

# Climate and Habitability of Earth-like Extrasolar Planets

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NASA Virtual Planetary Laboratory (VPL)

**Sun-Climate Symposium**  
**Lake Arrowhead, CA**  
**March 19-23, 2018**

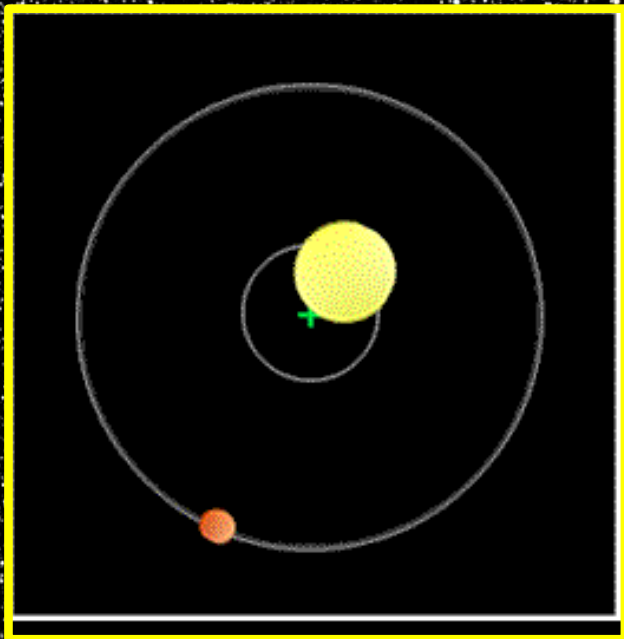


# As of today there are... 3706 Confirmed Exoplanets

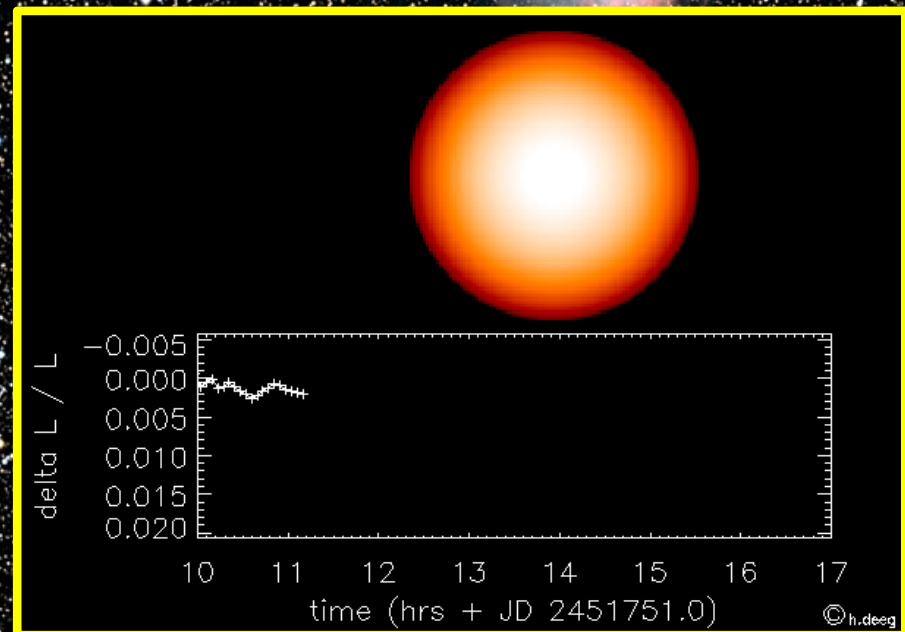
NASA Exoplanet Archive

<http://exoplanetarchive.ipac.caltech.edu>

radial velocity



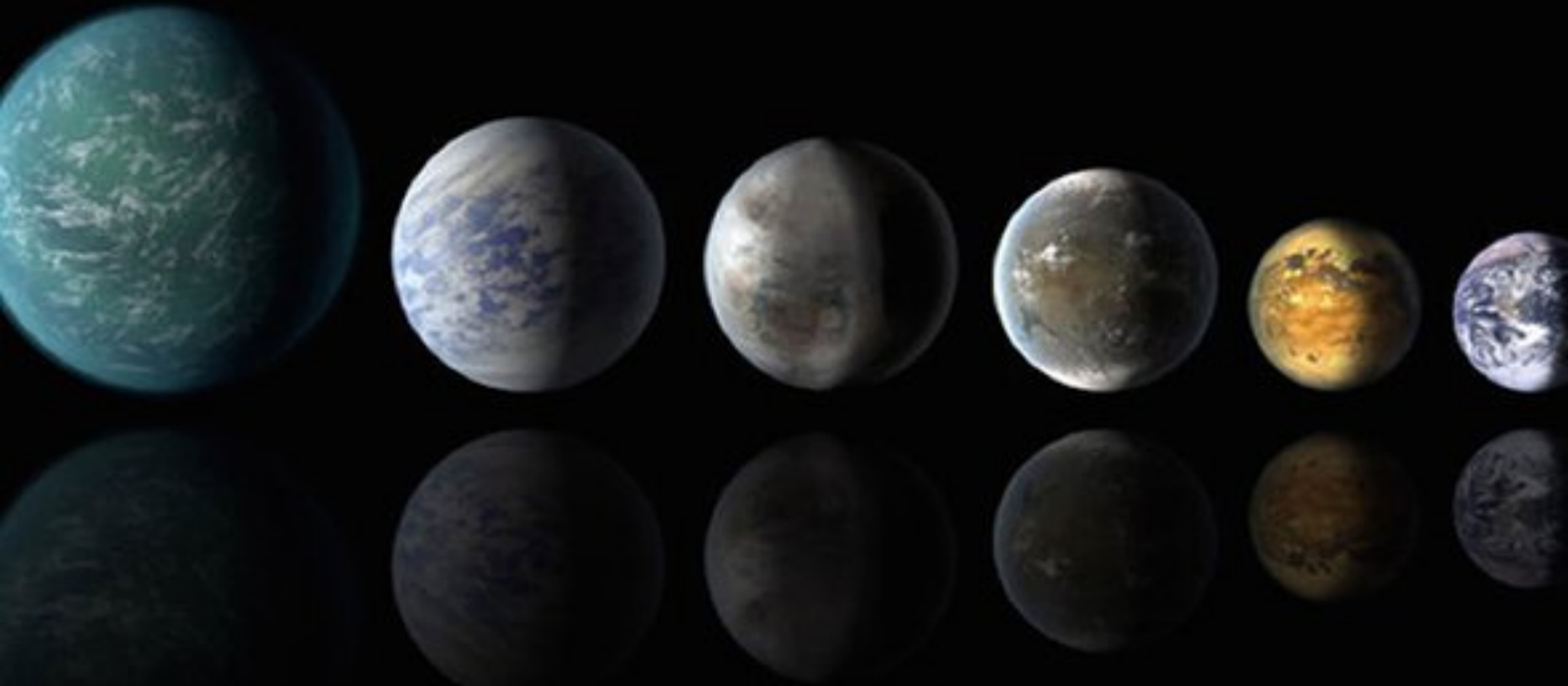
transit photometry





# Potentially habitable exoplanets:

***53 – optimistic assumptions***  
***13 – conservative assumptions***



Habitable Exoplanets Catalog

Planetary Habitability Lab, University of Puerto Rico, Acribo

<http://phl.upr.edu/projects/habitable-exoplanets-catalog>

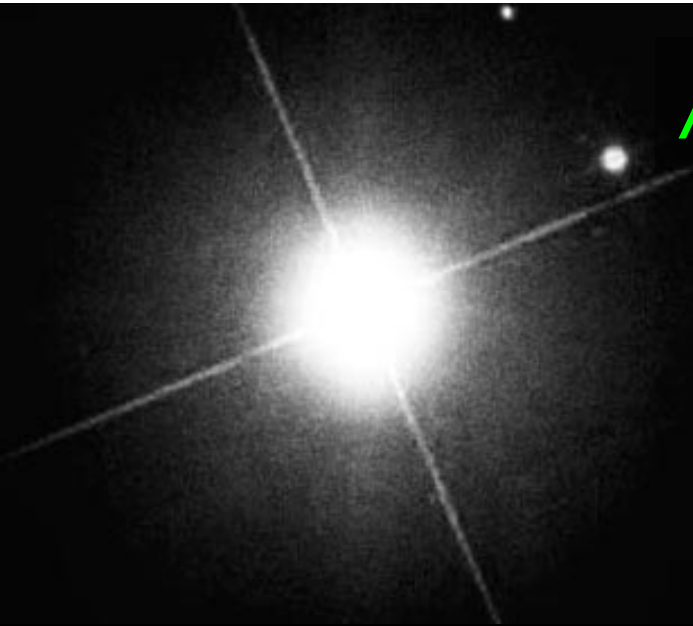


**What do we mean by a “habitable” planet?**





A habitable planet is ...  
*ocean covered.*





A habitable planet is ...  
ocean covered.

This provides significant  
constraint on surface  
temperature.

freezing

273 K

0° C

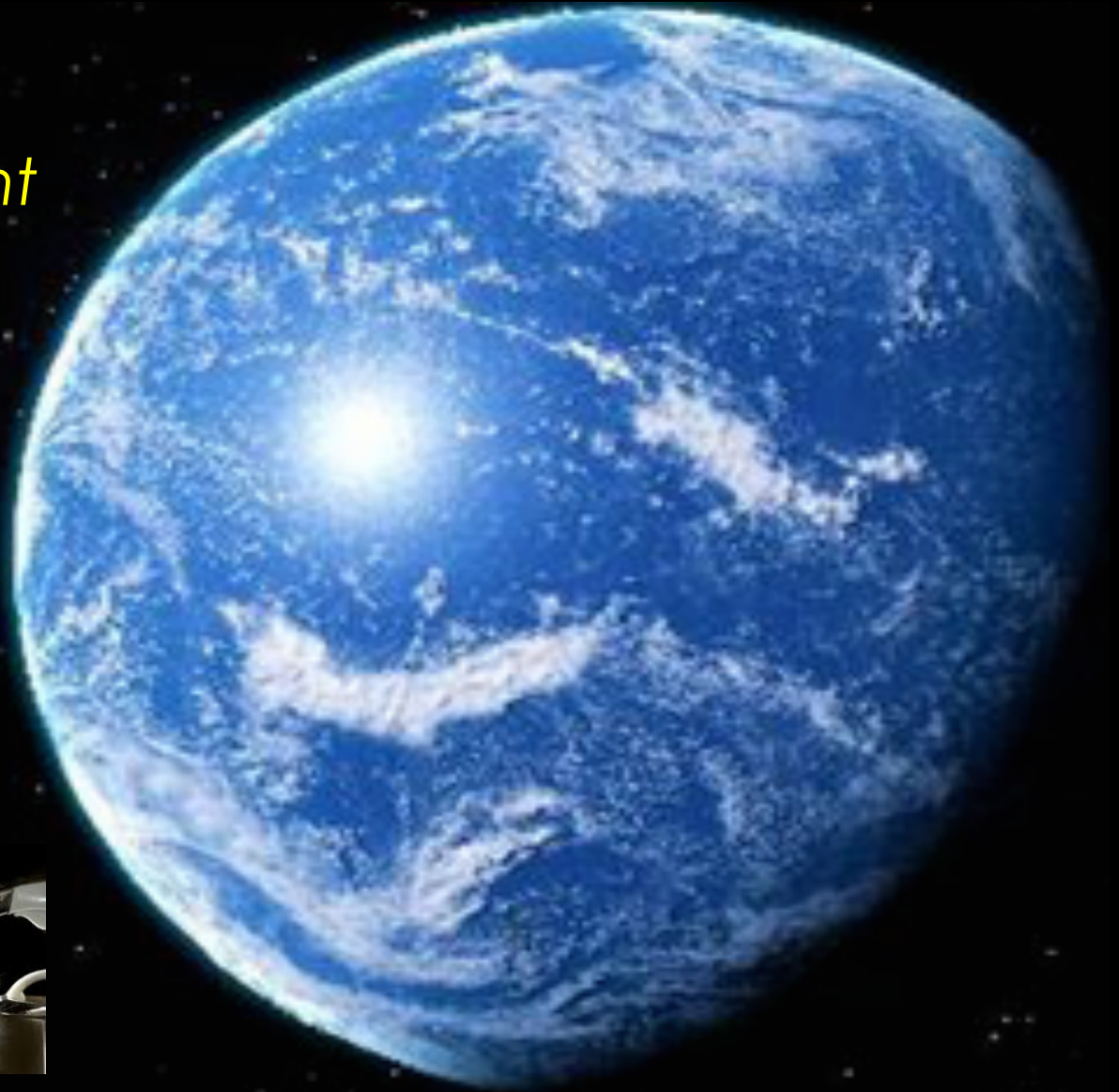
32° F

boiling

373 K

100° C

212° F





A habitable planet is ...  
*ocean covered.*

*H<sub>2</sub>O dominates climate*

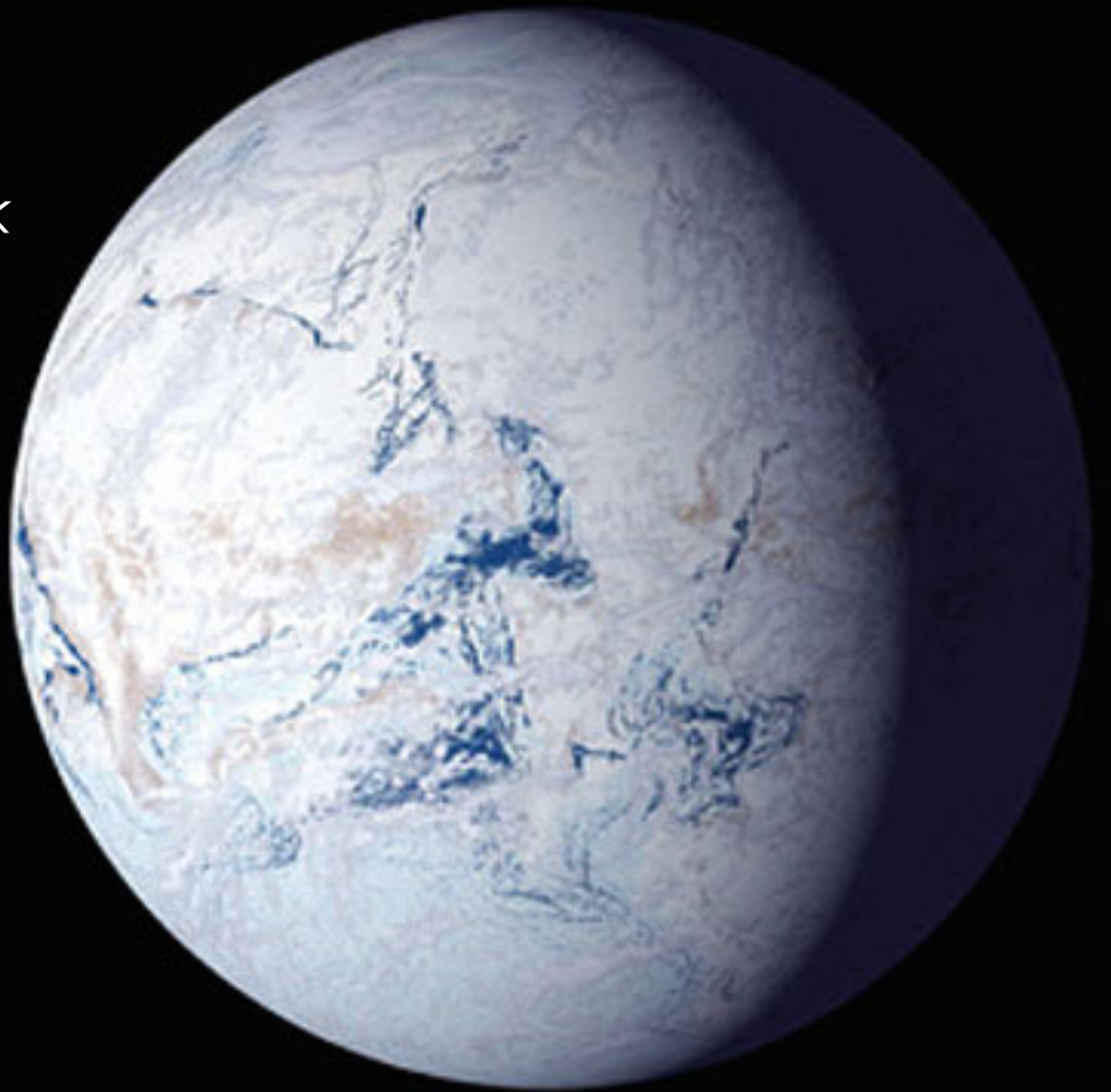
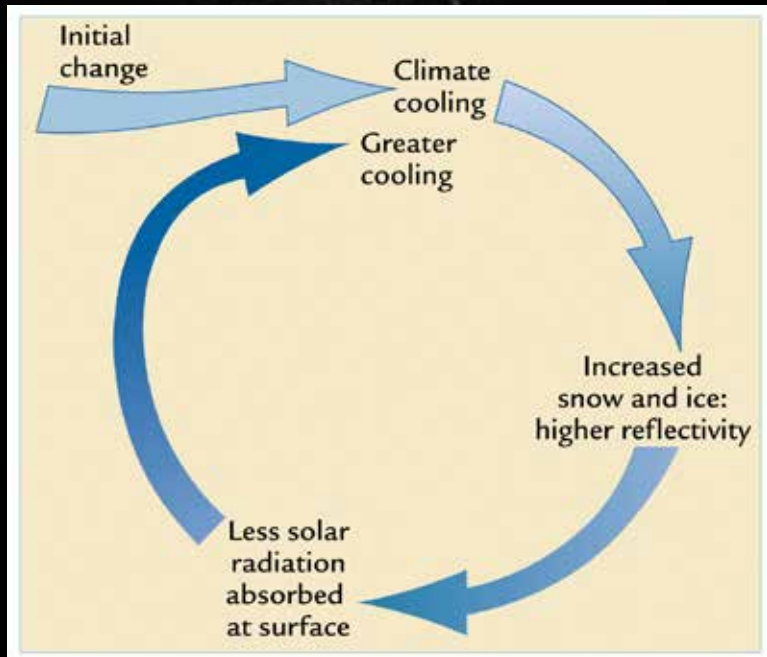
*H<sub>2</sub>O causes strong  
positive climatic  
feedbacks.*





# A habitable planet is *not* ... ...a snowball Earth

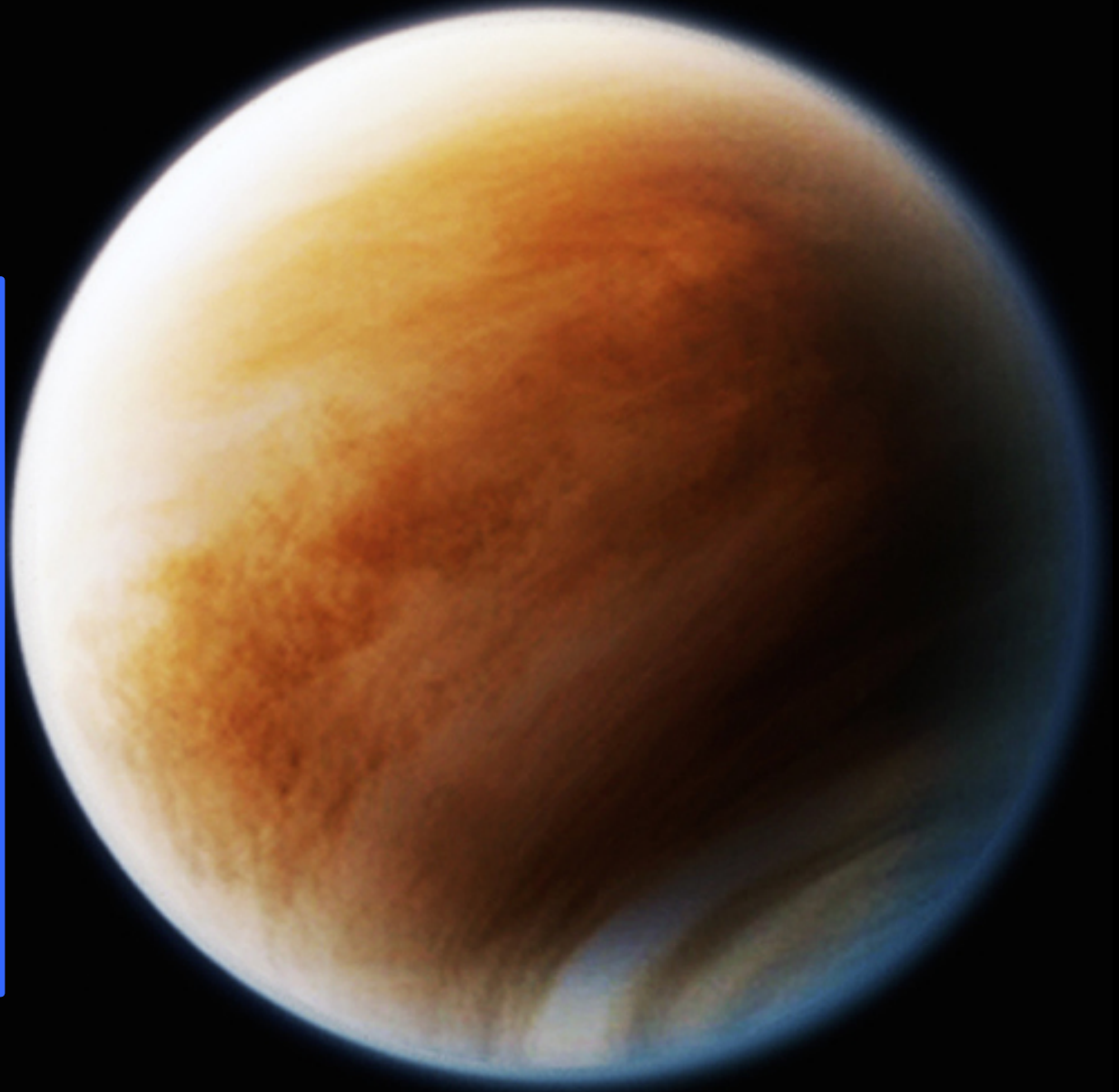
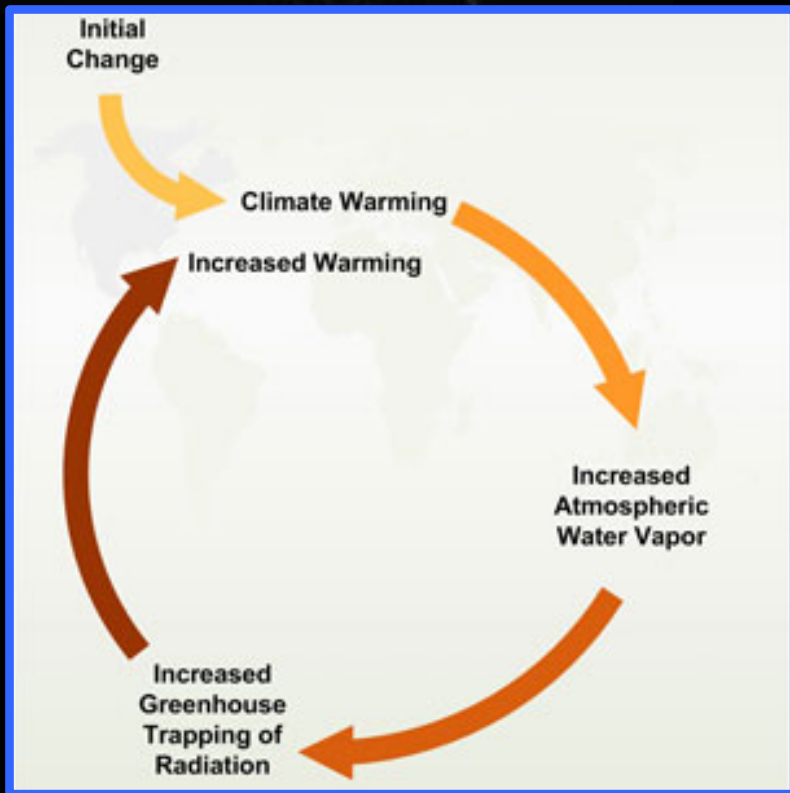
*Sea ice albedo feedback*





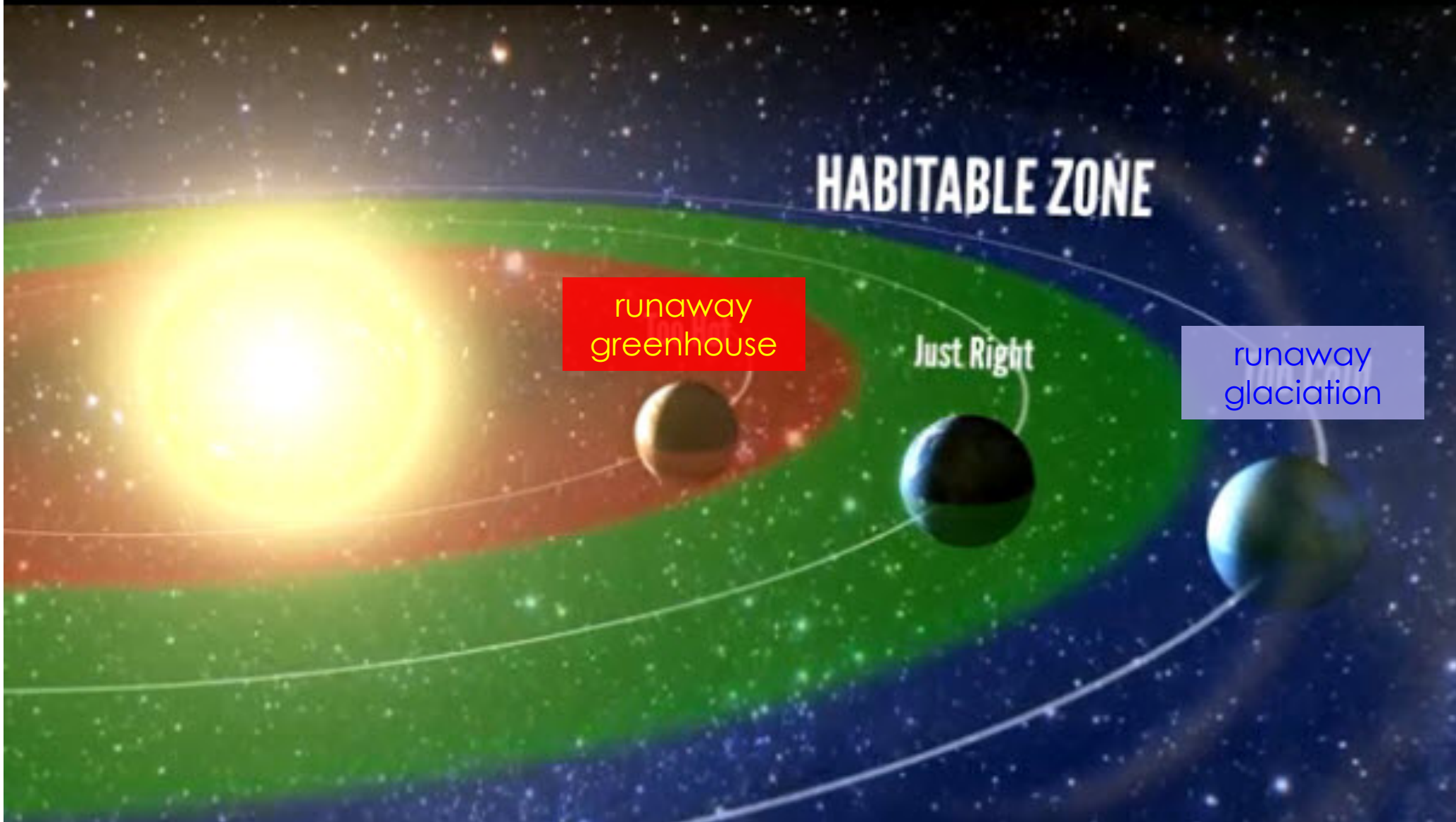
A habitable planet is *not* ...  
...a runaway greenhouse.

