

Laboratory for Atmospheric and Space Physics  
University of Colorado **Boulder**



# INTERPRETING DUST IMPACT SIGNALS DETECTED BY THE STEREO SPACECRAFT

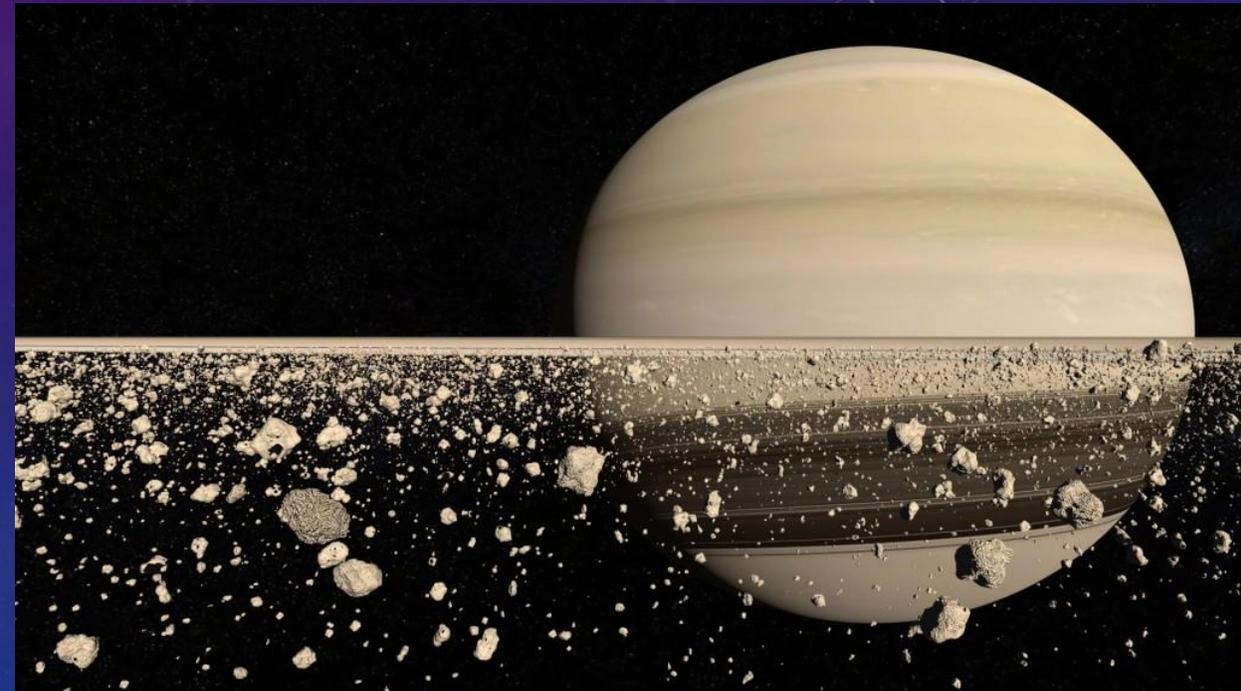
EVIN O'SHEA

MENTOR: ZOLTAN STERNOVSKY

Image: [jpl.nasa.gov](http://jpl.nasa.gov)

