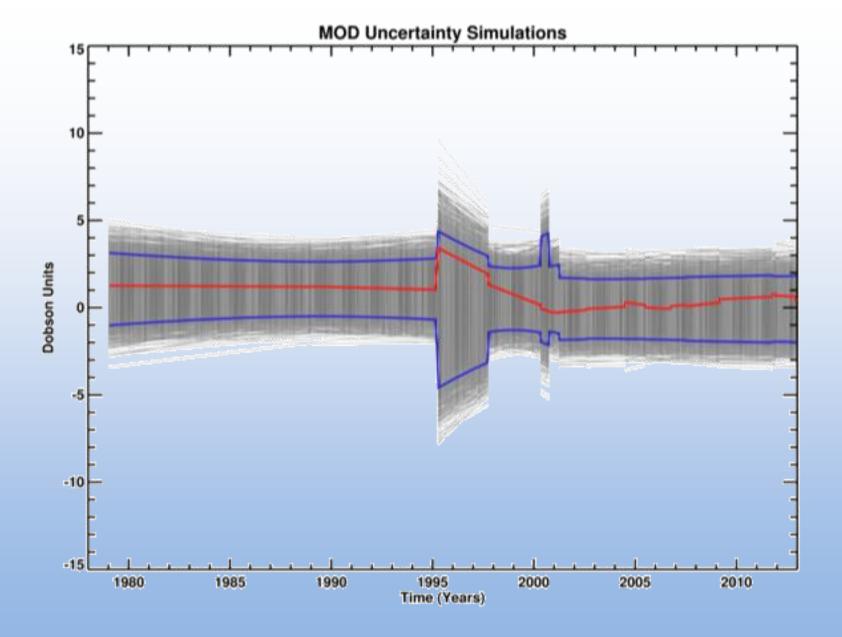
Anomalies



Merging Uncertainty

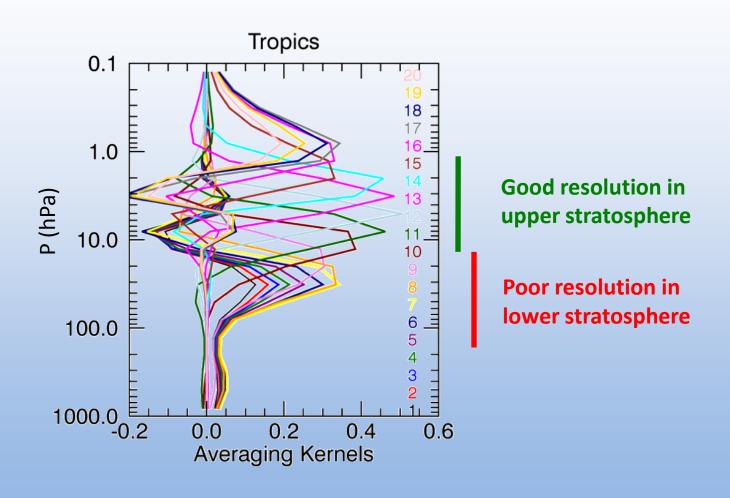






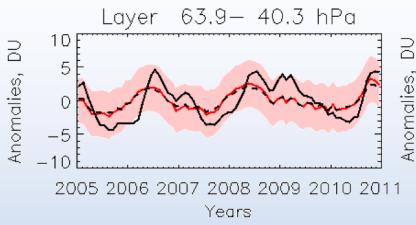
SBUV Altitude Profiles

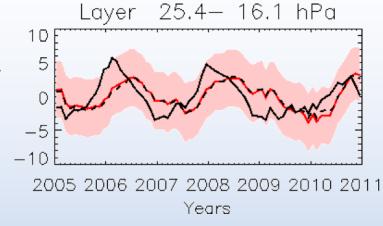
Instrument uses wavelength to scan in altitude



SBUV Lower Stratospheric Measurements

Kramarova, N. et al. Atmos. Meas. Tech. 6, 2089-2099, 2013





Black line: MLS anomalies Red line: SBUV anomalies

Red shaded: SBUV smoothing error

Black dashed line: MLS with SBUV
Averaging Kernel applied



Conclusion: SBUV measurements, integrated over a broad vertical layer, provide an excellent data record for the lower stratosphere

