## Predator-Prey Analogs for Saturn's Non-Linear Ring Dynamics

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# Transient ring structures appear where the rings are perturbed strongly

- Equinox objects at the Mimas 2:1 resonance
- Straw between density wave crests
- Excess variance increases between wave crests
- Gap edges when a moon passes by
- Solitary waves where the Janus resonance falls on Epimetheus density wave every 8 years



#### 'Straw' seen between density wave crests must form in less than 10 hours



Straw from Cassini Grand Finale



Alpha Virginis Rev 34I

Particle statistics show larger structures between density wave crests



#### Daphnis Edge Wake shows downstream effect



Ring Edge Shears and Separates









Solitary wave propagating through A ring, seen every 8 years

### How to explain this dynamic structure?

- Solitary waves and the large amplitude, rapidly growing transient structures indicate non-linear phenomena
- N-body simulations are too slow, and don't include all the physics
- Use a simpler model with an ecological analogy: *Predator–Prey*
- Track the mass and velocity dispersion: Relative velocity is stirred up by clumps, velocity disrupts clumps
- Force the system by the moon's gravity driving the surface mass density and the velocity dispersion
- Allow for disk instability, using Toomre's dispersion formula
- Use numerical simulation results for outcomes of stochastic collisions (Hyodo & Ohtsuki; Leinhardt & Stewart)

#### Predator-Prey Equations for Ring Clumping (Esposito *et al* 2012) M=∫n(m) m<sup>2</sup> dm / <M>; V<sub>rel</sub><sup>2</sup>=∫n(m) V<sub>rel</sub><sup>2</sup> dm / N

 $dM/dt = M/T_{acc} - V_{rel}^2/v_{th}^2 M/T_{coll}$ [accretion] [fragmentation/erosion]

 $\frac{dV_{rel}^{2}}{dt} = -(1-\epsilon^{2})V_{rel}^{2}/T_{coll} + (M/M_{0})^{2} V_{esc}^{2}/T_{stir}$ [dissipation] [gravitational stirring] - A\_{0} cos(\omega t) [forcing by streamline crowding]

The aggregate mass, M is the 'prey'; The dispersion  $V_{rel}^2$  is the 'predator': It feeds off of the mass by grav stirring; The predator reduces the prey by erosion

#### Predator Prey Model with Logistic Growth

*Basic equation* :

$$T_{S} = \frac{T_{orb}}{2\pi\tau'_{S}}$$

$$\tau_{s}' = \tau_{s} \left(1 - \frac{M}{M_{L}}\right)$$

$$\frac{dM}{dt} = \left[\frac{M}{T_{S}}\frac{S}{R}\frac{\rho_{0}}{\rho} + \frac{M}{T_{DI}} - \frac{M}{T_{S}}\frac{S}{R}\frac{\rho_{0}}{\rho}\frac{V_{rel}^{2}}{V_{th}^{2}}\right]$$

$$\frac{M}{T_{DI}} = 0 \text{ when } \omega^{2} \ge 0$$

$$\frac{dV_{rel}^{2}}{dt} = -\frac{V_{rel}^{2}}{T_{S}}\left[(1 - \varepsilon^{2}) + \frac{\tau_{B}}{\tau'_{S}}\right] + \frac{V_{esc}^{2}(M)}{T_{S}}\frac{\tau_{B}}{\tau'_{S}}$$

**Motivation:** When their Mass reaches a limiting value, the aggregates cannot grow further by ring particle sweep up, since we have only a finite number of ring particles to stick to the growing aggregates. This is modelled by adding a *logistic growth* term limiting the optical depth of smaller aggregates.

Thus, the closer the aggregate mass M to M\_limit, the slower the growth rate.

**Note:** After the limiting mass is reached, the aggregates change in mass only due to stochastic collisions which yield accretions and disruptions.

