

2018 Sun-Climate Symposium, 19-23 Mar. 2018 TSIS SIM

TSIS SIM: First Light and Early Observations





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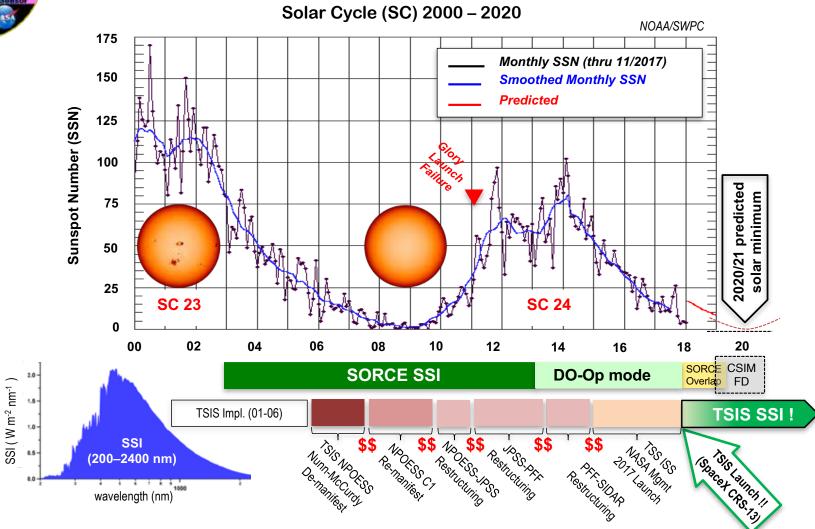
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"Success is not final, Failure is not fatal: it is the courage to continue that counts" - Winston Churchill



"Per ardua ad astra"

(By striving we reach the stars)



TSIS SIM Timeline

Launch	15 December 2017
Turn-on	3 January 2018
Commissioning	4 Jan – 1 Mar. 2018
First Light	3-5 March 2018
Normal Ops	14 March 2018



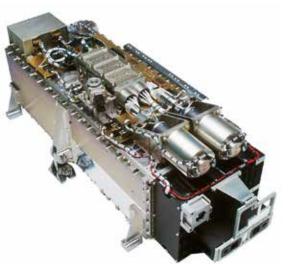


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Passing the SSI "Baton"







SORCE SIM (launched 1/25/2003)

- Two channel instrument (duty-cycled for stability corrections)
- Absolute ESR detector (NiP bolometer)
 - First generation (nominal performance)
 - Diamond substrate
 - NiP black absorber
 - Kapton™ thermal link
- Abs. accuracy: 2-10% wavelength dependent (no-SI validation)
- "...15 years into a 5 year mission"

TSIS SIM (launched 12/15/2017)

- ✓ <u>Three</u> channel instrument
 - For long-term stability validation of duty-cycling
- ✓ Absolute ESR detector (NiP bolometer)
 - Second gen. (improved noise performance)
 - Diamond substrate
 - NiP black absorber
 - Kapton™ thermal link
- ✓ Abs. accuracy 0.2 % (SI-traceable validation)

"...~2 weeks into a 5+ year mission"



TSIS Spectral Irradiance Monitor



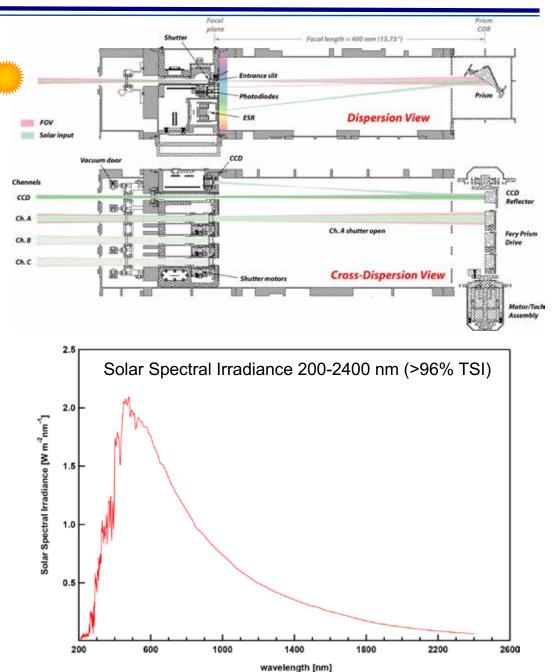
3-channel SSI radiometer

Each channel contains:

- Féry prism for dispersion
- 3 primary photodiode detectors
- Absolute ESR detector

Lessons learned from SORCE SIM

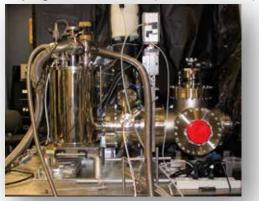
- Establish consistent prism exposure plan
- Maintain constant exposure ratio between channels (target 10% duty cycle based on 7-year plan)
 - Expose B channel **daily** to experience same solar activity & contam. env't.
 - Scan ESR over limited wavelength regions for A/B (&C) comparisons (avoid disparate point scans)
 - Expose Channel C to same "optical" conditions (twice annually)



