

Recent results on Chariklo's rings

CHARIKLO



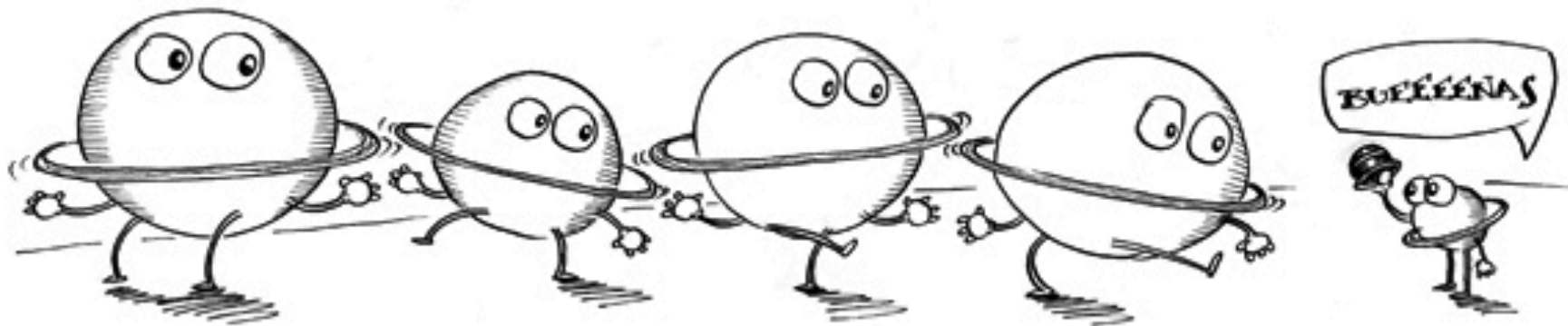
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Cornell University**

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(Cornell Univ.), Daniel Tamayo (Cornell Univ), Philip D. Nicholson (Cornell
Univ.), Joseph A. Burns (Cornell Univ.)**

Outline

- **The discovery**
- **Geometry and data processing**
- **Rings photometry and spectrometry**
- **Model**
 - **Photometry**
 - **Dynamics**
- **Results**
- **Origin**
- **Conclusion**