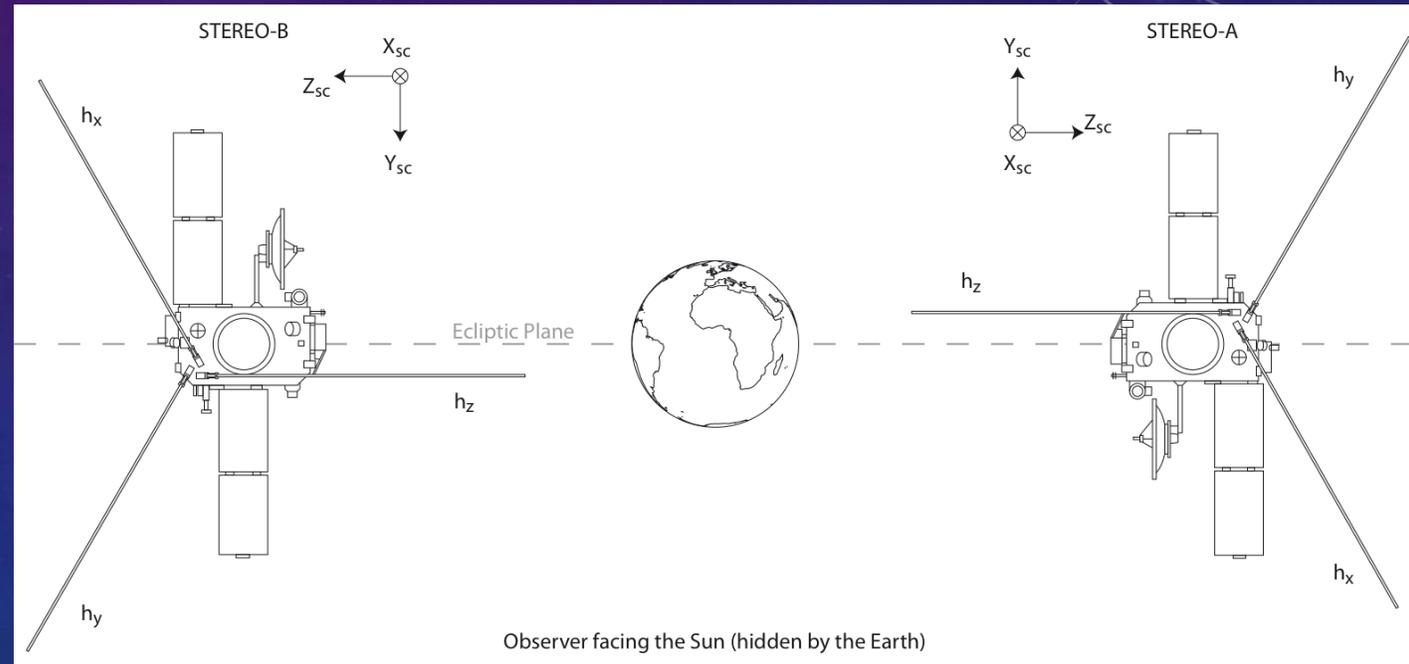
# DUST IN SPACE

- There are many types of dust in space: Interstellar, Interplanetary, etc
- Mostly dealing with interplanetary dust
- The spacecraft measure dust signals (Voyager, Cassini, STEREO, etc)



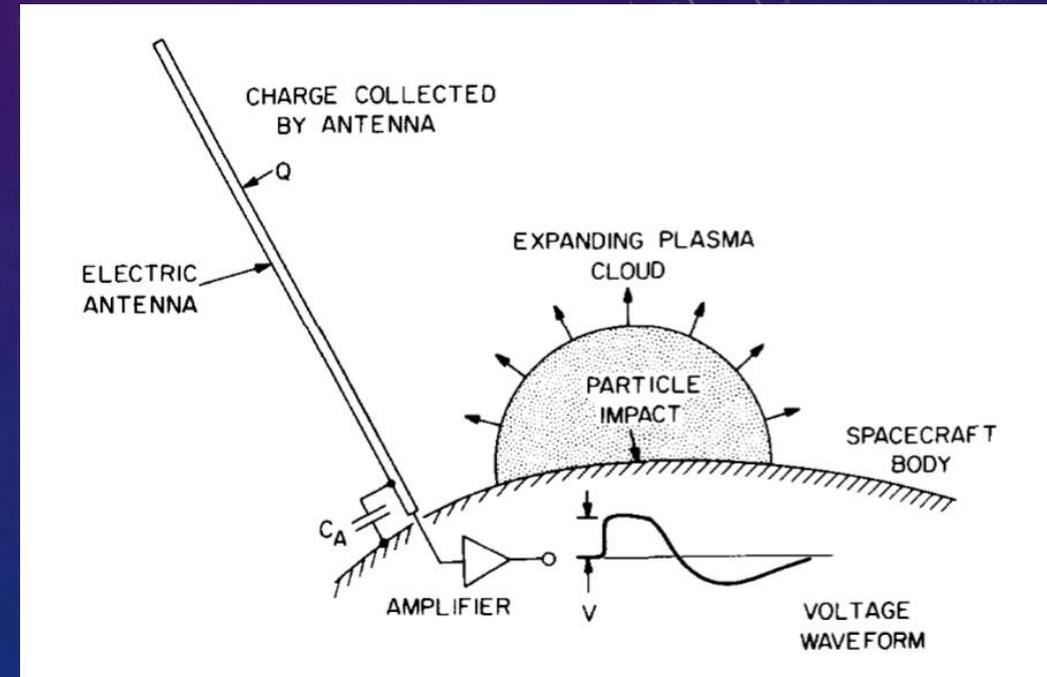
# THE STEREO MISSION AND SOLAR WIND CONDITIONS

- Two spacecraft: Ahead (A) and Behind (B)
- Placed at 1 AU (in Earth-like orbit)
- Designed to measure solar wind conditions
- The S/WAVES instrument has three antennas.
- The spacecraft and antennas have a floating potential due to solar wind conditions.



# BASICS OF DUST SIGNAL GENERATION

- High-velocity dust that impacts spacecraft partially ionizes and creates a plasma cloud.
- Charge collection and image charging effects are recorded by the antennas.
- It was assumed that the plasma cloud was large and there was a large amount of collection.
- The internal electric field will be small when the plasma cloud is a few centimeters large.
- Electrons and ions will follow single particle trajectories.