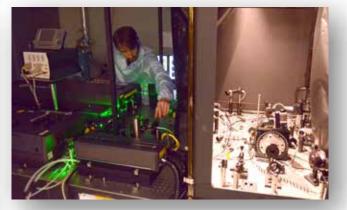


Absolute Irradiance Scale (LASP-SRF)

L-1 Cyrogenic radiometer (NIST traceable)



SRF "SIRCUS" Laser system (206 – 3000 nm coverage)

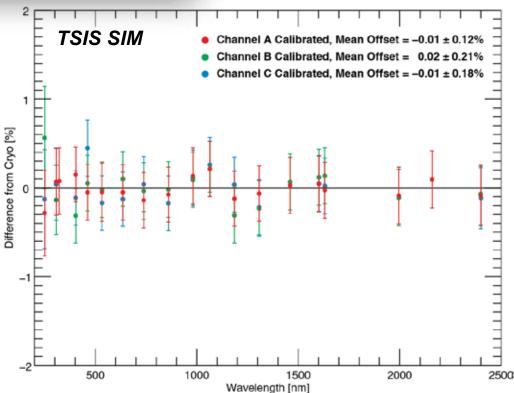


TSIS SIM



TSIS SIM absolute calibration in the LASP SRF

- SRF allows us to calibrate the instrument absolutely relative to the cryogenic radiometer and evaluate instrument optical performance as a function of wavelength
- ESR Calibration against the cryogenic radiometer is also part of the process as it provides the vacuum environment optimized for ESR noise testing
 - Get ESR vs Photodiode response
 - Get ESR noise floor performance
- Designed to achieve < 1% (0.2% goal) absolute accuracy uncertainty validation



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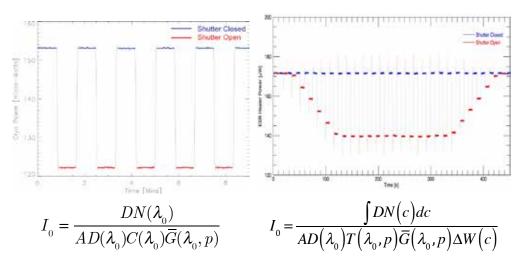


Absolute Irradiance Scale (LASP-SRF)



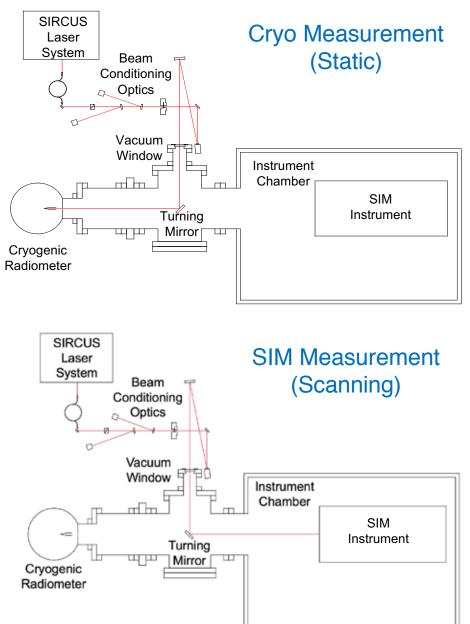
Cryogenic Radiometer Uncertainty Budget

Parameter	Unit	% Effect	% Unc. (k=1)
Power	W	100	0.015
Cavity Reflectance	-	0.01	0.004
Cavity Non-Equiv.	-	0	0.01
Slit Area: Measured	m ²	100	0.05
Slit Area: Contraction	-	0.04	0.01
Slit Area: Cosine effect	-	0.02	0.01
Slit Diffraction Loss	-	0.13	0.02
Total			0.07

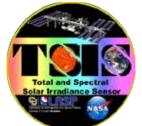




Parameter	Unit	% Effect	% Unc. (k=1)
Cryo Measurement	W/m ²	100	0.07
Turning Mirror Repeatability	-	0	0.004
Laser: Stability	-	0	0.060
Laser: Pattern Uniformity	-	0	0.023
Path Length Correction	-	0	0.0002
CSIM Spectral Integration	W/m ²	100	0.1
Total			0.14