CONCURSO DE HULA-HULA



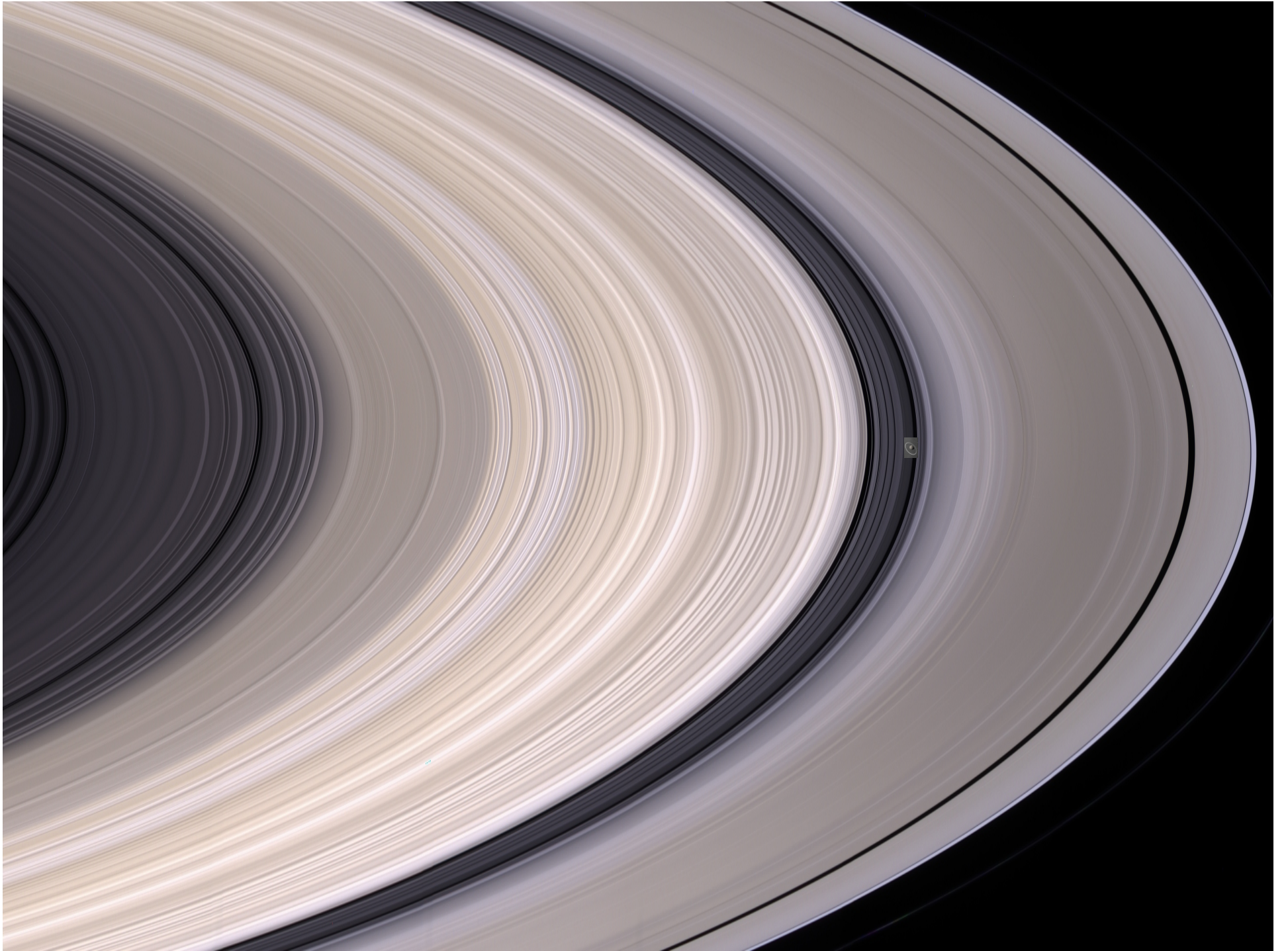
JUPITER
1979

NEPTUNO
1989

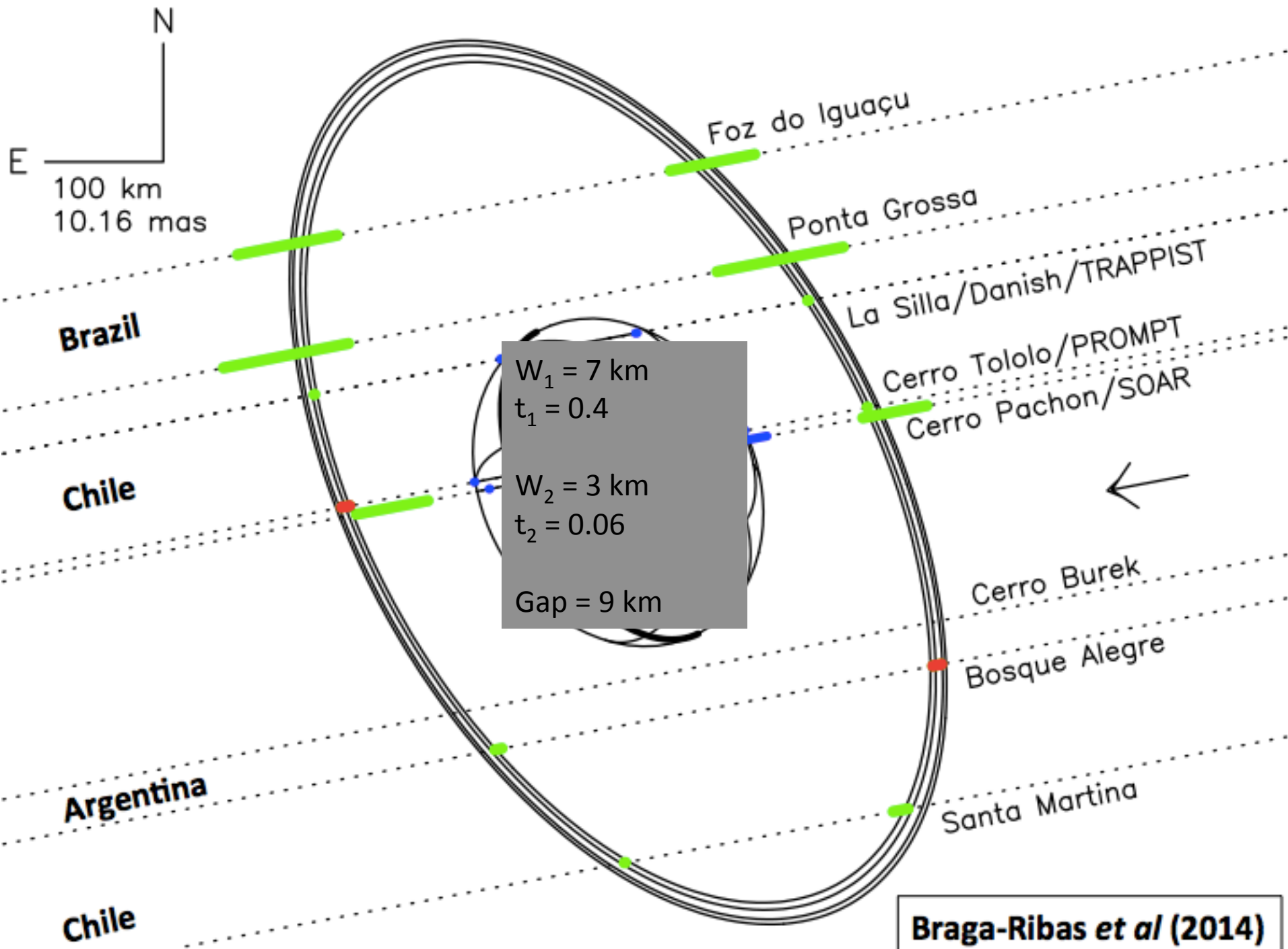
URANO
1977

SATURNO
1610

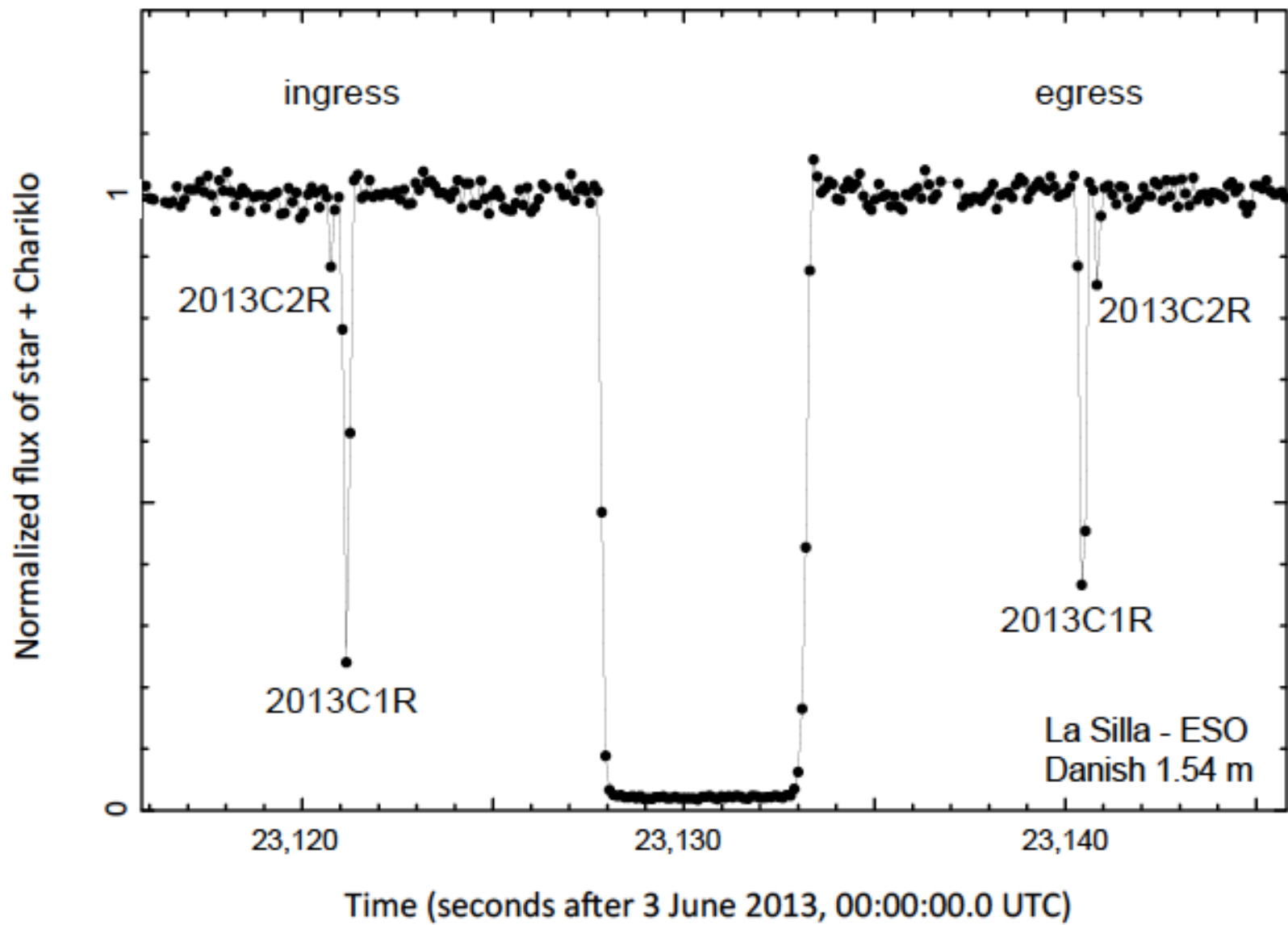
CHARIKLO
2013



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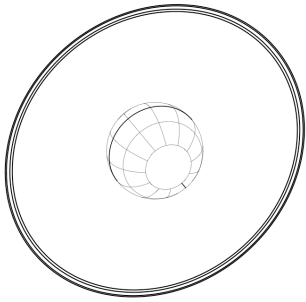


Braga-Ribas et al (2014)

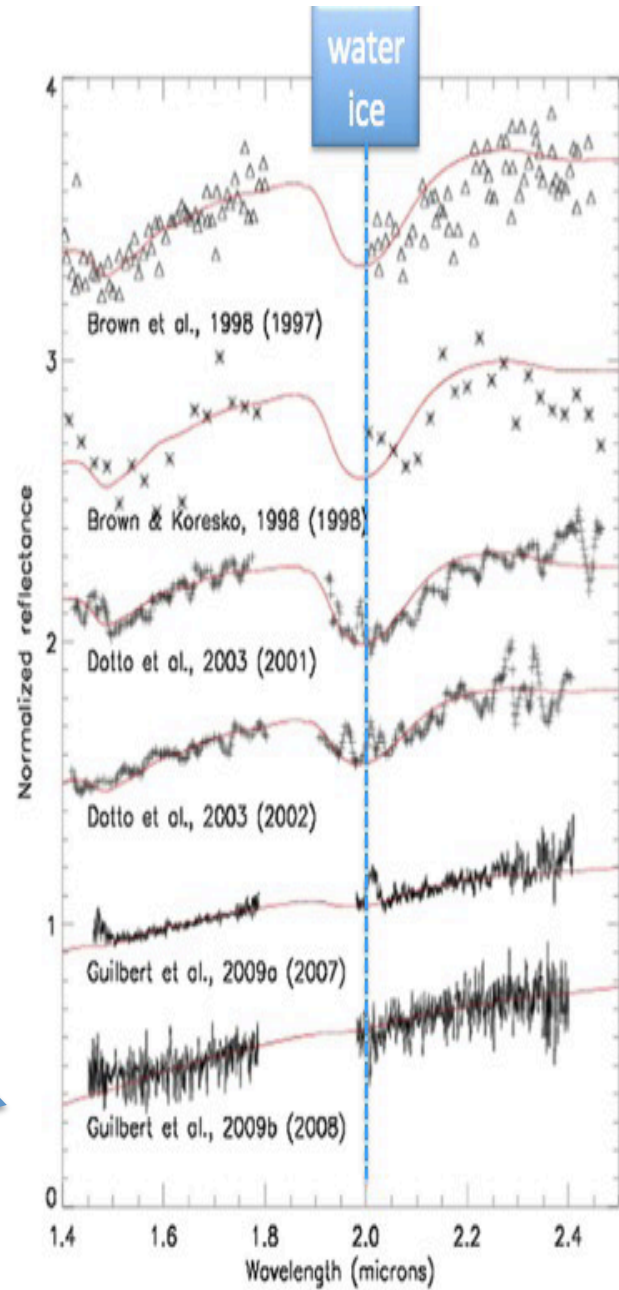
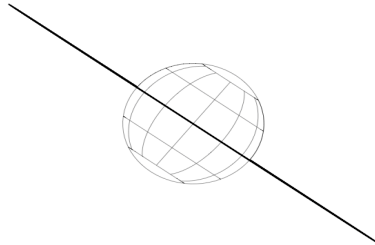