Water vapor greenhouse  
feedback





“Habitable Zone” -- the region in space surrounding a star where  $\text{N}_2$ - $\text{CO}_2$ - $\text{H}_2\text{O}$  composed atmospheres can sustain surface liquid water.





# What do we know about habitable zone extrasolar planets?





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# What do we know about habitable zone extrasolar planets?

## *Things we know*

mass  
radius  
orbital period  
incident stellar flux  
stellar spectra



## *Things we don't know*

atmospheric composition  
surface properties  
rotation rate





# What do we know about habitable zone extrasolar planets?

## *Things we know*

mass  
radius  
orbital period  
incident stellar flux  
stellar spectra



## *Things we don't know*

atmospheric composition ~ **Earthy\***  
surface properties ~ **ocean covered**  
rotation rate ~ **24 hour diurnal cycle**



\*For determining the habitable zone we assume an  $\text{N}_2 + \text{H}_2\text{O} + \text{CO}_2$  composition



# What do we know about habitable zone extrasolar planets?

## Things we know

mass

radius

orbital period

incident stellar flux *Large changes in the  
stellar spectra solar input!!*

## Things we don't know

atmospheric composition ~ **Earthy\***

surface properties ~ **ocean covered**

rotation rate ~ **24 hour diurnal cycle**

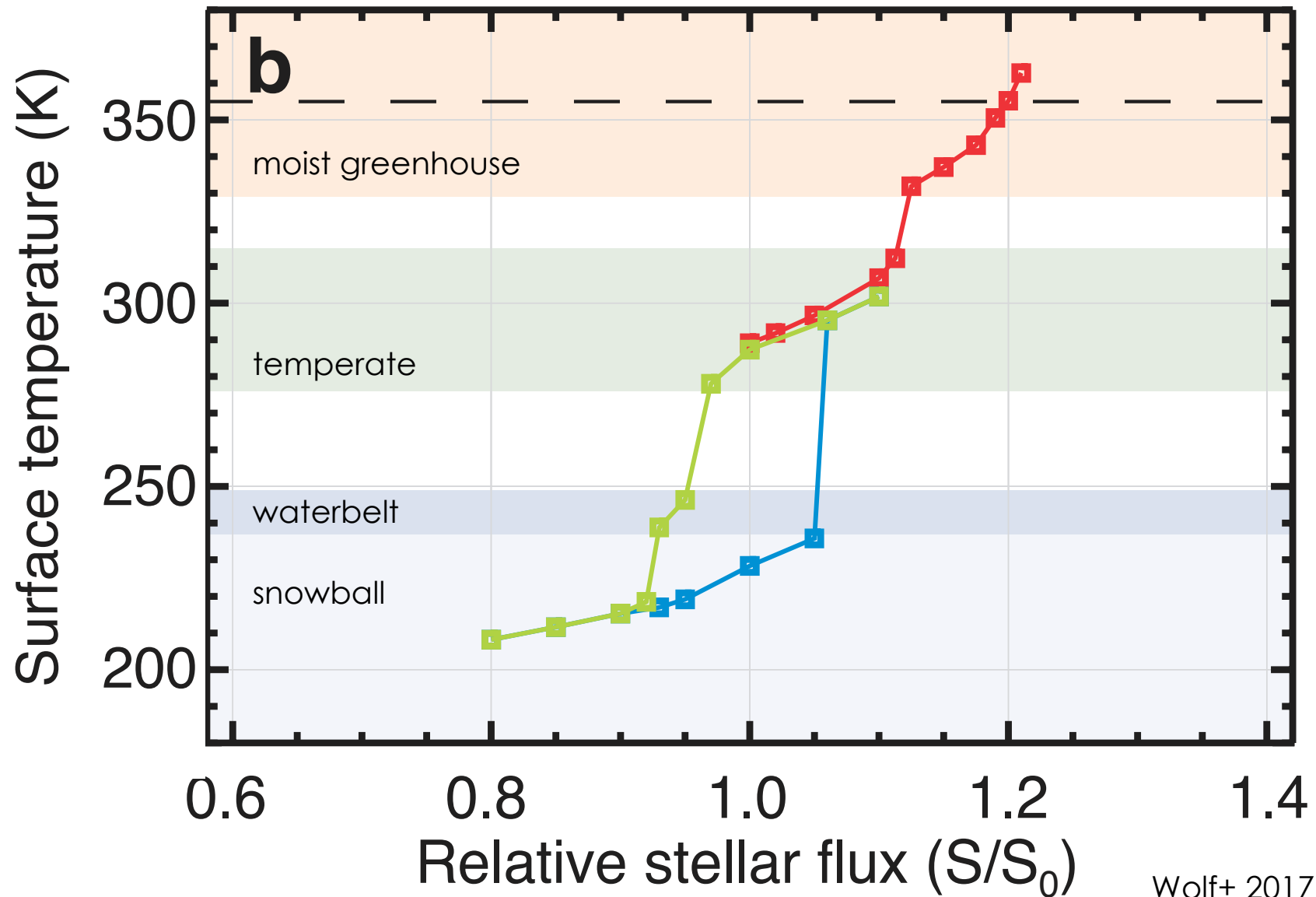
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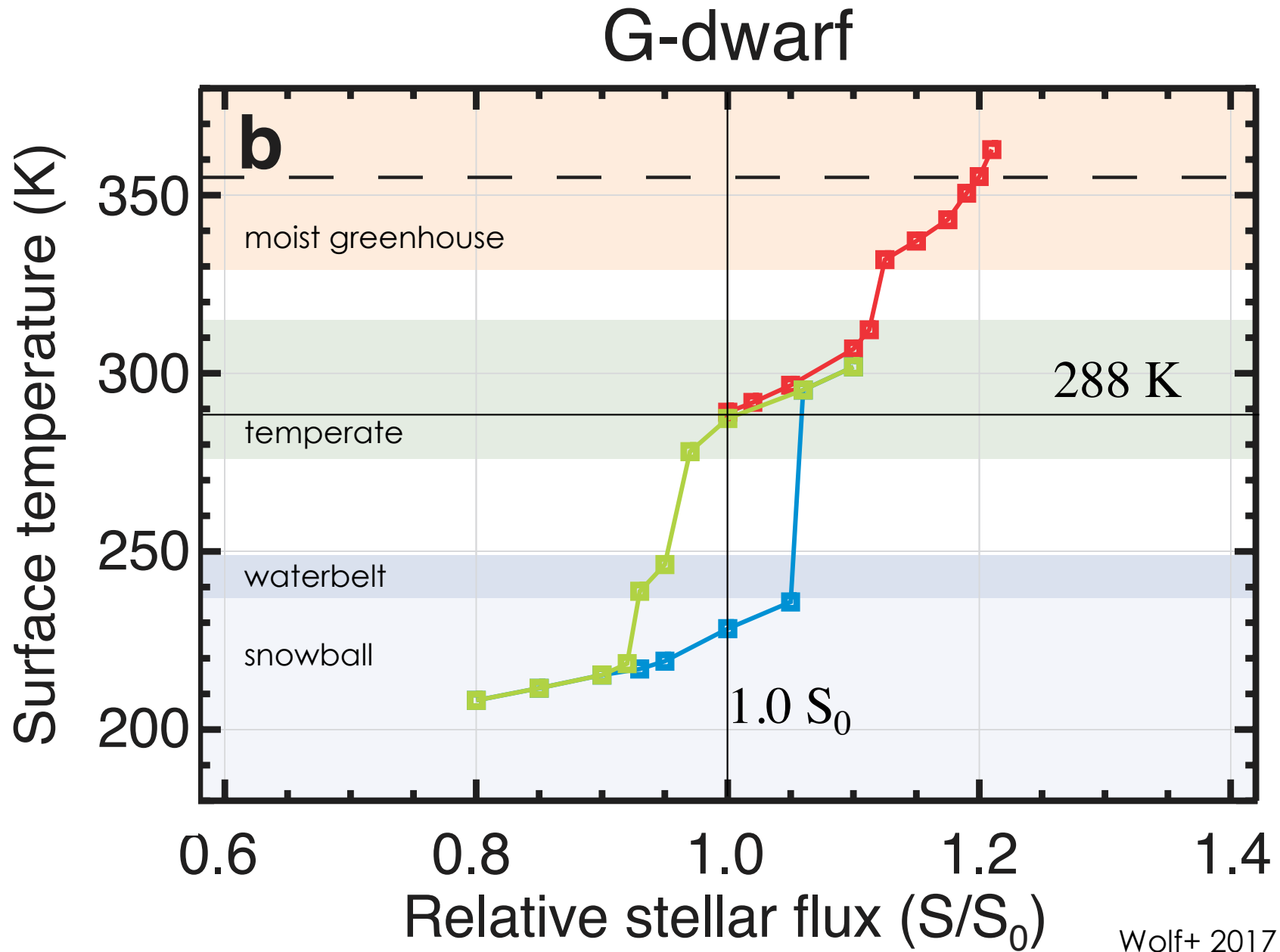
Let's start with the simplest possible problem...  
Earth, modern CO<sub>2</sub>, around the Sun, at various TSI

## G-dwarf





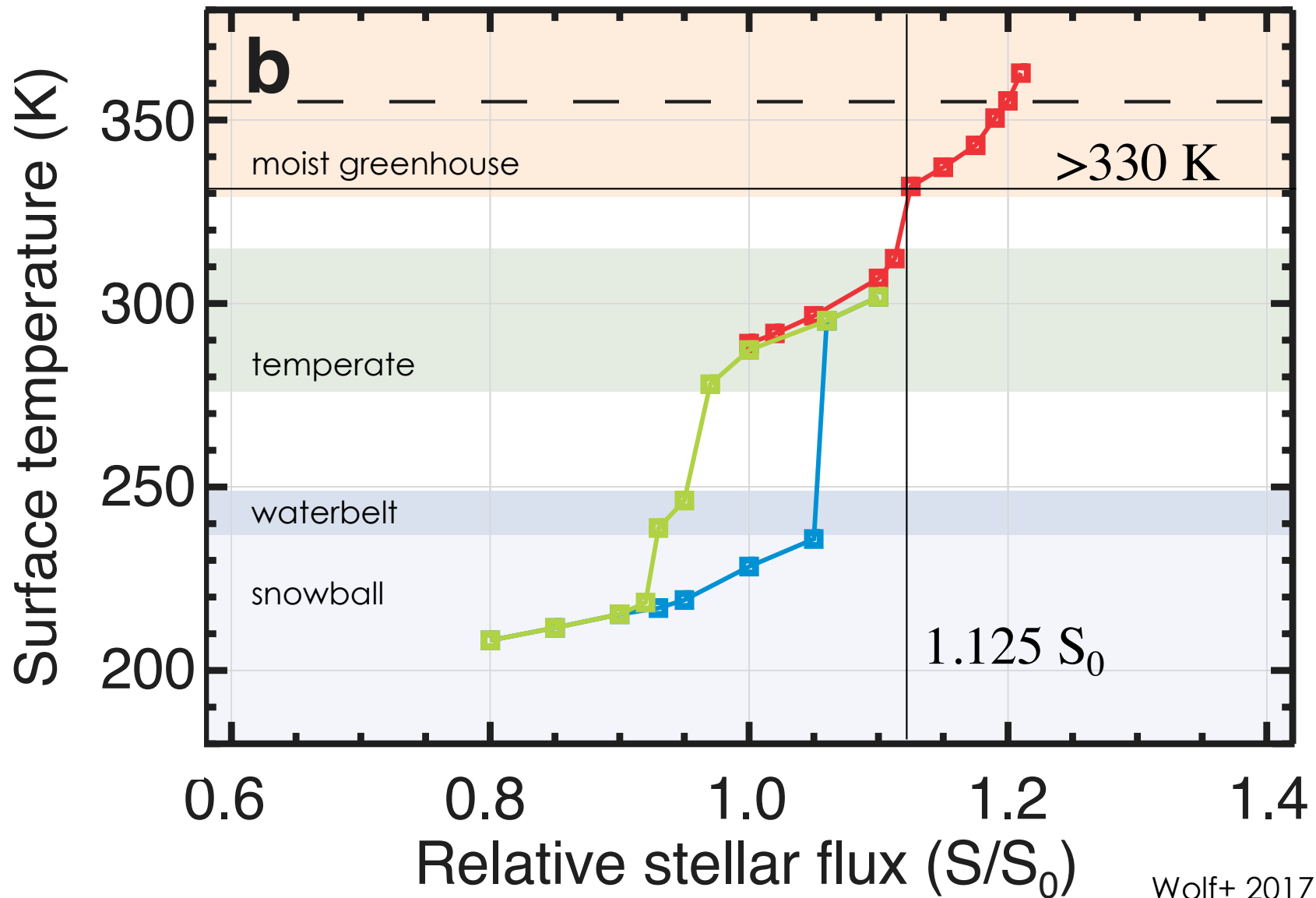
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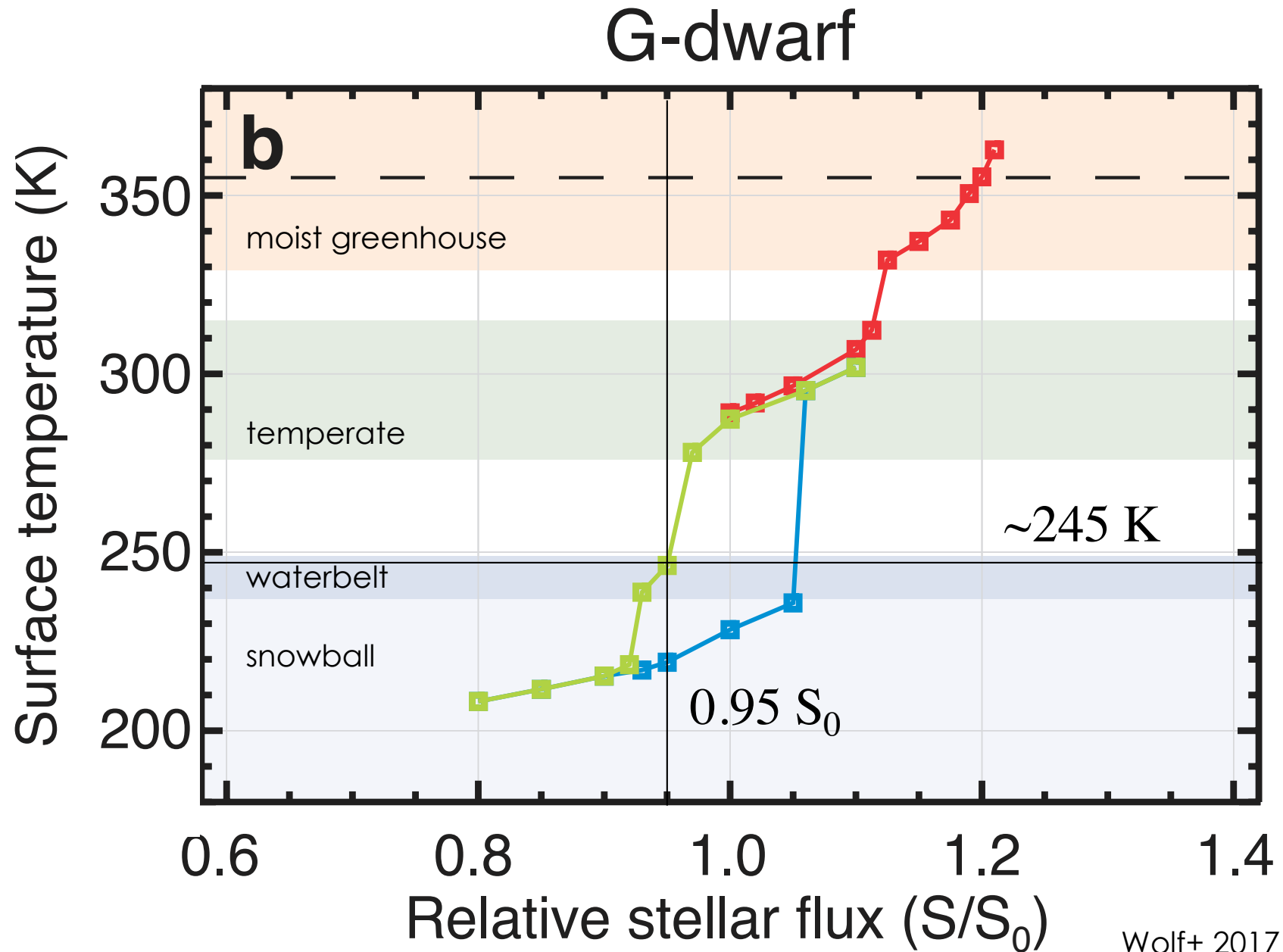
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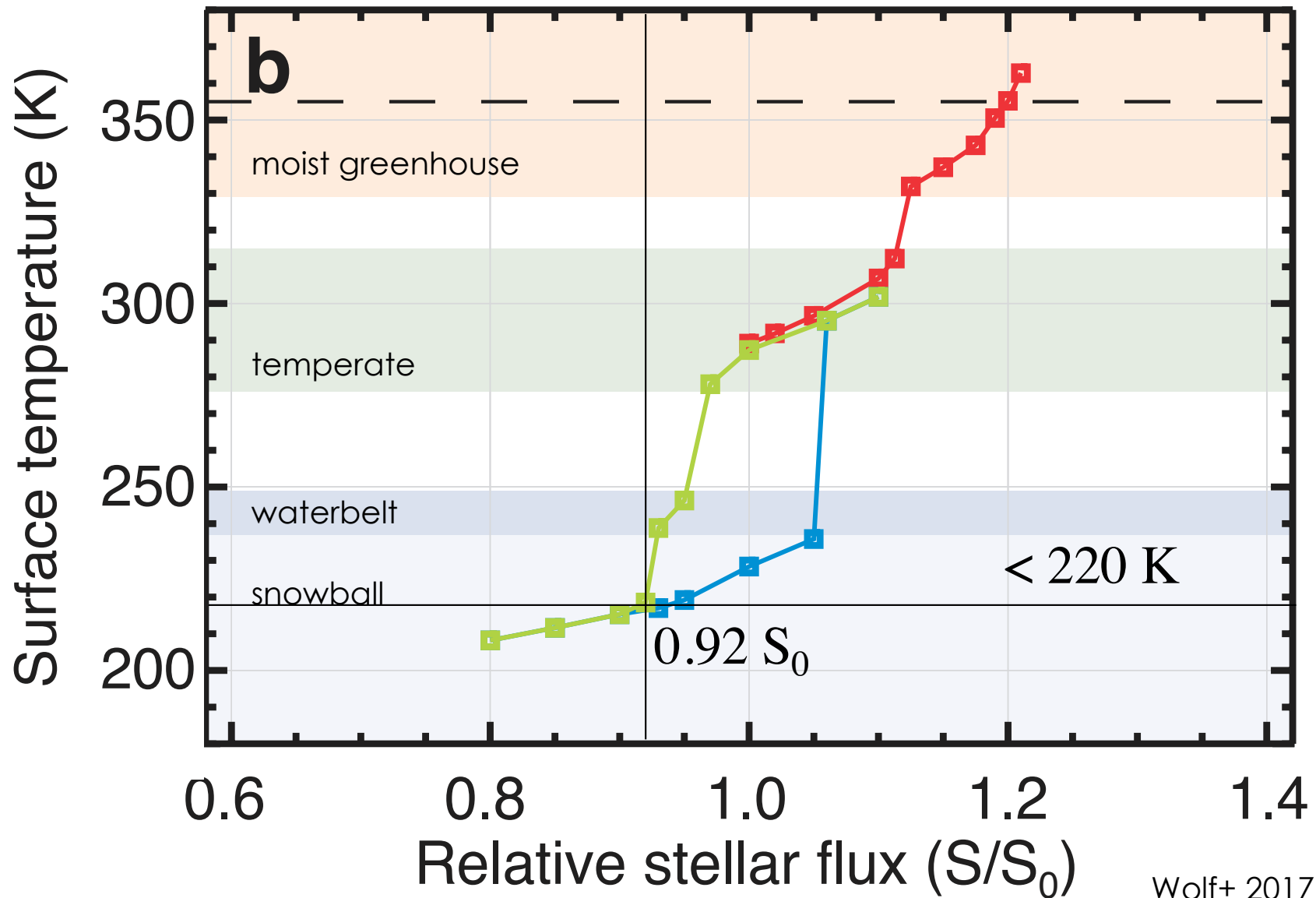
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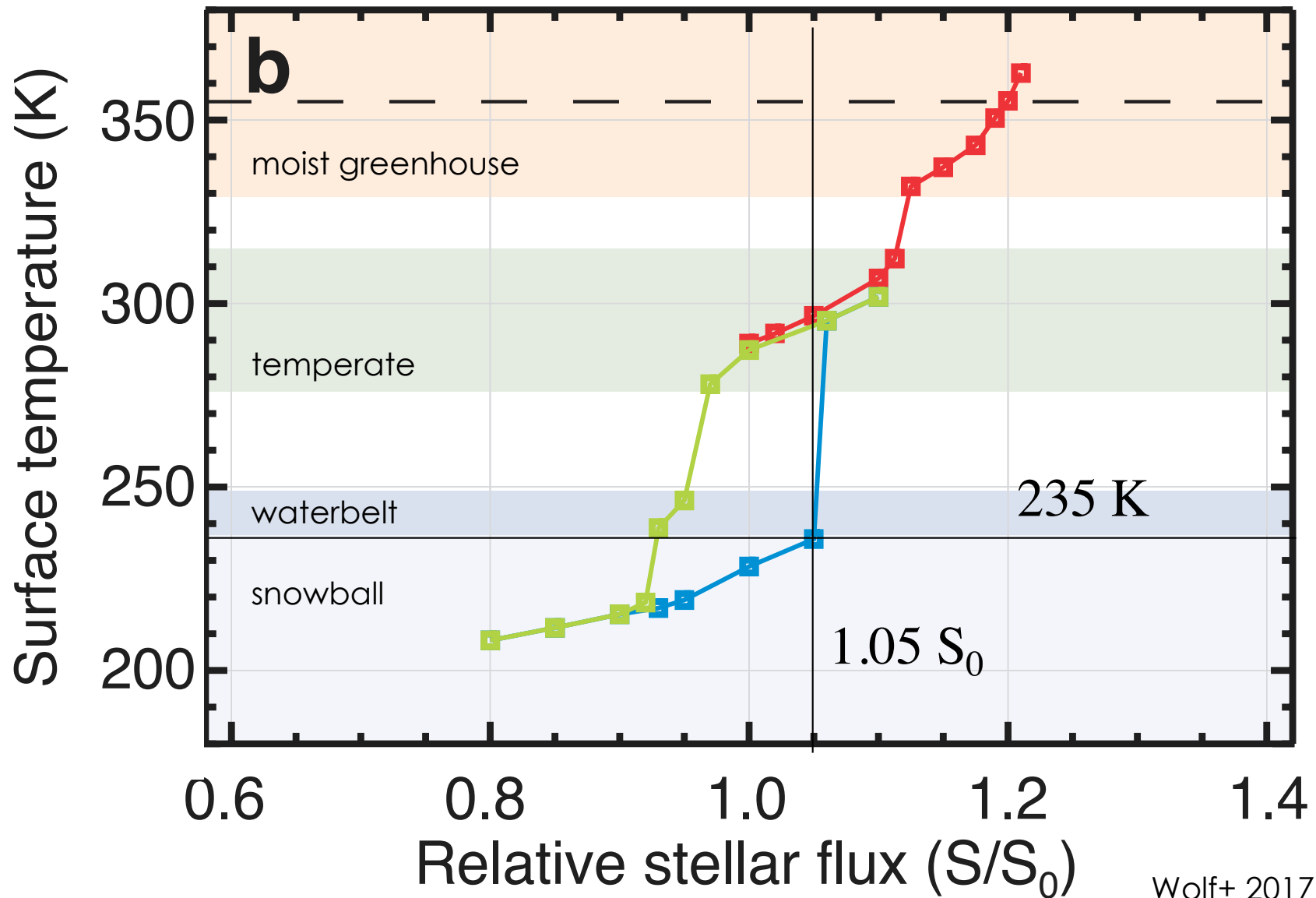
## G-dwarf