Column Ozone (DU) Ozone Change (DU) Icanic Effects Ozone Change (DU) 1976 1982 1988 1994

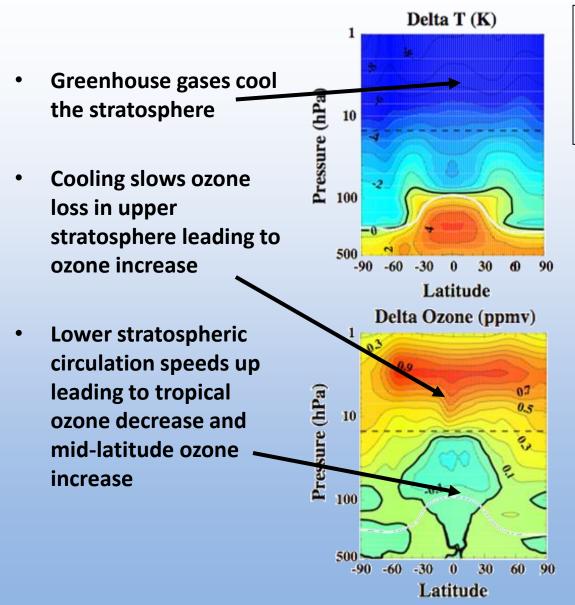
The Regression Problem

Solar term and volcanic aerosol term are similar for two cycles

Third cycle is different at same time chlorine term changes direction

In addition, CO₂ affects ozone and has varied nearly linearly over time. We have the problem of separating this signal from the chlorine signal.

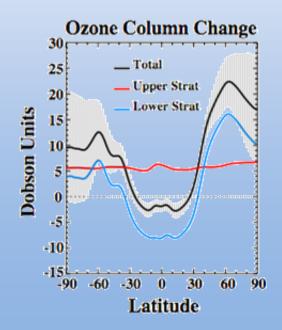
The Impact of GHGs on Stratospheric Ozone



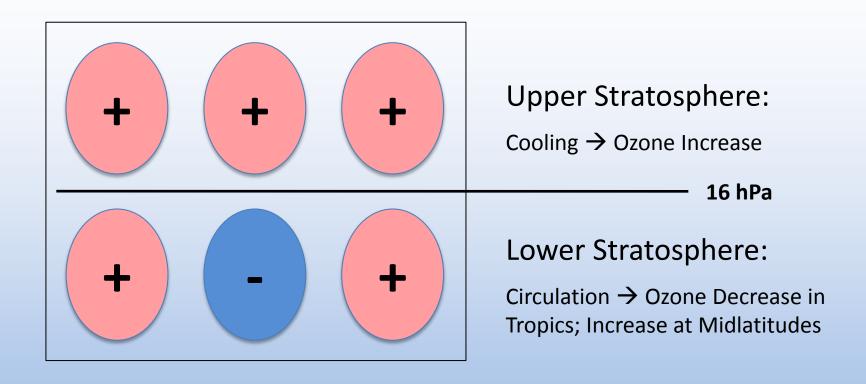
Results from the GEOS CCM 2065-1980

Li, F., et al. (2009), Stratospheric ozone in the post-CFC era, *Atmos. Chem. Phys.*, *9*(6), 2207–2213.

Net result is a column ozone increase at mid to high latitudes and almost no change near the equator



Expected Pattern for GHG Impact on Ozone

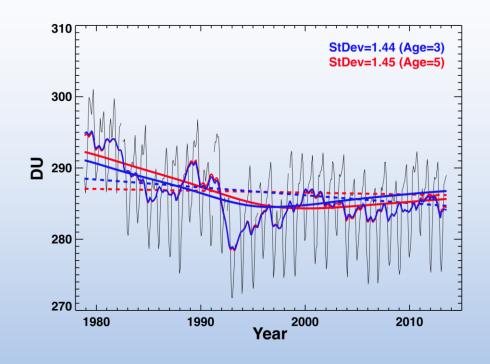


Can we separate ozone change due to ODSs from that due to GHGs?