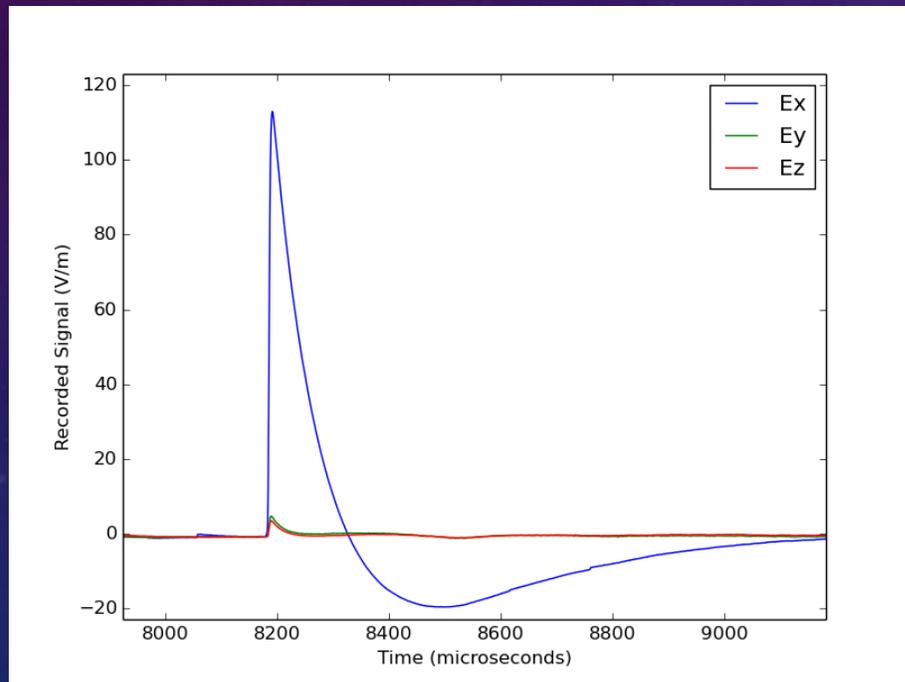


# OUTLINE OF RESEARCH PROJECT

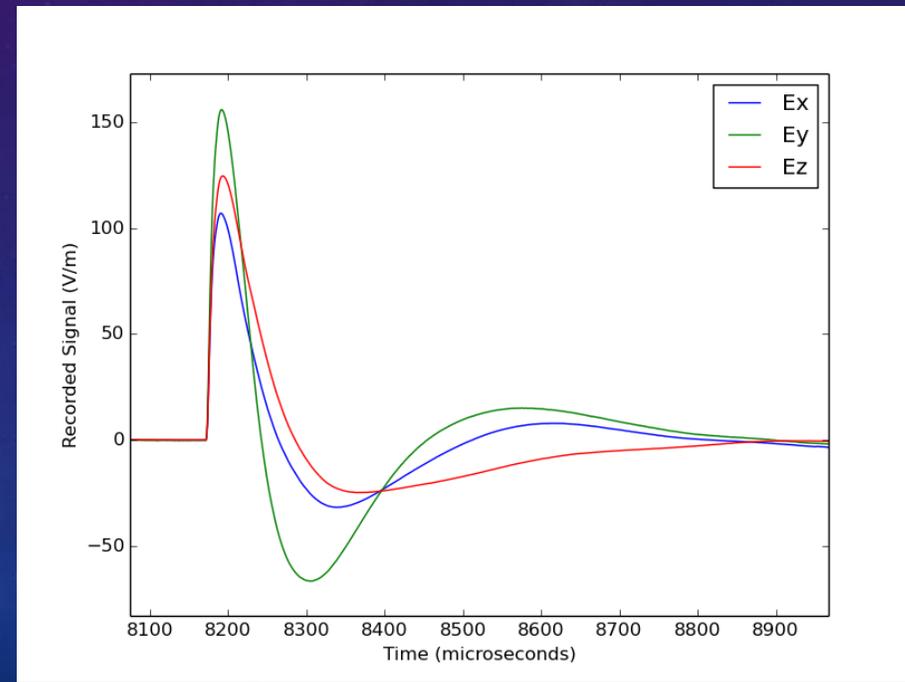
- There was a model proposed by Zaslavsky [2015] to explain dust signals.
- The charge collection by antennas was assumed to be significant.
- Run SIMION<sup>®</sup> simulations to test spacecraft/antenna collection.
- Investigate what types of signals occur and how frequently.
- Use findings to investigate the proposed model.

# THE TYPICAL DUST SIGNALS <sup>1</sup>

## Single Hit



## Triple Hit



# FLOATING POINT PERTURBATION MODEL<sup>2</sup>

- Want to find a steady state solution to:

$$\frac{dq}{dt} = I_{sw} + I_{ph} + \dots$$

- Adding a perturbation to the steady-state solution gives:

$$\frac{d\delta\phi}{dt} + \frac{1}{\tau}\delta\phi = \frac{I_{dust}}{C}$$

- $\tau$ ,  $C$ , and  $I_{dust}$  are different for spacecraft and antenna
- Spacecraft:

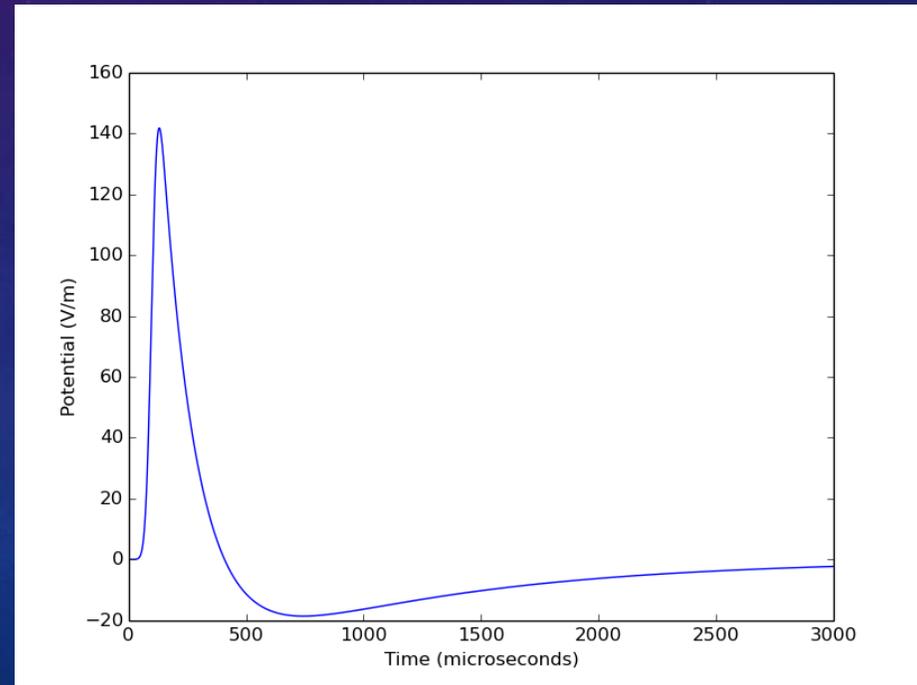
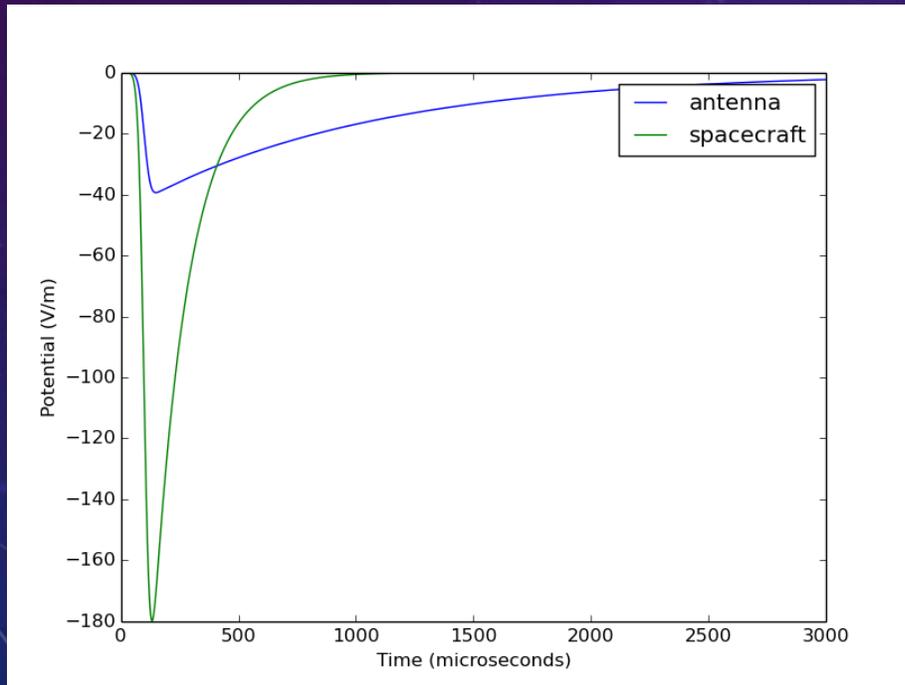
$$I_{dust} = (1 - p_{ant}) \frac{Q}{\sqrt{2\pi\Delta t}} e^{-\frac{t^2}{2\Delta t^2}}$$

- Antenna:

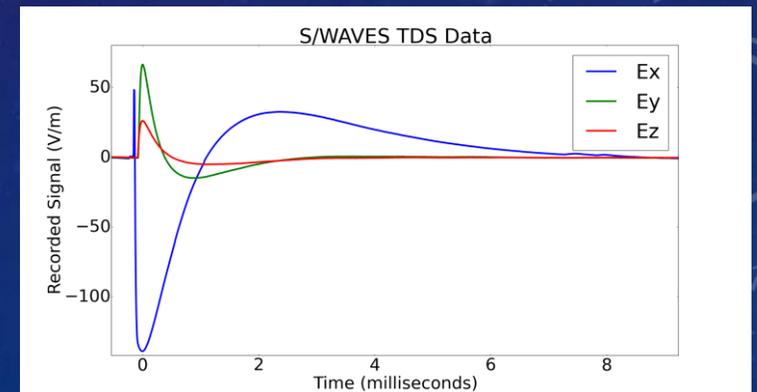
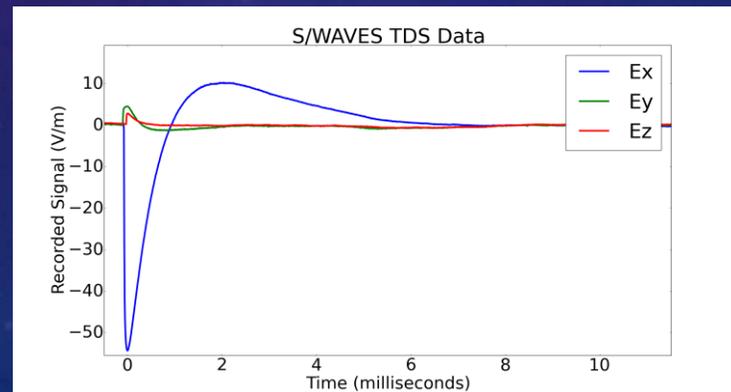
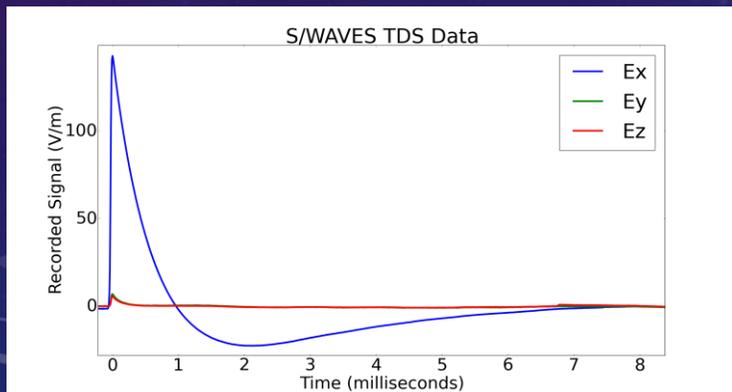
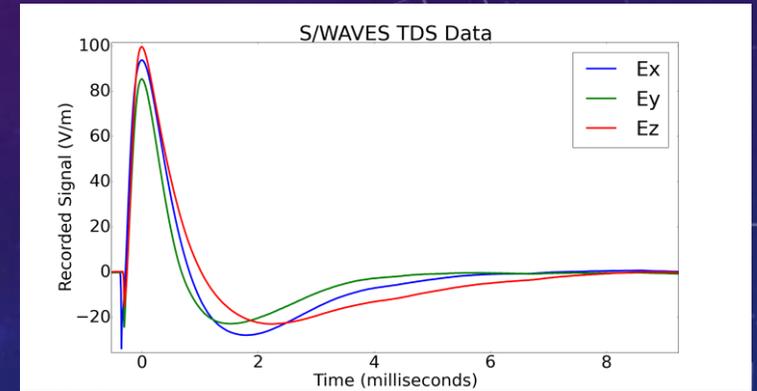
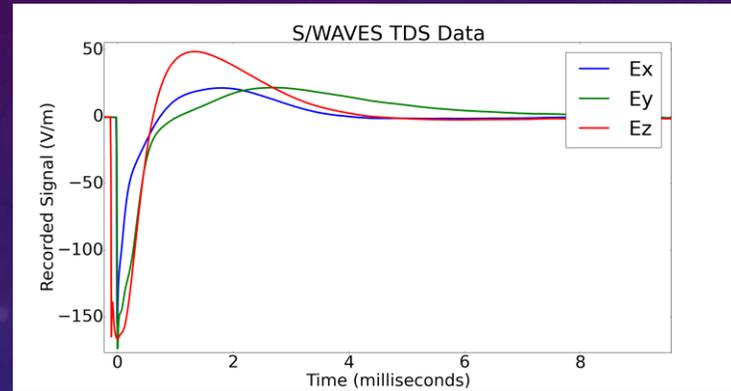
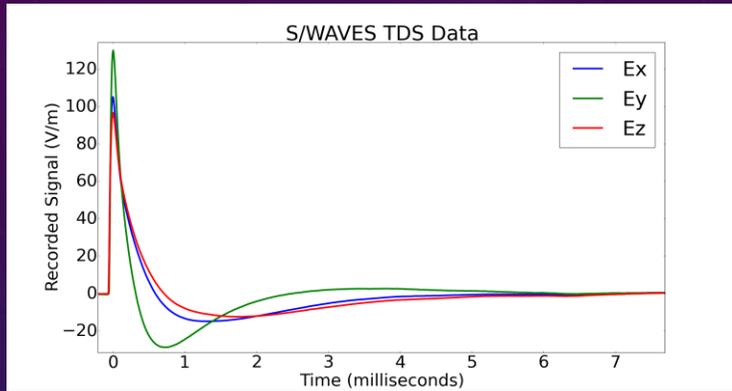
$$I_{dust} = p_{ant} \frac{Q}{\sqrt{2\pi\Delta t}} e^{-\frac{t^2}{2\Delta t^2}}$$