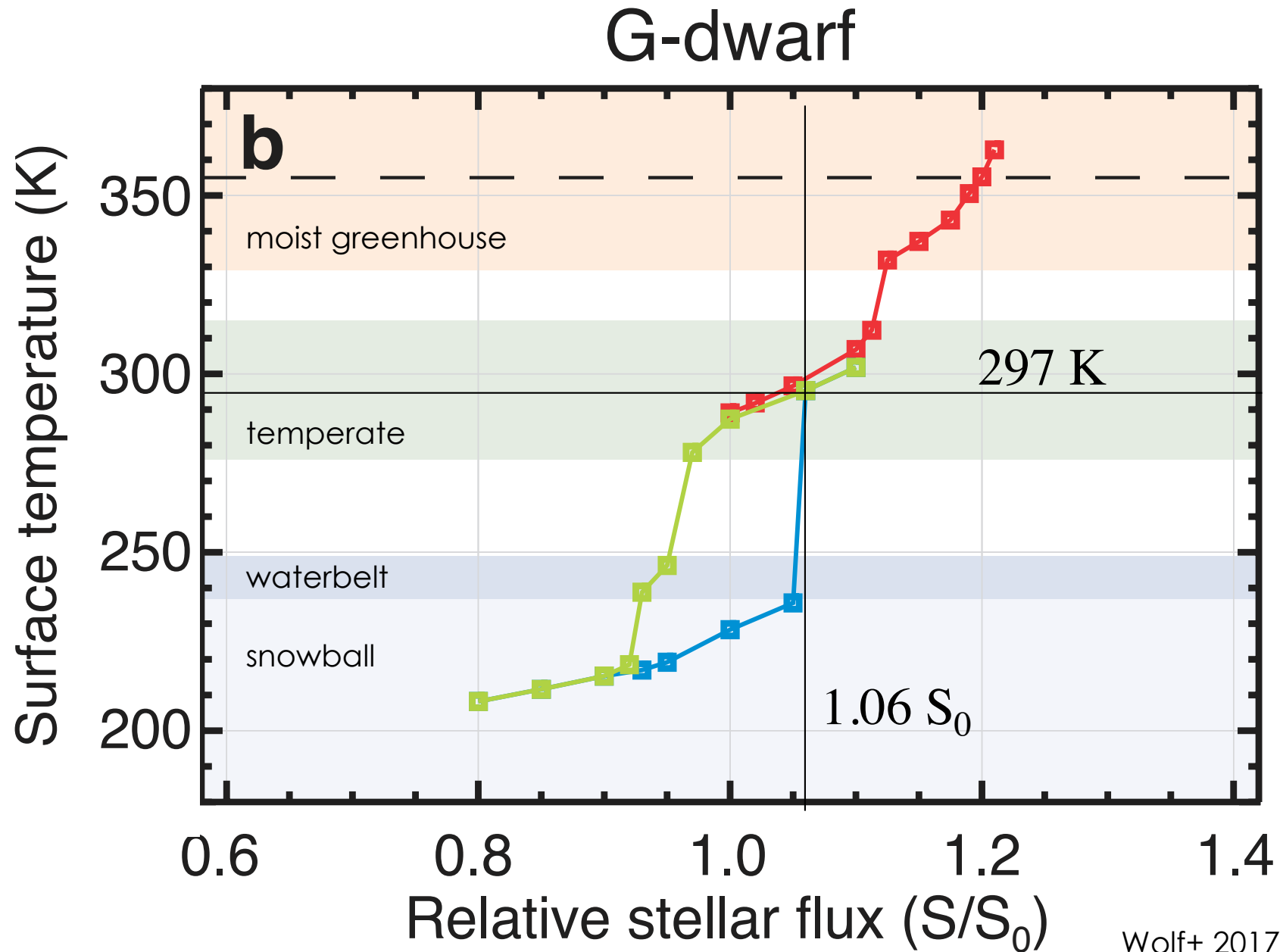
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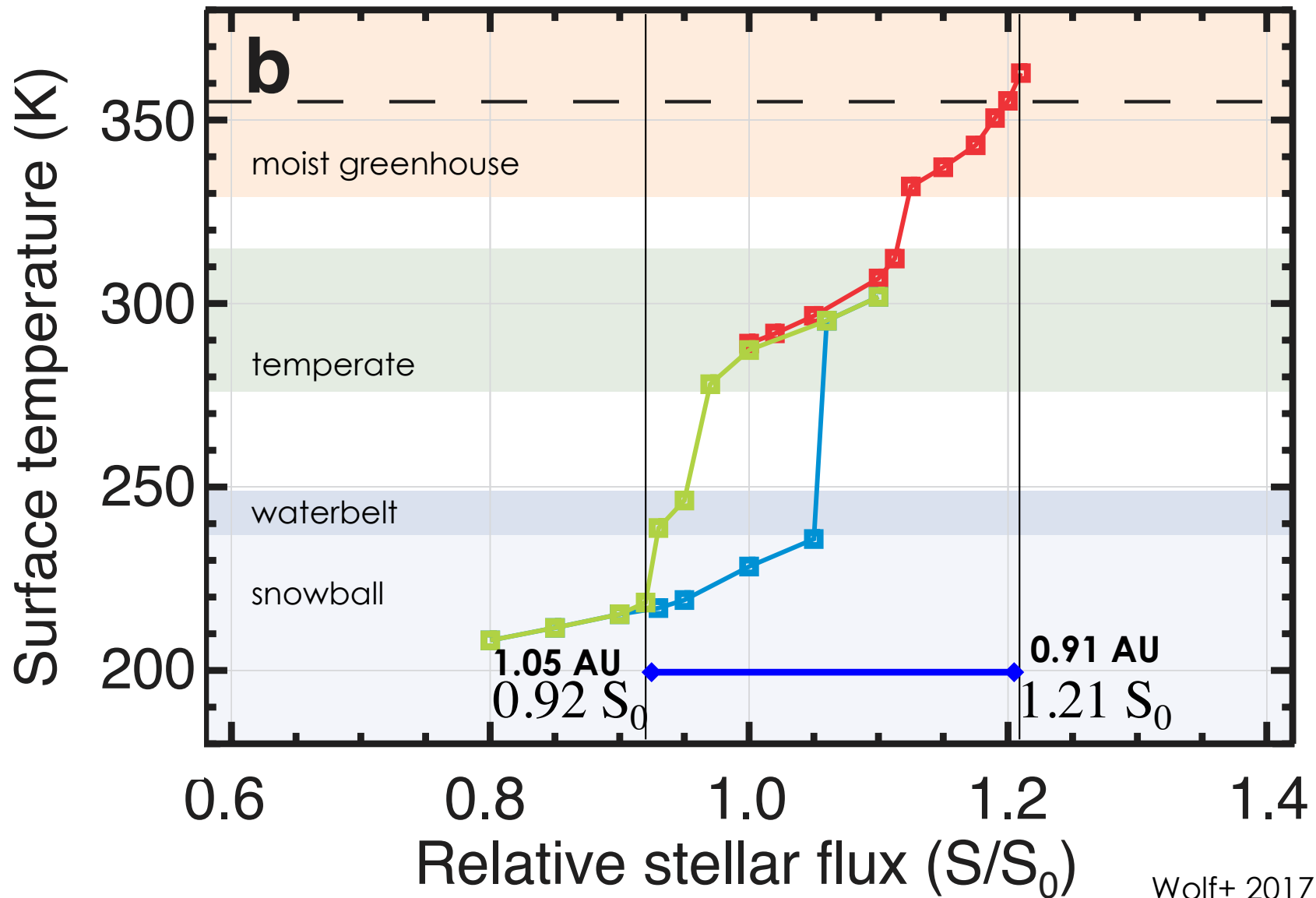
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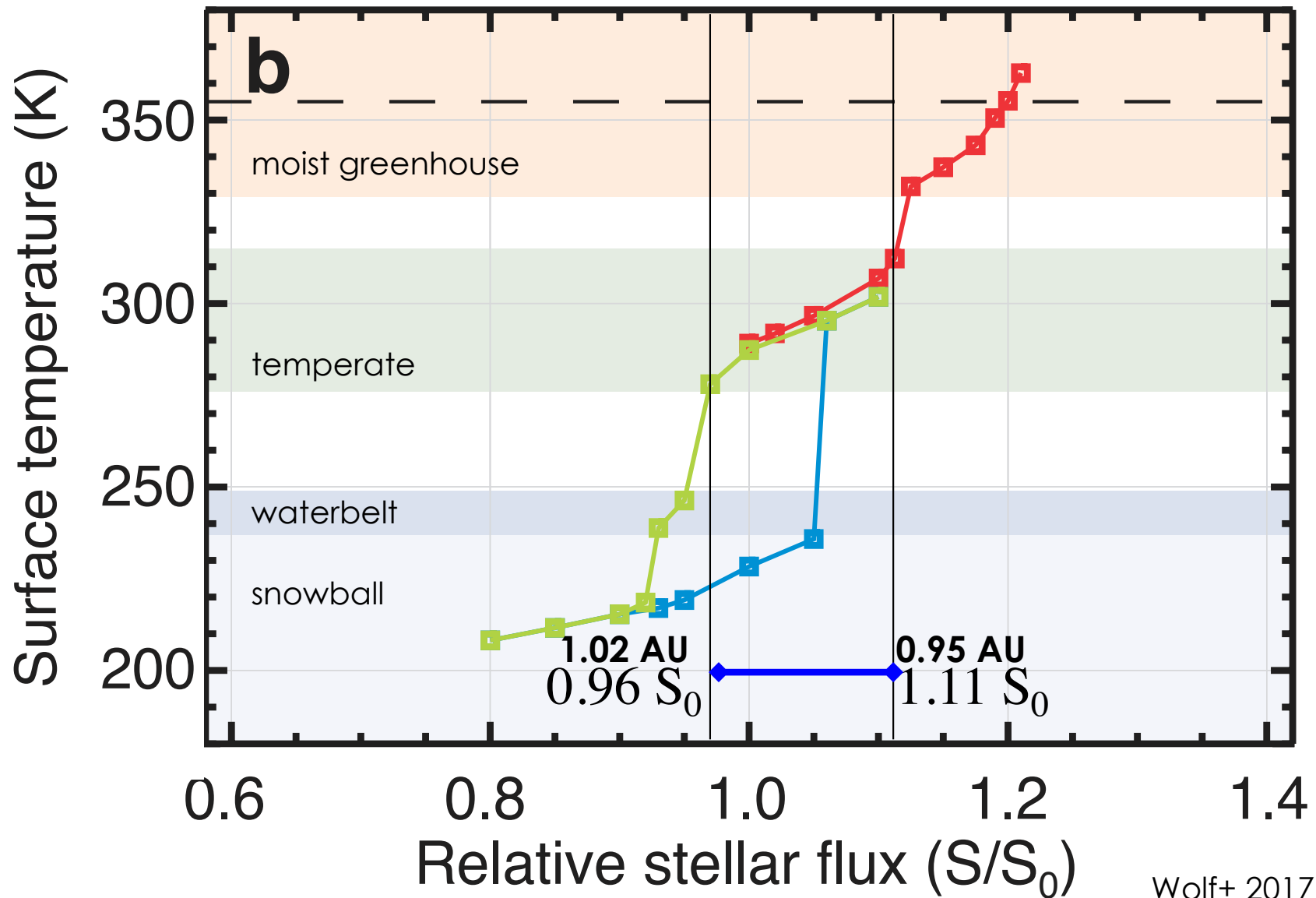
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## G-dwarf





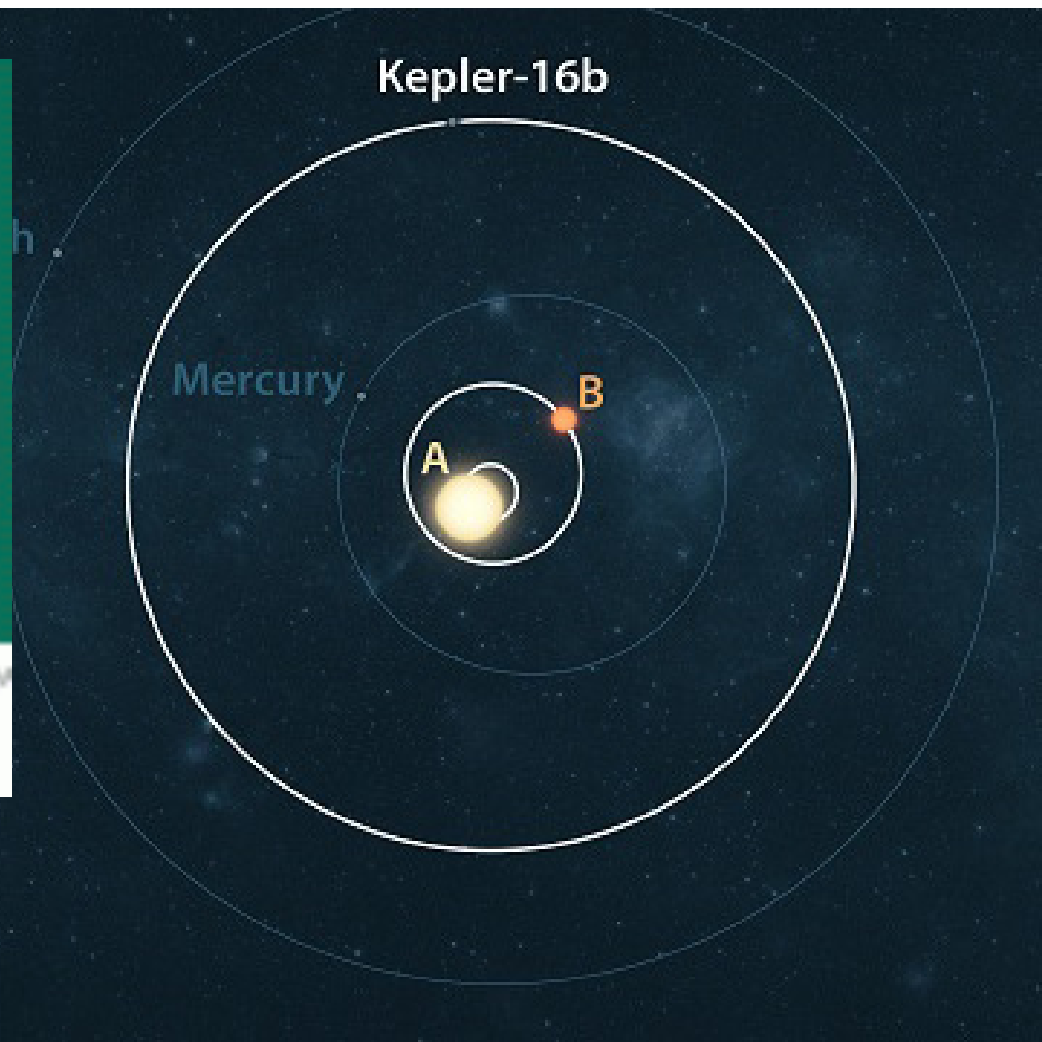
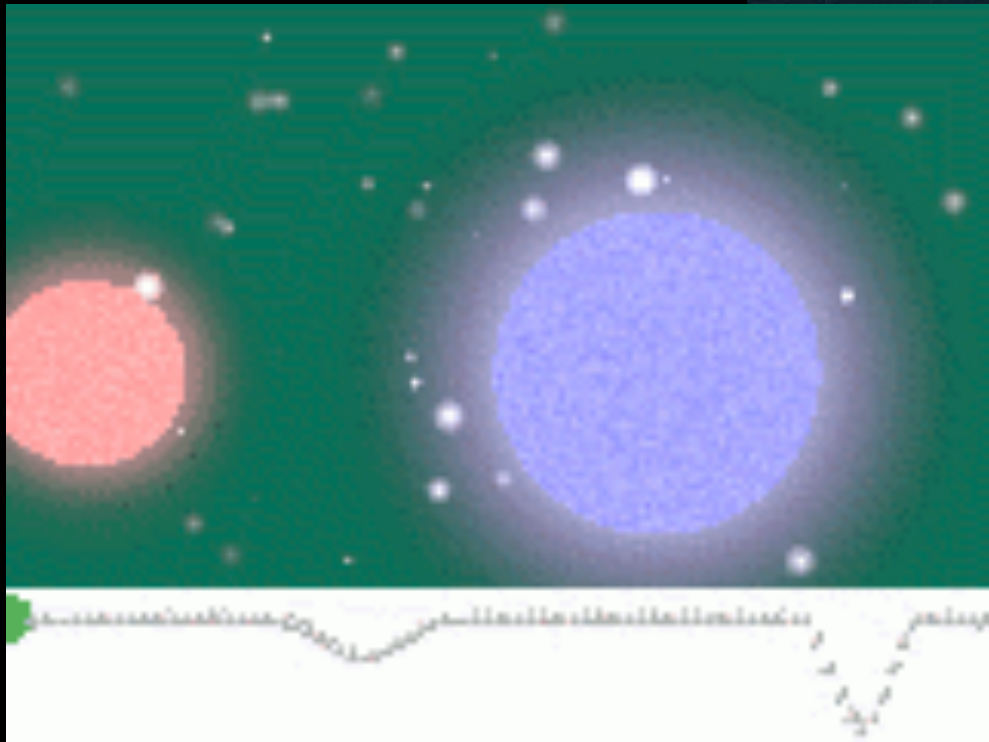
A digital illustration of a binary star system. Two bright, yellow-orange stars are positioned in the upper center, appearing to be in the process of merging or in a very close orbit. The stars have a glowing, hazy atmosphere. Below them, the curved horizon of the Earth is visible, showing blue oceans and green landmasses. The background is a deep black space filled with numerous small, distant stars and a faint, diffuse nebula or interstellar dust cloud in shades of blue and purple.