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TSIS-1 on ISS





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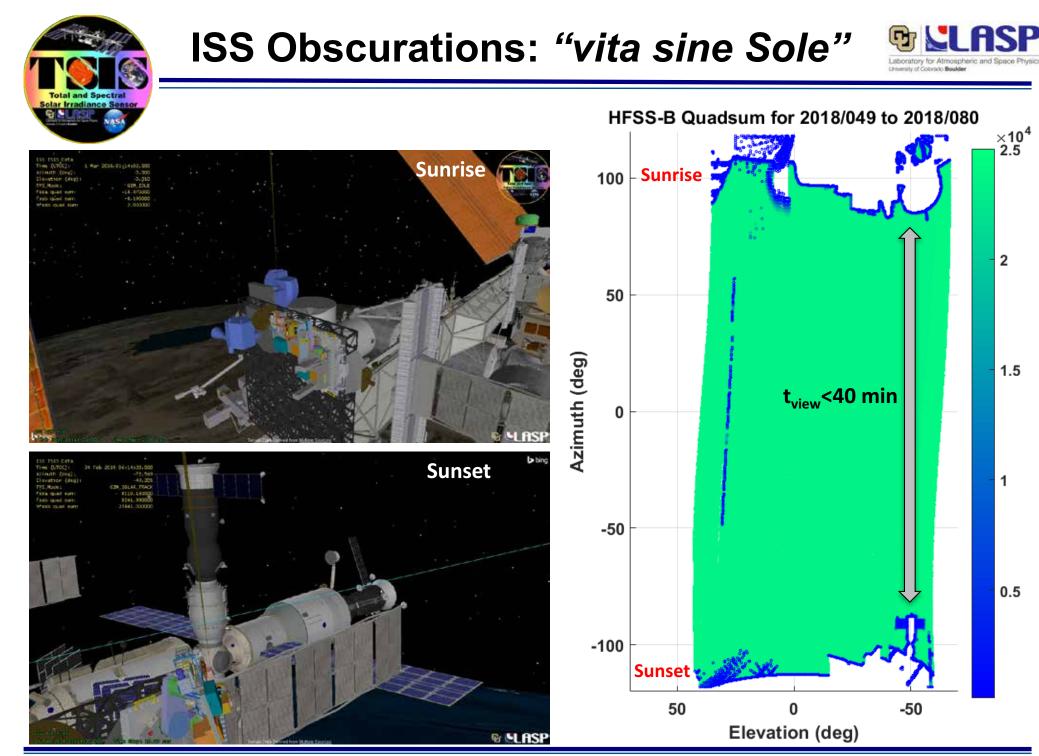


TSIS-1 on ISS



Deployed 31 Dec. 2017





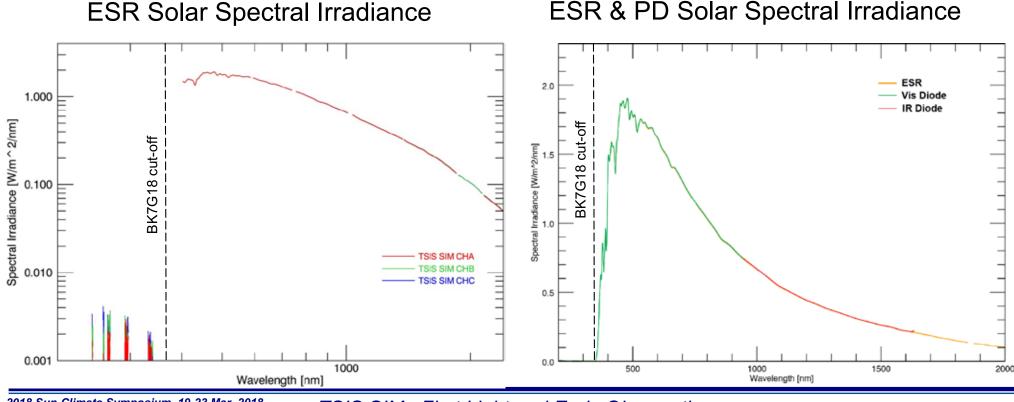
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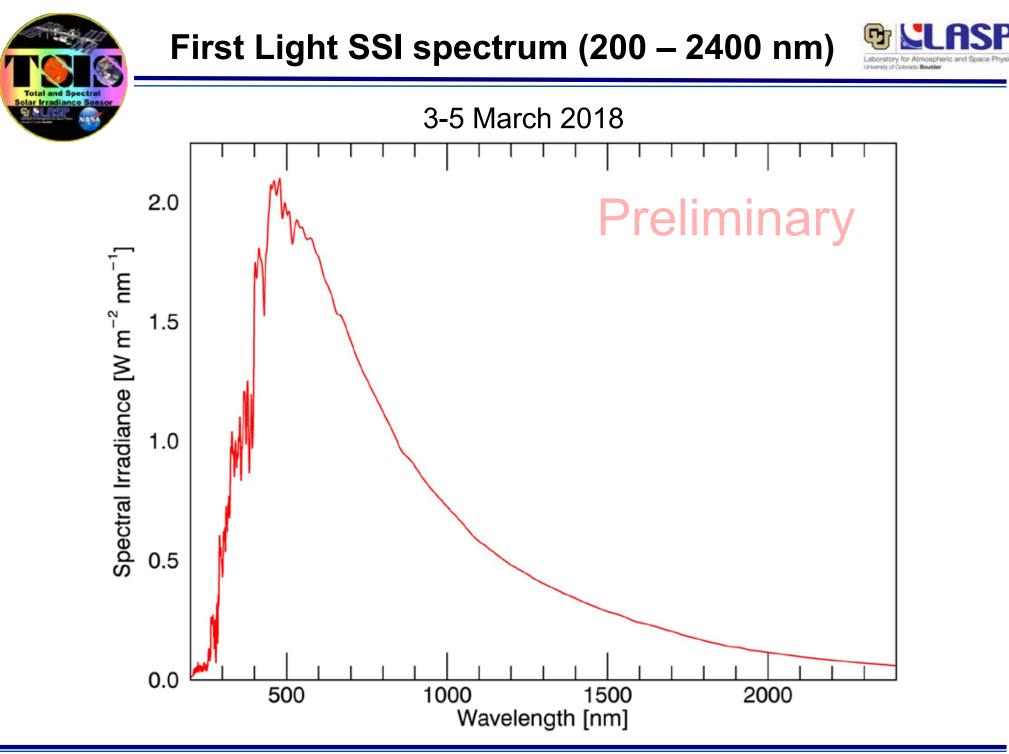
Performance Summary: SSI Spectrum