# FPP MODEL CONT.

- If  $I_{dust}$  is a simple Gaussian we get the results ( $\delta\phi_{meas} = \delta\phi_{ant} - \delta\phi_{sc}$ ):



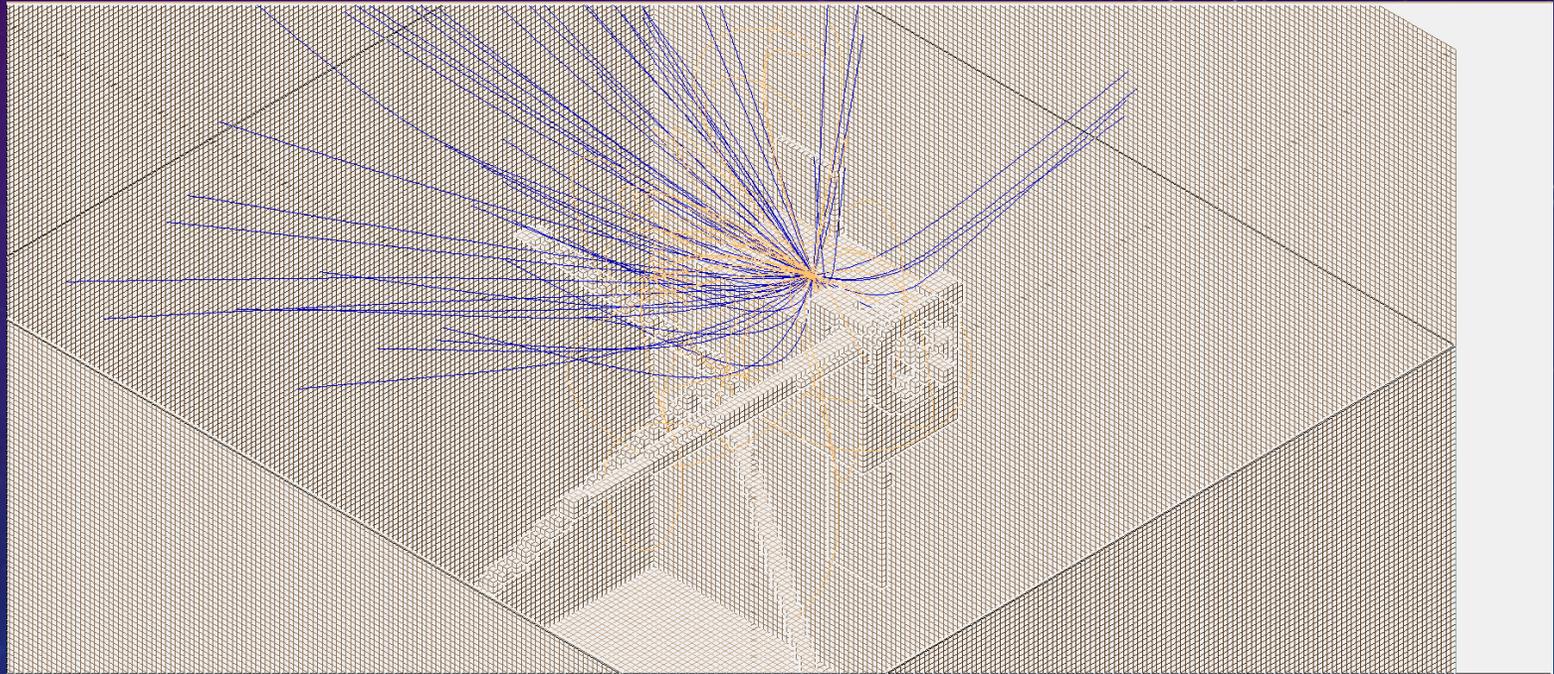
# REVIEW OF 2008 STEREO A WAVES DATA



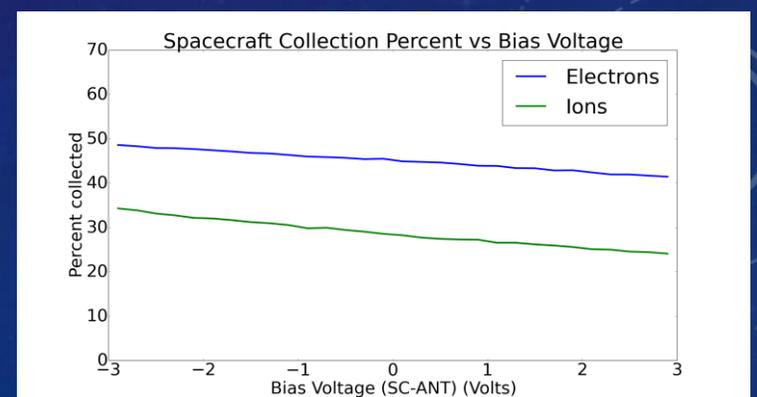
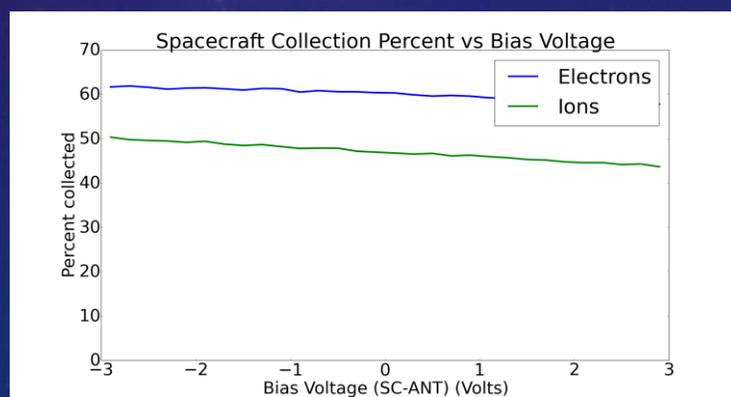
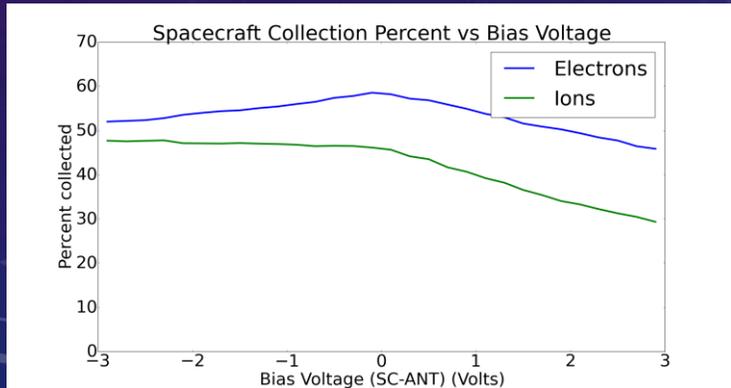
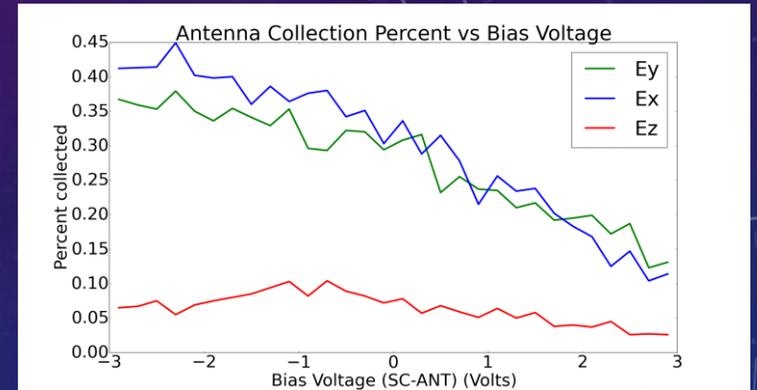