**Many Sun-like stars exist in binary pairs.**

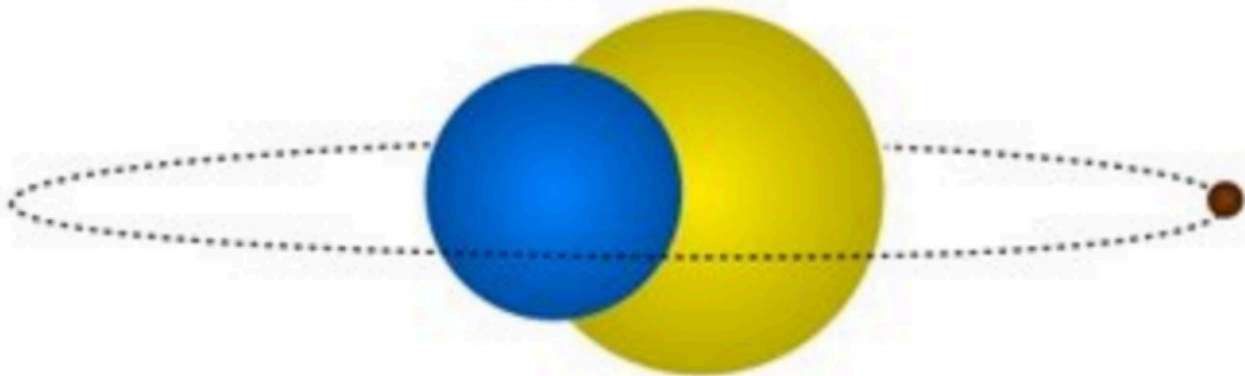




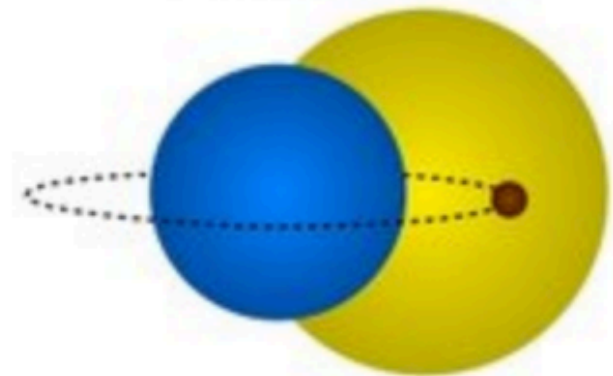




P-type planet



S-type planet



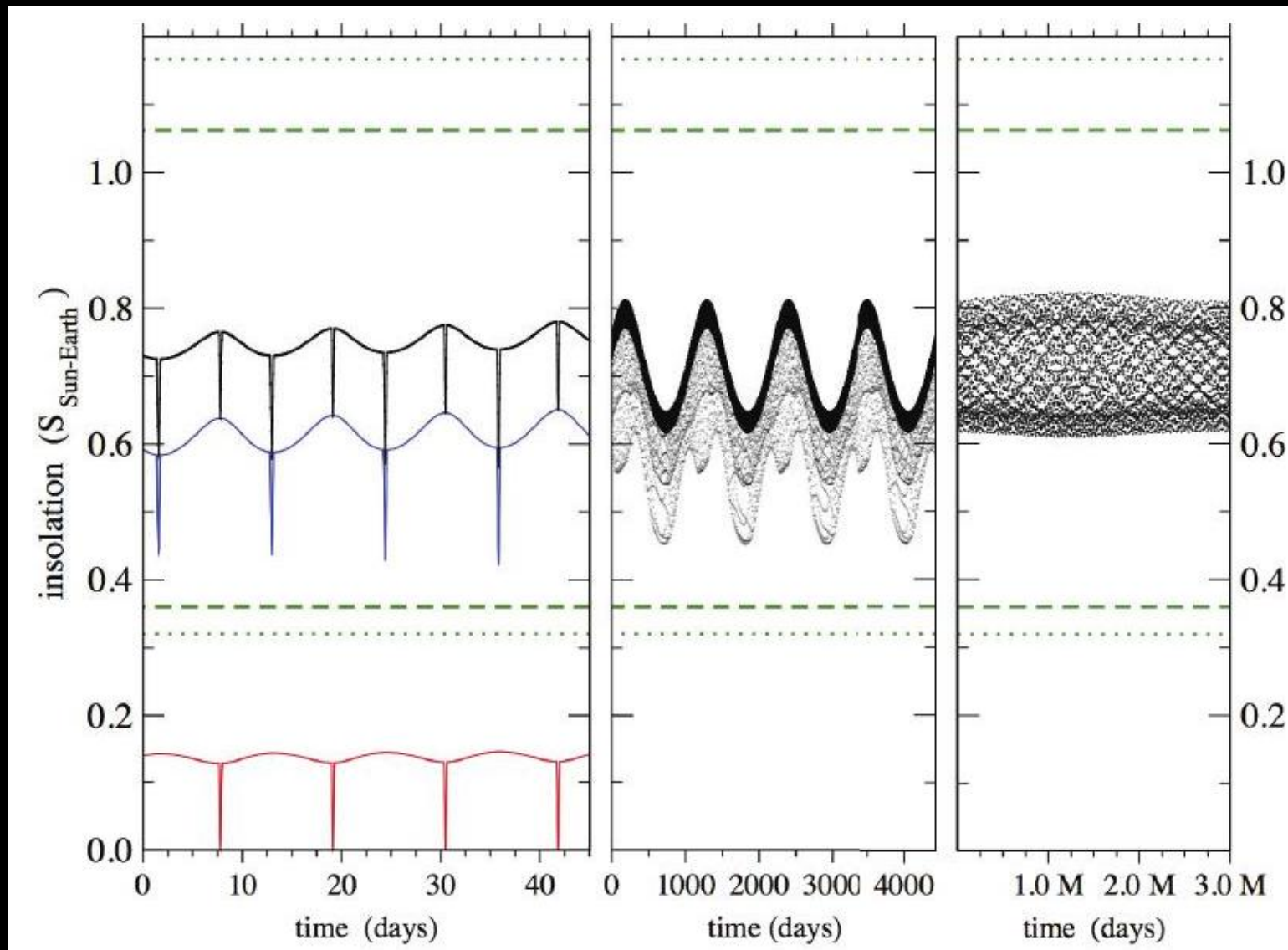


# Variability of TSI in Kepler-1647 System (P-type)

Short term

Medium term

Long term



Kostov+ 2016

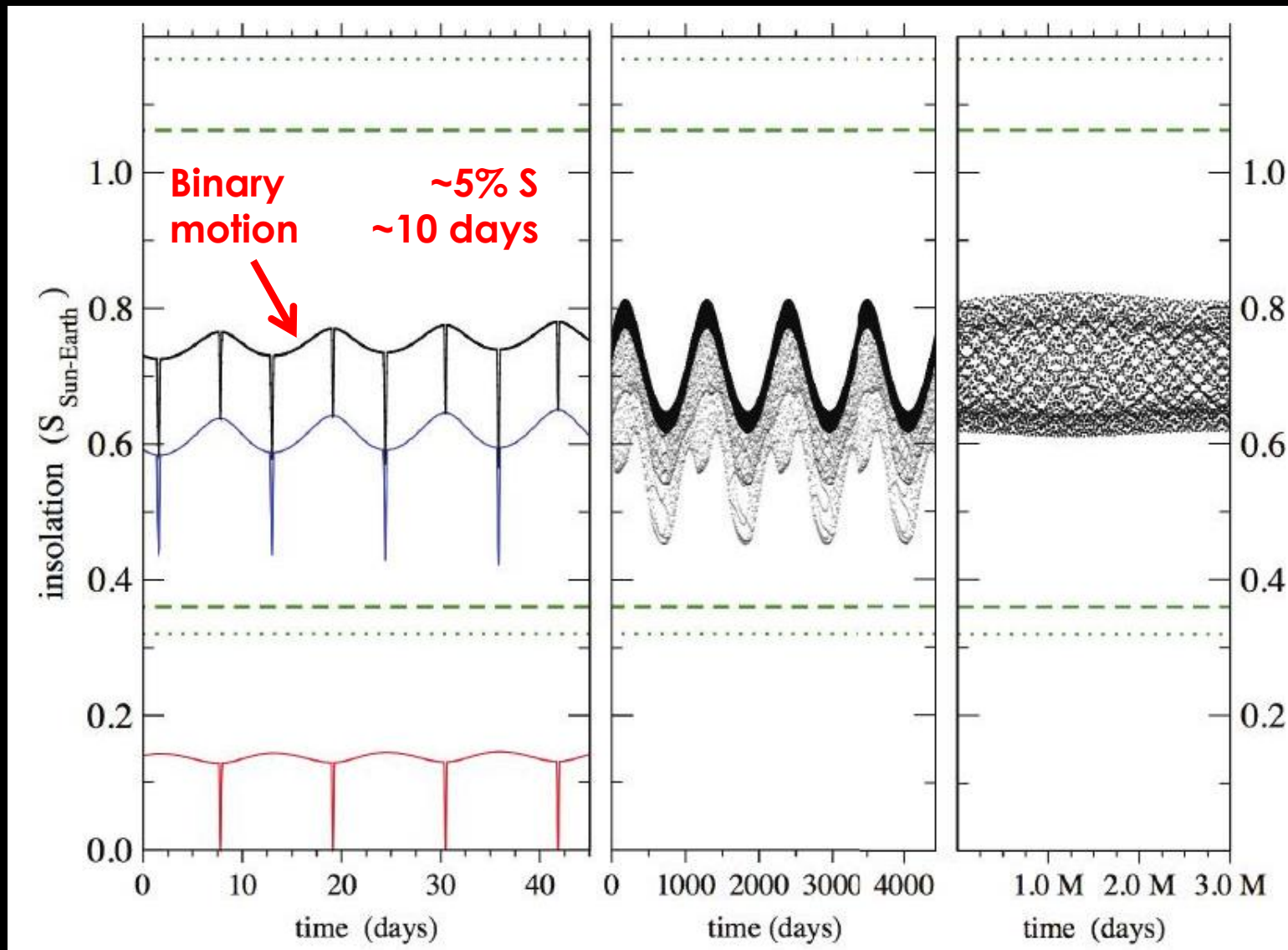


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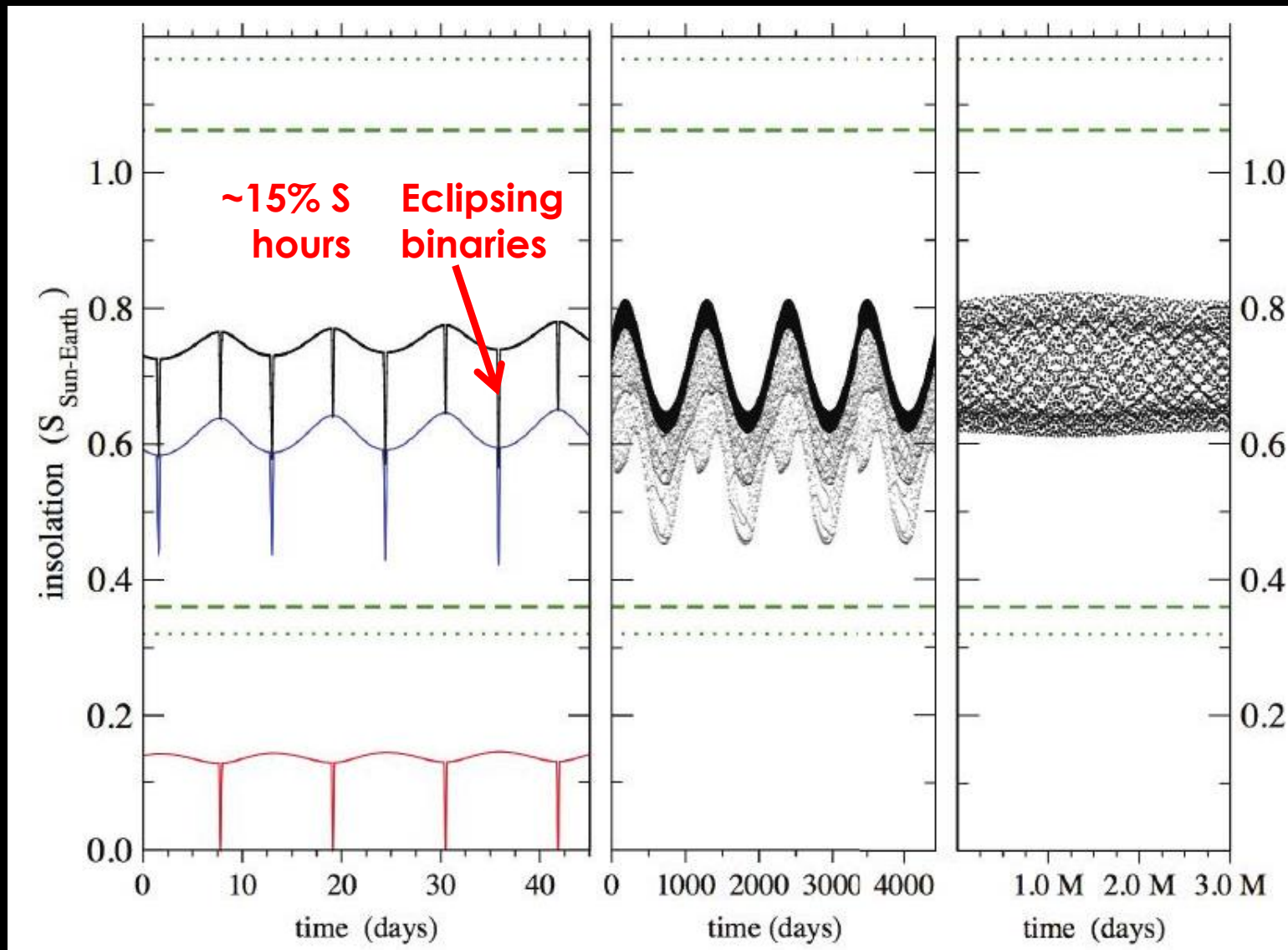


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Kostov+ 2016

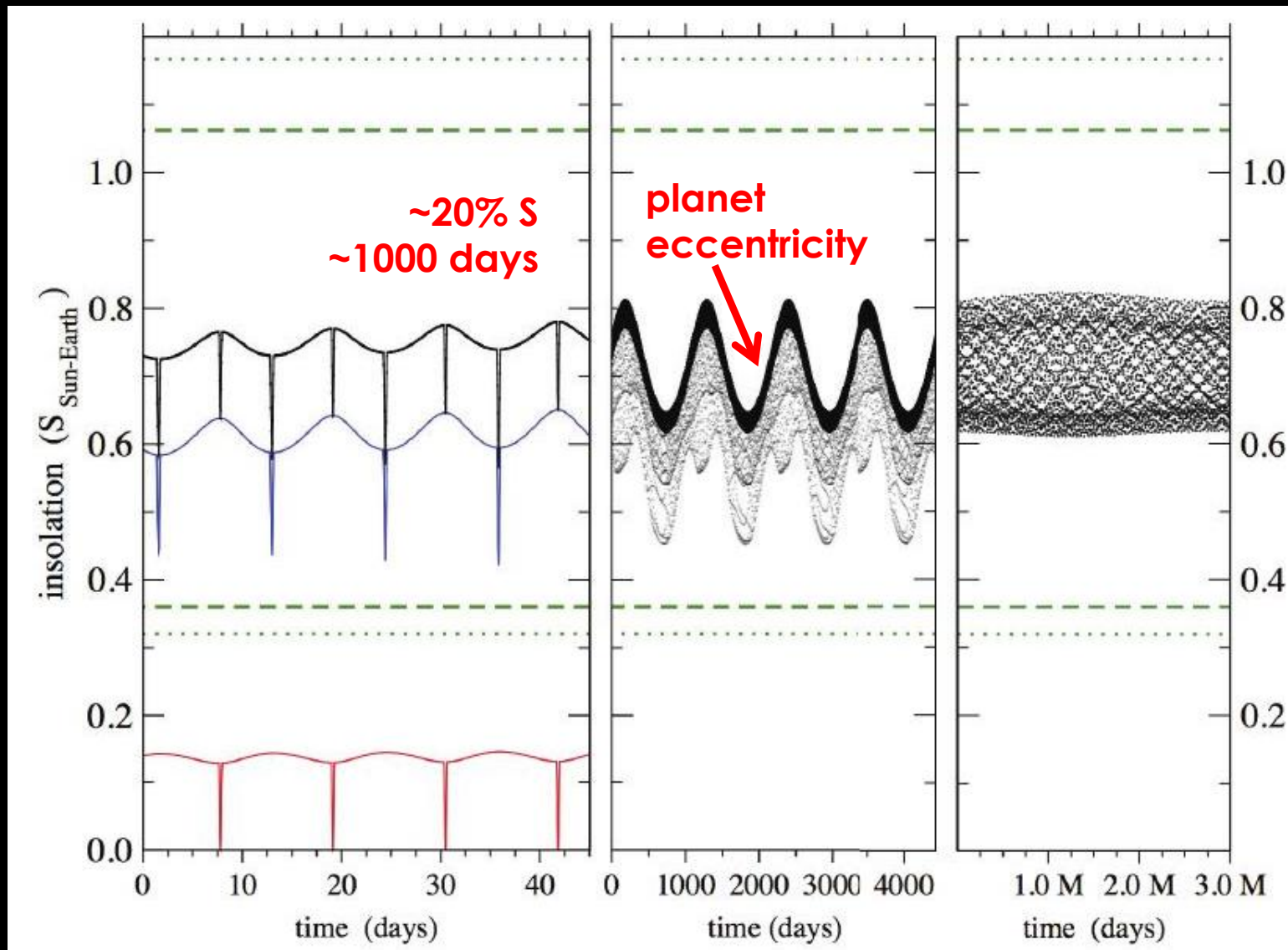


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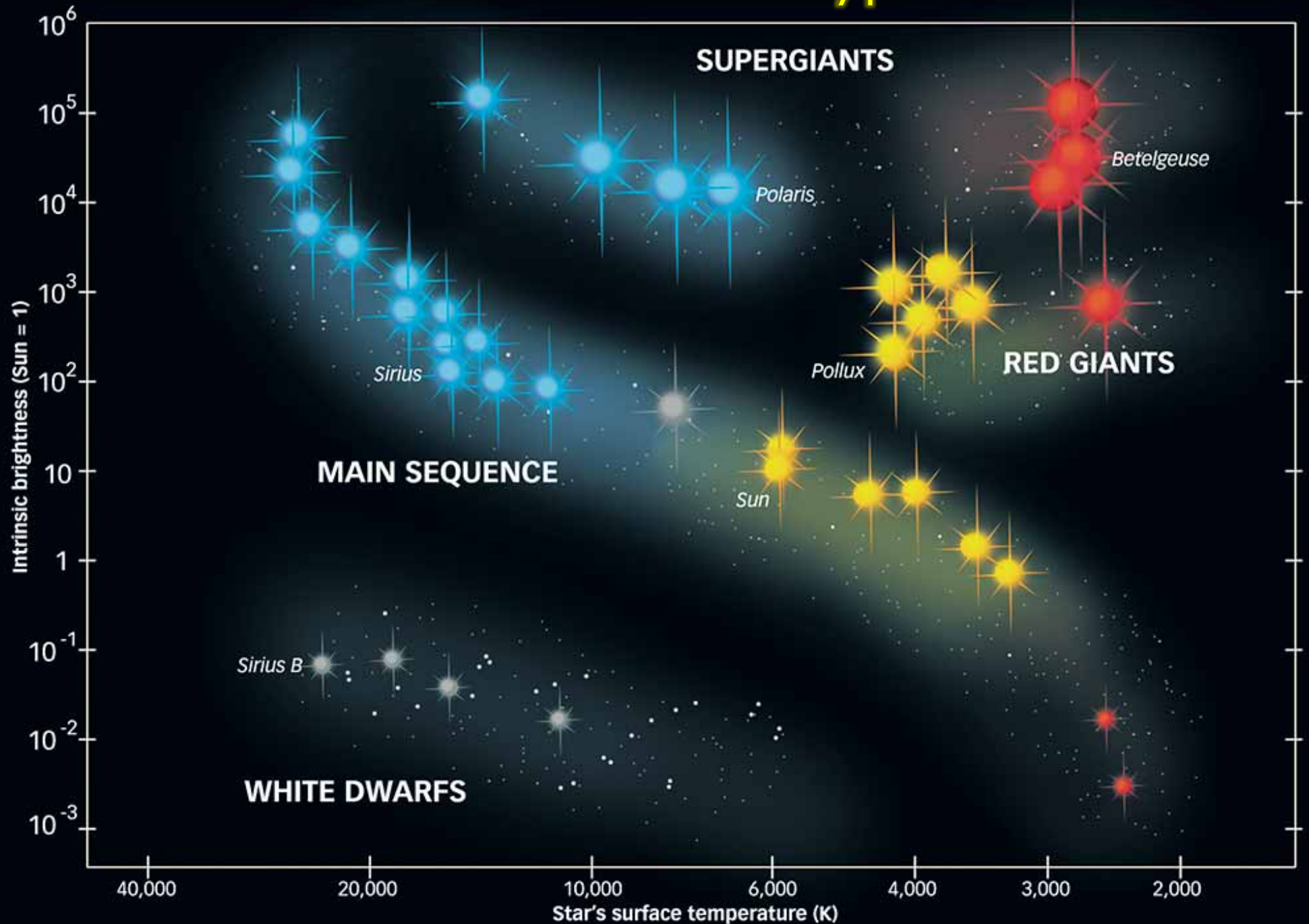
Long term