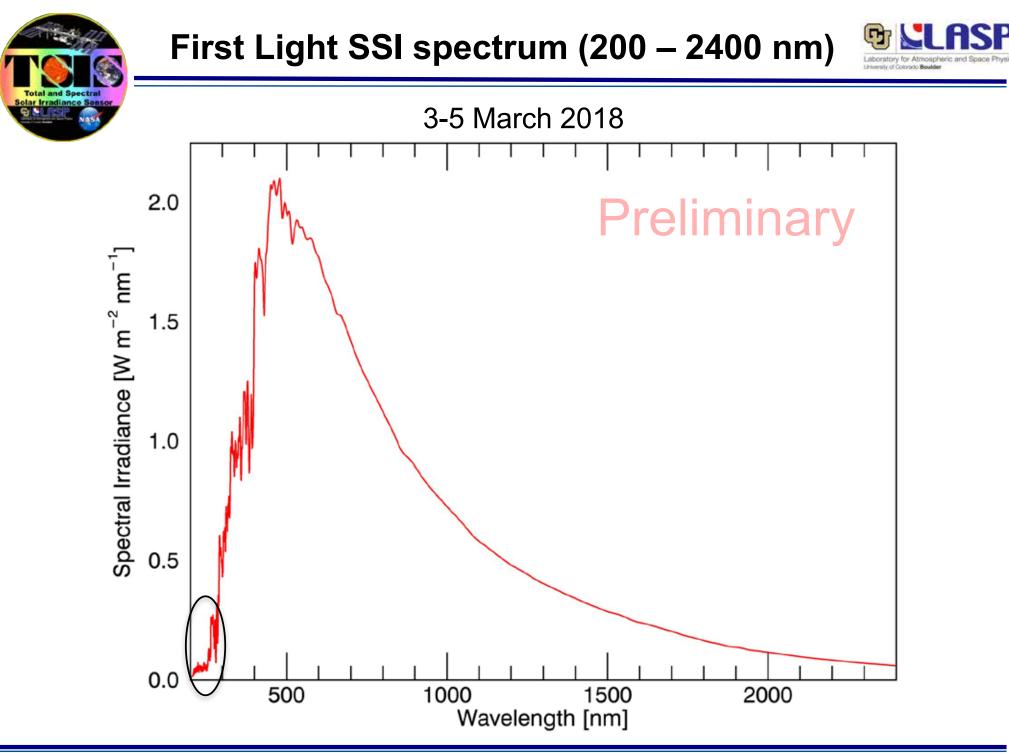
Commissioning

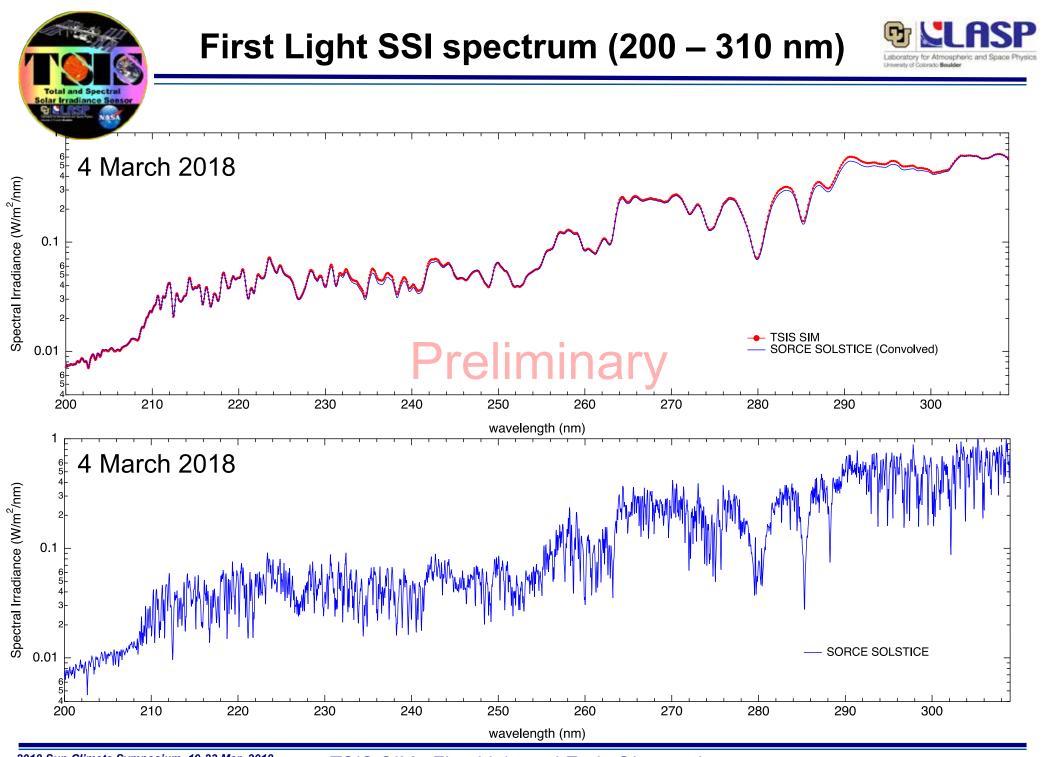
- Performed dry-run Full Scans (vac. door closed, through BK7G18 window). ESR & PD scans
- No UV solar signal (< 350 nm) through BK7G18, allows for quantification of background signal (stray & scattered light, In-Field/Out-of-Band)