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maximum opening plane: end of 1995



ring plane crossing: end of 2007



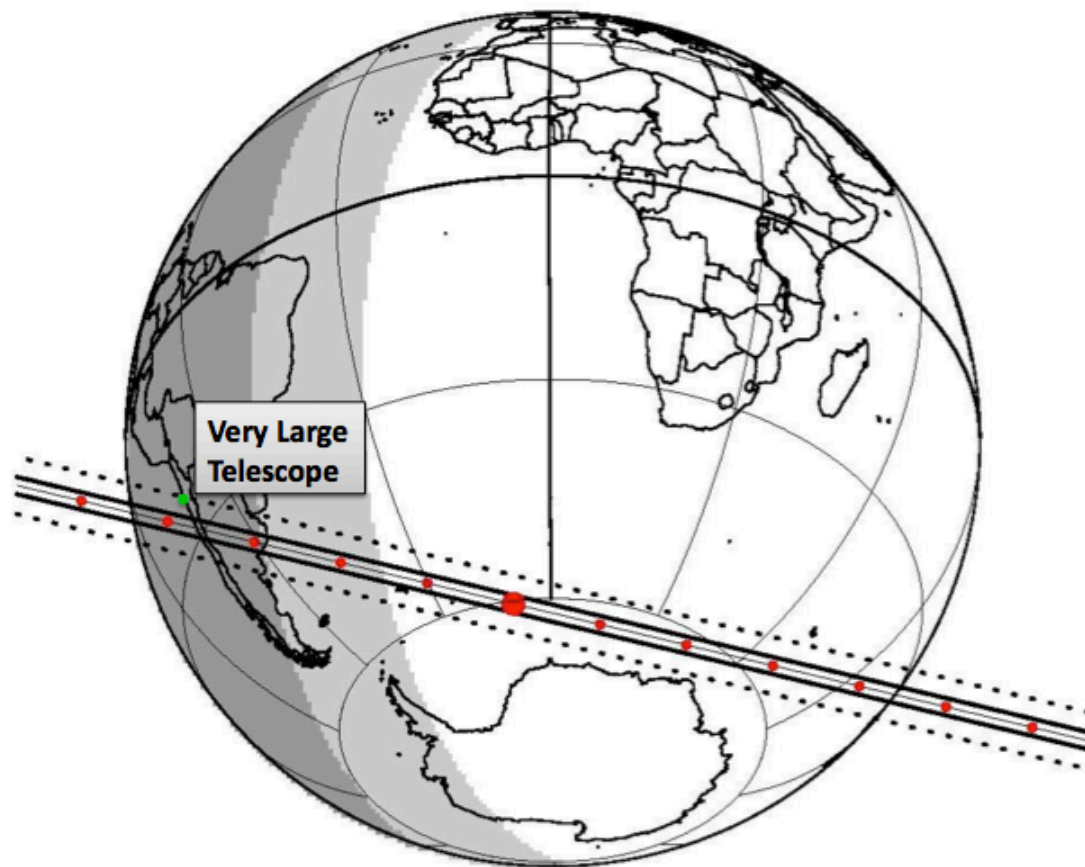
rings wide-open ($B \sim 50^\circ$)

rings edge-on ($B \sim 0^\circ$)

Guilbert-Lepoutre 2011

Chariklo: Star OPD08sep13_T60cm, NIMAv2 ephem.