Kostov+ 2016

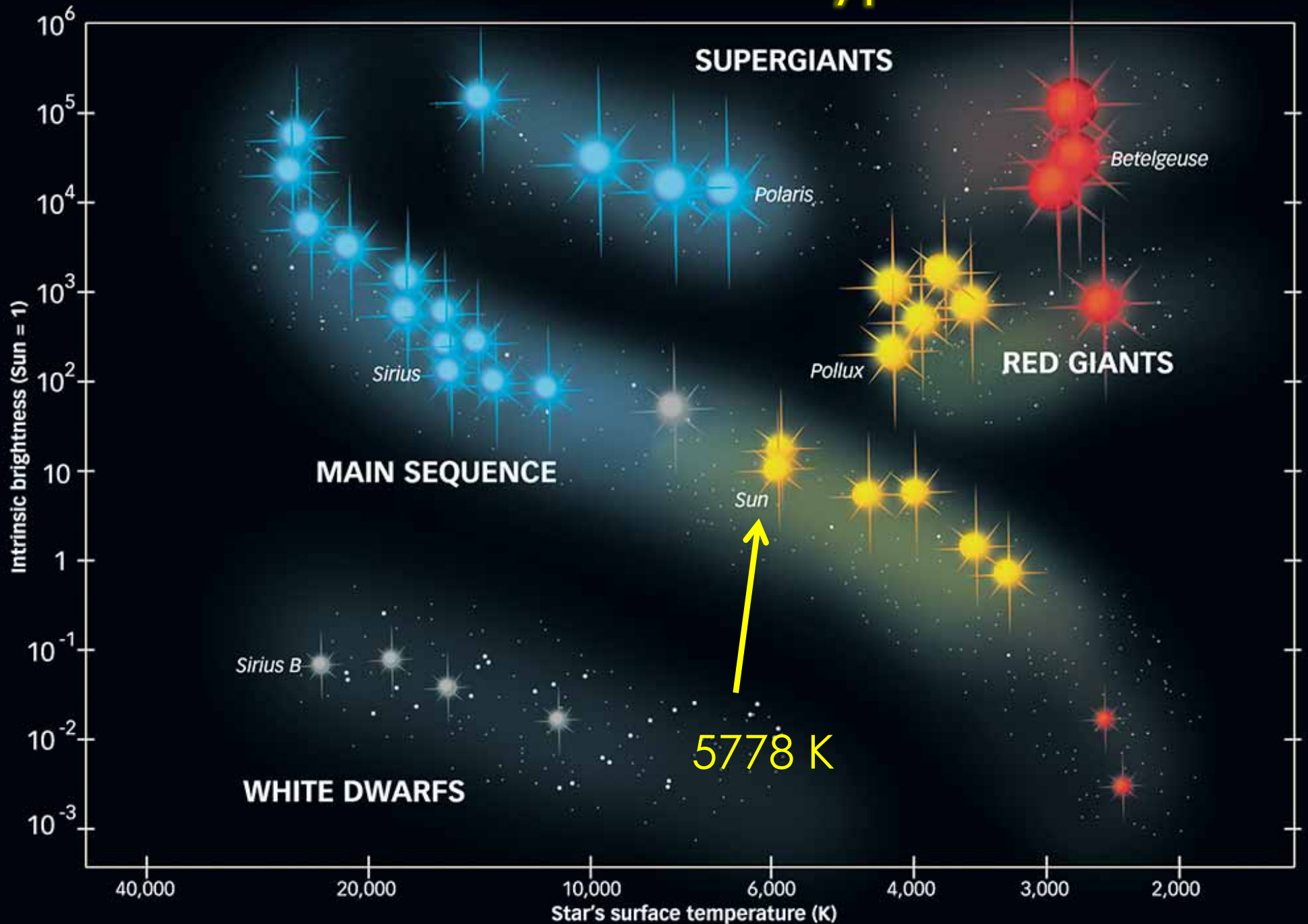


# What about other stellar types?



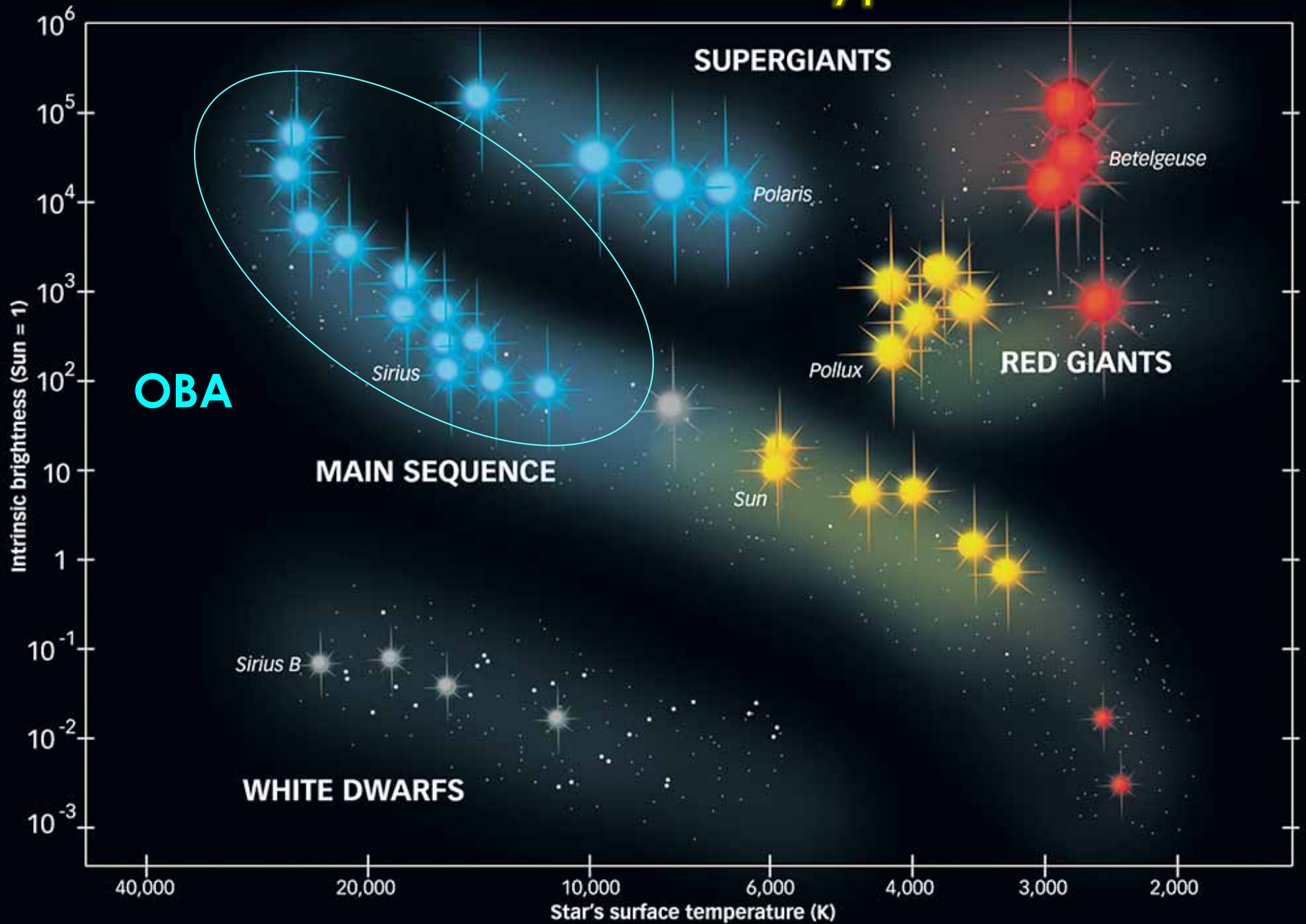


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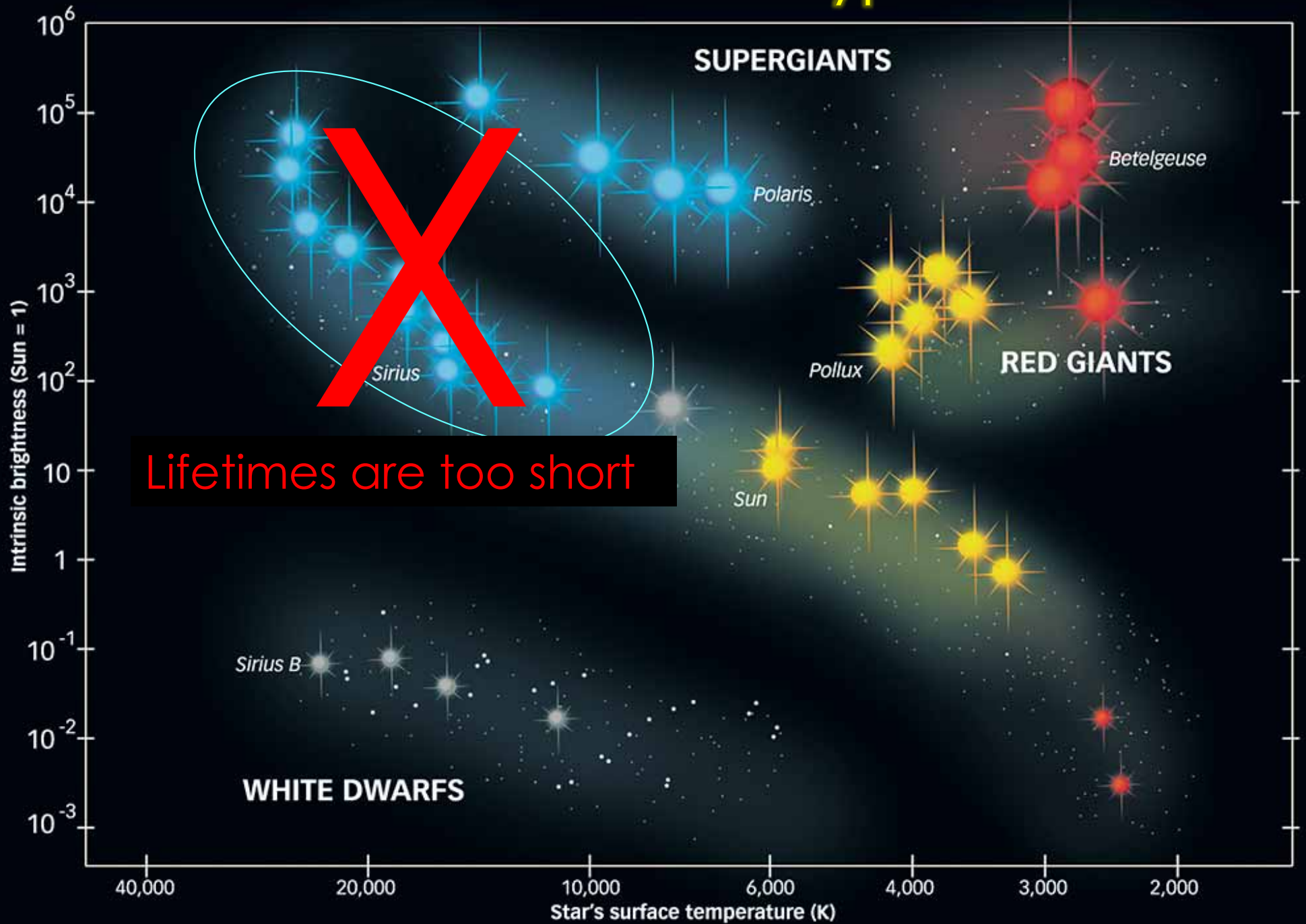


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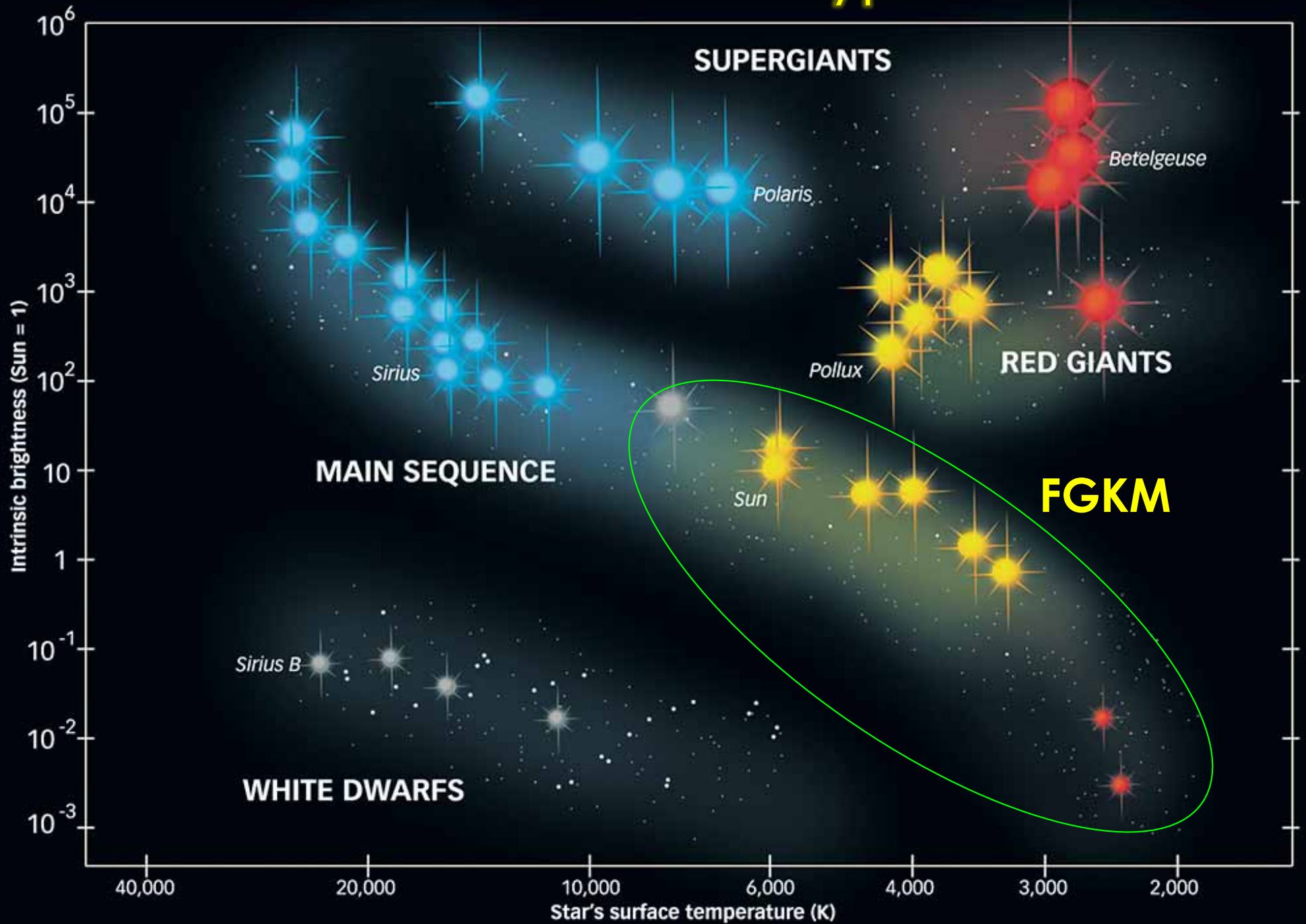


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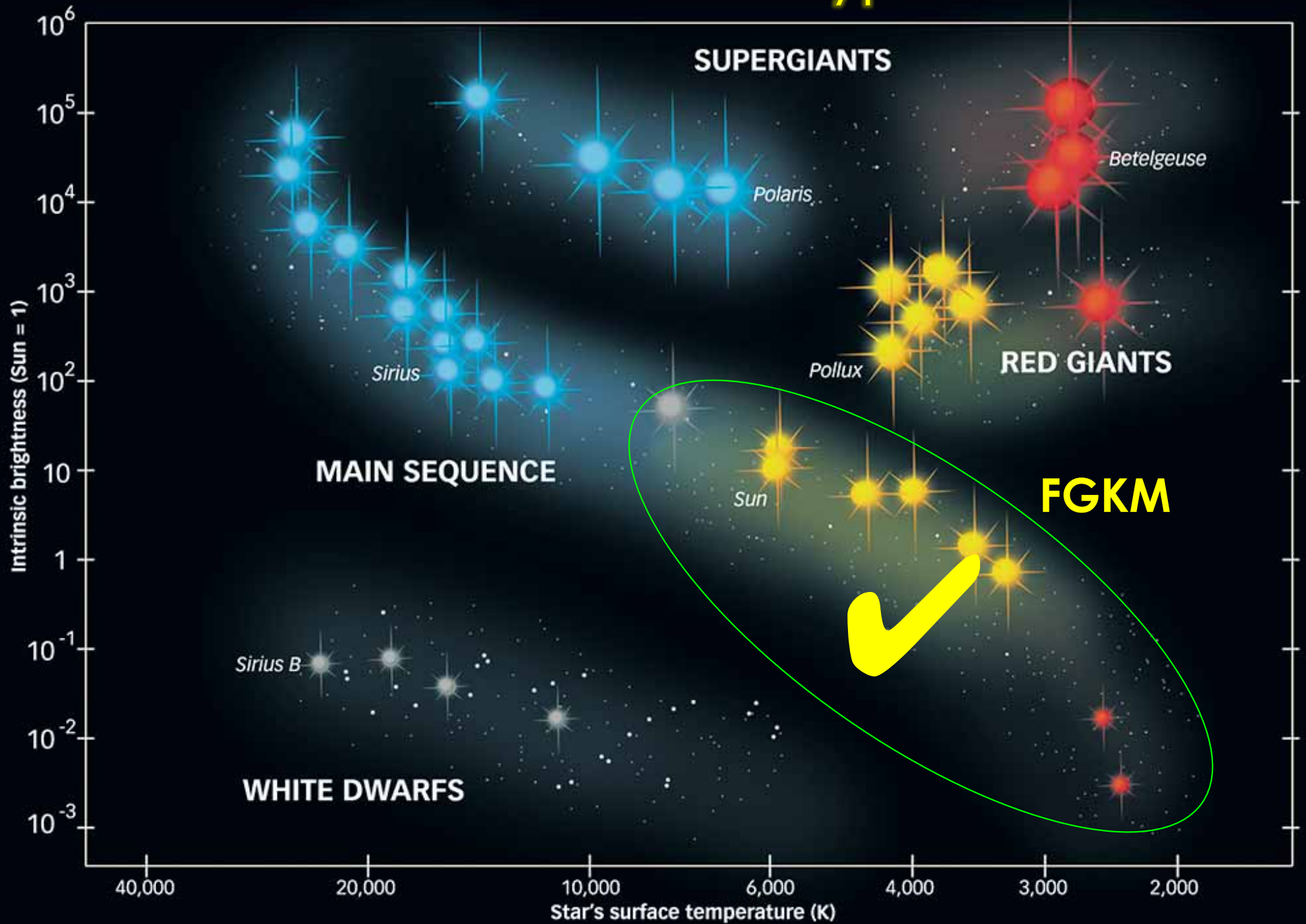


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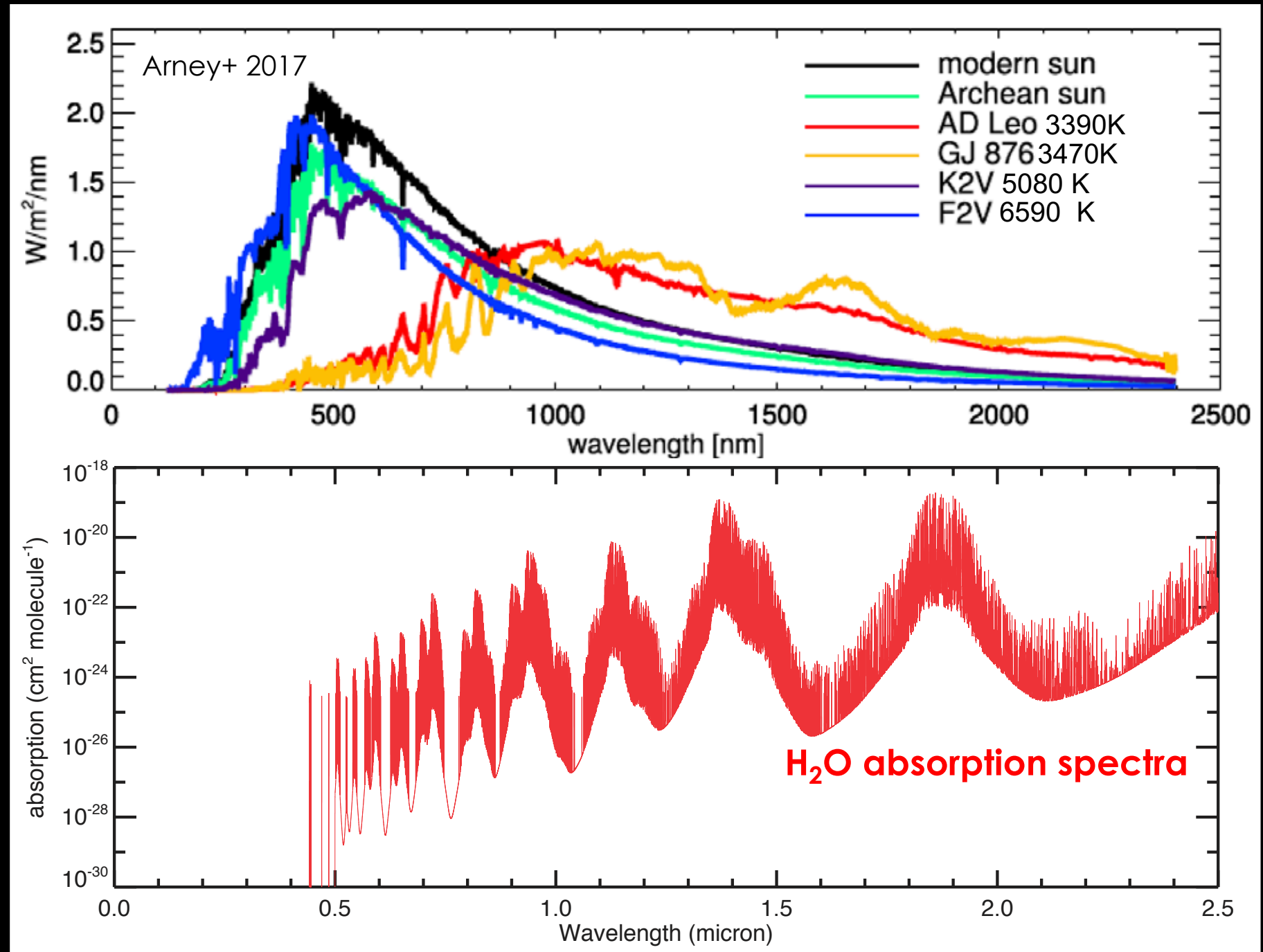


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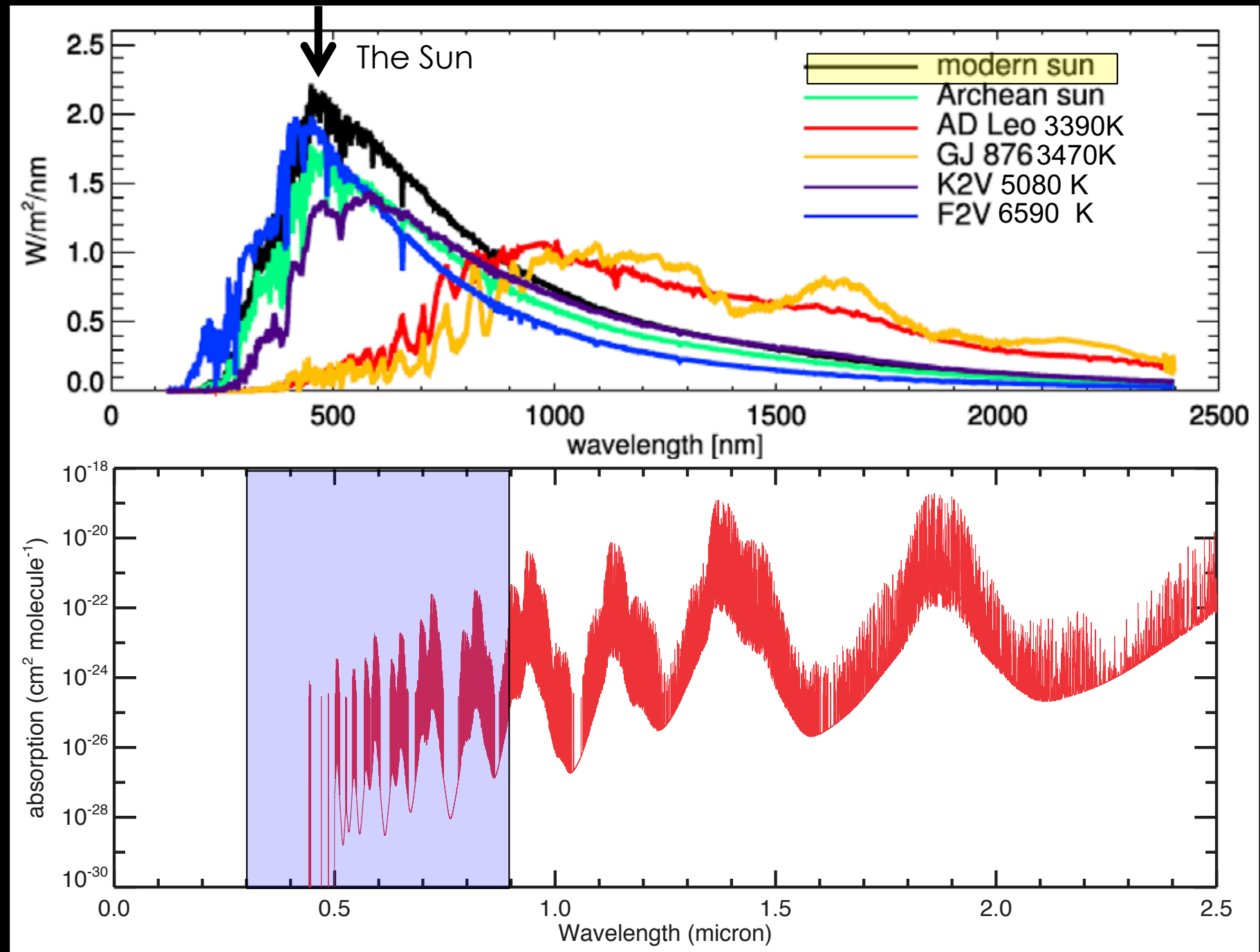


# Shifts stellar spectra affects near-IR absorption by H<sub>2</sub>O



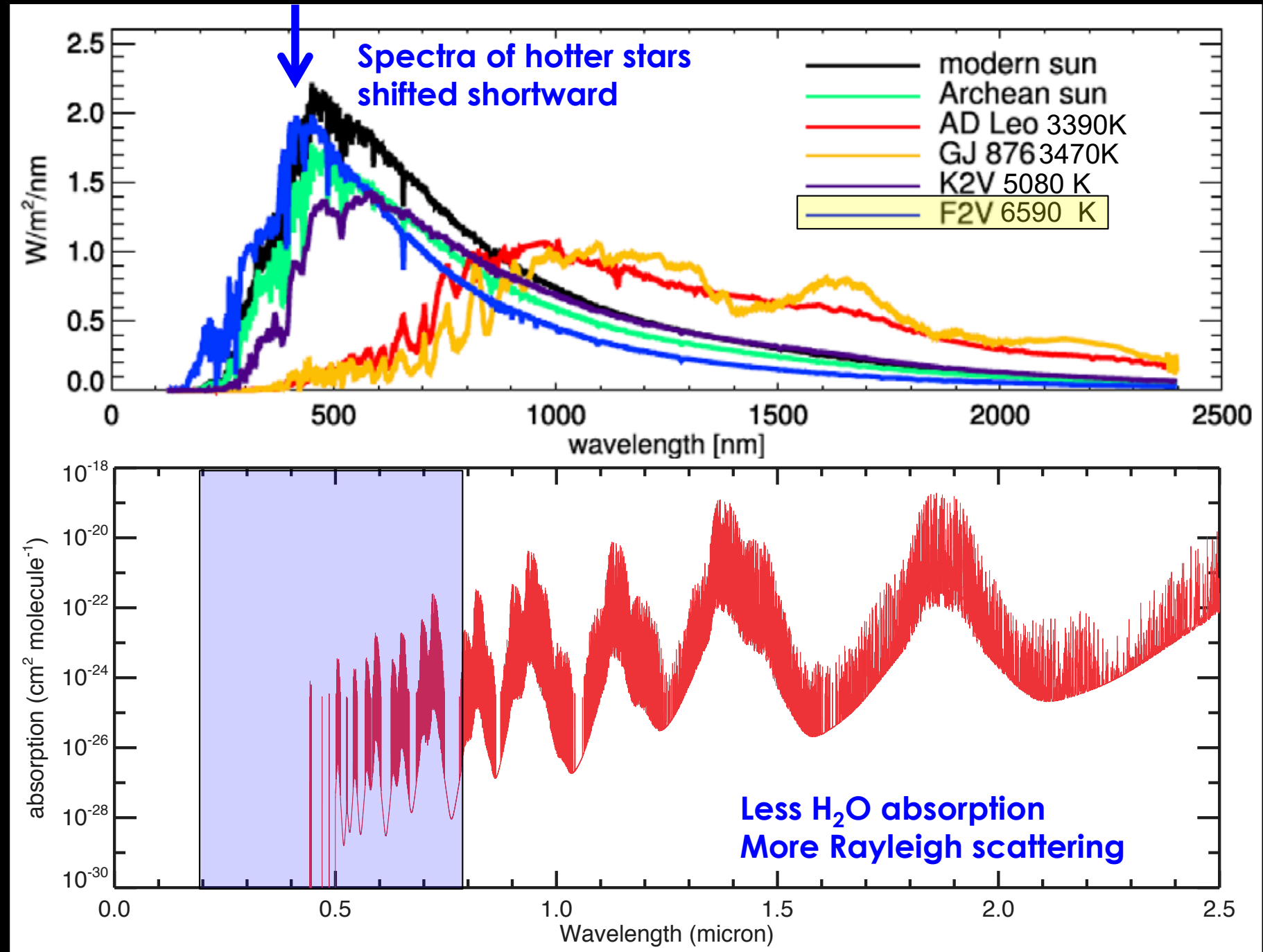


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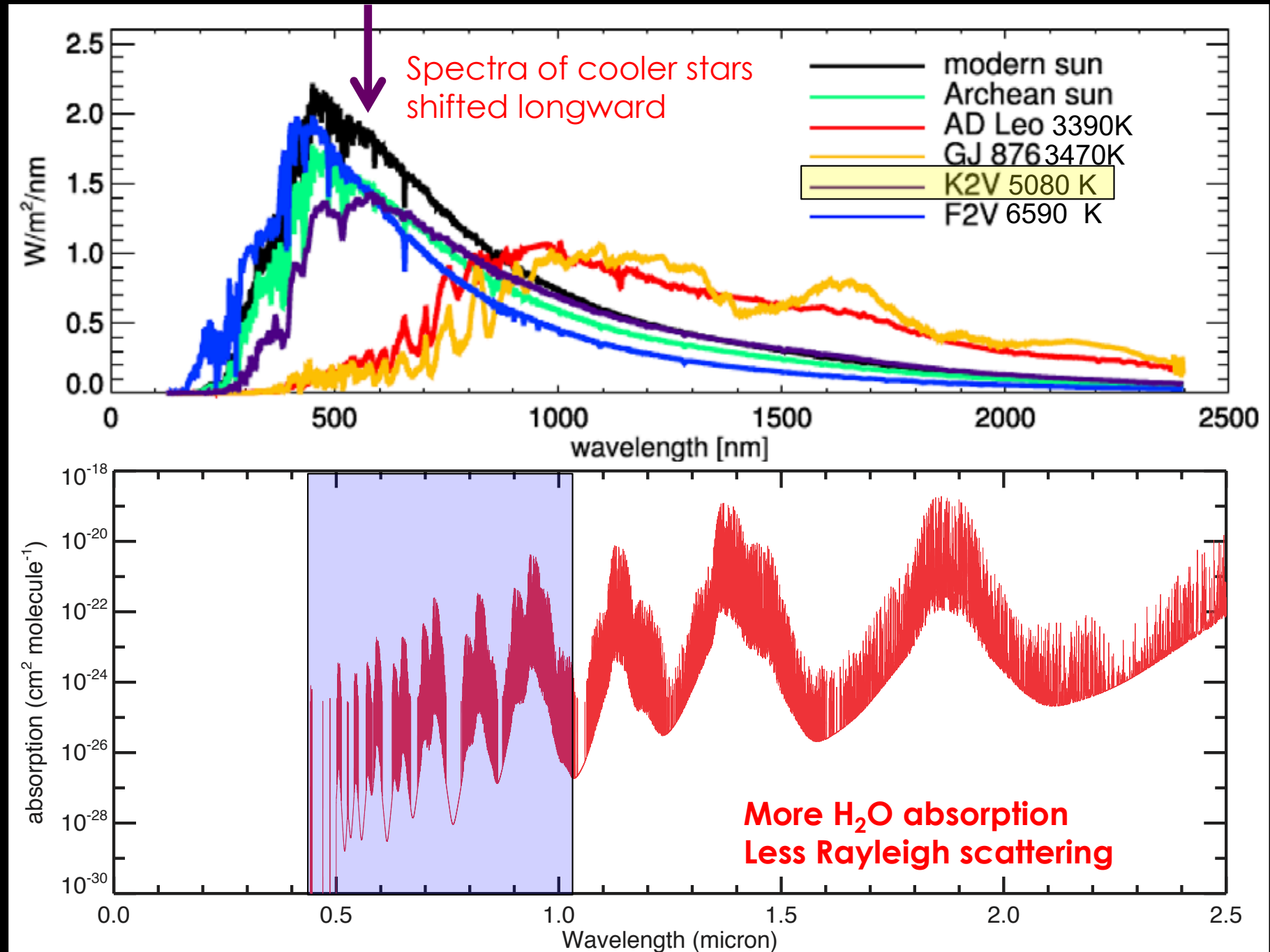


