# Fluxes within lo Oval: North vs. South 

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## Introduction

- The internal magnetic field models of Jupiter seem to have stronger magnetic fields in the northern hemisphere than in the south
- Flux, through the ovals created by Io's footprints, should be equal for the north and south



# Step 1: Trace Io Footprints Onto Sphere 

- Io Footprints come from Bertrand Bonfond and is reproduced in Hess et al. (2011)
- For each model, we trace along the magnetic field lines that pass through the lo footprints from an oblate spheroid to a spherical surface
- 1 /oblateness = 15.41
- Step Size of Trace $=0.0001 \mathrm{R}_{\mathrm{J}}$
- Easier to integrate over a sphere than an oblate spheroid

Figure
Exaggerated for Emphasis

# Step 2: Rotate Ovals to the Equator 

- Differential areas are more constant along the equator of a sphere than the poles
- In spherical coordinates, a rotation of $90^{\circ}$ is:

$$
\begin{aligned}
& \theta^{\prime}=\arccos (\sin \theta \sin \phi) \\
& \phi^{\prime}=-\arctan (\cot \theta \sec \phi)
\end{aligned}
$$

- Each point must be rotated back to the poles to calculate the magnetic field using:

$$
\begin{aligned}
& \theta^{\prime}=\arccos (-\sin \theta \sin \phi) \\
& \phi^{\prime}=\arctan (\cot \theta \sec \phi)
\end{aligned}
$$

Must correct for ambiguity of arctangent

# Step 3: Interpolate Between lo Footprints 

- To determine the boundaries of the surface integral, must interpolate between the 36 Io footprints
- Used IDL's INTERPOL command with a Spline fit
- Cubic spline to the nearest four neighbors
- Done twice. Once for the upper half of the oval and once for the lower half of the oval
- Used five times +1 the number footprints, spread evenly over longitude


## Step 4: Numerically Integrate

- Solve for the flux by multiplying $B_{r}$ by the differential area $r^{2} \sin \theta d \theta d \phi$ over the entire oval - $r=1 R_{j}$ on the sphere
- We use a differential angle of $0.01^{\circ}$ for both $\mathrm{d} \theta$ and d $\phi$



## Results

|  | VIP4 | VIT4 | VIPAL |
| :---: | :---: | :---: | :---: |
| North (TWb) | 4.062 | 3.885 | 3.856 |
| South (TWb) | -4.010 | -4.086 | -3.680 |
| -North/South | $\mathbf{1 . 0 1 3}$ | $\mathbf{0 . 9 5 0 9}$ | $\mathbf{1 . 0 4 8}$ |

## Conclusions

- There is a discrepancy between the north and south of:
- VIP4: 1.3\%
- VIT4: $4.9 \%$
- VIPAL: $4.8 \%$
- Removing interpolation only creates a $0.2 \%$ increase in discrepancy
- Adding the atmosphere reduces the discrepancy by $0.5 \%$
- Maybe errors in footprint locations can account for this?
- Not likely to be a big enough factor
- This method could be another constraint on future magnetic field models