#### Strength regime surface mass density forcing phase plots



4% variation

$$\Sigma = \Sigma_0 \frac{1}{\left(1 + m \sin\left(\omega_{syn}t\right)\right)}, m = 0.97$$

Disk instability when:  

$$\omega^{2} = \kappa^{2} - 2\pi G \Sigma |k| + v_{s}^{2} k^{2}$$

$$\omega^{2} < 0$$

# Equilibrium distribution of aggregates from stochastic collisions is a power law



### Conclusions

- Ring structure shows transient clumping in perturbed regions: We conclude that forcing by the moon triggers aggregation. This also increases the relative velocity, liberating small particles
- The structure forms rapidly, on orbital time scales, out of phase with the moon... and downstream of the moon's wake initiation
- We find that growth by sweep-up is too slow to explain the excess structure observed in between density wave crests
- Gravitational disk instability can act on orbital time scales; We use Toomre's stability parameter Q to estimate the growth rate for clumps
- We achieve rapid growth by modulating the surface mass density, decreasing the velocity dispersion or by decreasing the shear
- Aggregates from stochastic collisions have a power-law size distribution

Take away message: Moon forcing drives accretion, triggers disk instability, producing transient clumps downstream: A continuing process of Cosmic Recycling

## Back-Up



We can force the Predator-Prey model by surface mass density or by velocity variations, which give similar outcomes.



### Clump mass M, from $V_{rel}=0$ for > 0.3T<sub>orb</sub>



Mass variations around the fixed points: 0.7torb = 64.4% 0.6torb = 60% 0.5torb = 51% 0.4torb = 48% 0.3torb = 31%

### Using Numerical simulations results

- What happens when equal sized object collide randomly?
- Can Numerical simulations be used to base the statistics of random events?

## Using the results of Hyodo and Ohtsuki:

- The outcomes of the stochastic events are based on the ratio of Impact/Escape velocity and the direction of collision.
- The direction of collision can be radial, azimuthal and vertical. Direction is chosen with equal probability.





#### Hyodo and Ohtsuki: 140K km case simulation

Random Event Outcomes:

- Accretion : Green region: This event doubles the current mass.
- Hit and run : Blue region: This event does not change the mass.
- disruption : Red region: This event halves the mass.

Note: This simulation considers presence of strong tidal waves. (Distance from Saturn : 140k km)

# Limiting mass calculation: computed based on cell size

$$M_{L} = 500m \times 10^{4} m \times \Sigma_{0}$$

$$M_{L} = 5 \times 10^{4} cm \times 10^{6} cm \times 100g / cm^{2}$$

$$M_{L} = 5 \times 10^{12} g \text{ or } 5 \times 10^{9} kg$$

$$M_{0} = 1.0472 \times 10^{6} kg$$

$$M_{L} = 4.7746 \times 10^{3} M_{0} \sim 5 \times 10^{3} M_{0}$$

Input parameters:

tauS = 1

tauB = 0.1; tauS/tauB = 10.

epsilon = 0.1 Coefficient of restitution

 $rho0 = 0.25g/cm^3$ . Uncompressed density of ring particle aggregates.

m0 =  $1.05 \times 10^9$  g, mass of R0=10m sphere with rho0=0.25g/cm<sup>3</sup>. Reference mass.

S = 300cm, small particle radius, from mass density rho0=0.25 g/cm<sup>3</sup>, and optical depth tauS=0.1.

Vthresh(M0) = 1 cm/sec

# Equilibrium distribution of Mass of aggregates:



The Power Law for index radius distribution was found to be : -0.3386

## Final Plots (Distance from Saturn 140K km, presence of tidal environment)

Power Law Index: -0.3386

x(t) - Mass (M/M\_)

Power Law Index: -1.0158 Equilibrium probabilities of Mass statespace 10<sup>0</sup> Equilibrium probabilities of Radius of aggregates 10<sup>0</sup> 10<sup>-1</sup> 10<sup>-1</sup> Probabilities Probabilities 10<sup>-2</sup> 10<sup>-2</sup> 10<sup>-3</sup> 10<sup>2</sup> 10<sup>3</sup> 10<sup>4</sup> 10<sup>5</sup> R in m 10<sup>-3</sup> 10<sup>6</sup> 10<sup>8</sup> 10<sup>10</sup> 10<sup>4</sup>

#### Conclusion:

- The Predator Prey model can include the outcomes of random collisions in the presence of tidal environment by using the results of numerical simulations.
- The power law index of mass distribution was found to be : -0.3386
- The power law index of radius of aggregates distribution was found to be : -1.0158
- The power law index of the mass distribution obtained from the simulation match well with results obtained from observations. (more explanation might be needed for this point)
- The Mass distribution can be computed for different settings of tidal environment.
- The Long term behavior of the rings can be statistically predicted using the equilibrium mass distributions using Predator Prey model, which could otherwise be very time consuming.
- Though there is a strong presence of tidal environment (140k km, there is still a possibility of finding
  aggregates with high masses, this could explain the presence of Straws in F ring ?(Not very sure about this
  point)