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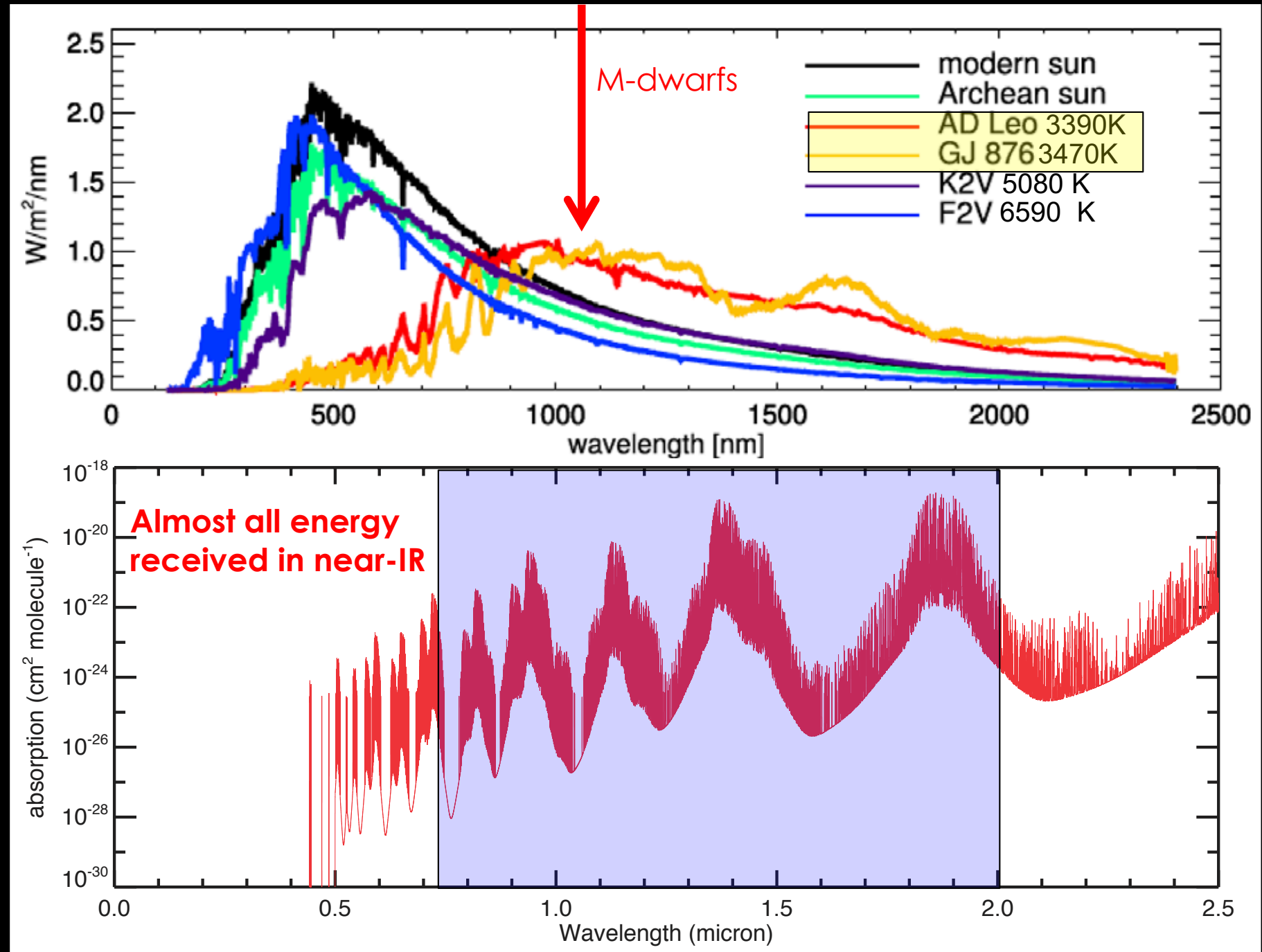


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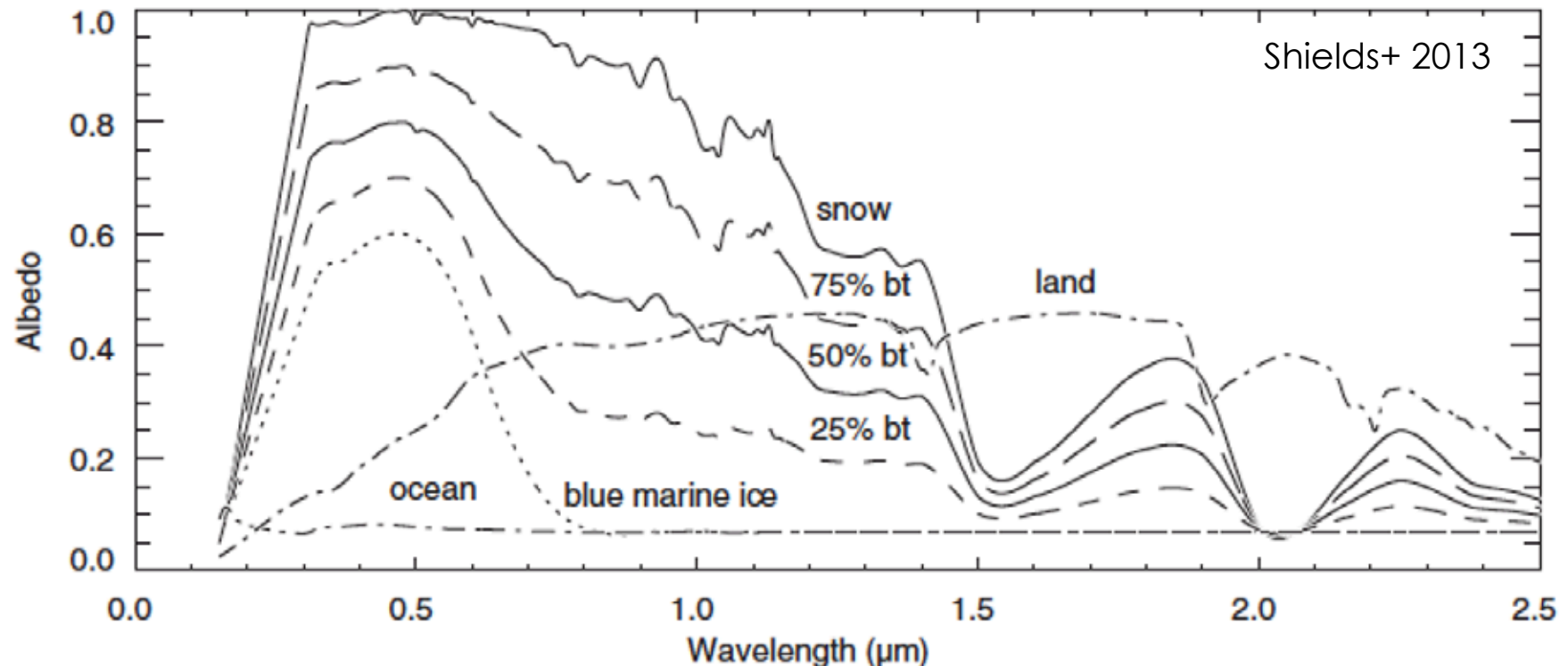
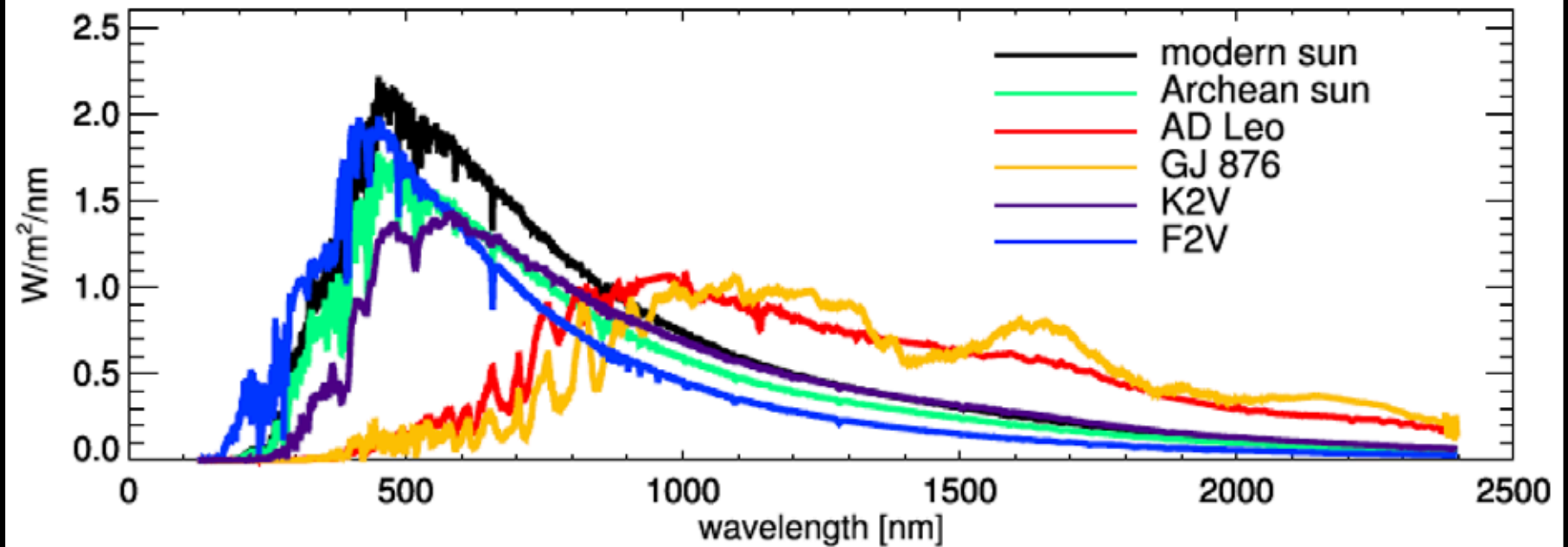


