Predator-Prey Model for Haloes in Saturn's Rings

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Cassini Observed 'Haloes' in Saturn's A Ring

- Annuli of increased brightness were seen by VIMS and UVIS at Saturn Orbit Insertion
- Found at strongest density waves, but not at Mimas
 5:3 bending wave



Halo Morphology

- UVIS & VIMS: centered about 100km outside the resonance, a bright region extending about 450 km away from the center;
- VIMS sees a region of increased grain size, with a half-width of about 100km, and smaller grain sizes in symmetric bands about 500 km wide; smallest grain size about 200 km away;
- UVIS sees smaller correlation lengths at the wave location, the lows are coincident with the density wave crests: size distribution changes in hours!
- ISS sees 'straw' in the troughs of strong density waves;
- Radio occultations also show a signature (Marouf, private communication).

UVIS SOI (150 km resolution elements)









We See Both Smaller and Larger Correlation Lengths in Density Waves

- Inferred sizes from excess variance are anticorrelated with optical depth, as predicted in the Predator-Prey model for moonlettriggered accretion: mass aggregates and peak collision velocity are out of phase
- Straw in SOI images between the wave crests
- The Janus 2:1 density wave region shows both larger and smaller particles than in nearby comparable regions in the B ring



What's to Explain?

- Connection to strongest resonances [Predator-prey model of moon-triggered accretion]
- Halo morphology: Width, shape, why they are centered outside resonance [Stirring by aggregates, transport and diffusion]
- Photometry: Brighter in UV, visible, IR [Production of new grains on exposed surfaces]
- Spectroscopy: Larger particles in VIMS spectra, surrounded by smaller particles at greater distances [Aggregate stirring, grain production and transport]

Predator-Prey Model: Use this to simplify ring dynamics

- Periodic forcing from the moon causes streamline crowding
- This damps the relative velocity, and allows aggregates to grow
- About a quarter phase later, the aggregates stir the system to higher relative velocity
- The limit cycle repeats each orbit, with relative velocity ranging from nearly zero to a multiple of the orbit average: 2-10x is possible

Predator-Prey Equations for Ring Clumping

$$M=\int n(m) m^2 dm / ;$$
$$V_{rel}^2 = \int n(m) V_{rel}^2 dm / N$$

$$dM/dt = M/T_{acc} - V_{rel}^{2}/v_{th}^{2} M/T_{coll}$$
[accretion] [fragmentation/erosion]

 $\frac{dV_{rel}^{2}}{dt} = -(1-\epsilon^{2})V_{rel}^{2}/T_{coll} + (M/M_{0})^{2}V_{esc}^{2}/T_{stir}$ [dissipation] [gravitational stirring] $-A_{0}\cos(\omega t)$ [forcing by streamline crowding]

Upgrades to Predator-Prey Model: Collisions among Ring Particles

- Add stochastic forcing to simulate aggregate collisions: Random outcome doubles or halves aggregate mass. Previously, no collisions.
- Add threshold for gravity-bound aggregates: above this it is harder to disrupt aggregates. Previously, erosion of aggregates from Blum (2007). Results in higher collisions speeds!
- This allows us to find the fixed points, their stability, basins of attraction, and asymptotic behavior, not easy for N-body codes



Transport Model: Collisions cause thermal diffusion; Model this, following Feller (1971)

- Random walk on the line, with particle jumps distance Δx_i , every time step ΔT .
- In the production region (where we observe the density wave damping, $\pm x_D^{-100}$ km), mean Δx_j = 200km, ΔT = 100 T_{orb}
- Outside the production region, we set mean $\Delta x_j = 0.5$ km
- The jumps are drawn from a Gaussian distribution

Effects of Transport

- The higher collision velocity arises from stirring by aggregates, whose formation is triggered by streamline crowding in the wave, as hypothesized in the Predator-Prey model
- Gravitationally bound aggregates can form, which provide even higher peak velocities, up to 10 m/s near them
- This preferentially removes the regolith in the production region, exposing fresh ice surfaces to collisions and meteoritic bombardment, as in the model of Elliott and Esposito



Production Model

- Meteorites produce new regolith material if the regolith is thin enough, h(r,t) < h_{threshold}.
- This fresh material is then transported by the Brownian motion random walk transport
- We assume an initial regolith depth of 10cm, with pollution of 5% (Elliott and Esposito)
- We assume complete vertical mixing after each transport, and track the pollution to predict the reflectivity profile of the halo

Model results: We can fit the UVIS and VIMS morphology if

- Perturbation centered 100km outside the ILR: this is the location of max damping, see X_D from Esposito 1983
- Jump distance for Brownian motion is 200km, corresponding to $V_{ei} = 10m/sec$
- Production rate is 10⁻⁴ cm/yr
- Diffusion in A ring back ground region is small: 100cm²/sec, corresponds to jump distance 200m in ΔT=100T_{orb}= 4.8x10⁶sec



Summary

- A predator-prey model for ring dynamics produces transient structures like 'straw' that can explain the halo structure and spectroscopy:
 - Cyclic velocity changes cause perturbed regions to reach higher collision speeds at some orbital phases, which preferentially removes small regolith particles
 - Surrounding particles diffuse back too slowly to erase the effect: this gives the halo morphology
- This requires energetic collisions (v ≈ 10m/sec, with throw distances about 200km, implying objects of scale R ≈ 20km)

Ring dynamics and history implications

- Moon-triggered clumping at perturbed regions in Saturn's rings creates both high velocity dispersion and large aggregates at these distances, explaining both small and large particles observed there
- This confirms the triple architecture of ring particles: a broad size distribution of particles; aggregate into temporary rubble piles; coated by a regolith of dust
- Aggregates can explain many dynamic aspects of the rings and can renew rings by shielding and recycling the material within them