Offset (mas): 0.0 0.0

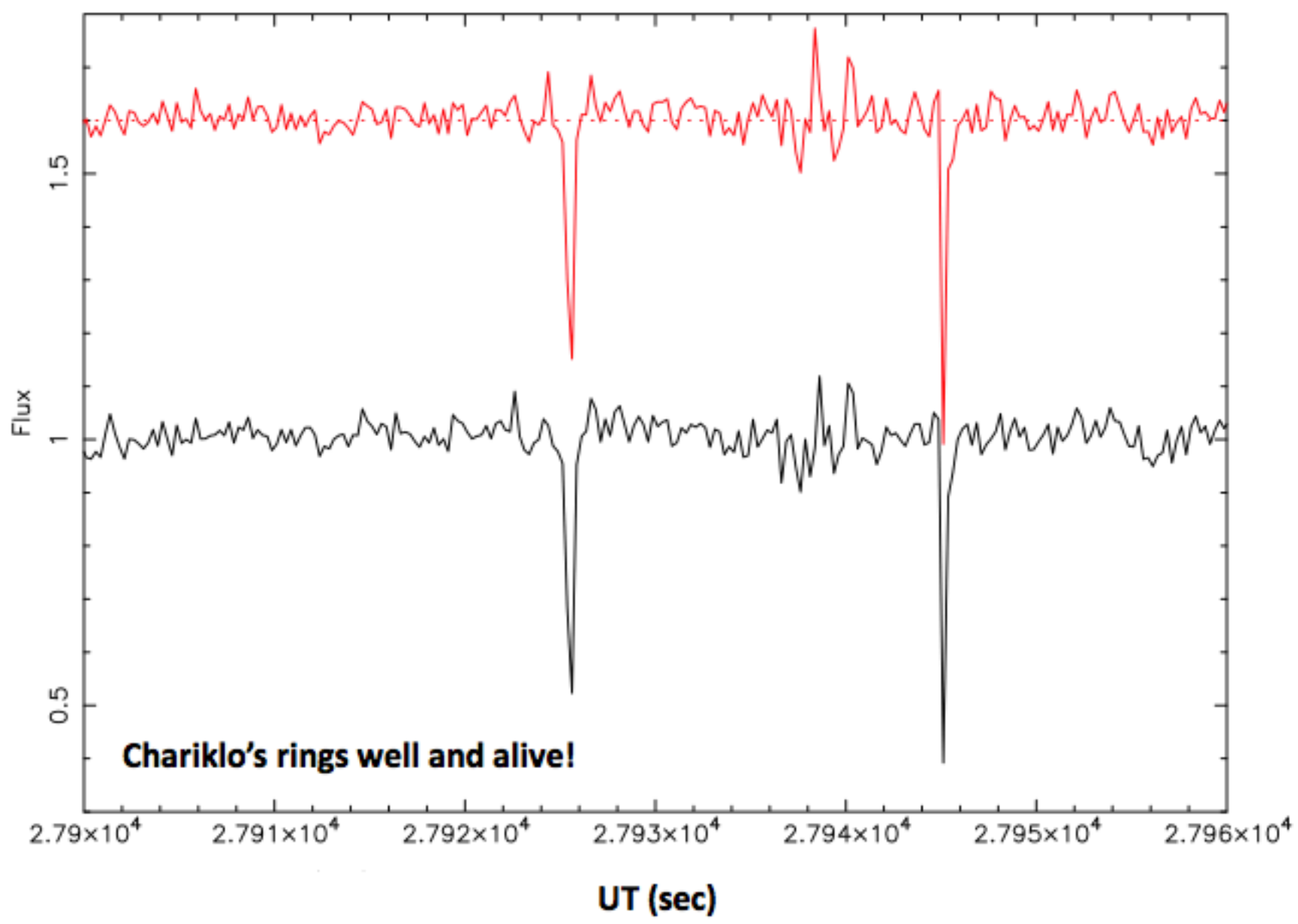


16 February 2014 Chariklo stellar occultation

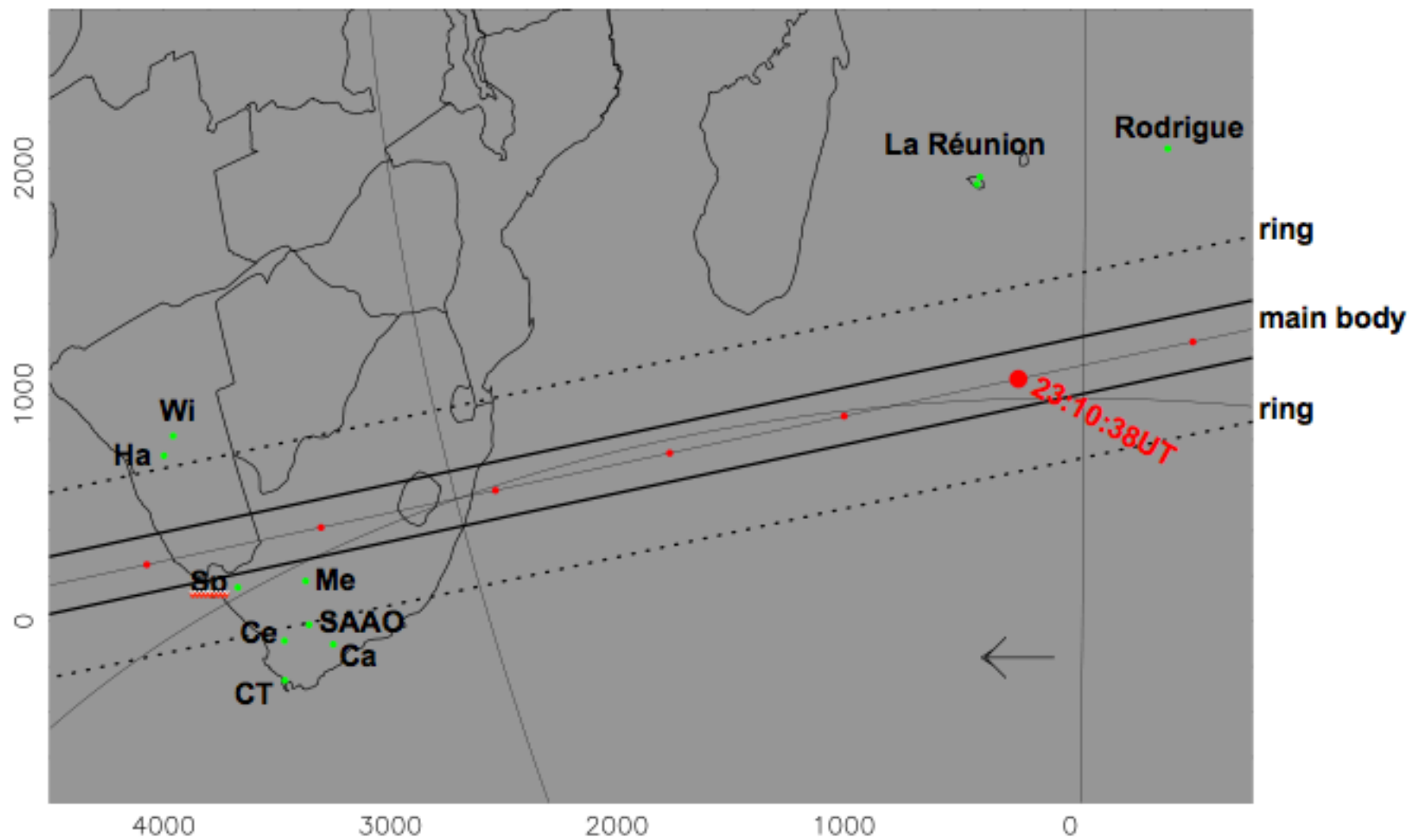


d	m	year	h:m:s UT	ra	dec	J2000_candidate	C/A	P/A	vel	Delta	R*	K*	long
16	02	2014	07 49 13.	17 35 55.3333	-38 05 17.184		0.230	193.40	22.12	15.15	16.9	12.0	0.

Very Large Telescope – HAWK-I 16 February 2014

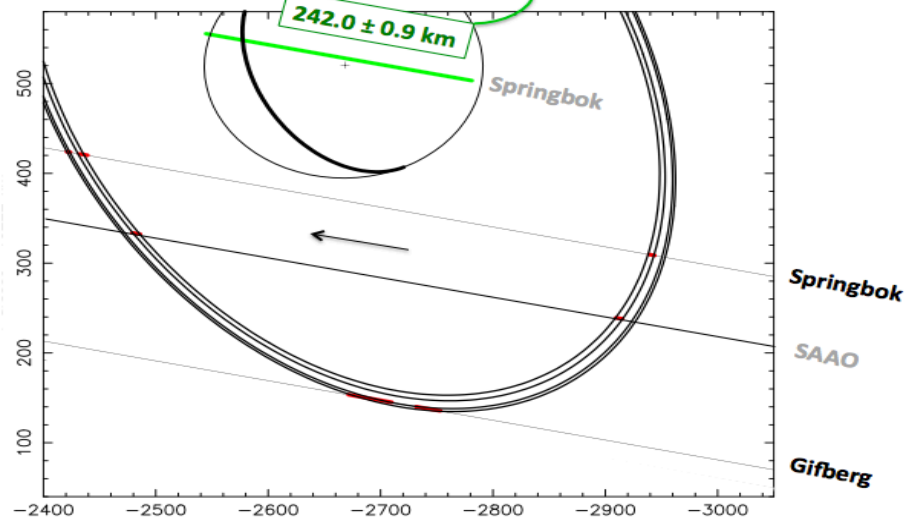


the 29 April 2014 Chariklo stellar occultation

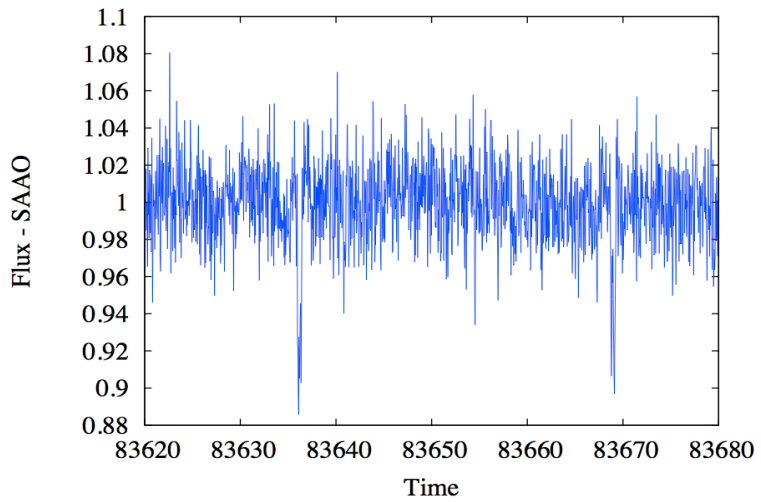


Chariklo 29 April 2014:
a double star !

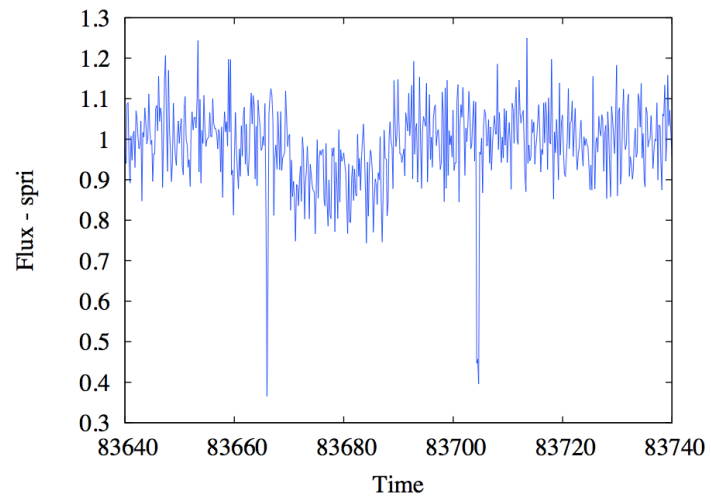
vs. 248 ± 18 km
(Fornasier et al. 2013)



Saao



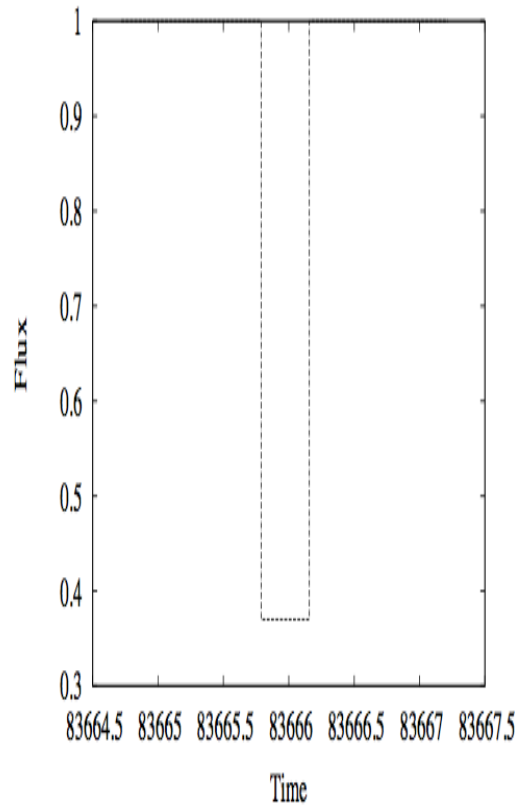
Springbok



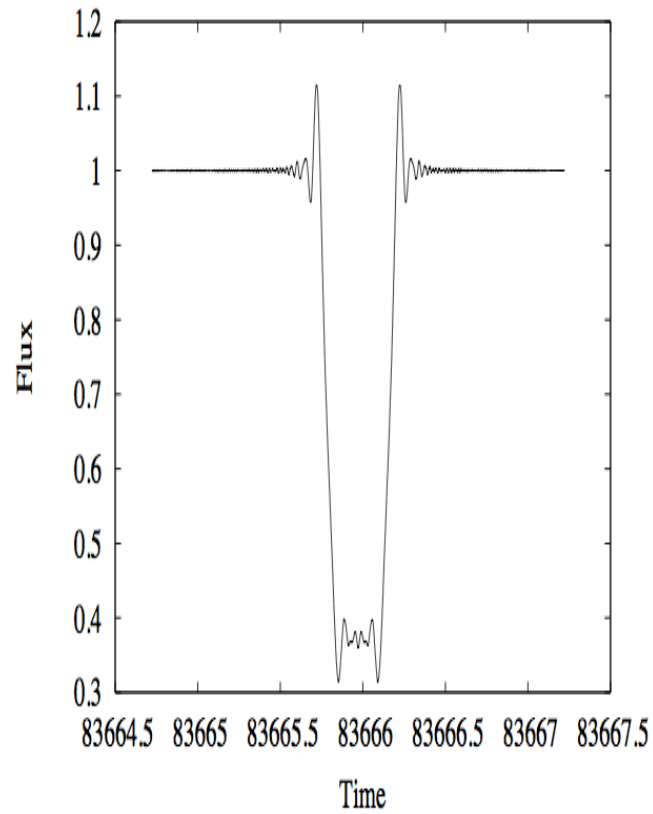
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Data analysis