# Shifts stellar spectra affects near-IR absorption by H<sub>2</sub>O



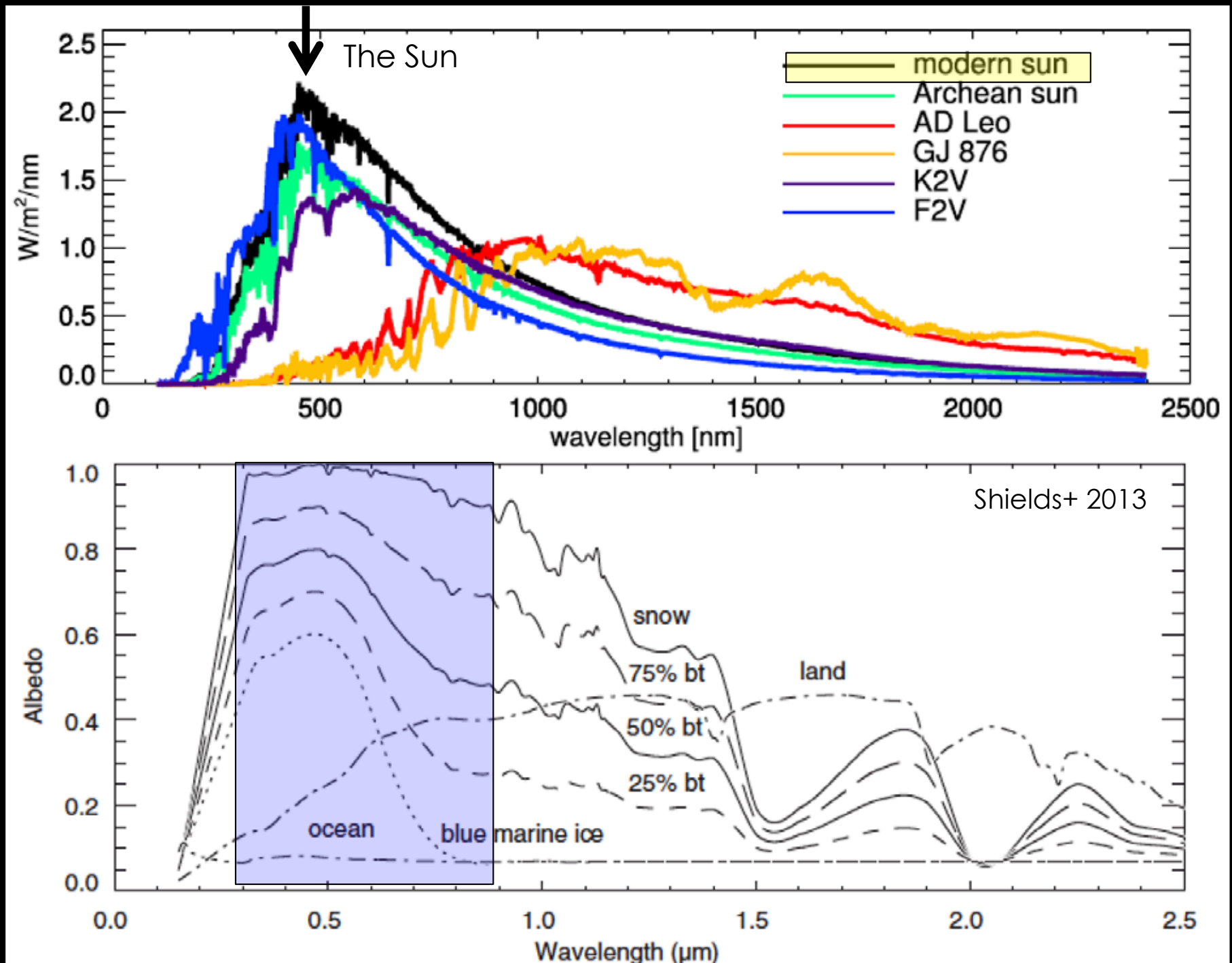


# Shifts stellar spectra affects integrated surface albedo



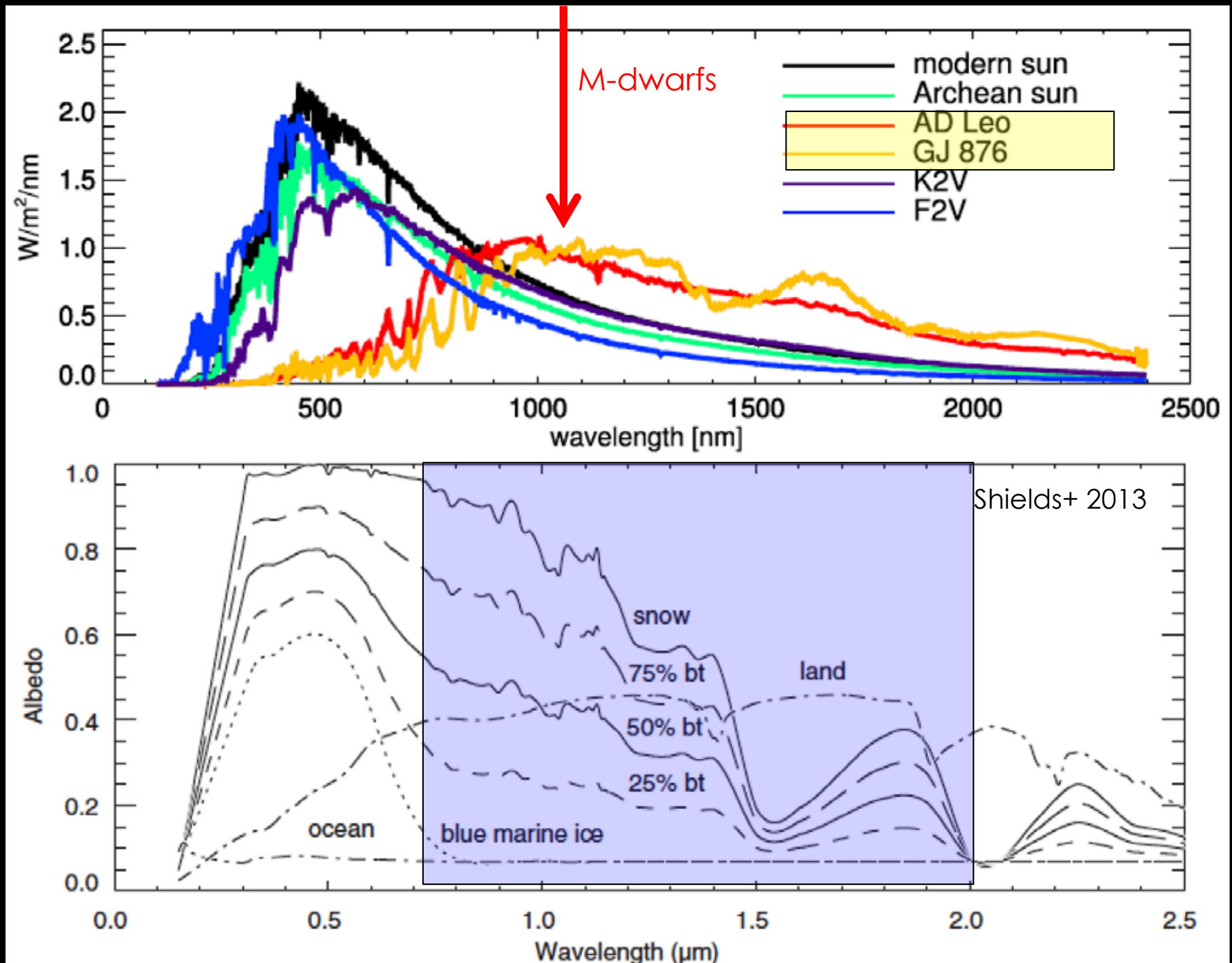


# Shifts stellar spectra affects integrated surface albedo





# Shifts stellar spectra affects integrated surface albedo





# Atmospheric Compositions

1 bar surface  
pressure

~400 ppm CO<sub>2</sub>

H<sub>2</sub>O

Clouds

KB



# Atmospheric Compositions Extending the HZ

1 bar surface  
pressure

$\text{CO}_2$

$\text{H}_2\text{O}$

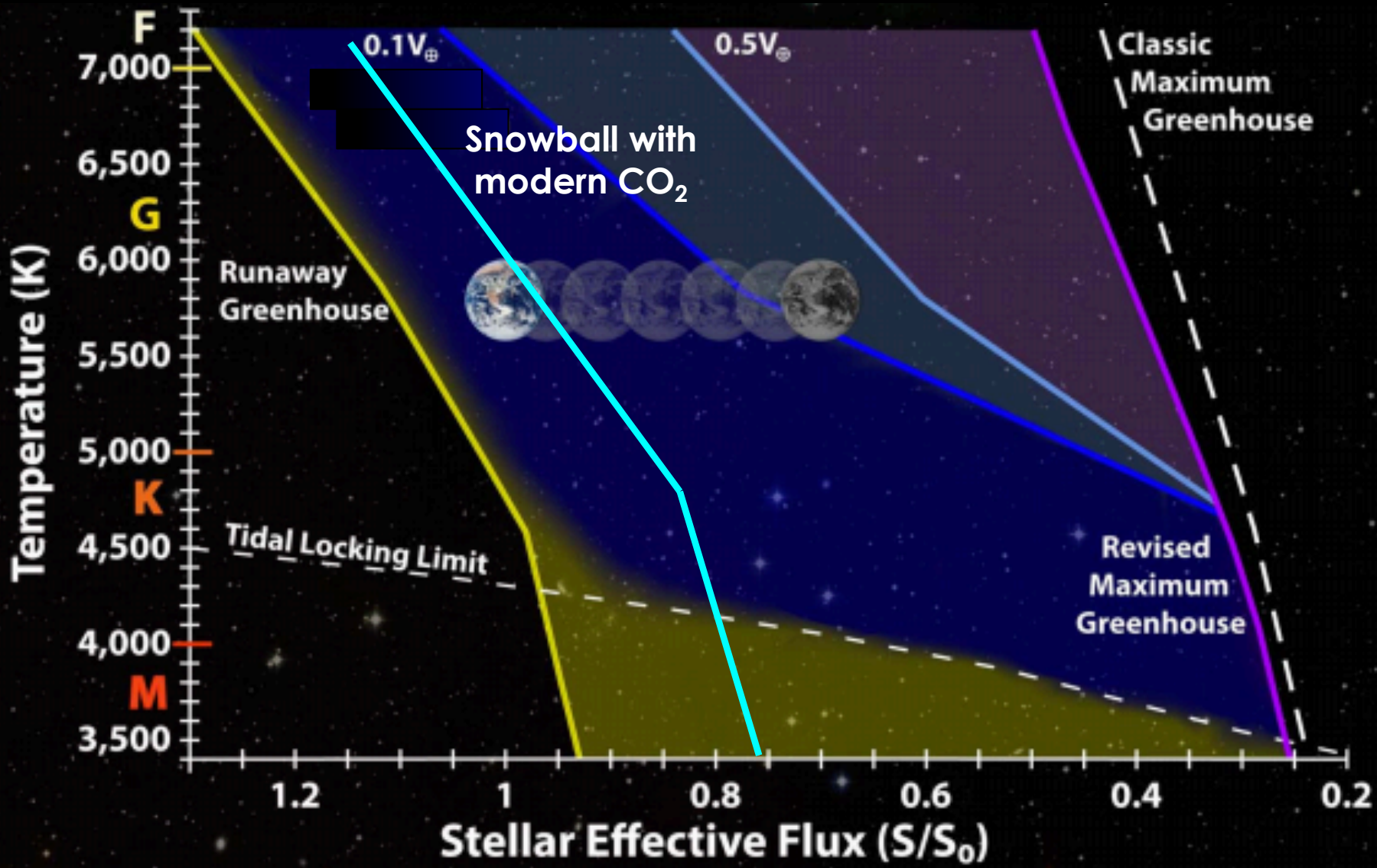
Clouds

KB



The outer edge of the habitable zone

# The Maximum CO<sub>2</sub> Greenhouse Limit

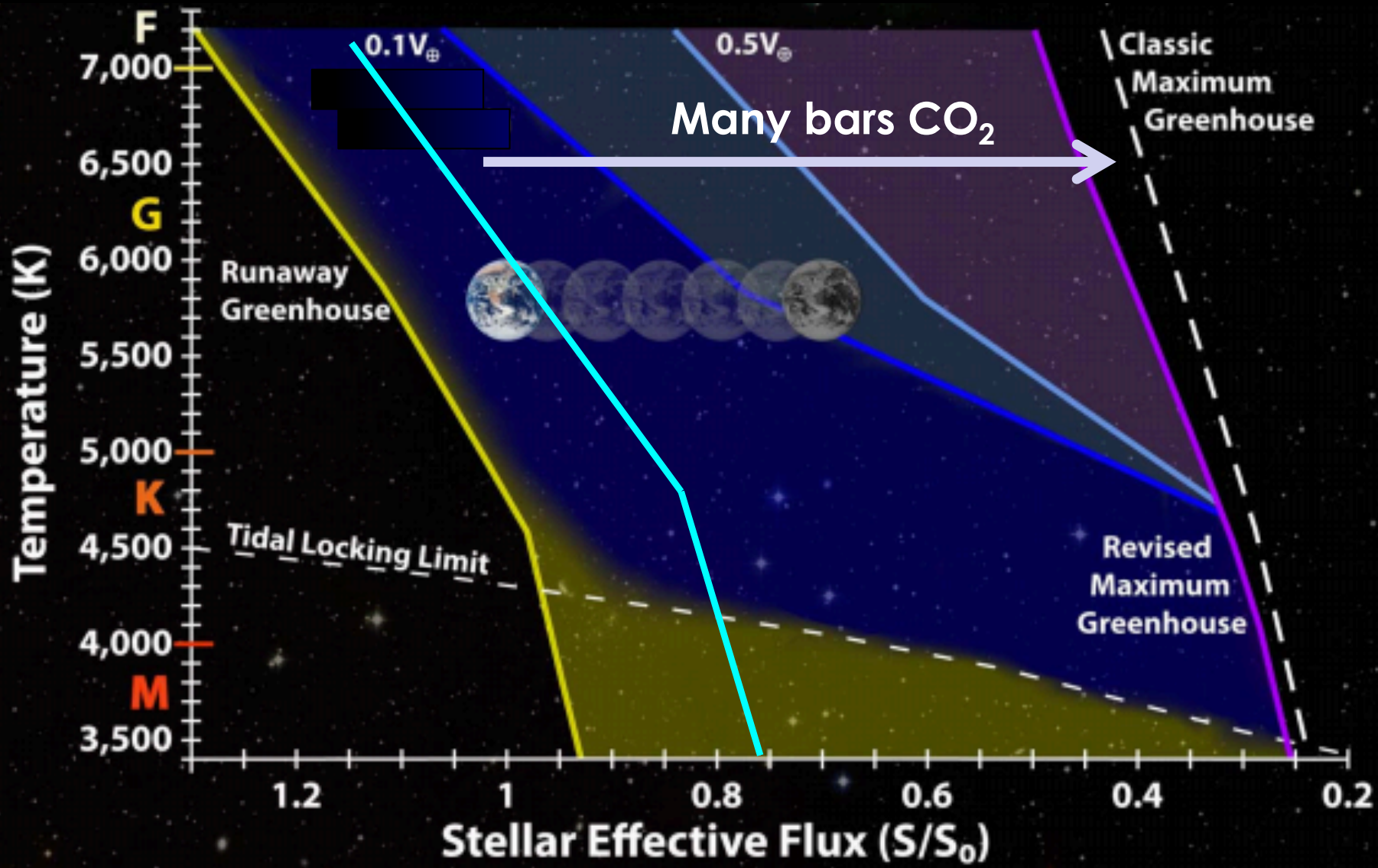


Haqq-Misra+ 2016



The outer edge of the habitable zone

# The Maximum CO<sub>2</sub> Greenhouse Limit

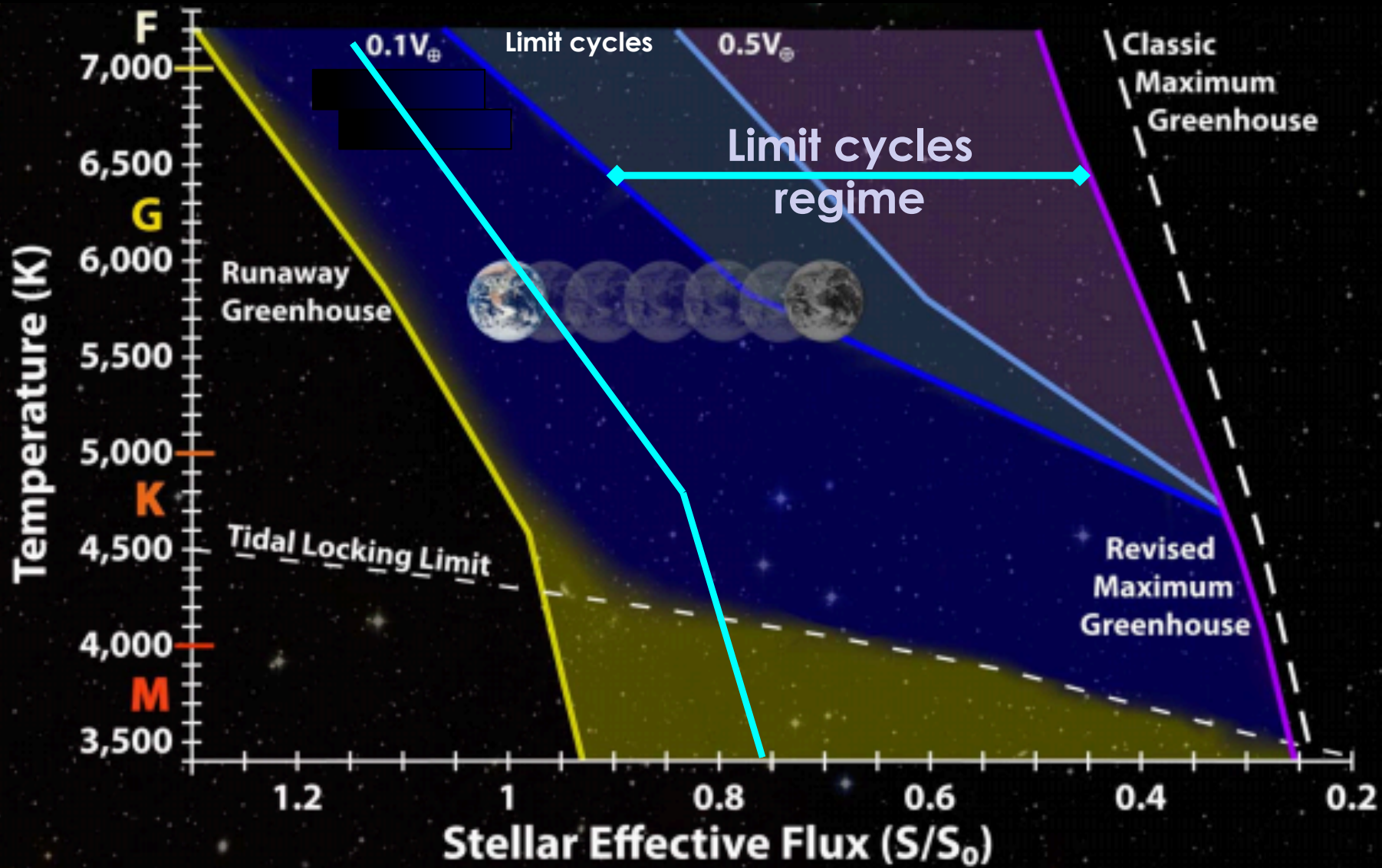


Haqq-Misra+ 2016



The outer edge of the habitable zone

# Limit Cycles

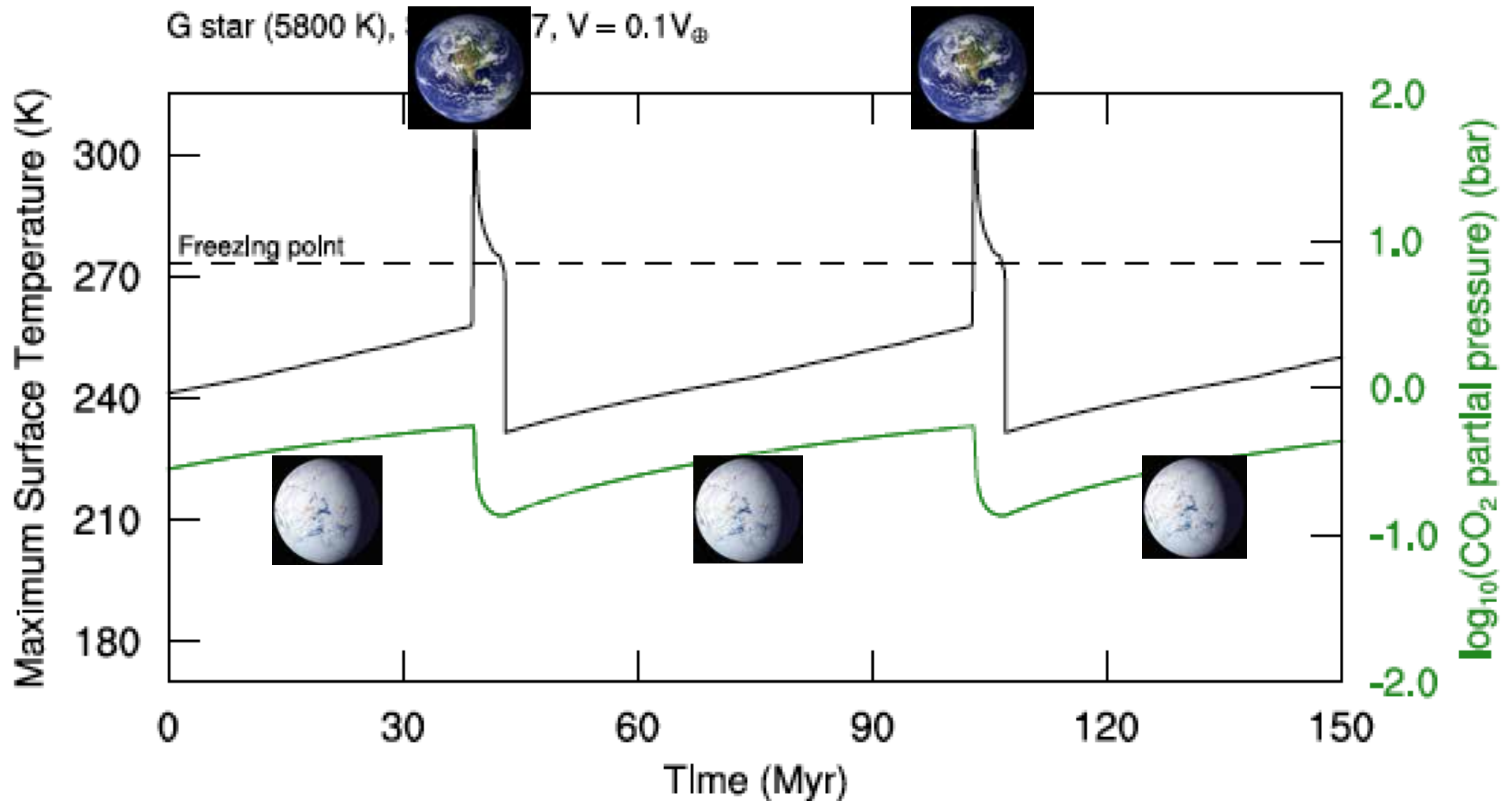


Haqq-Misra+ 2016



The outer edge of the habitable zone

# Limit Cycles







# ***M-dwarf systems are compelling targets.***

~70% of the stellar population

small size means planet signal is relatively larger

short period planetary orbits means many transits possible

Proxima Cen b, TRAPPIST-1 b–h, LHS1140b, Ross-128b

*interesting physics and dynamics of the terrestrial planet  
atmospheres*

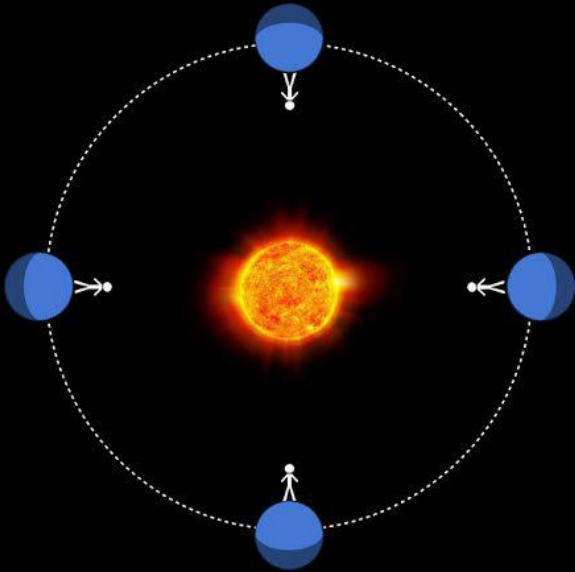


M-dwarf stars are the smallest and most common in the universe



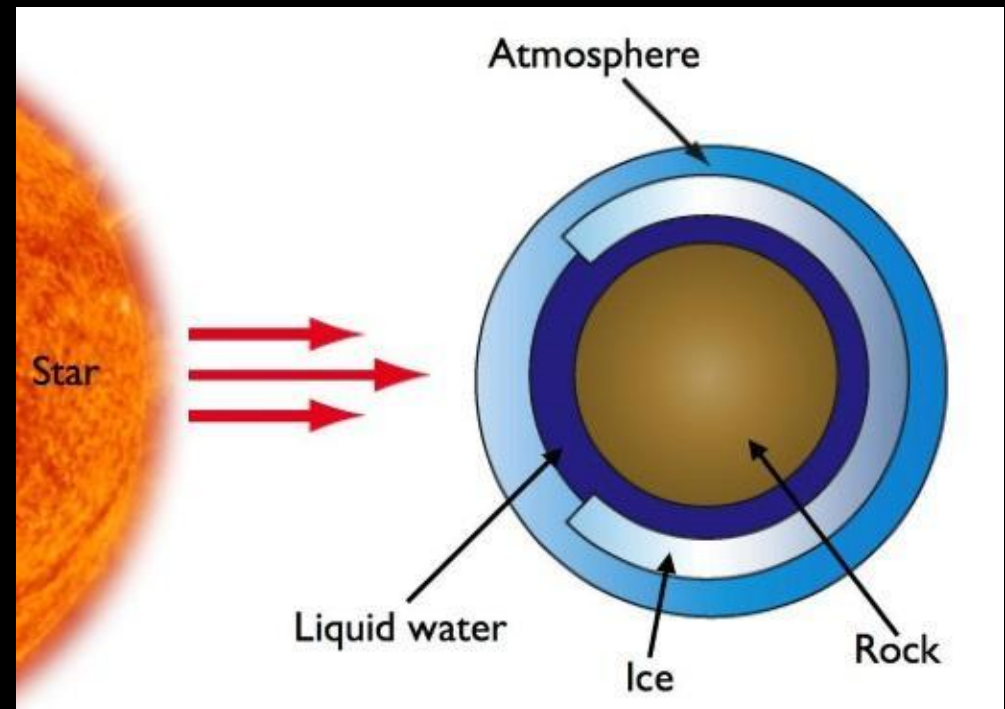


# Consequences of tidal locking



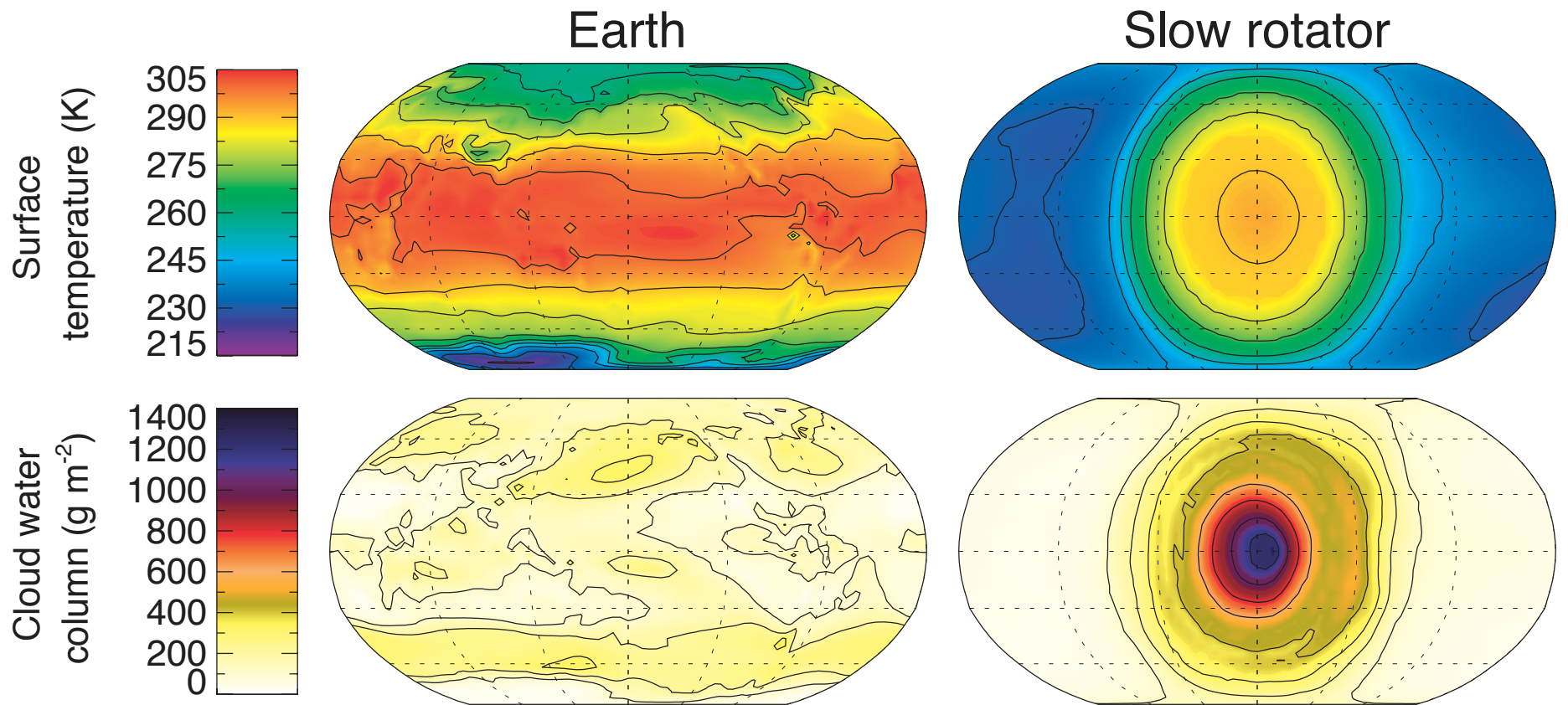
- rotation rate = orbital period
- spatial variations in solar insolation
- Strongly affects atmospheric dynamics and clouds

## The "Eyeball" Climate State





# The "Eyeball" Climate State

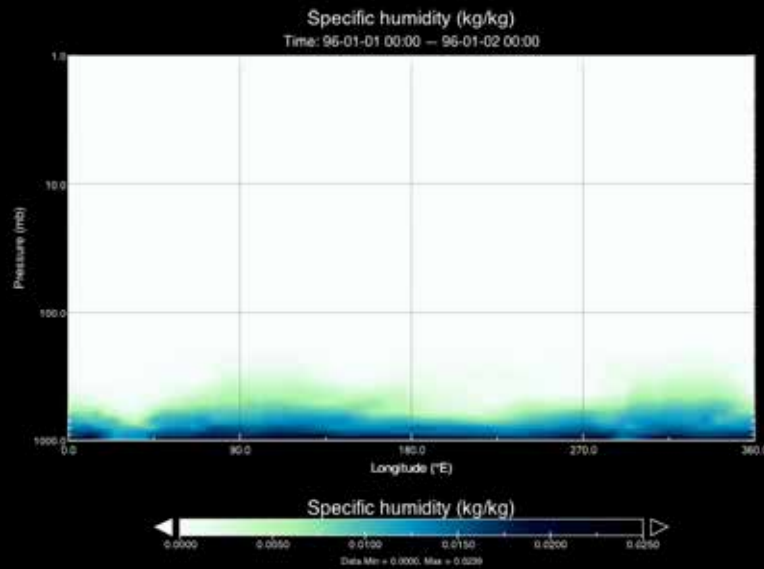


The albedo of slow rotators is greatly increased due to substellar clouds

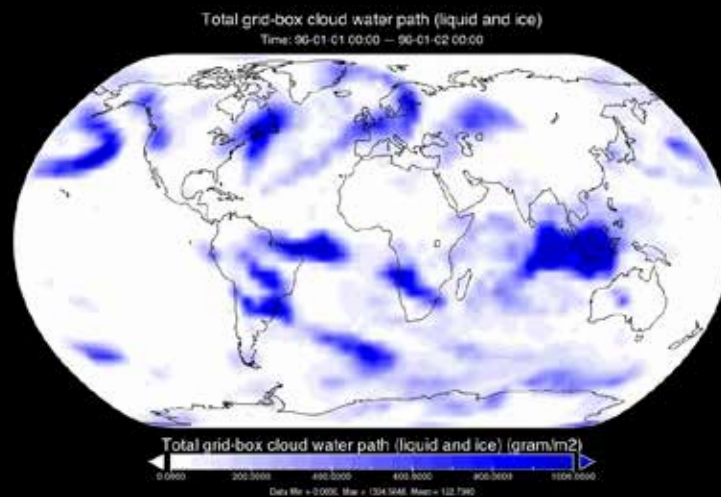


rapid rotator  
24 hours non-synchronous  
(translation... *Earth*)