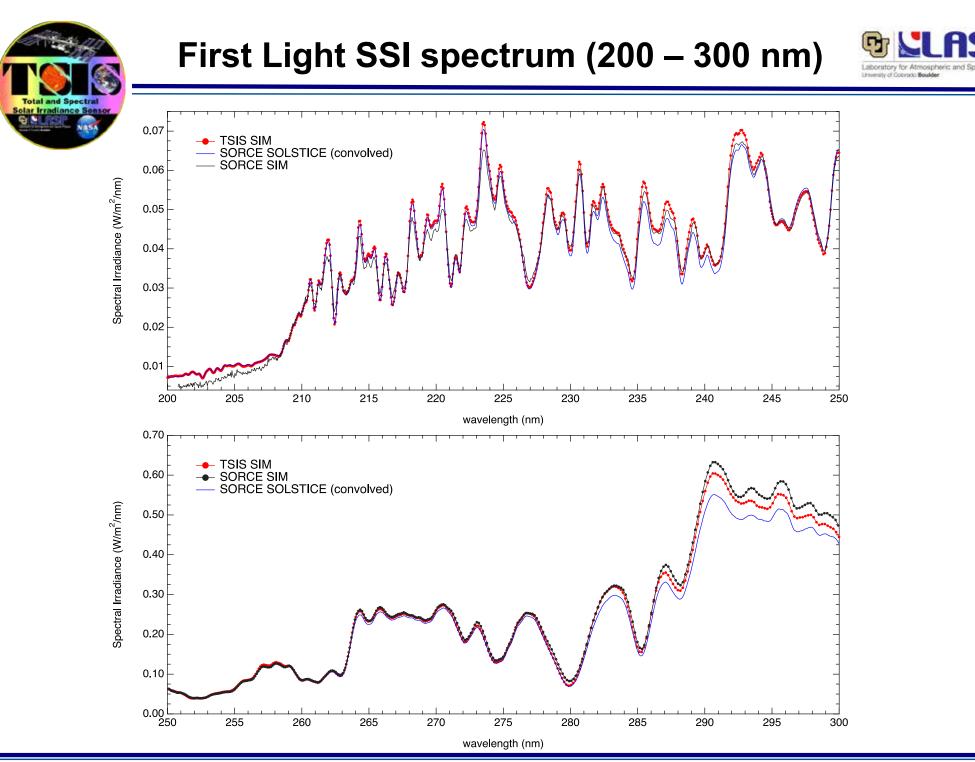
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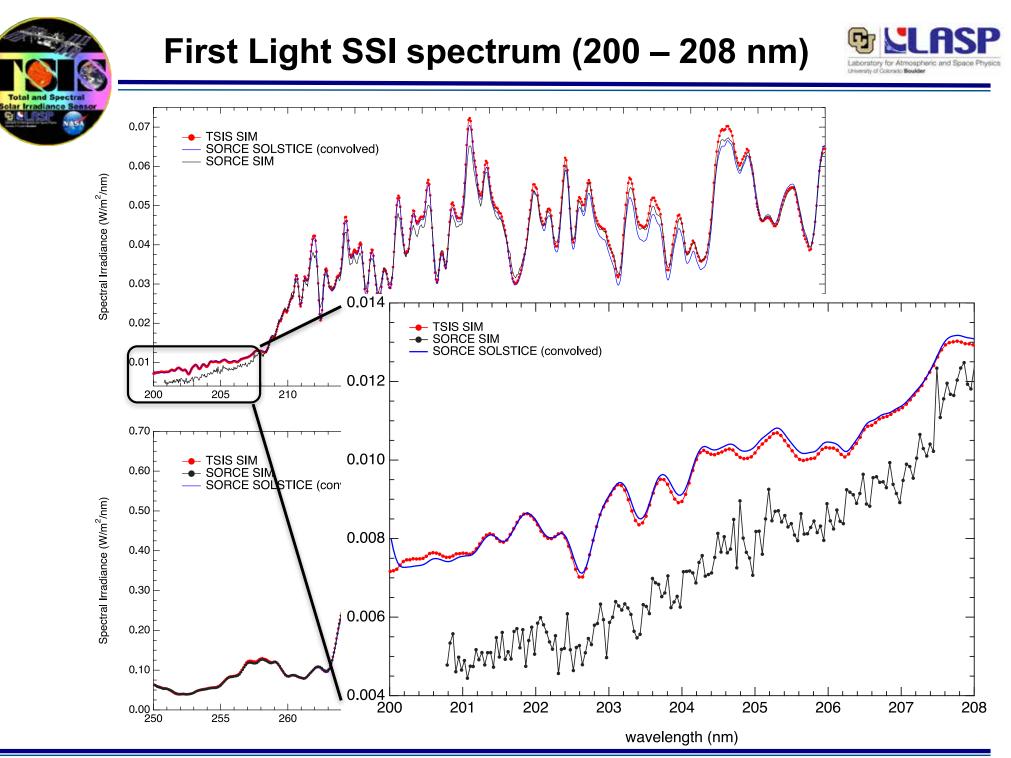