Fit to EESC + Linear Trend
(plus Solar, volcanos, QBO, and ENSO)

Use Nash/Newman EESC (2 examples; Age=3 years and Age=5 years)

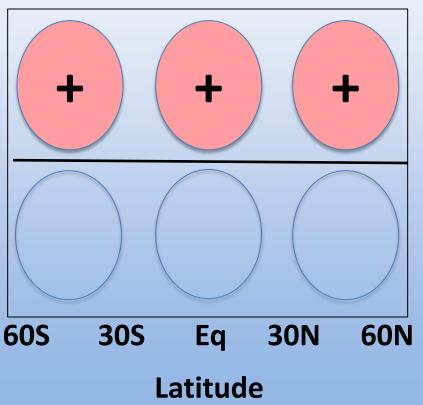
Linear trend represents GHGs and is expected to have a positive coefficient

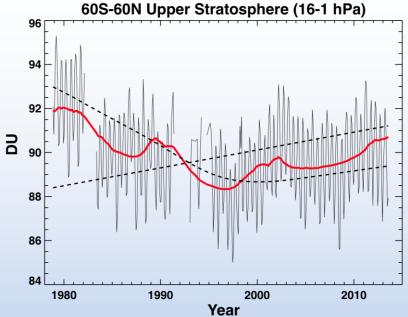


	EESC trend pre- 1993	EESC trend post- 2000	Linear trend
Age = 3 years	- 4.5 ± 1 DU/dec	+ 1.3 ± 0.3 DU/dec	- 1.1 ± 0.5 DU/dec
Age = 5 years	- 4.7 ± 1 DU/dec	+ 1.3 ± 0.3 DU/dec	- 0.2 ± 0.7 DU/dec

Upper Stratosphere (16-1 hPa)

Fit to EESC, Solar Cycle, and Linear Trend

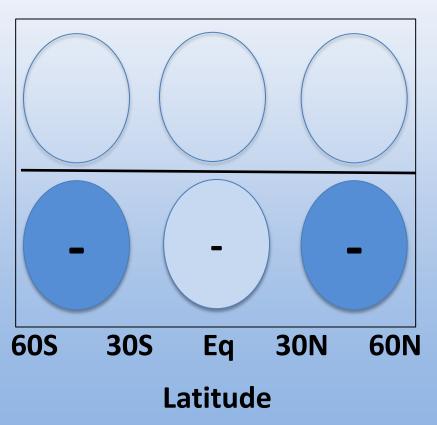


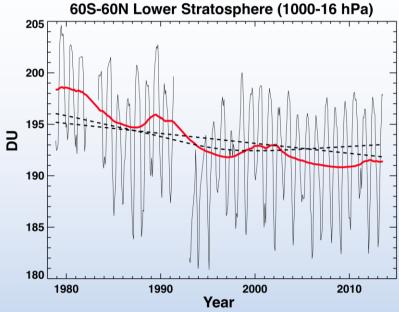


Upper Strat (16-1 hPa)	Trend (DU/decade)
60S-60N	+0.7 ± 0.4
60S-30S	+0.5 ± 0.3
30S-30N	+0.9 ± 0.6
30N-60N	+0.6 ± 0.4

Lower Stratosphere (1000-16 hPa)

Fit to EESC, Solar Cycle, and Linear Trend + QBO, Volcanos, ENSO

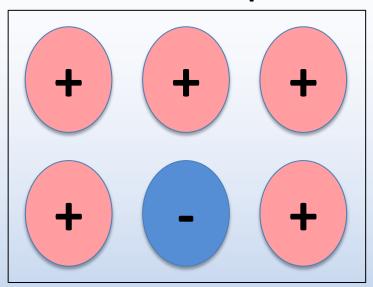




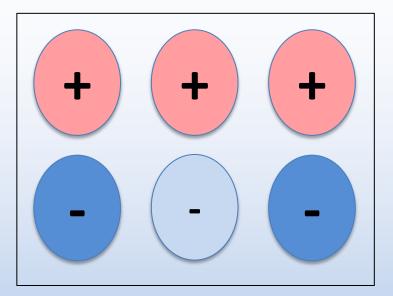
Lower Strat (1000-16 hPa)	Trend (DU/decade)
60S-60N	-1.4 ± 0.5
60S-30S	-2.3 ± 1.6
30S-30N	-0.8 ± 0.9
30N-60N	-2.0 ± 1.8