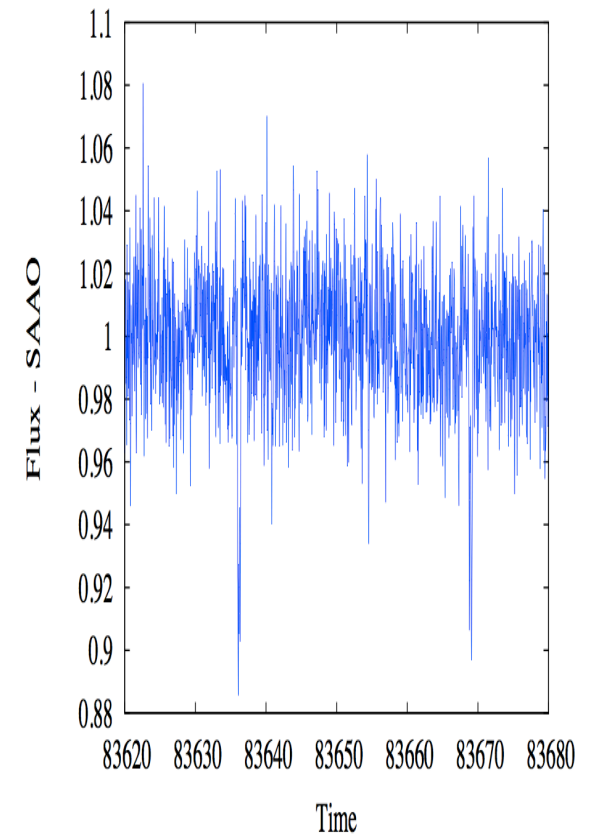
Square model



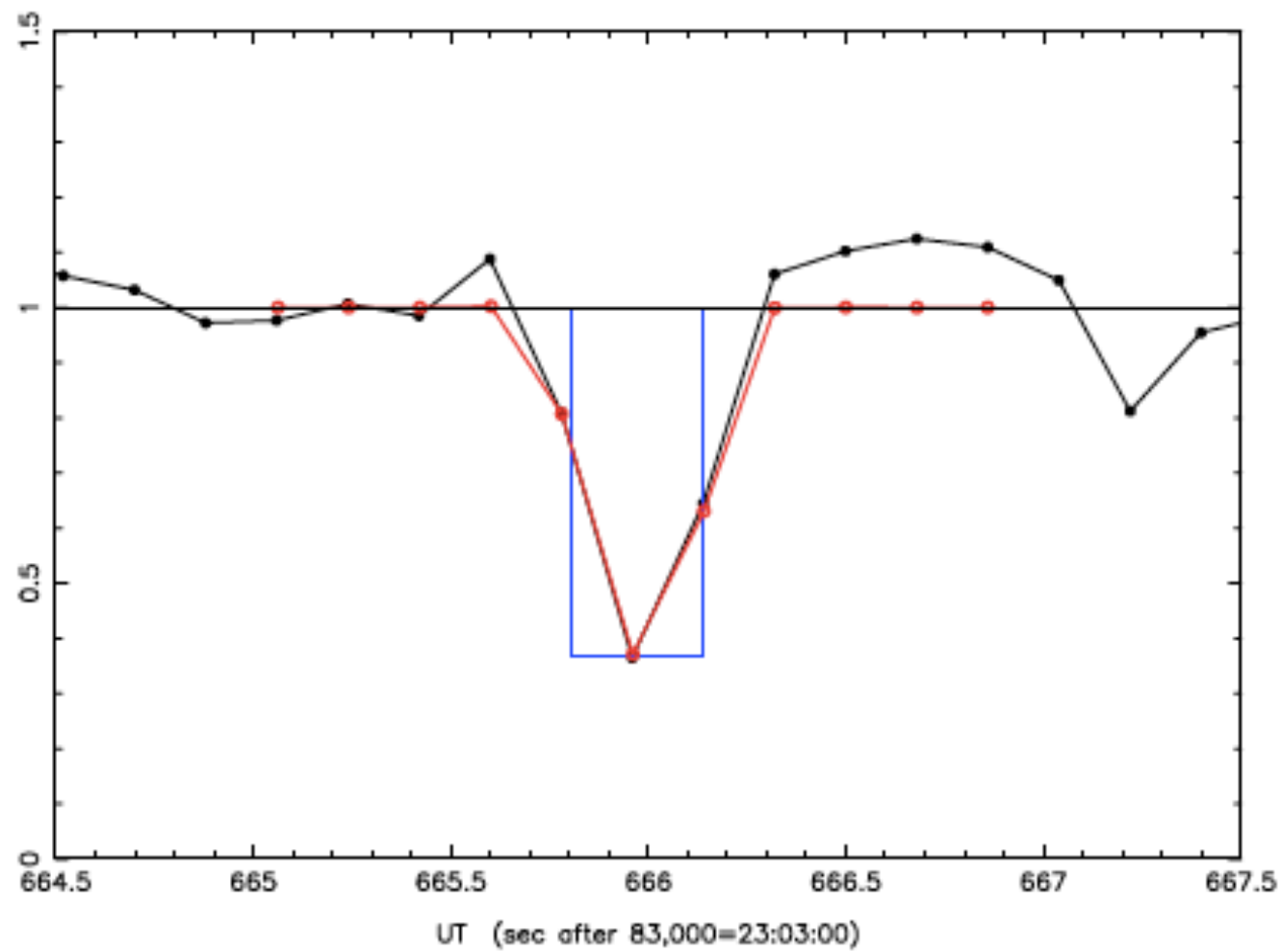
Diffraction model

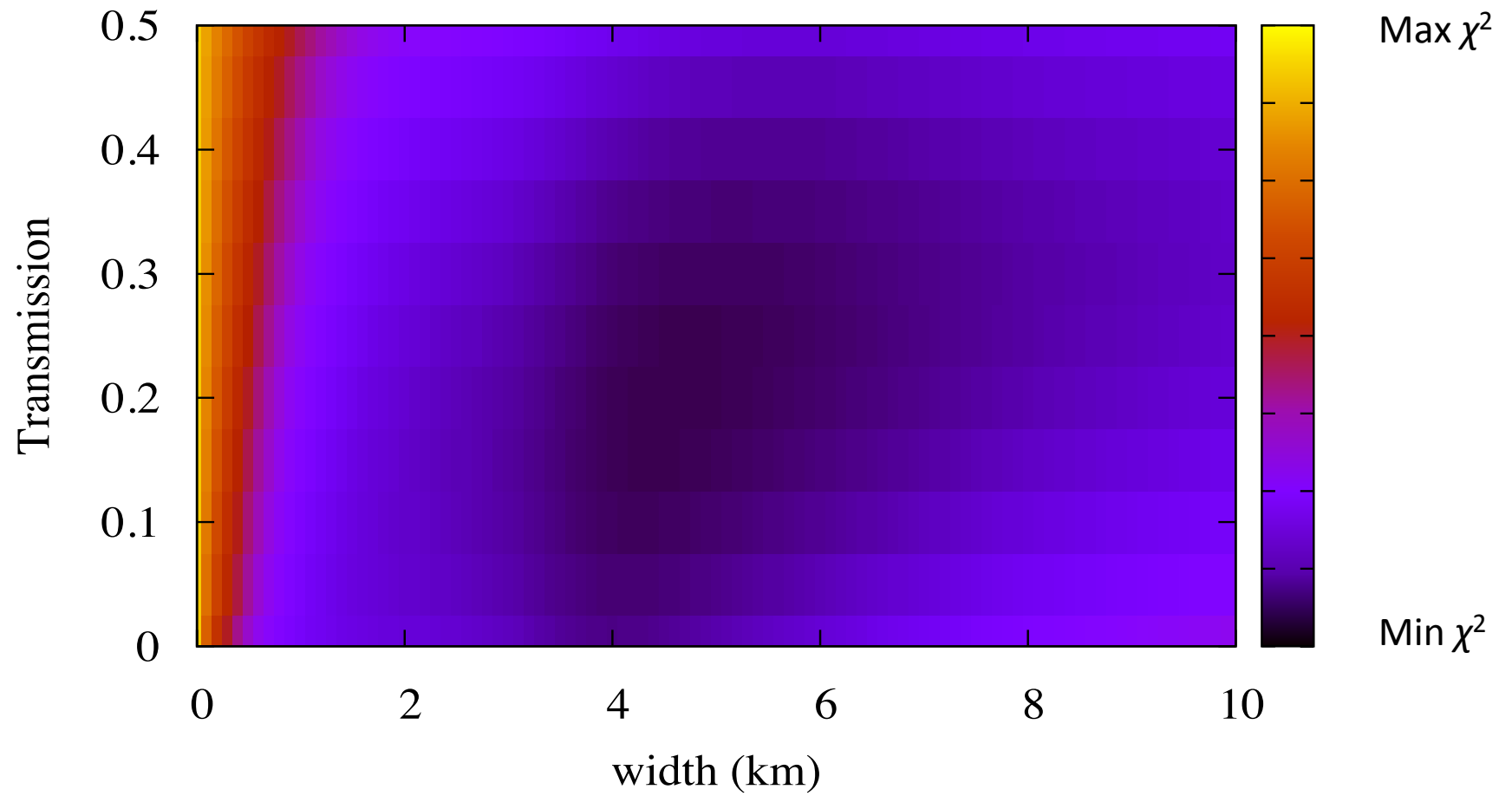


Data