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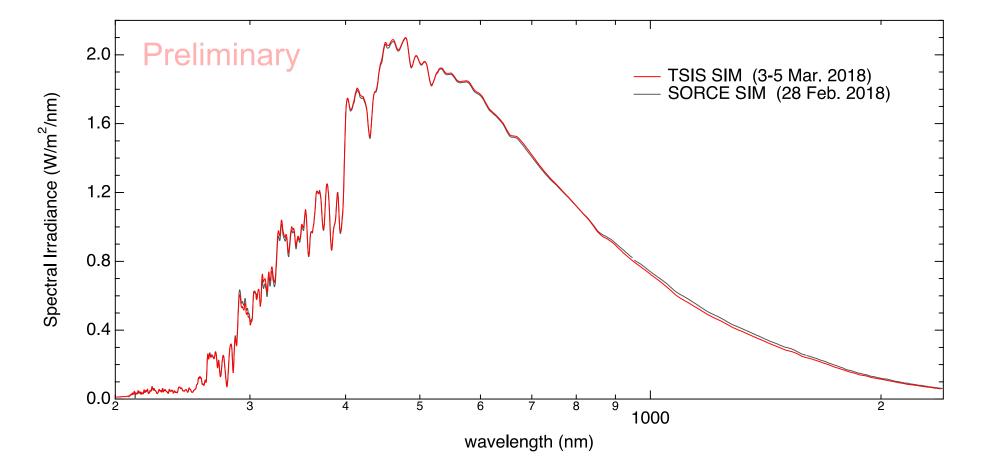




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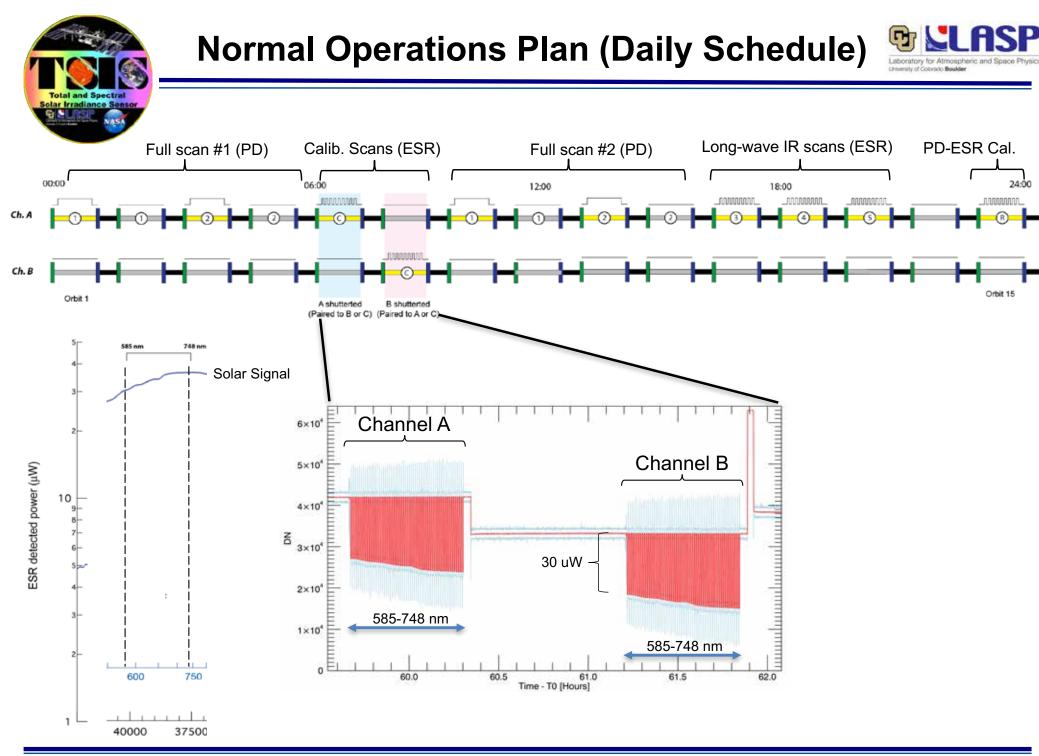


Uncorrected reference spectra integrals (relative comparison)

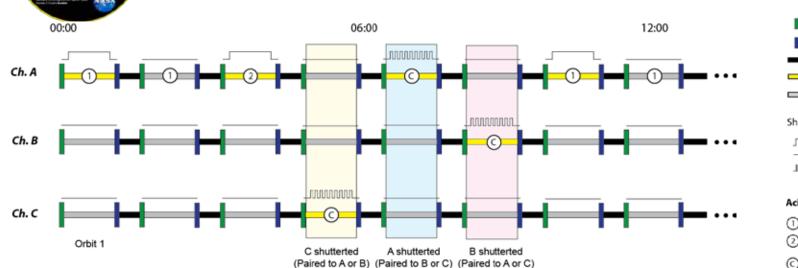
Spectrum	205-2390 (W/m²) (96% TSI)	+ 52 (W/m²)*	TIM TSI (W/m²)	% Diff.
ATLAS-3	1333	1386	1362-1360	+1.76-1.88
SIRS-WHI	1323	1375	1362-1360	+0.95-1.1
TSIS SIM	1307.6	1359.6	1360.6	-0.08