Summary

What we expect

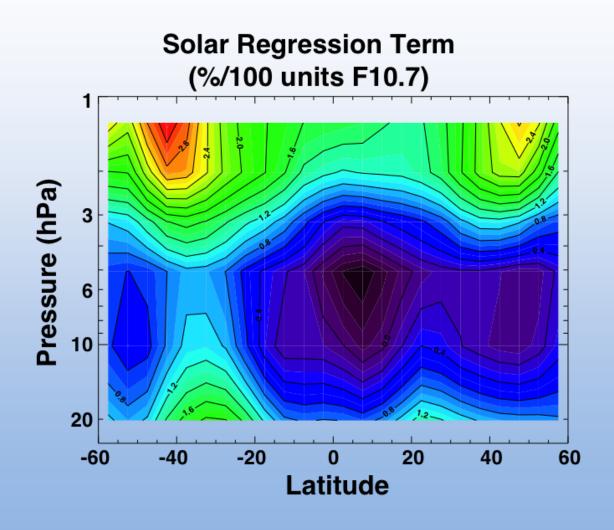


What we see



- Upper stratospheric cooling shows positive ozone response as expected
- Lower stratospheric ozone does not show evidence of circulation speed-up

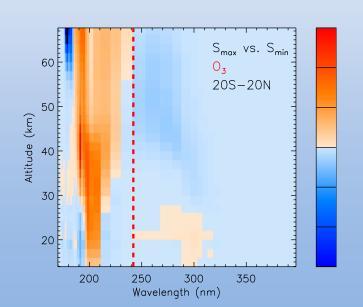
SBUV Data 1979-2013

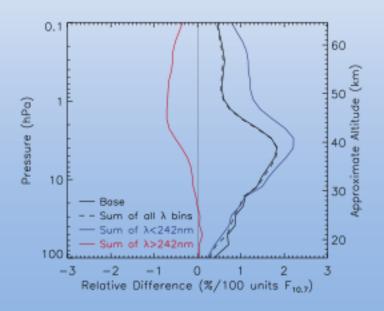


Solar Cycle Impact on Ozone: Theory (Top Down)

$$O_2$$
 + hv (λ < 242 nm) \rightarrow O + O + k.e.
 O_3 + hv (λ < 308 nm) \rightarrow O_2 + O + k.e.

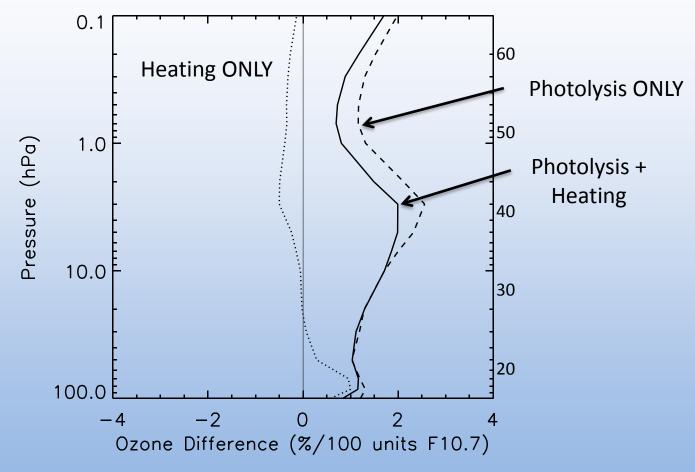
- O₂ photolysis produces ozone (positive response)
- O₃ photolysis produces atomic oxygen that recombines ozone (negative response)





Photolysis accounts for virtually the entire ozone response

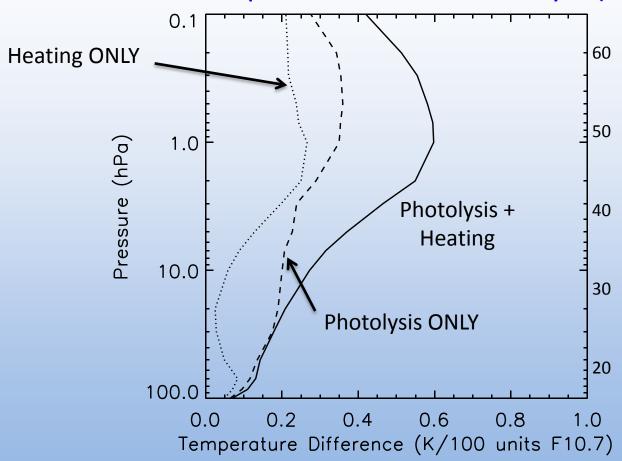
SMax-SMin Ozone: Timeslice 25 yrs (60S-60N)



Chemistry—transport models (CTMs) should get the ozone response just about right.