Chariklo 29 April 2014, Springbok C1R imm.

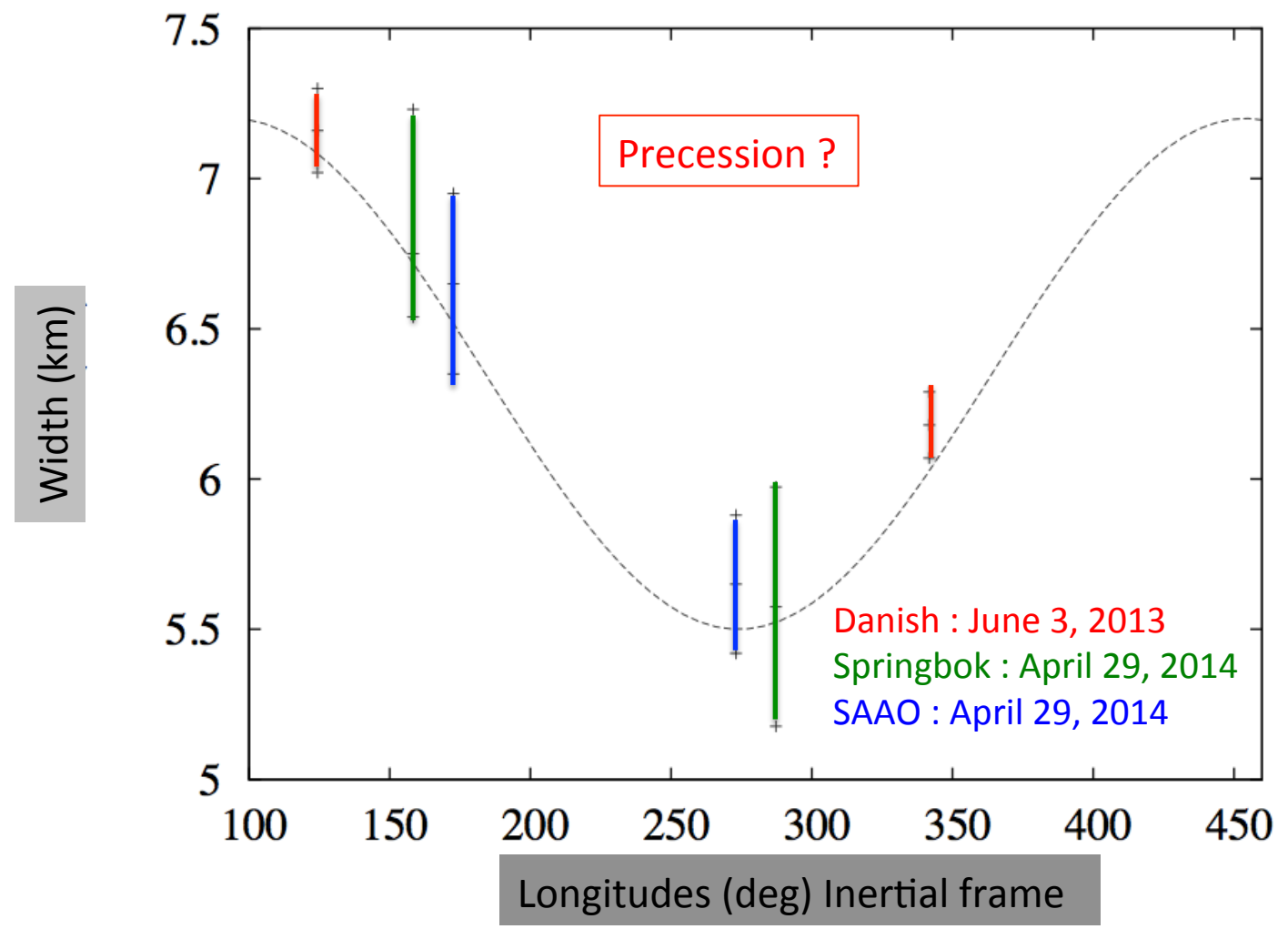




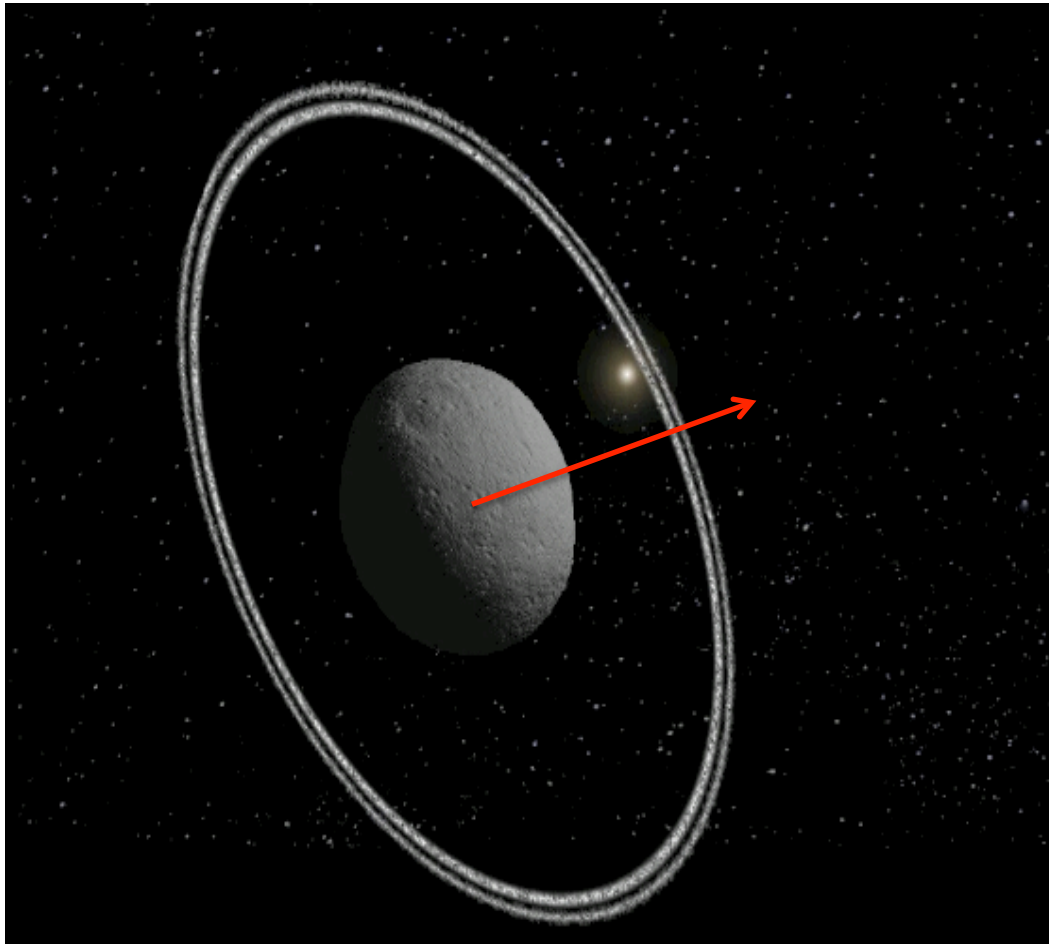
---> The **width** and the **transmission**