*Integrated SSI contribution outside 205-2390 nm

L. Dame, *"New Solar Reference Spectrum SOLAR-ISS"* Session 2: 5:30



Timeline for Channel C Calibrations

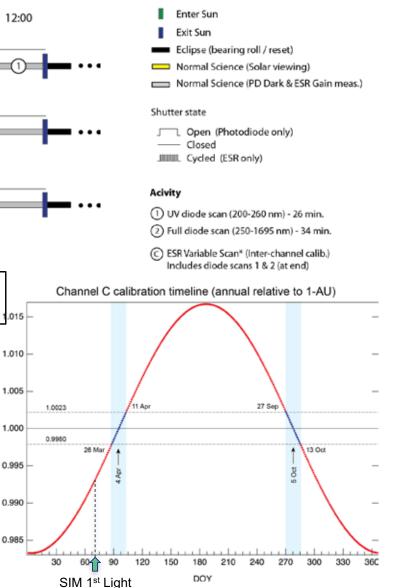


Note: Addition of Ch. C interleaved calibration does not affect the nominal Ch. A & B operational timeline (exposure cadence undisturbed

The Channel C calibration activities occur <u>twice per</u> <u>calendar year</u> and require 17 days (like channel B) centered at common 1-AU times (4/4 & 10/5).

The reason for this timing relates to guaranteeing common:

- Field-of-view (similar solar image in prism)
 - Want to match degradation spot on prism
- Solar flux (similar distance correction)
 - 6.7% irradiance change over 6 months, therefore different correlation to exposure time between 17 days in January and 17 days in July



olar Irradiance S

Distance [AU]



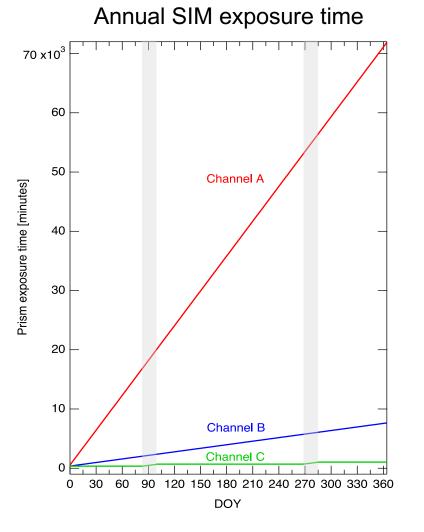


Total and Spectral Solar Irradiance Sensor

Total prism exposure for all SIM channels

Calibration	Prism – Solar exposure totals			
Totals	Daily (min.)	17-day Cal. (min.)	Annual (min.)	Annual (days)
Channel A	196	3332	71540	49.68
Channel B	~20	345	7245	5.03
Channel C	-	345 (every 6- months)	690	0.48

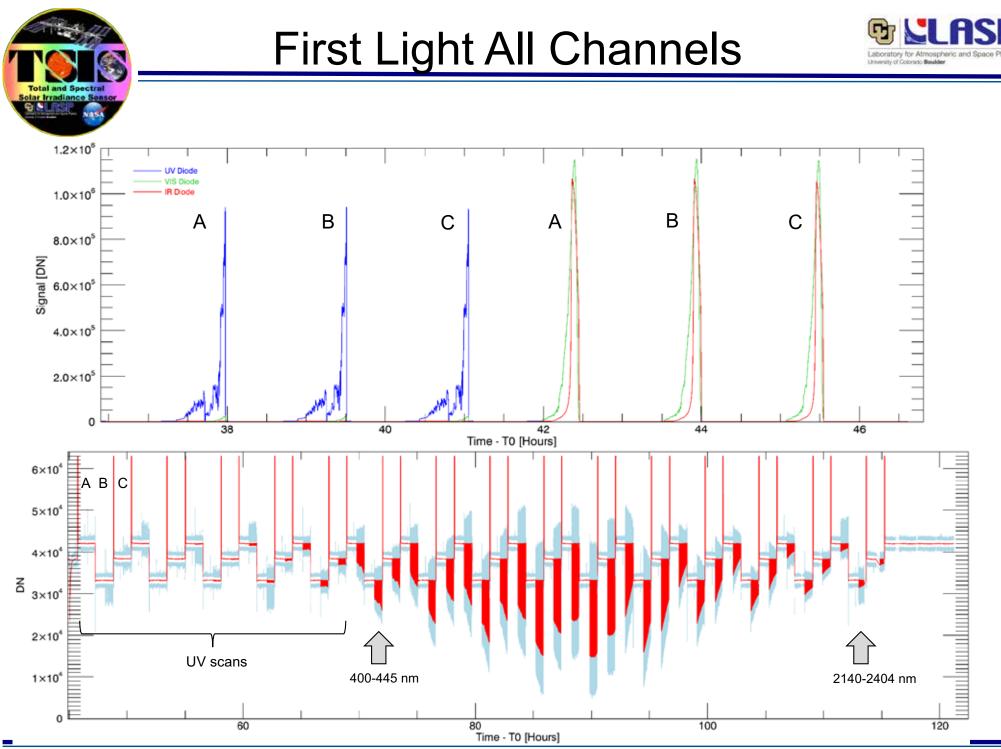
Channel B-to-A duty cycle = 10.1% Channel C-to-B duty cycle = 9.5%