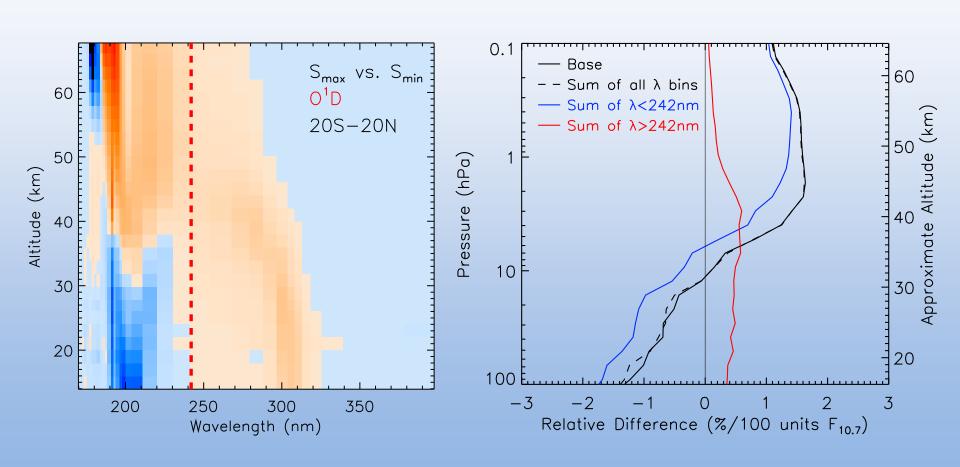
Photolysis accounts for more than half the temperature response

SMax-SMin Temperature: Timeslice 25 yrs (60S-60N)

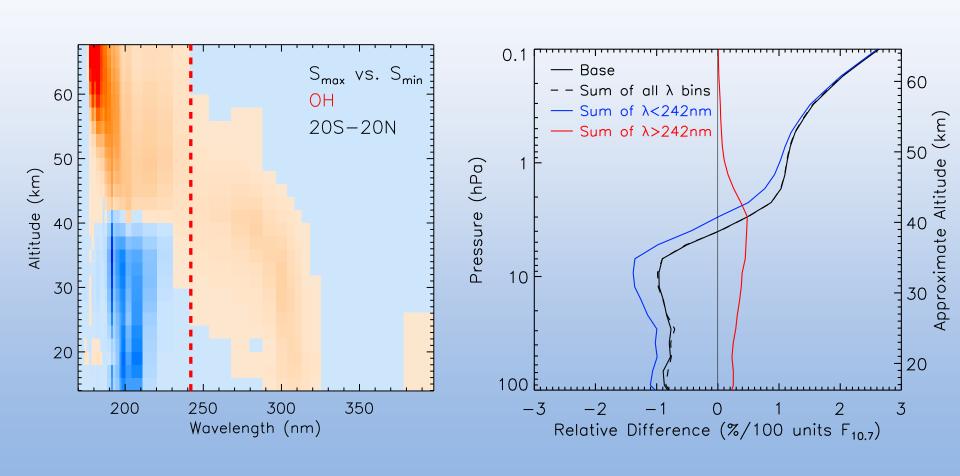


General circulation models (GCMs) without coupled chemistry may capture only half (or less) of the solar cycle impact on temperature.