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Variation of width versus longitude



Shape and rotation of Chariklo



Fornasier et al 2014, A&A

$$\Omega = 7.004 \pm 0.002 \text{ hours}$$

Braga-Ribas et al 2014, Nature

$$e = 0.213 \pm 0.002$$

Chariklo:

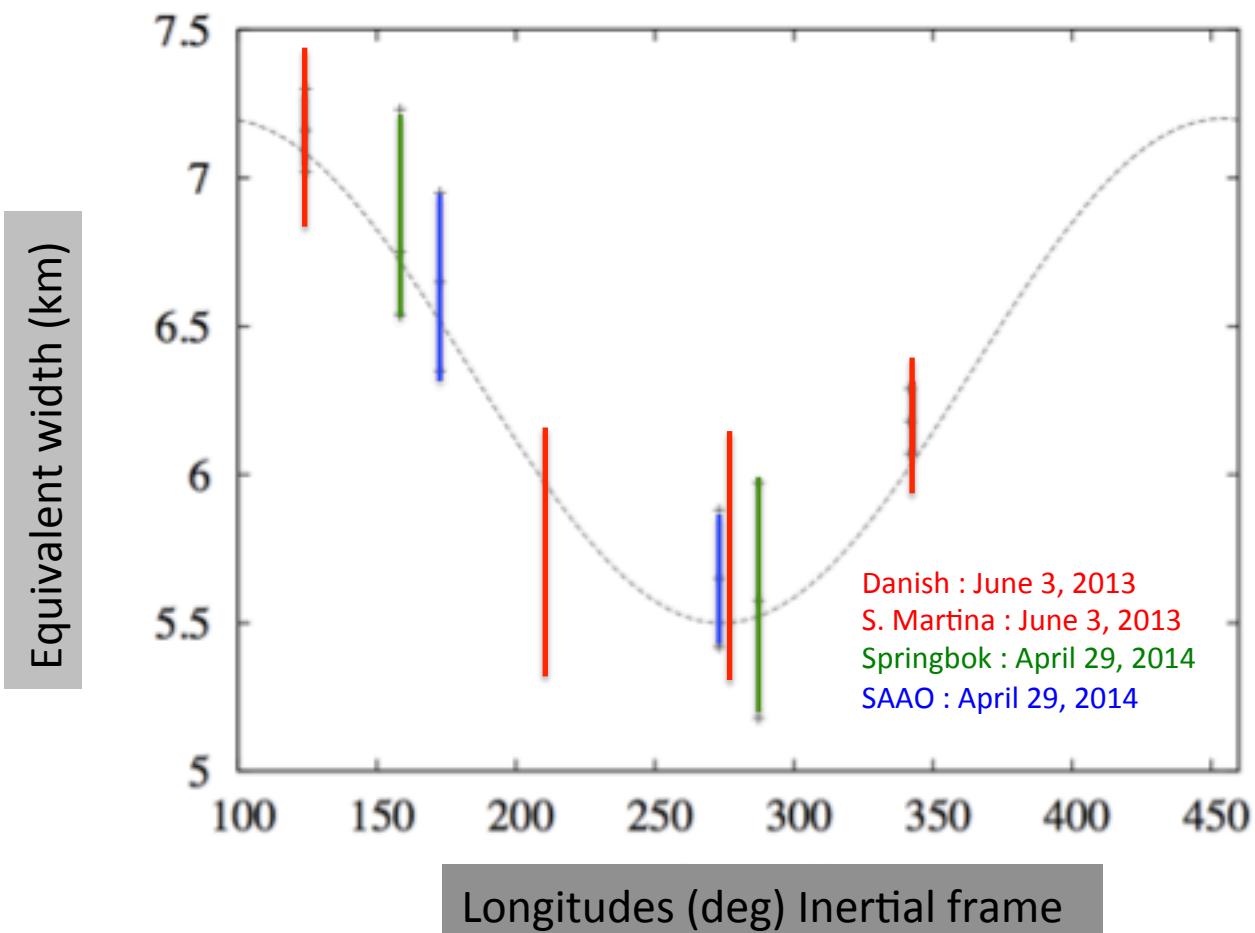
- Hydrostatic equilibrium
- Homogeneous

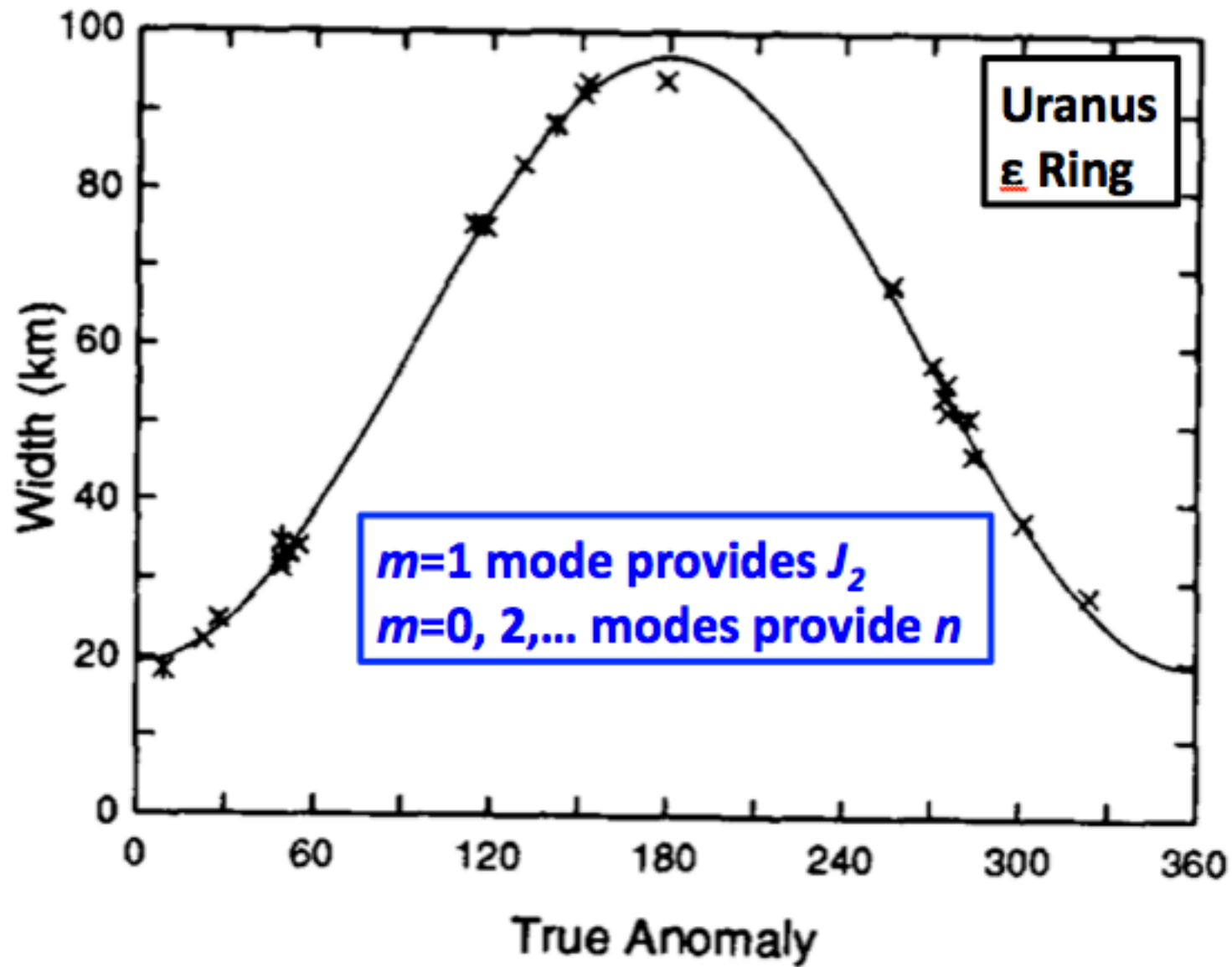


$$2.13 \times 10^{19} \text{ kg} < M < 2.17 \times 10^{19} \text{ kg}$$

$$J_2 = 0.044 \quad \text{???? very large !!}$$

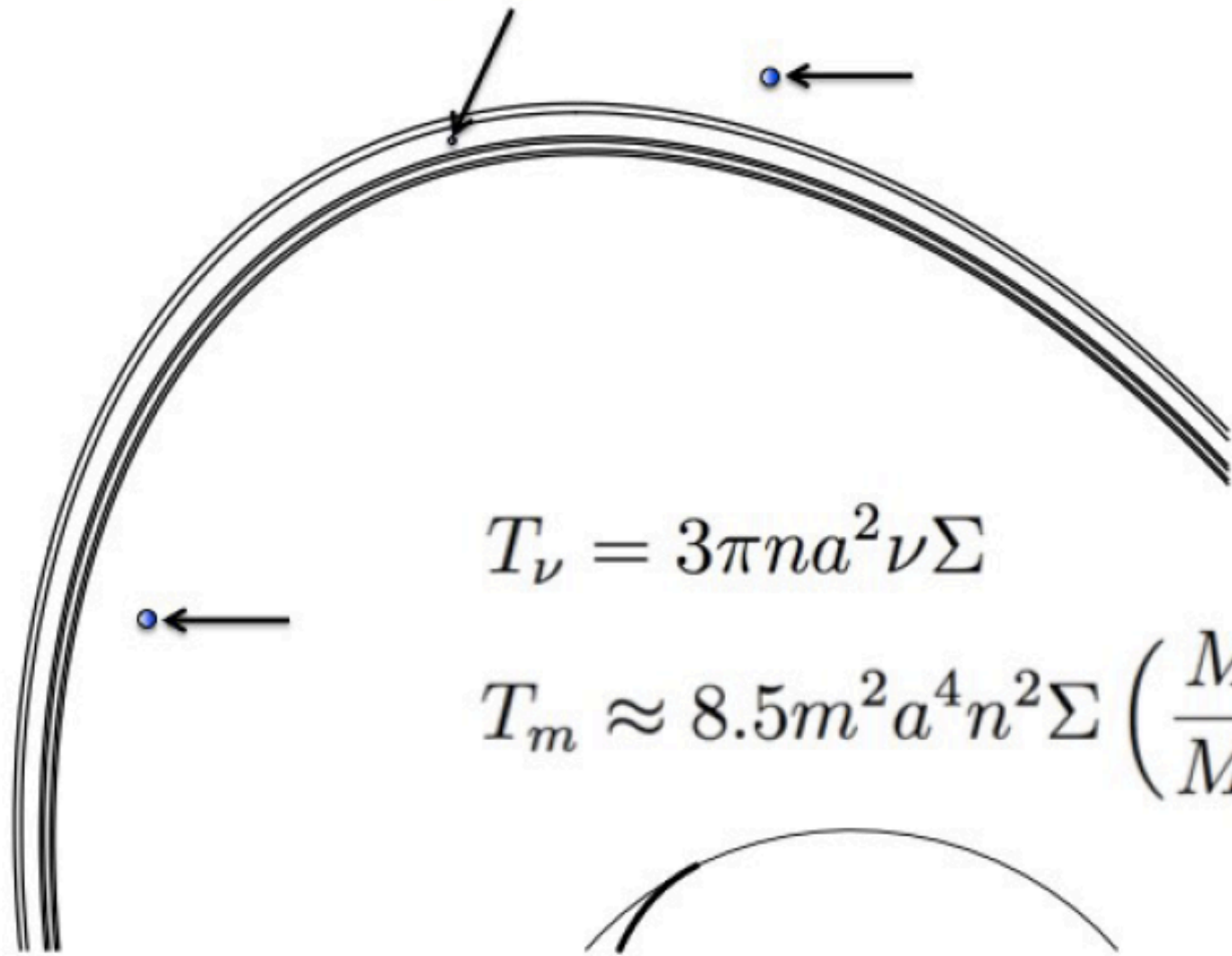
Variation of **equivalent width** versus longitude





French, Elliot & Levine
Icarus 67, 134 (1986)

shepherds ~ few-km-across moonlet
« gap opener » ~ sub-km moonlet



$$T_\nu = 3\pi n a^2 \nu \Sigma$$

$$T_m \approx 8.5 m^2 a^4 n^2 \Sigma \left(\frac{M_s}{M_C} \right)^2$$

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Origin ?

- **Exogenous?**