BACK UP SLIDES

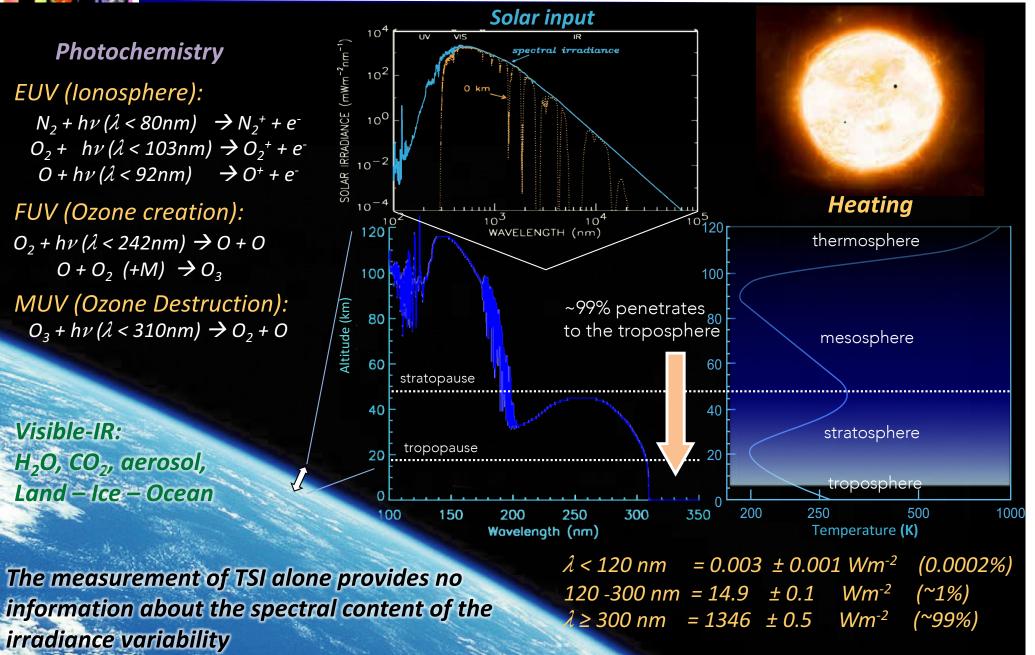


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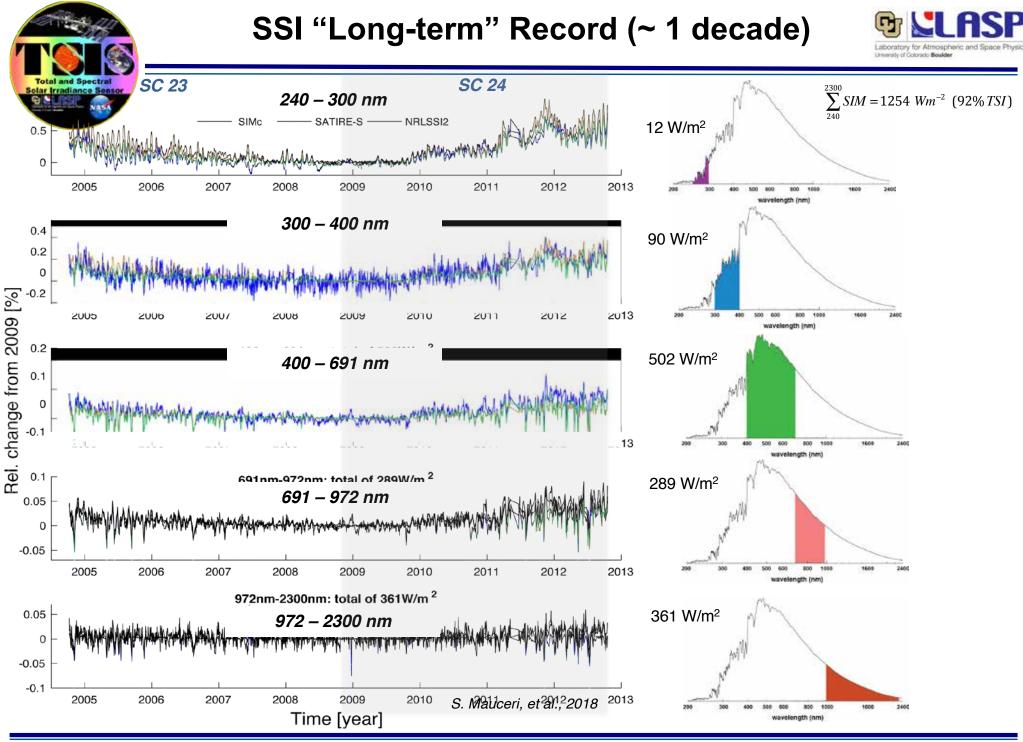


Wavelength Dependent Responses





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