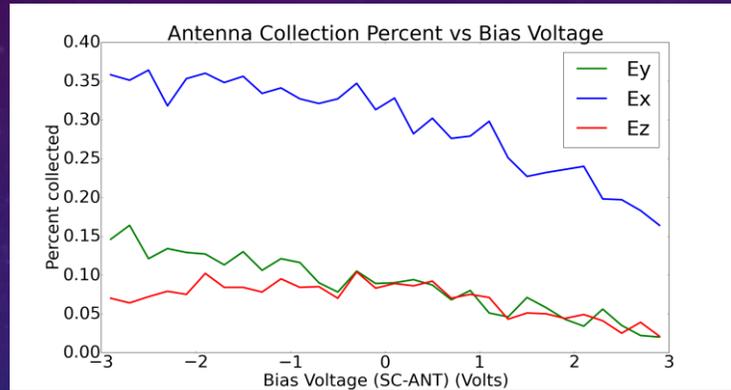
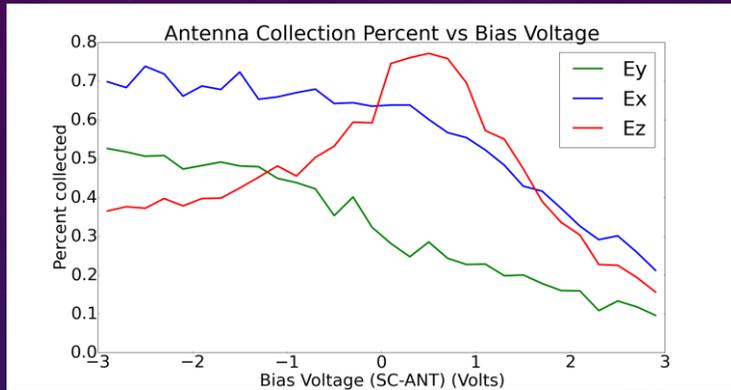
# SIMION<sup>®</sup>

## SIMULATIONS

- Particles were initially in a Maxwell-Boltzmann distribution of energy
- 5eV electrons and 30eV protons (ions) <sup>3</sup>
- Placed on different spacecraft locations
- Spacecraft was set to 5V and antennas were varied
- Output was TOF for each particle that hit the spacecraft or antenna