- Collision on Chariklo → debris+moonlets → rings
- Collision onto a pre-existing satellite → rings

- **Endogenous?**

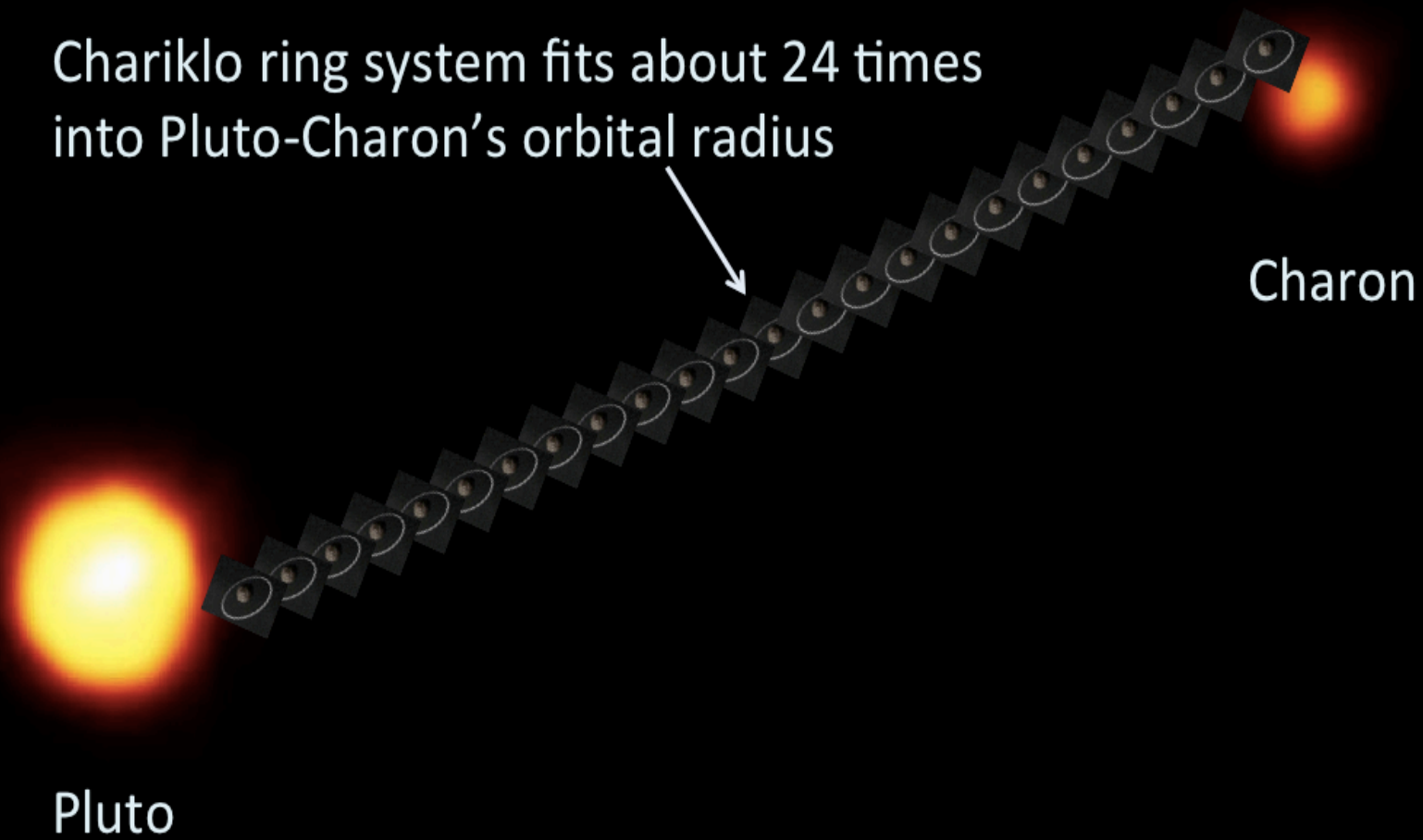
- Inter-satellite collision → rings
- Satellite spiralling inward → tidal breakup

Conclusion

- First detection of ring system around an asteroid
 - Small primary body (compared to giant planets)
 - Narrow rings (5 - 7 km) with high optical depth ($\tau = 0.4$)
- Confirmation by many other stellar occultations in 2014
- Similarities with Saturn rings (composition and optical depth)
- Similarities with Uranus rings (optical depth and width)
- Determination of width and transmission by square and diffraction model
- First determination of J_2 of Chariklo (Hydrostatic equilibrium), we need more data to check this value.
- Determination of size of shepherding satellites



Chariklo ring system fits about 24 times
into Pluto-Charon's orbital radius



Credit : Bruno Sicardy