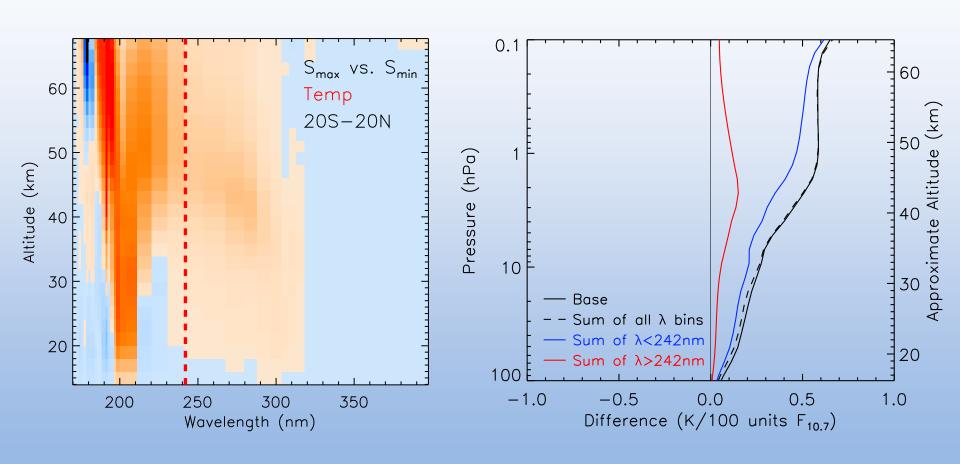
O^1D response = sum of the λ bins



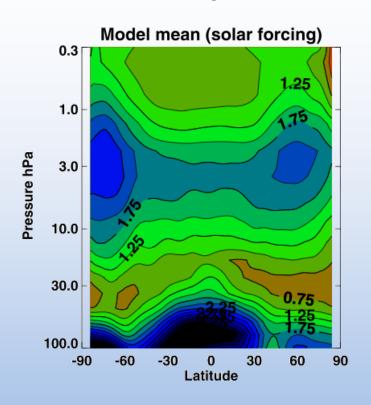
OH response = sum of the λ bins

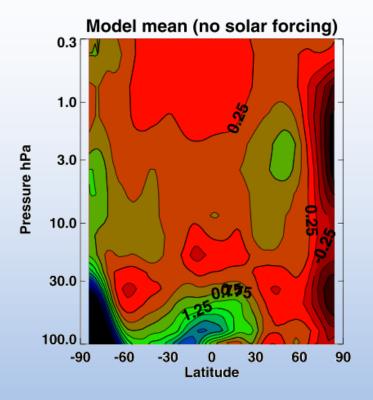


Temperature response = sum of the λ bins



Model (3D CCM) results for ozone response to solar UV forcing



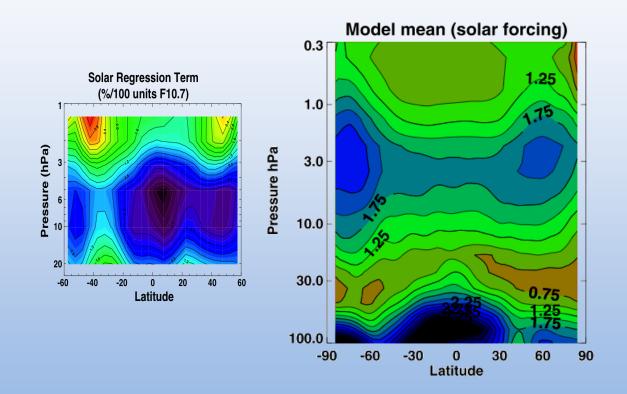


Models obtain maximum ozone response centered at about 3 hPa; larger at mid to high latitudes than at equator

Models without solar cycle forcing can appear to have a solar signal in polar regions and in lower stratosphere

Austin, J. et al. (2008), Coupled chemistry climate model simulations of the solar cycle in ozone and temperature, *J. Geophys. Res.*, 113(D11), doi:10.1029/2007JD009391.

Observed solar signal in upper stratosphere has expected latitude signature, but occurs at higher altitude than expected (at least in SBUV data)



Conclusions

- Separation of chlorine signal from greenhouse gas signal in ozone data presents a problem
 - Upper stratosphere where GHG cool seems to work
 - Lower stratosphere where circulation speed-up is expected doesn't work
- Separation of solar cycle signal from volcanic aerosol signal may get confused with chlorine/GHG separation
 - Works pretty well in upper stratosphere
 - Give result in lower stratosphere for solar signal, but extension of data set could easily modify answers