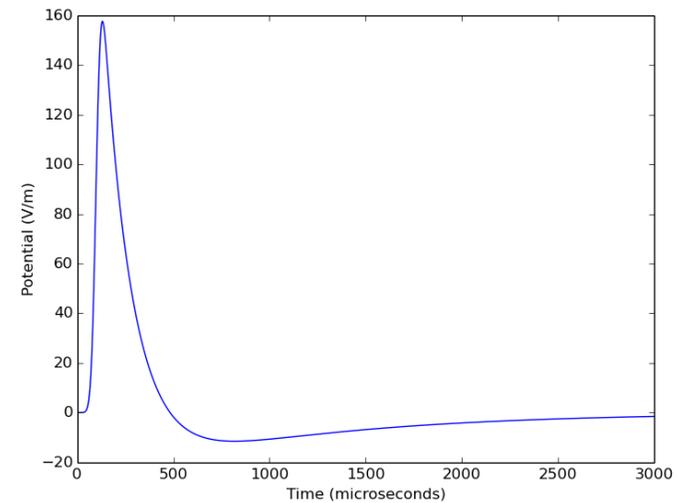
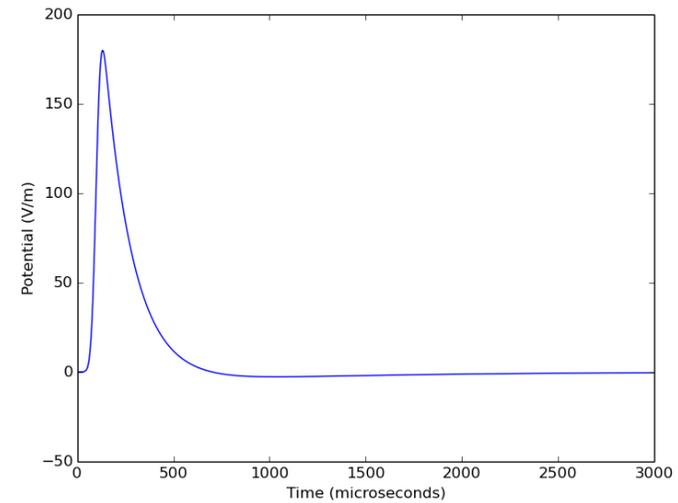


# ESTIMATES OF $P_{ANT}$ ON THE ORDER OF 0.01



# FPP MODEL USING $p_{ant}$ FROM SIMULATIONS

- Top:  $p_{ant} = \frac{1}{100}$
- Bot:  $p_{ant} = \frac{1}{30}$
- These overshoots are not large enough.



# CONCLUSIONS

- Many of the assumptions of the model hold such as electron collection times.
- There may not be enough charge collection to support FPP Model.
- The model cannot explain more than one signal at a time. It must be extended or amended.

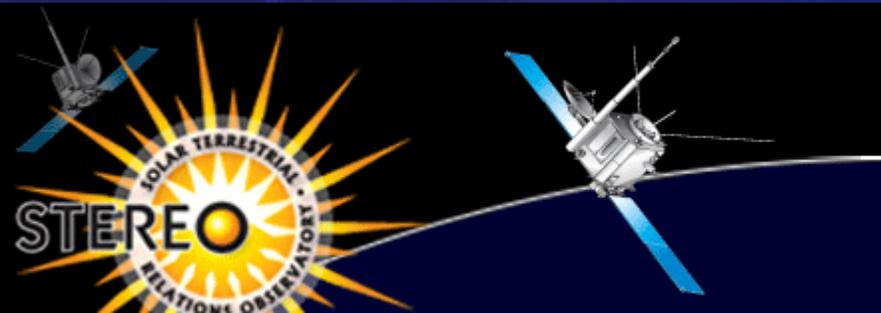
# REFERENCES AND ACKNOWLEDGEMENTS

- Thank you to Zoltan Sternovsky for the project, help and fun time throughout the REU program.
- I acknowledge the support from NASA through the IMPACT node of the Solar System Exploration Research (SSERVI) program. I participated in this project through NSF's Research Experience for Undergraduates (REU) program.
- Thank you also to David Malaspina, Ted, and others who helped with the project in anyway.
- References:

Meyer-Vernet, N., et al. (2009), Dust detection by the wave instrument on STEREO: Nanoparticles picked up by the solar wind?, *Sol. Phys.*, 256, 463–474, doi:10.1007/s11207-009-9349-2.

Zaslavsky, A. (2015), Floating potential perturbations due to micrometeoroid impacts: Theory and application to S/WAVES data, *J. Geophys. Res. Space Physics*, 120, 855–867, doi:10.1002/2014JA020635.

Collette, A., G. Meyer, D. Malaspina, and Z. Sternovsky (2015), Laboratory investigation of antenna signals from dust impacts on spacecraft, *J. Geophys. Res. Space Physics*, 120, doi:10.1002/2015JA021198.



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