water vapor

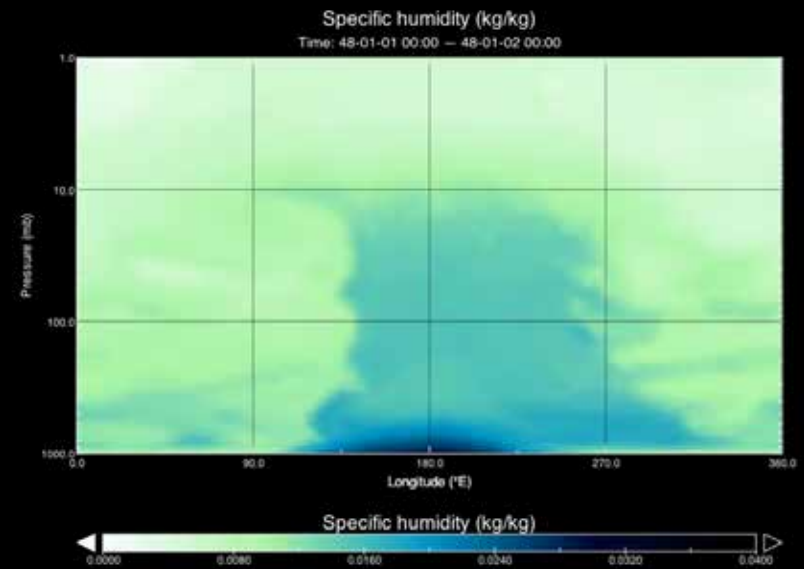


Cloud water column

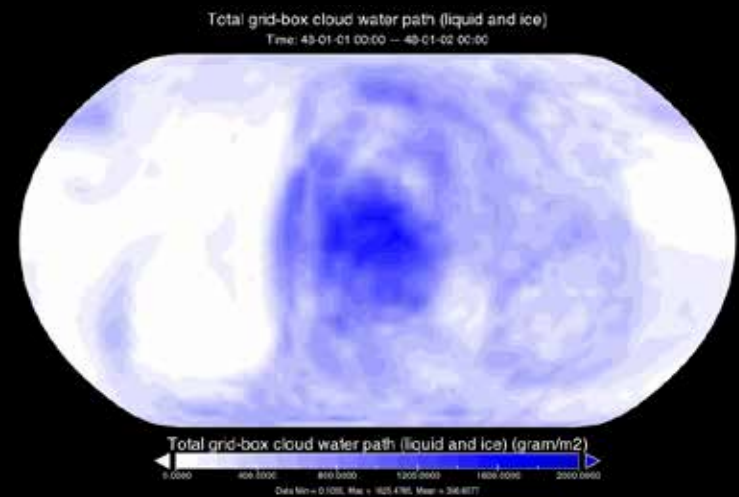


slow rotator  
84.7 day, synchronous

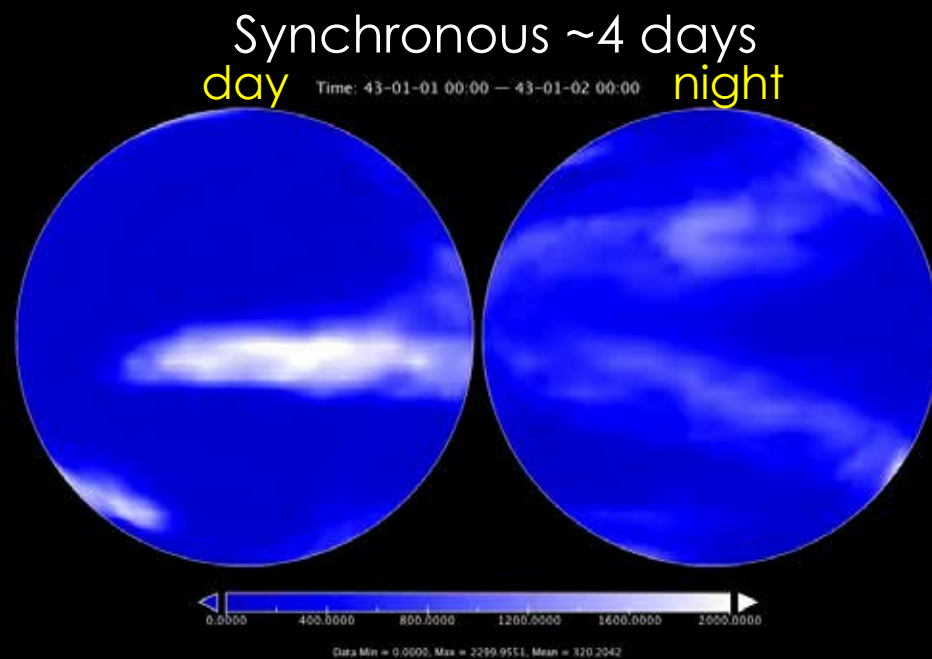
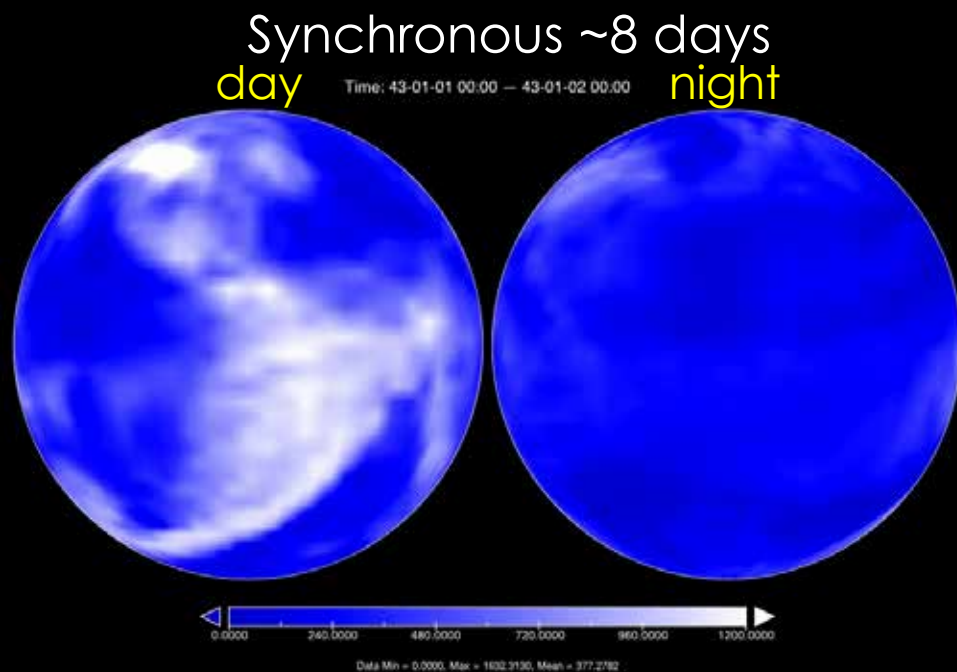
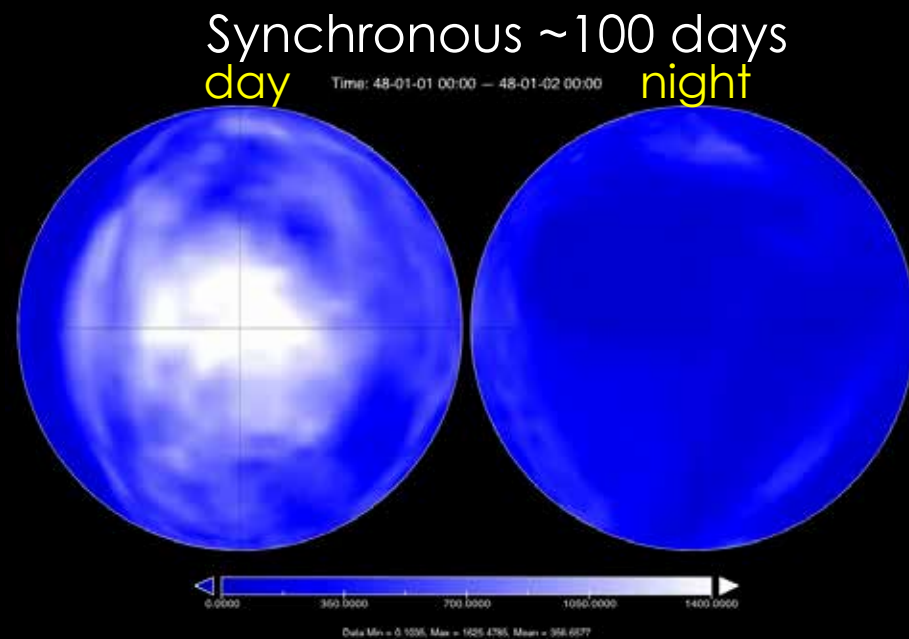
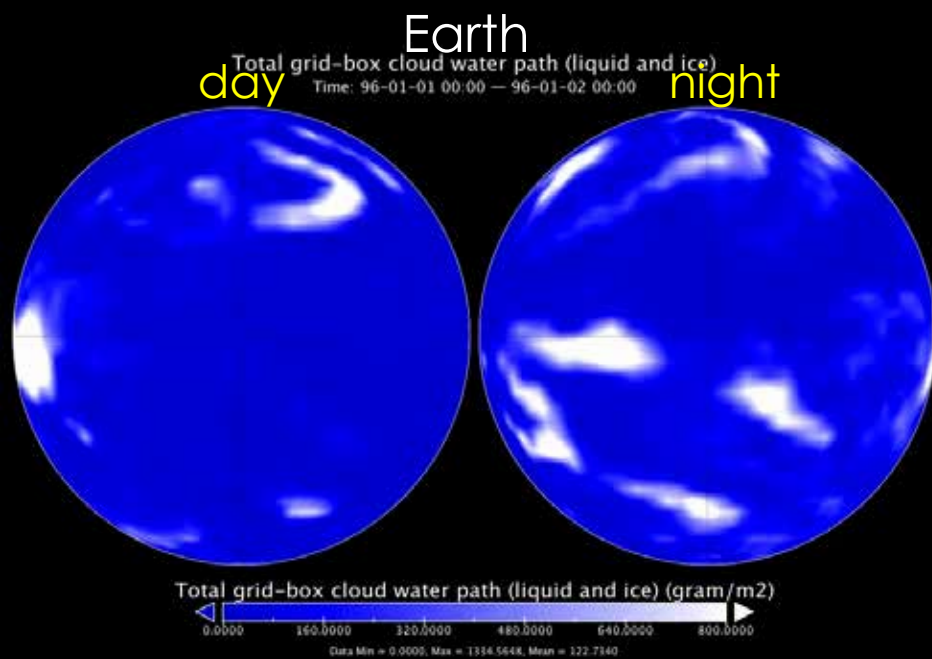
water vapor



Cloud water column









# Can planets around M-dwarfs be habitable?

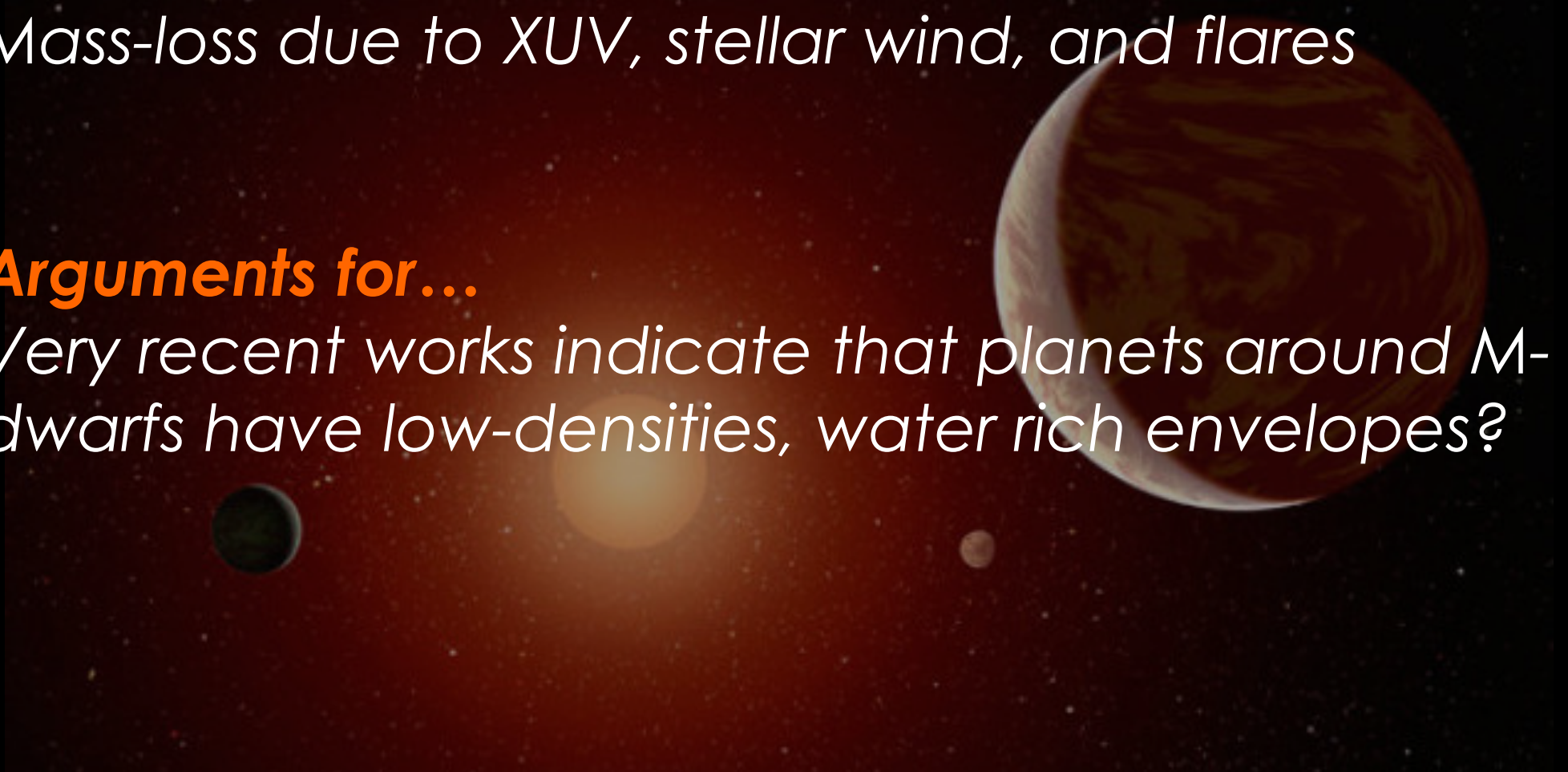
## Arguments against...

Super-luminous pre-main sequence phase

Mass-loss due to XUV, stellar wind, and flares

## Arguments for...

Very recent works indicate that planets around M-dwarfs have low-densities, water rich envelopes?





# There are many things I didn't get to...

- Habitability of non “Earth-like” planet-types

“Dune” planets, photochemical hazes, Intense volcanism on tidally locked planets, Super-Earths

- Different greenhouse gases

$\text{CH}_4$ ,  $\text{N}_2$ - $\text{H}_2$  CIA

- Atmospheric chemistry under different SEDs

- Surface UV radiation environments

- Different evolutionary paths

planetary migration, abiotic  $\text{O}_2$  build up

- **Relevance of climate states to transmission spectra, and thermal phase curves**

Can we discriminate between them?





# *Transiting Exoplanet Survey Satellite* *NASA*

*Launch date April 18<sup>th</sup>, 2018*



*The successor to Kepler*



TESS: Delivered to Kennedy  
Space Center Feb. 19<sup>th</sup>, 2018



# James *W*ebb *S*pace *T*elescope

## NASA

Launch date Spring 2019



Infrared  
0.6 to 28.5  $\mu\text{m}$

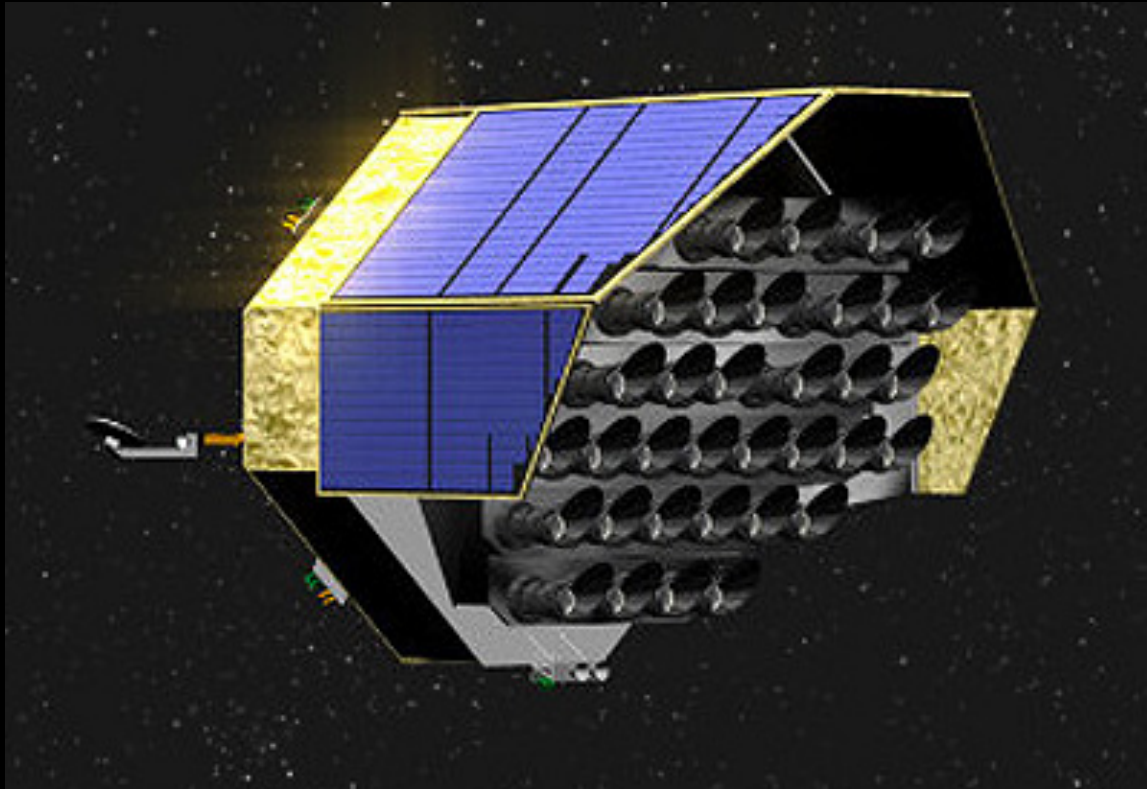
Designed to study the  
early universe, galaxies  
... and also exoplanets



# **PLA**netary **T**ransits and **O**scillations of stars

ESA

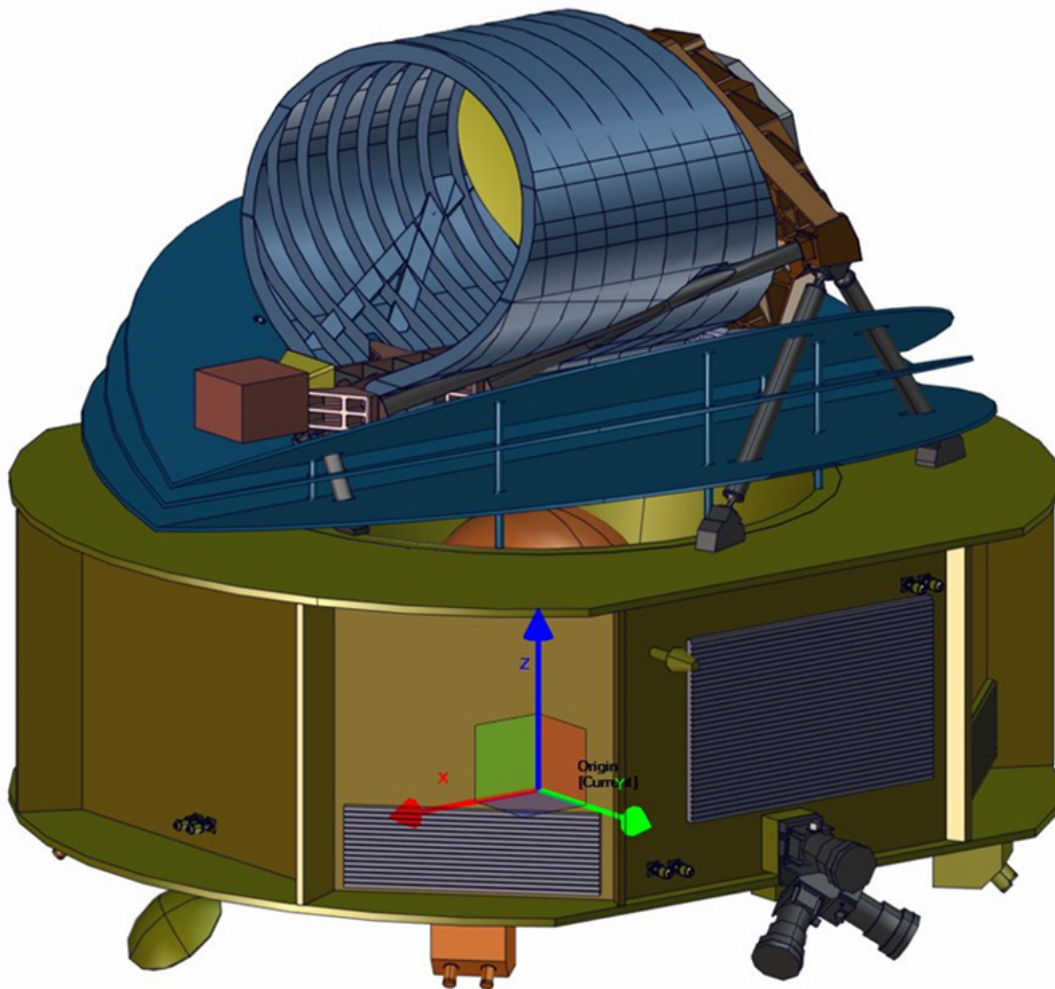
**Launch date 2026**



Designed to discover and characterize Earth-like planets around solar-type stars.



# **A**tmospheric **R**emote-sensing **I**nfrared **E**xoplanet **L**arge-Survey ESA Launch date 2026



Designed to probe  
atmospheres of hot  
exoplanets (500 K and  
hotter)

*No habitable planets.*



# Thirty-Meter Class Ground Based Telescopes Late 2020s?

**T**hirty **M**eter **T**elescope  
**E**uropean **E**xtra **L**arge **T**elescope





# Decadal Survey Mission Concept Studies Late 2030s?

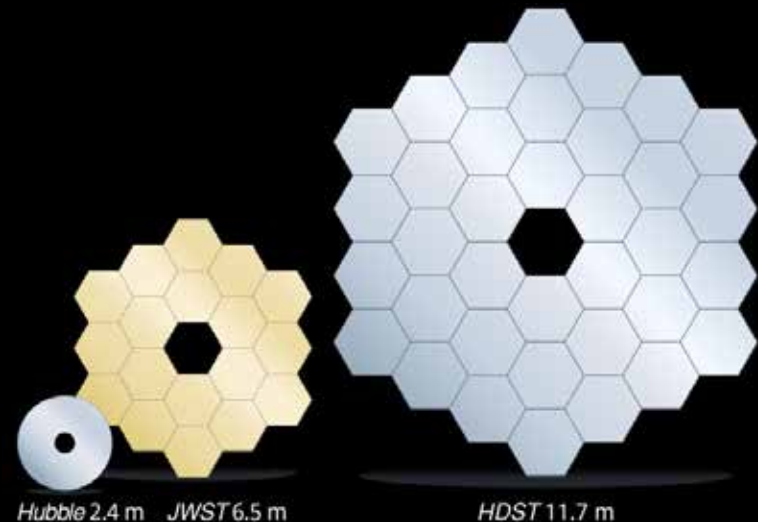
**L**arge **UV** **O**ptical **I**nfra**R**ed Surveyor (LUVOIR)

**O**rigins **S**pace **T**elescope

**Hab**itable **Ex**oplanet Imaging Study (HabEx)

**Lynx** X-ray Observatory

*For 3 out of the 4 have characterization terrestrial exoplanet atmospheres as the central plank in their design specifications.*





***Take home points ...***

***The Universe is teeming with planets.***

***Even “Earth-like” exoplanets can be pretty weird.***

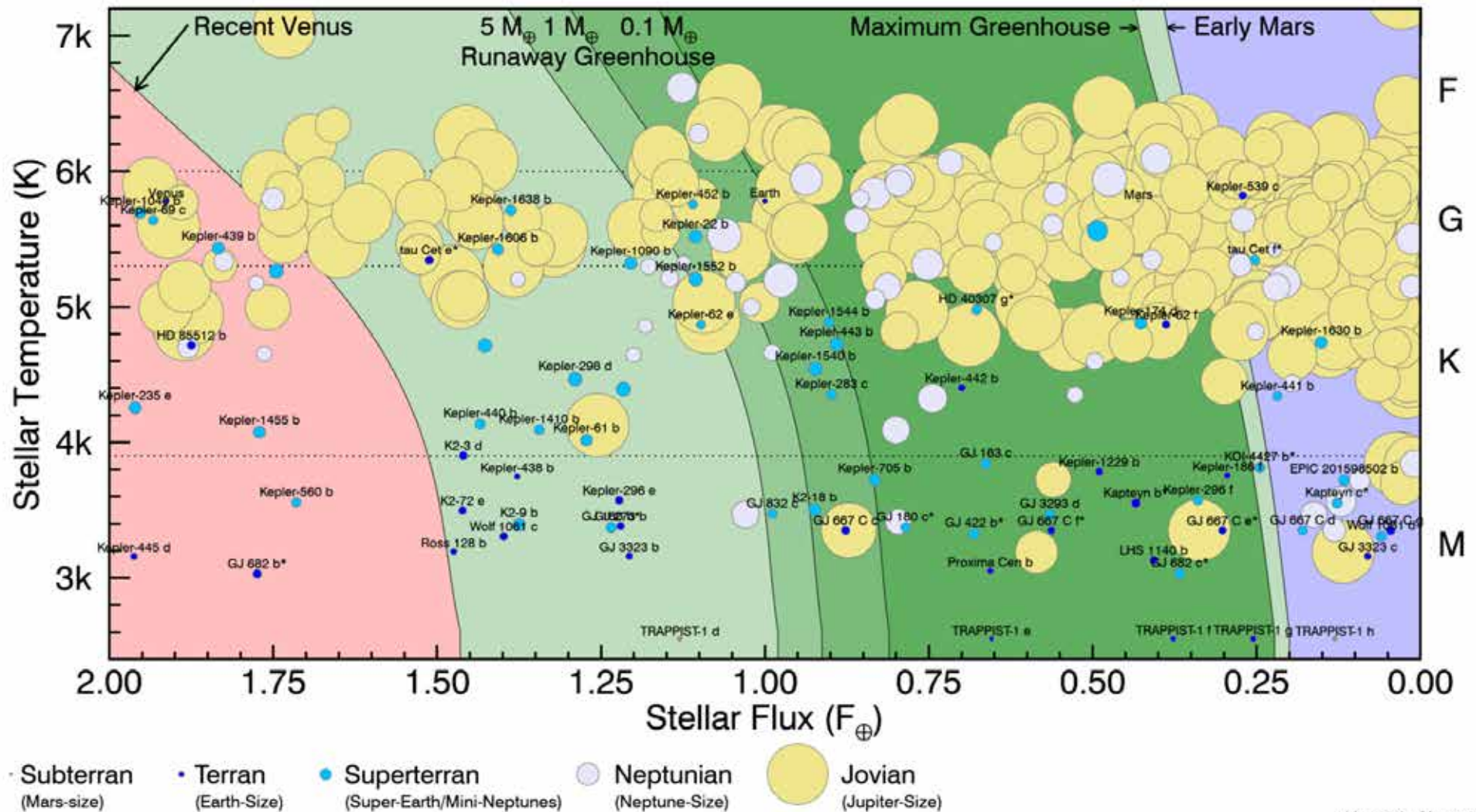
***Habitability is a confluence of complex astro-  
geophysical and climatological processes.***

***Undoubtedly, we will be surprised at what we find.***

***Our voyage has just begun.***



# The Habitable Zone



phl.upr.edu, Mar 2018

Habitable Exoplanets Catalog  
Planetary Habitability Lab, University of Puerto Rico, Acribo  
<http://phl.upr.edu/projects/habitable-exoplanets-catalog>